

Title: Black hole evaporation in random matrix theory (RMT) and statistical CFTs

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Collection/Series: Quantum Information

Subject: Quantum Information

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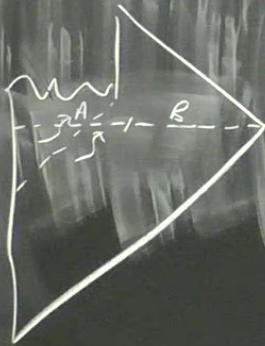
Abstract:

In this talk, with two parts, I will first show how to capture both Hawking's non-unitary entropy curve and density matrix-connecting contributions that restore unitarity, in a toy RMT quantum system modelling black hole evaporation. The motivation is to find the simplest possible dynamical model that captures this aspect of gravitational physics. In the model, there is a dynamical phase transition in the averaging that connects the density matrices in a replica wormhole-like manner and restores unitarity in the entropy curve. In the second half of the talk, I will discuss ongoing follow-up work describing black hole evaporation and unitarity restoration in statistical descriptions of holographic CFTs.

Black hole evaporation in random matrix theory and statistical CFTs

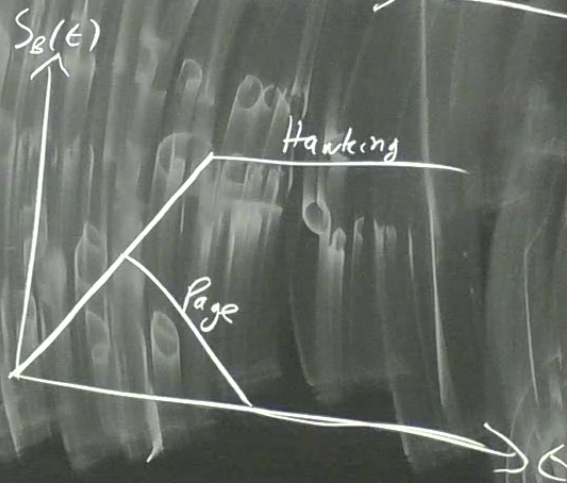
w/ Jan de Boer & Jildou Hollander
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Intro



$$S_B = S_{\text{elom}}$$

Radiation entropy curves

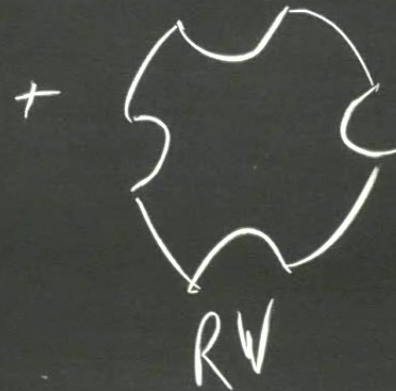


Unitarity \Rightarrow Page \Leftrightarrow ?

Replica wormholes

$$S(\rho) = - \partial_n \text{Tr}(\rho^n) |_{n=1}$$

$$\text{Tr}(\rho^n) = \begin{array}{cc} \rho^{(1)} & \textcircled{1}^{\rho} \\ \rho^{(1)} & \textcircled{1}^{\rho} \end{array} +$$



Ensemble gravity

$$Z(\beta)_{\text{can.}}^2 = \text{Diagram} \Rightarrow \text{Var}(Z) \neq 0$$


Messages

- 1) Ensembles \rightarrow entropy
- 2) Haar averaging \rightarrow Replica wormholes
- 3) Semiclassical gravity = Statistical CFT

CAUTION

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PROCEED WITH CARE IN THE USE OF THE BOARD.

IT IS ESSENTIAL TO AVOID
ANY ACCIDENTS OR INJURIES.

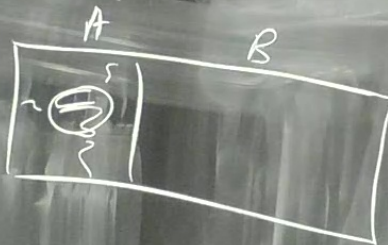
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Black hole evaporation in random matrix theory and statistical CFTs

RMT model

Hilbert space

$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$$

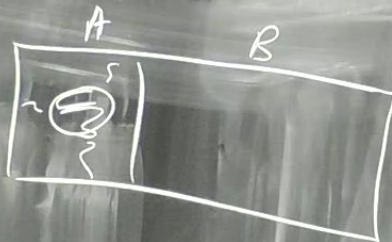


and statistical CFTs

RMT model

Hilbert space

$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$$



$$\mathcal{H} = \bigoplus_E \mathcal{H}_{E, S, E}$$

$$\mathcal{H}_E = \bigoplus_{E'} (\mathcal{H}_{A, E-E'} \otimes \mathcal{H}_{B, E'})$$

\triangleq BH state $|\psi_1\rangle$

$$\approx \left(|E\rangle_A \otimes |0\rangle_B \right) \oplus \left(|0\rangle_A \otimes \mathcal{H}_{B, E} \right)$$

$N-1$ radiation states $|\psi_2\rangle, \dots, |\psi_N\rangle$

$N-1$ Rad.

Hamiltonian

$$H = H_A + H_B + H_{int}$$

$$|\psi(t)\rangle = e^{iHt} |\psi_i\rangle$$

Ensemble averaged state \rightarrow Mixed Hawking radiation

$$\frac{d}{dt} \text{Tr}(\rho^2(t))$$

Ensemble

$$H = u^\dagger \Lambda u$$

$$P(H) = P_{\text{Haar}}(u) P(\Lambda)$$

Averaged density matrix

$$\bar{\rho}_B(t) = \int dH P(H) \text{Tr}_A(\rho(t))$$

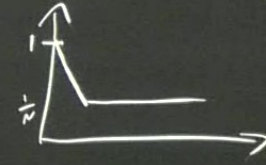
$$= \frac{1 + \bar{g}(t)N}{N+1} |0\rangle\langle 0|_B + \frac{N - \bar{g}(t)N}{N+1} \sum_{i=2}^N |i\rangle\langle i|_B$$

$$\bar{g}(t) := \int dH P(H) \frac{|\text{tr} e^{iHt}|^2}{N^2}$$

$$= |0\rangle\langle 0|_B \quad \text{at } t=0$$

$$= \frac{1}{N+1} \sum_{i=2}^N |i\rangle\langle i|_B \quad \text{as } t \rightarrow \infty$$

$$S(\bar{\rho}_B(t)) = -\bar{g} \log \bar{g} - (1 - \bar{g}) \log \left(\frac{1 - \bar{g}}{N} \right)$$



CAUTION

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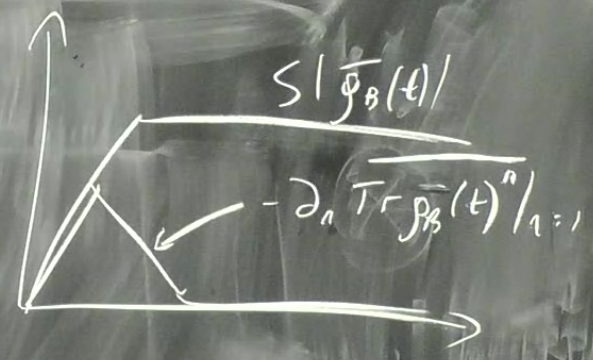
Rescuing unitarity

$$\begin{aligned} S(\rho) &\neq -\partial_n \overline{\text{Tr}(\rho^n)} \Big|_{n=1} \\ \text{Unitarity Page} & \parallel \\ &= \bar{g} \log \bar{g} - (1-\bar{g}) \log(1-\bar{g}) \\ &= S(\bar{\rho}_S(t)) + (1-\bar{g}) \log N \end{aligned}$$

Black hole evaporation in random matrix theory and statistical CFTs

Rescuing unitarity

$$\begin{aligned}
 S(\rho) &\neq -\partial_n \overline{\text{Tr}(\rho^n)} \Big|_{n=1} \\
 \text{Unitarity Page} &\quad \parallel \\
 &= \bar{g} \log \bar{g} - (1-\bar{g}) \log(1-\bar{g}) \\
 &= S(\bar{\rho}_S(t)) + (1-\bar{g}) \log N
 \end{aligned}$$



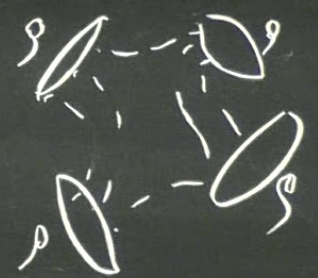
Replica wormholes from Haar averaging

$$g(t) = U e^{i\lambda t} U^\dagger \chi_1 \times \chi_1 U e^{-i\lambda t} U^\dagger$$

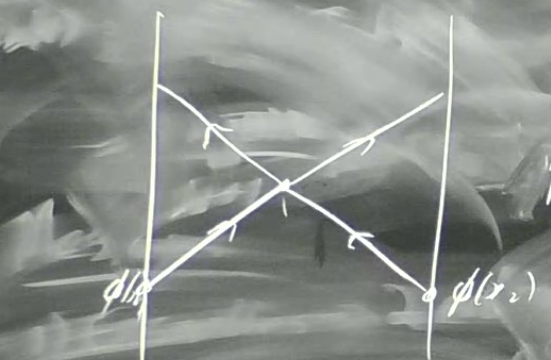
$$\overline{\text{Tr } \rho_B^n} = \overbrace{U^\dagger U \dots U}^{4n}$$

$$\overline{U_{ab} U_{cd}^\dagger} \sim \frac{\delta_{ad} \delta_{bc}}{N}$$

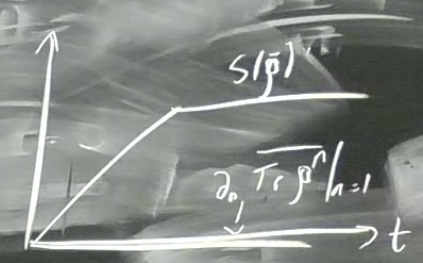
$$\overline{\text{Tr}(\rho_B^2)} = \left(\text{disc.} + \dots \right) + \left(\text{conn.} + \dots \right)$$



BHs in stati. holo. CFTs



$$|\psi(t)\rangle = e^{iHt} \rho(x_1) \phi(x_2) |0\rangle$$



$$\frac{1}{2} C_{\phi\rho H} = \frac{\Delta_H^{-1}}{f(\Delta_H)}$$

