

Title: Lecture - Strong Gravity, PHYS 777

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Collection/Series: Strong Gravity (Elective), PHYS 777, February 24 - March 28, 2025

Subject: Strong Gravity

Date: March 20, 2025 - 10:15 AM

URL: <https://pirsa.org/25030050>

Generalized Harmonic Formulation

$$R_{ab} = 0 \quad \rightarrow \quad R_{ab} - \nabla_{(a} C_{b)} = 0$$

$$C^a = H^a - \square x^a$$

Maxwell Eqns

$$\partial_a F^{ab} = \partial^a \partial_a A_b - \partial^a \partial_b A_a = 0$$

+ component

$$-\cancel{\partial_+^2 A_+} + \partial_+ \partial^+ A_+ + \cancel{\partial_+^2 A_+} - \partial^+ \partial_+ A_+ = 0$$

$$\partial^+ (\partial_+ A_+ - \partial_+ A_+) = 0$$

$$\nabla \cdot E = 0$$

Lorenz Gauge

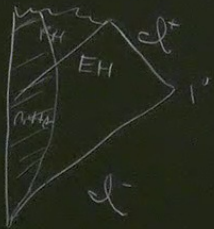
$$\partial_a A^a = 0$$

$$\partial_a \partial^a A_b = 0$$

$$\partial^b (\partial_a \partial_a A_b) = \partial_a \partial^a (\partial^b A_b) = 0$$

Event horizon - the boundary of the causal past of future null infinity

$$\text{BH: } \mathcal{B} = \mathcal{M} - \mathcal{J}^-(\mathcal{I}^+) , \quad \text{EH is } \partial\mathcal{B}$$



- Requires knowledge at to \mathcal{I}^+
- "Teleological" anticipates future

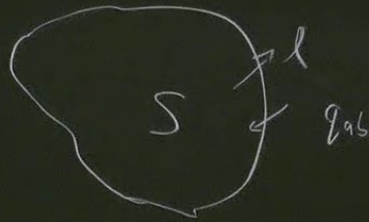
\mathcal{M}

the null infinity
is $\partial\mathcal{B}$

$$\frac{2M}{r} \leq 1$$

Apparent horizons

$\Sigma_+ - 3 \text{ dim}$



Expansion

$$\Theta = \mathcal{L}_l (\log(\sqrt{g})) = g^{ab} \nabla_a l_b$$

Take l to be null

Outward/inward null

$$l^\pm = n^a \pm s^a$$

\nearrow unit normal to Σ_+ \nwarrow unit spacelike vector normal to Σ_+

$$n_a n^a = -$$

$$-1 \quad S_a S^a = 1 \quad l_{\pm}^a l_a^{\pm} = 0 \quad K_{\pm}^a l_a^{\pm} = -2$$

Outward null expansion

$$\sigma_{l^+} = g^{ab} \nabla_a l_b^+ = D_i S^i + K_{ij} S^i S^j - K \quad (\text{in } 3+1 \text{ variables})$$

$$\text{If } \sigma_{l^+} \leq 0 \text{ on } S \quad \frac{1}{r^2} \partial_r (r^2) = \frac{2}{r}$$

$\sigma_{l^+} < 0$ S is outer trapped surface

$\sigma_{l^+} = 0$ S is marginally " " "

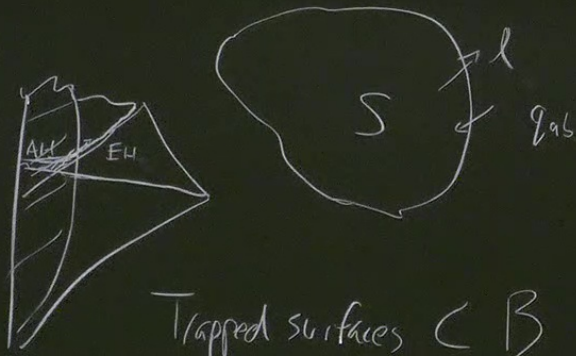
(MOTS)

Outermost MOTS = Apparent horizon

Apparent horizons

$$n_a n^a = -1$$

$\Sigma_+ - 3 \text{ dim}$



Trapped surfaces $\subset B$
(assume (C, NEC))

Expansion

$$\Theta = \mathcal{L}_l (\log(\sqrt{|g|})) = g^{ab} \nabla_a l_b$$

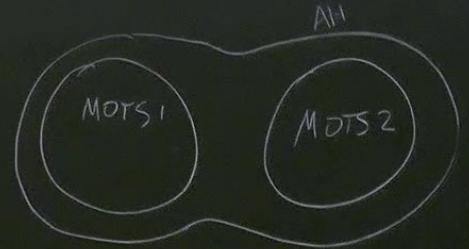
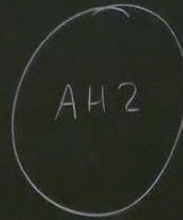
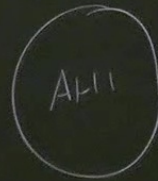
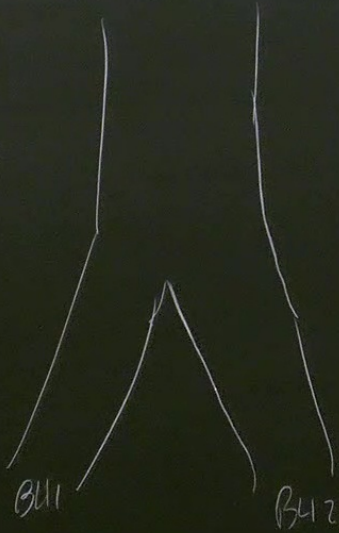
Take l to be null

Outward/inward null

$$l^\pm = n^a \pm s^a$$

\nearrow unit normal to Σ_+ \nwarrow unit spacelike vector normal to Σ_+

Event horizon



Gravitational Waves

ref. arXiv gr-qc/0501041

Linear perturbations of metric (vacuum $R_{ab} = 0$)

$$g_{ab} = \hat{g}_{ab} + h_{ab}$$

$$|h_{ab}| \ll \hat{g}_{ab}$$

$$g^{ab} = \hat{g}^{ab} - h^{ab}$$

(to linear order)

$$g^{ab} = \hat{g}^{ab} - h^{ab} \quad (\text{to linear order})$$

$$\Gamma_{bc}^a = \hat{\Gamma}_{bc}^a + \frac{1}{2} \hat{g}^{ad} (\partial_b h_{cd} + \partial_c h_{bd} - \partial_d h_{bc}) - \frac{1}{2} h^{ad} (\partial_b \hat{g}_{cd} + \partial_c \hat{g}_{bd} - \partial_d \hat{g}_{bc})$$

$$\hat{g}^{bc} \delta \Gamma_{bc}^a = \Gamma_{bc}^a - \hat{\Gamma}_{bc}^a = \frac{1}{2} \hat{g}^{ad} (\hat{\nabla}_c h_{bd} + \hat{\nabla}_c h_{bd} - \hat{\nabla}_d h_{bc})$$

$$(-\hat{\nabla}_a x^a = -\nabla x^a + h^{bc} \hat{\Gamma}_{bc}^a = \hat{\Gamma}_{bc}^a + \hat{\nabla}_b (h^{ab} - h^a b)) \quad , h = \hat{g}^{ab} h_{ab}$$

$$R^s_{bcd} = \partial_c \Gamma_{bd}^a - \partial_d \Gamma_{bc}^a + \Gamma_{ce}^a \Gamma_{bd}^e - \Gamma_{de}^a \Gamma_{bc}^e$$

$$= \partial_c \hat{\Gamma}_{bd}^a - \partial_d \hat{\Gamma}_{bc}^a + \partial_c \delta \Gamma_{bd}^a - \partial_d \delta \Gamma_{bc}^a$$

$$= \hat{R}^s_{bcd} + \hat{\nabla}_c \delta \Gamma_{bd}^a - \hat{\nabla}_d \delta \Gamma_{bc}^a \quad \text{in general}$$

$\Gamma_{bc}^a = 0$
in locally flat coordinates