

**Title:** Lecture - Strong Gravity, PHYS 777

**Speakers:** William East

**Collection/Series:** Strong Gravity (Elective), PHYS 777, February 24 - March 28, 2025

**Subject:** Strong Gravity

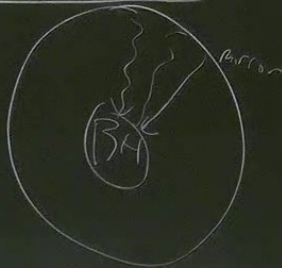
**Date:** March 06, 2025 - 10:15 AM

**URL:** <https://pirsa.org/25030046>

# Superradiance

$$0 < \omega < m\Omega_H$$

Press & Teukolsky



# Massive Scalar Field

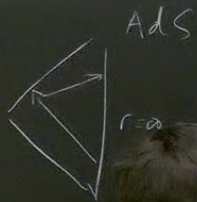
$$\square \phi = \mu^2 \phi$$

mass of boson (up to  $\pm$ )

$$\mu', r \gg M_{BH}$$

non-relativistic

$$\square \phi = \frac{1}{\sqrt{-g}} \partial_a (\sqrt{-g} g^{ab} \partial_b \phi), \quad \sqrt{-g} =$$



## Massive Scalar field

$$\square \phi = \mu^2 \phi$$

mass of boson (up to  $\pm$ )

$\mu^2, r \gg M_{\text{Pl}}$  non-relativistic limit

$$\square \phi = \frac{1}{\sqrt{-g}} \partial_a (\sqrt{-g} g^{ab} \partial_b \phi), \quad \sqrt{-g} = r^2 \sin \theta$$

$$g^{tt} \partial_t^2 \phi + \partial_i \partial^i \phi - \mu^2 \phi = 0, \quad g^{tt} \approx - \left( 1 + \frac{2M}{r} \right)$$



Ansatz  $\phi = \frac{1}{\sqrt{2\mu}} \left[ \psi(x) e^{-i\omega t} + \psi^*(x) e^{i\omega t} \right]$

$$\left[ -\left(1 + \frac{2M}{r}\right) (-\omega)^2 + \partial_r \partial_r - \mu^2 \right] \psi = 0$$

$$(\omega^2 - \mu^2) \psi = \left( -\partial_r \partial_r - \frac{2M\omega^2}{r} \right) \psi$$

$$(\omega - \mu) \psi = \left( -\frac{1}{2\mu} \partial_r \partial_r - \frac{\mu M}{r} \right) \psi$$

$$\omega \approx \mu$$

"fine"

$$\partial_i \partial^i \phi - \mu^2 \phi = 0$$

$$g^{++} \approx - \left( 1 + \frac{2M}{r} \right)$$

"fine structure constant"  $\alpha = M\mu \approx 0.1 \left( \frac{\mu}{10^{-12} \text{eV}} \right) \left( \frac{M_{\text{BH}}}{10 M_{\odot}} \right)$

Energy  $E = \omega - \mu \approx -\frac{\alpha^2}{2n^2}$

Bohr Radius  $r_B = (\mu\alpha)^{-1}$

$$\phi \sim e^{-r/2r_B} Y^{l=1}(\theta) e^{-i(\omega t - m\phi)}$$

$$|\text{Im}(\omega)| \ll |\text{Re}(\omega)|$$



# 3+1 Decomposition of Spacetime

GR, Wald Chap 10.2  
3+1 Formalism Gourgoulhon

arXiv:gr-qc/0703035

## Hypersurfaces of Spacetime

4-dim spacetime manifold  $M$

3-dim manifold  $\Sigma$

Embedding of  $\Sigma$  in  $M$

$$\Phi: \Sigma \rightarrow M, \quad \Phi, \Phi^{-1}$$

$\Sigma = \Phi^{-1}(\Sigma)$  is called a hypersurface

Level Set  $\Sigma = \{x^a \in M \mid F(x) = 0\}$

$\partial_a F$  is normal to  $\Sigma$

$\Leftrightarrow (\partial_a F)v^a = 0$  for any  $v^a$  tangent to  $\Sigma$

$(\partial_a F)$ timelike	, call $\Sigma$ spacelike
$\partial_a F$ spacelike	, " " timelike
$\partial_a F$ null	, " " null

---

Restrict to  $\Sigma$  spacelike



Choose coordinates  $x^a = (t, x^i)$  with  $x^i$  coordinates  
 $\uparrow$  3-index  $\sum$

$$\Sigma = \{x^a \in M \mid t=0\}$$

$$\Phi: \Sigma \rightarrow M$$

$$x^i \rightarrow (0, x^i)$$

Unit normal to  $\Sigma$

$$n_a = \frac{-\nabla_a t}{\sqrt{-\nabla_a t \nabla^a t}}, \quad n_a n^a = -1$$

$$n^a = (1, \vec{0}) \text{ in flat space}$$



$x = (t, x^i)$  with  $x^i$  coordinates  
 $\uparrow$  3-index

$x^a \in M \mid t=0$

$\mathbb{R}^4 \rightarrow M$

$x^i \rightarrow (0, x^i)$

$n_a = \frac{-\nabla_a t}{\sqrt{-\nabla_a t \nabla^a t}}, \quad n_a n^a = -1$

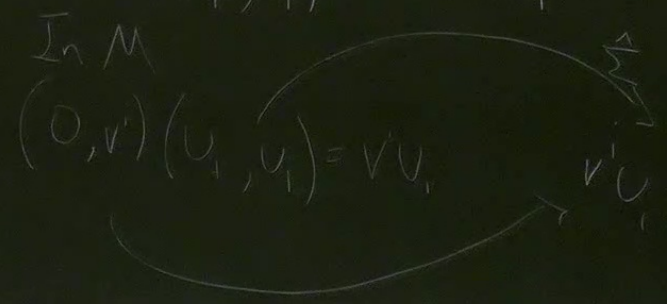
lat space

$\Phi: \mathbb{R}^4 \rightarrow M$   
 $v^i \rightarrow (0, v^i)$

push-forward mapping

$\Phi^*: M \rightarrow \mathbb{R}^4$   
 $(u_+, u_-) \rightarrow u_+$

pull back



$$n^a = (1, \vec{0}) \text{ in flat space} \quad \nabla_a n^b = \nabla_a \delta^b$$

For  $g_{ab}$  metric in  $M$ , induced metric  $\gamma_{ij}$  on  $\Sigma$   
 $\gamma_{ij} = g_{ij}$  in our adapted coordinates

Diff hats  $\Sigma \hookrightarrow \Delta$  tangent

$$v^a = \underbrace{-v^b n_b}_{\text{orthogonal}} n^a + \underbrace{(v^a + v^b n_b n^a)}_{\text{tangent}} = \gamma^a_b v^b$$

$$n_a \gamma^a_b v^b = n^b v_b - n_b v^b = 0$$

$$\gamma^a_b = \delta^a_b + n_b n^a$$



$$(U_+, U_-) = V U_+$$

$$V U_+$$

$$M \rightarrow \Sigma$$

$$V^a \rightarrow \gamma^i_b V^b$$

$$\gamma_{ab} = \gamma^i_a \gamma^j_b \gamma_{ij} = g_{ab} + n_a n_b$$

$$-n_s V^b = 0$$

$$U^a n_a = V^a n_a = 0, \quad \gamma_{ab} U^a U^b = g_{ab} U^a U^b$$

$$U^a = \lambda n^a, \quad \gamma_{ab} U^a V^b = 0$$



# Intrinsic Curvature

Since  $(S_{\text{sub}}, M) \rightarrow (\delta_{ij}, \tilde{\Sigma})$

$$D_i \delta_{jk} = 0$$

(also assume torsion free)

$$\begin{matrix} (3)M \\ | \\ jk \end{matrix}$$

$$(D_i D_j - D_j D_i) v^k = {}^{(3)}R^k{}_{lij} v^l$$

$${}^{(3)}R_{ij} = {}^{(3)}R^k{}_{ikj}$$

scalar  ${}^{(3)}R = \delta^{ij} {}^{(3)}R_{ij}$

$$\Sigma) \quad D_a T^{b_1 b_2} c_1 c_2 = \delta_{d_1}^{b_1} \delta_{c_1}^{e_1} \delta_a^f \nabla_f T^{d_1} e_1$$

$$D_a \delta_{bc} = \delta_b^d \delta_c^e \delta_a^f \nabla_f (g_{de} + n_d n_e) = 0$$

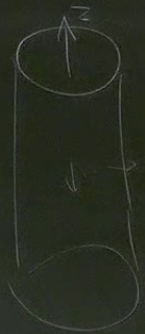


$$\gamma_b^a = \delta_b^a + \kappa_b^a n^a$$

## Extrinsic Curvature



$$K_{ij} = -\nabla_i n_j \quad (\text{symmetric})$$



Cylinder as hypersurface on  $\mathbb{R}^3$

$$\Sigma = \{ (x, y, z) \in \mathbb{R}^3 \mid \rho^2 = x^2 + y^2 = 1 \}$$



$$n_a v_b - n_b v_a = 0$$

$$U^a n_a = V^a n_a = 0, \quad \gamma_{ab} U^a U^b = g_{ab} U^a U^b$$

$$U^i = \lambda \eta^a$$

$$\gamma_{ab} U^a V^b = 0$$

$$n_a = \left( \frac{x}{\rho}, \frac{y}{\rho}, 0 \right)$$

$$\nabla_b n_a = \begin{pmatrix} \frac{y}{\rho^2} & -\frac{x}{\rho^2} & 0 \\ -\frac{x}{\rho^2} & \frac{y}{\rho^2} & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$J_i^a = \frac{dx^a}{dx^i}$$

$$i = \{\phi, z\}$$

$$a = \{x, y, z\}$$

$$\tan \phi = \frac{y}{x}$$

$$K_{ij} = -\nabla_b n_a J_i^a J_j^b$$

$$K_{\phi\phi} = \frac{-(x^2 + y^2)}{\rho^3} \Big|_{\rho=r} = -\frac{1}{r^2}$$

$$U^a n_a = V^a n_a = 0$$

$$U^i = \gamma n^i$$

$$\gamma_{ab} U^a U^b = g_{ab} U^a U^b$$

$$\gamma_{ab} U^a V^b = 0$$

$$n_a = \left( \frac{x}{\rho}, \frac{y}{\rho}, 0 \right)$$

$$\nabla_b n_a = \begin{pmatrix} \frac{y}{\rho^2} & -\frac{x}{\rho^2} & 0 \\ -\frac{x}{\rho^2} & \frac{y}{\rho^2} & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$J_i^a = \frac{dx^a}{dx^i}$$

$$i = \{t, z\}$$

$$a = \{x, y, z\}$$

$$\tan \phi = \frac{y}{x}$$

$$K_{ij} = -\nabla_b n_a J_i^a J_j^b$$

$$K_{\phi\phi} = \frac{-(x^2 - y^2)}{\rho^3} \Rightarrow \dots$$

$$K_{ab} = \delta_a^i \delta_b^j K_{ij}$$

$$= -\nabla_b n_a - (n^c \nabla_c n_a) n_b$$

=  $a_a$   
acceleration