

Title: Lecture - IR S-matrix b

Speakers: Hofie Sigridar Hannesdottir

Collection/Series: Celestial Holography Summer School 2024

Date: July 25, 2024 - 10:00 AM

URL: <https://pirsa.org/24070017>

Infrared Divergences in QFT

(Some) References

S. Caron-Huot, M. Giroux, HSH, S. Mizera
2308.02125

HSH, M. Schwartz 1911.06821

A. Strominger 1703.05448

D. Kapec, M. Perry, A.-M. Raclariu,

A. Strominger 1705.04311

Review from NRQM

- Short-range potentials

$$\psi^{\text{in}} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p} \cdot \vec{x}} + \underbrace{f(p, \cos\theta)}_{\text{Scattering amplitude}} \frac{e^{ipr}}{r}$$

- Coulomb potential:

$$\psi^{\text{in}} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p} \cdot \vec{x}} + \frac{i\delta}{p} \log[pr(1-\cos\theta)] - \check{f}(p, \cos\theta) \frac{e^{ipr - \frac{i\pi}{2}}}{r}$$

ances in QFT

Giroux, HStl, S. Mizera

2308.02125

1911.06821

1703.05448

ny, A. M. Radwin,

omingel 1705.04311

Review from NRQM

- Short-range potentials

$$\psi_{in} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p}\cdot\vec{x}} + \underbrace{f(p, \cos\theta)}_{\text{Scattering amplitude}} \frac{e^{ipr}}{r}$$

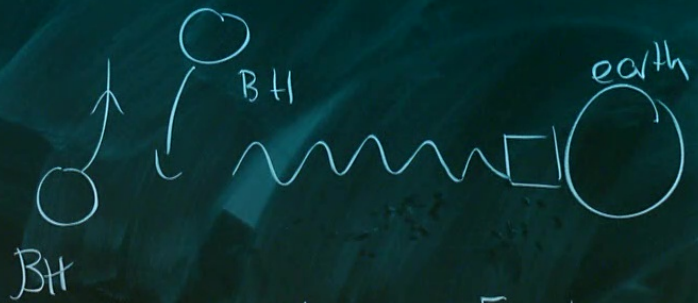
- Coulomb potential:

$$\psi_{in} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p}\cdot\vec{x}} + \underbrace{\frac{i\gamma}{p} \log[pr(1-\cos\theta)]}_{\text{causes time delay}} - \underbrace{f(p, \cos\theta)}_{\checkmark} \frac{e^{ipr - \frac{i\gamma}{p} \log(pr)}}{r}$$

$f(p, \cos\theta) \frac{e^{ipr}}{r}$
 scattering amplitude

$+ \frac{i\delta}{p} \log[pr(1-\cos\theta)]$ causes time delay
 $f(p, \cos\theta) \frac{e^{ipr - \frac{i\delta}{p} \log(pr)}}{r}$

Waveform:



$\hat{M} = \hat{M}^{\text{tree}} \exp[i(\#_1 + \#_2) \frac{1}{\epsilon_{\text{IR}}}]$
 waveform

Key point: IR divergences
 ↔ physical imprint of long-range dynamics.

Complications in QFT:

- IR divergence is not just a phase.
- IR divergences do not need to exponentiate.
- IR divergences not only from low-energy particles.

$$\frac{1}{(p+k)^2 - m^2 + i\epsilon} \rightarrow \frac{1}{2pk + k^2 + i\epsilon}$$

$k \rightarrow 0$ } divergence.

propagator: $\frac{1}{(p+k)^2 + i\epsilon} \sim \frac{1}{2|k|(1 - \cos\theta) + i\epsilon}$

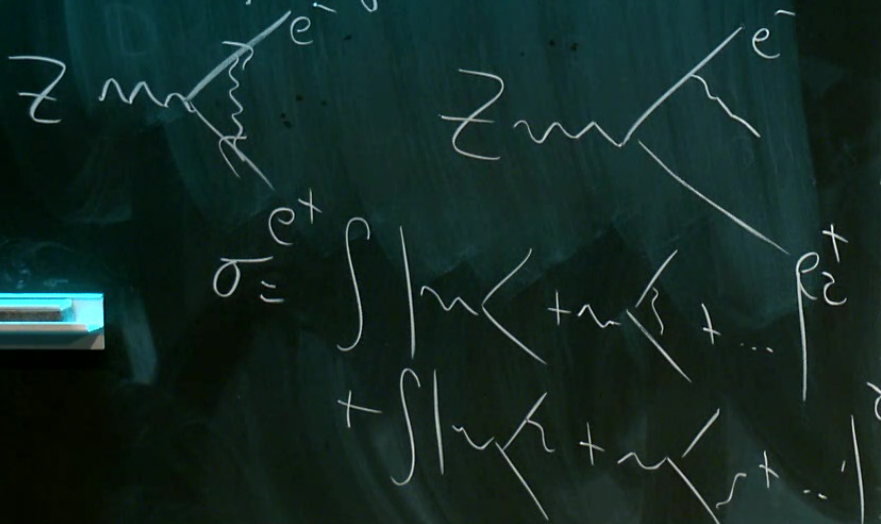
divergences $\left\{ \begin{array}{l} |k| \rightarrow 0 \text{ soft} \\ \theta \rightarrow 0 \text{ collinear} \end{array} \right.$

Infrared Divergences in QFT

What can we do?

① Define observables

that are agnostic to long-range effects.



Review from NRQM

• Short-range potentials

$$\psi_{in} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p} \cdot \vec{x}} + \underbrace{f(\vec{p}, \cos\theta)}_{\text{Scattering amplitude}}$$

• Coulomb potential:

$$\psi_{in} \xrightarrow{|\vec{x}|=r \rightarrow \infty} e^{i\vec{p} \cdot \vec{x} + \frac{i\alpha}{p} \log[pr]} - f(\vec{p}, \cos\theta)$$

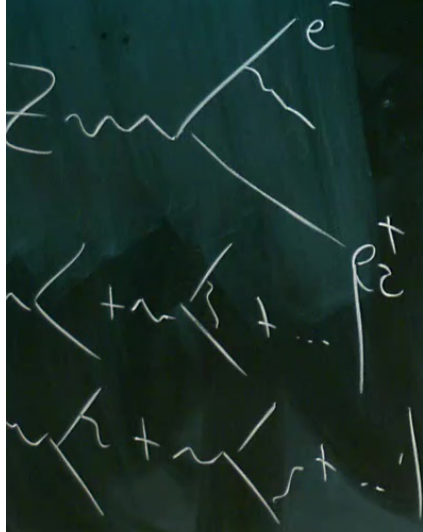
es in QFT

do?

observable's

stic to

effects.



More generally (QCD) more complicated

Idea: $\langle \hat{O} \rangle \equiv \langle \Psi | \hat{O} | \Psi \rangle$
in in

prove that is IR finite.

What if we like to preserve S-matrix properties?
Causality, analyticity, dispersion rel.

Wan

BH

wo
Ke

$|\psi\rangle_{in}$

te.

matrix properties?

dispersion rel.

II) Redefine / modify S-matrix scattering states.

Heisenberg picture

Schrodinger picture

Recall QM case:

$$\left. \begin{aligned} \psi_{in} &\underset{r \rightarrow \infty}{\sim} \frac{e^{ipr}}{r} f(p, \cos\theta) \\ \psi_{out} &\underset{r \rightarrow \infty}{\sim} e^{i\vec{p}' \cdot \vec{x}} \end{aligned} \right\} \begin{aligned} &\langle \psi_{out} | \psi_{in} \rangle \\ &= \# \dots f \end{aligned}$$

Complications

- IR divergences not just
- IR divergences need to
- IR divergences from low-

redefine / modify
matrix scattering states.

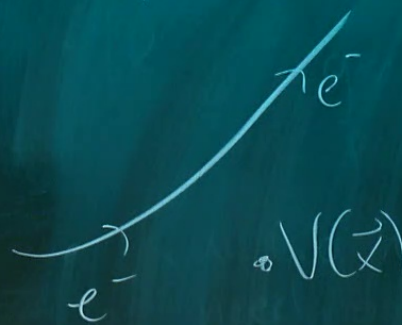
Sternberg picture

Schrödinger picture

in QM case:

$$\left. \begin{aligned} \psi_{in} &\sim \frac{e^{ipr}}{r} f(p, \cos\theta) \\ \psi_{out} &\sim e^{ip|x} \end{aligned} \right\} \begin{aligned} &\langle \psi_{out} | \psi_{in} \rangle \\ &= \delta(\dots) \end{aligned}$$

Anticipating
Away from potential scattering



State as $t \rightarrow -\infty$ $e^{-iH_0 t} |i\rangle$
 $t \rightarrow +\infty$ $e^{-iH_0 t} |f\rangle$

H_0 : Free Hamiltonian

Evolution op e^{iHt}

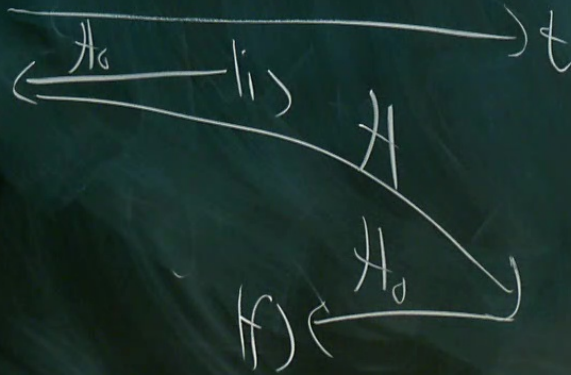
propagator
divergence

Infrared Divergences in QFT

What can we do?

S-matrix:

$$S_{fi} = \lim_{t_i \rightarrow -\infty} \langle f | e^{iH_0 t_f} e^{-iH t_f} e^{iH t_i} e^{-iH_0 t_i} | i \rangle$$



More generally (QCD) complicated!

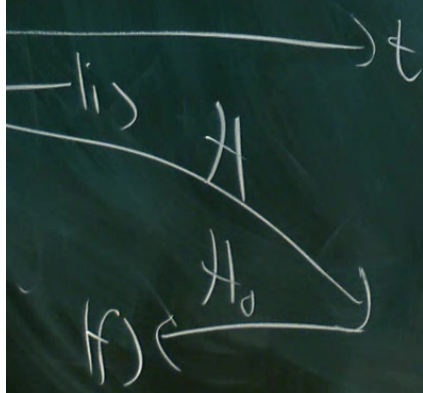
Idea: $\langle \hat{O} \rangle \equiv \langle \dots \rangle_{in}$
 Prove that \int is IR

What if we like to preserve Causality, analyticity

ferences in QFT

we do?

$$S = \lim_{t \rightarrow \infty} e^{-iHt} e^{iH_0 t} = \lim_{t \rightarrow \infty} e^{-i(H-H_0)t}$$



• Free theory: $H = H_0$

$$S = 1, \quad S_{fi} = \langle f | i \rangle \checkmark$$

• Short-range potential scattering \checkmark

• Constant potential: $H = H_0 + \frac{V_0}{i2V_0 t}$

$$S_{fi} = \langle f | i \rangle \lim_{T \rightarrow \infty} e^{-i2V_0 T}$$

$$\text{• QED: } S_{fi} = \left(1 - \frac{\alpha \#}{\epsilon} + \dots \right) \hat{S} = \exp \left[-\frac{\alpha \#}{\epsilon} \right] \hat{S}$$

$$\xrightarrow{\epsilon \rightarrow 0} 0$$

(II)

R

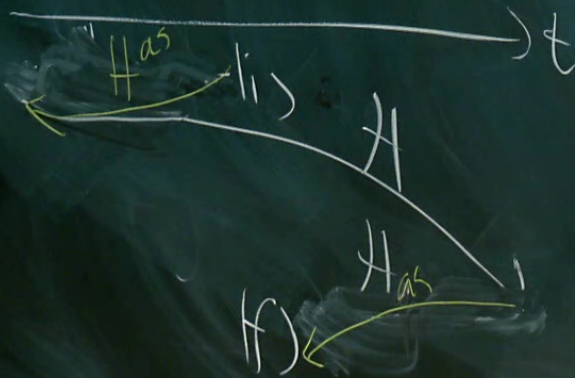
Infrared Divergences in QFT

What can we do?

S-matrix:

$$S_{fi} = \lim_{t_i \rightarrow -\infty} \langle f | e^{iH_0 t_f} e^{-iH t_f} e^{iH t_i} e^{-iH_0 t_i} | i \rangle$$

S



• Free theory: $H = H_0$

$$S = I, \quad S_{fi} = \langle f | i \rangle$$

• Short-range potential scattering

• Constant potential: $H = H_0 + V$

$$S_{fi} = \langle f | i \rangle \lim_{T \rightarrow \infty} e^{-iV T}$$

• QED: $S_{fi} = \left(1 - \frac{\alpha}{\epsilon} + \dots \right)$

Question: How to pick H_{as} ?

- \mathbb{R} finite.
- Incoming states are free of radiation.
- Calculable.
- Useful (i.e. represent something physical).
- Compatible with symmetries.



How to pick H_{as} ?

finite.

coming states are
free of radiation

calculable.

useful (i.e. represent
something physical).

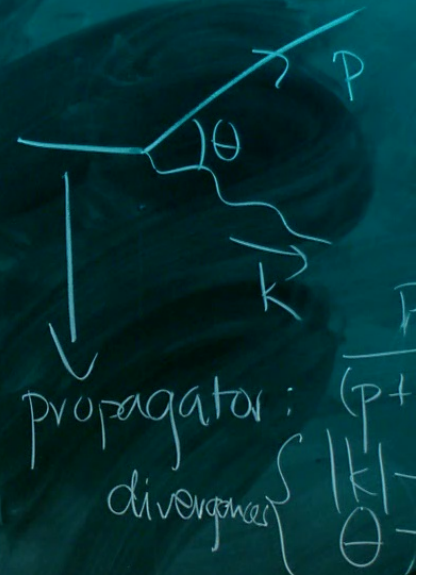
compatible with symmetries.

In NRQM (Dollard 1971)

$$H_{as} \sim H_{ot} + \frac{i\gamma}{\sqrt{-\nabla^2}} \log\left(\frac{|\vec{\nabla}|}{m}\right)$$

In QED:

$$H_{as} - H_{ot} = H_{int}$$
$$\Theta(\omega_k < \gamma)$$
$$\Theta(\theta < \gamma)$$



Examples:

(I) QED: divergences exponentiate



$$\mathcal{M}^{\text{ren}} = \exp \left[-\frac{\alpha}{2\pi\epsilon_{\text{IR}}} \left[(\beta - i\pi) \cosh\beta - 1 \right] \right] \times \mathcal{M}^{\text{finite}}$$

$$\cosh\beta = \frac{v_1 v_2}{N_{11} |v_2|} \quad v_1^M = \frac{p_1^M}{E_1}$$

Question Ho

- IR fin
- Incom free
- Calcul
- Usefu
- Comp

(Weinberg 1965)
 as exponentiate

$$\left[(\beta - i\pi) \coth(\beta - 1) \right]$$

$$M_{finite}$$

$$v_1 v_2$$

$$N_1 N_2$$

$$v_i^M = \frac{p_i^M}{E_i}$$

Allows for a redefinition of
 the S-matrix

(Faddeev & Kulish 1970)

$$S_{fi}^H = \lim_{t_{\pm} \rightarrow \pm\infty} \langle f | e^{iH_{ast} t_+} e^{-iH_{tot} t_+} | f \rangle$$

$$\langle f' | e^{iH_{tot} t} e^{-iH_{tot} t} e^{iH_{tot} t} e^{-iH_{tot} t} | f \rangle$$

$$\langle i | e^{iH_{tot} t} e^{-iH_{ast} t} | i \rangle$$

In NRQ

$$A_{ast} \sim$$

In QED

$$A_{ast}$$

(Weinberg 1965)
 as exponentiate

$$\left[(\beta - i\pi) \coth(\beta - 1) \right]$$

$$M_{finite}$$

$$v_i^M = \frac{p_i^M}{E_i}$$

$$\frac{v_1 v_2}{N_1 N_2}$$

Allows for a redefinition of
 the S-matrix

(Faddeev & Kulish 1970)

$$S_{fi}^H = \lim_{t_{\pm} \rightarrow \pm\infty} \langle f | e^{iH_{ast} t_+} e^{-iH_{tot} t_+} | f \rangle$$

$$\langle f' | e^{iH_{tot} t} e^{-iH_{tot} t} e^{iH_{tot} t} e^{-iH_{tot} t} | f \rangle$$

usual S

$$\langle f' | e^{iH_{tot} t} e^{-iH_{ast} t} | f \rangle$$

In NRQ

$$A_{ast} \sim$$

In QED

$$A_{ast}$$

for a redefinition of
 S-matrix
 (Faddeev & Kulish 1970)

$$\lim_{t \rightarrow +\infty} \langle f | e^{iH_{\text{as}} t} e^{-iH_{\text{tot}} t} | i \rangle$$

$$\langle f | e^{iH_{\text{tot}} t} e^{-iH_{\text{tot}} t} | i \rangle$$

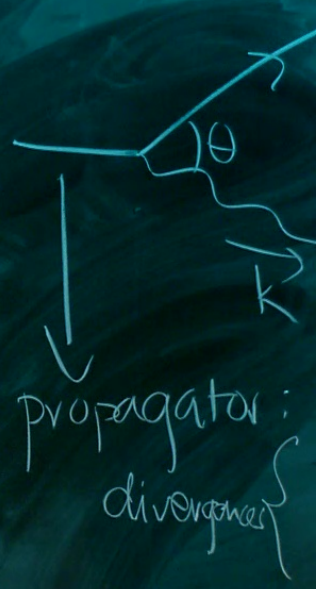
usual S

$$\langle f | e^{iH_{\text{tot}} t} e^{-iH_{\text{as}} t} | i \rangle$$

real rad. ← Coulomb phase

$$S^{\pm} = e^{-R - i\Phi} S e^{R + i\Phi}$$

$$R_f \sim \int d^3k d^3p (f^{\mu} a_{\mu}^{\dagger}(\vec{k}) - f^{*\mu} a_{\mu}(\vec{k})) \rho(\vec{p})$$



for a redefinition of
 S-matrix
 (Faddeev & Kulish 1970)

$$\lim_{t \rightarrow +\infty} \langle f | e^{iH_{\text{as}} t} e^{-iH_{\text{tot}} t} | i \rangle$$

$$\langle f | e^{iH_{\text{tot}} t} e^{-iH_{\text{tot}} t} | i \rangle$$

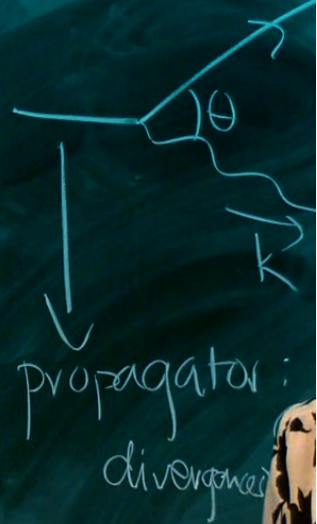
usual S

$$\langle f | e^{iH_{\text{as}} t} e^{-iH_{\text{as}} t} | i \rangle$$

real rad → Coulomb phase

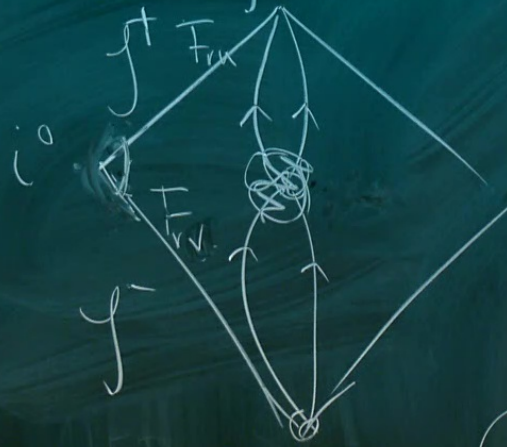
$$S^{\#} = e^{-R-i\Phi} S e^{R+i\Phi}$$

$$R_f \sim \int_0^{\Delta} d^3k d^3p (f^{\mu} a_{\mu}^{\dagger}(\vec{k}) - f^{*\mu} a_{\mu}(\vec{k})) \rho(\vec{p})$$



QFT

Also equivalent to imposing
Scattering of eigenstates
of asymptotic charge.



$$Q_{\xi}^{+} = \frac{1}{e^2} \int_{j^{+}} d^2z \delta_{z,\bar{z}} \mathcal{E}(z, \bar{z}) F_{in}$$

$$Q_{\xi}^{-} = \frac{1}{e^2} \int_{j^{-}} d^2z \delta_{z,\bar{z}} \mathcal{E}(z, \bar{z}) F_{out}$$

Conservation of Q_{ξ}

$$\langle out | Q_{\xi}^{+} S - S Q_{\xi}^{-} | in \rangle$$

Allows for
the

$$S^{\dagger} = S^{-1}$$

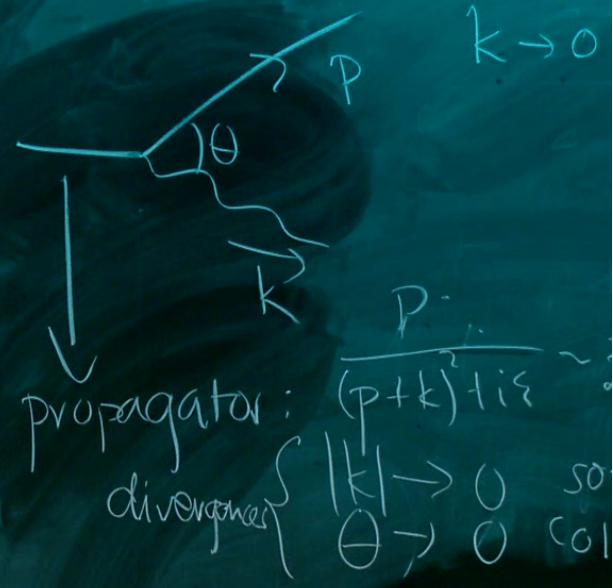
basis
 on the (Weinberg)
 eigenstates
 and Q^\pm
 & Kulish.

real rad. ←
 Coulomb phase ←

$$S^\pm = e^{-R - i\Phi} S e^{R + i\Phi}$$

$$R_f \sim \int_0^\Delta d^3k d^3p (f^\mu a_\mu^\dagger(\vec{k}) - f^{*\mu} a_\mu(\vec{k})) \rho(\vec{p})$$

$\frac{p \cdot k}{p \cdot k}$



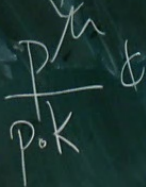
nberg)

tes

real rad Coulomb phase
 ↙ ↘
 $-R - i\Phi$ $R + i\Phi$

$$S^\# = e^{-R - i\Phi} S e^{R + i\Phi}$$

$$R_f \sim \int_0^\Delta \int_0^3 d^3k d^3p (f^\mu a_\mu^\dagger(\vec{k}) - f^{*\mu} a_\mu(\vec{k})) \rho(\vec{p})$$



What do we take as Δ and f^μ ? ($\phi \rightarrow 1$ $|\vec{k}| \rightarrow 0$)

- 1) Connect to physical scales?
- 2) Keep analyticity properties?

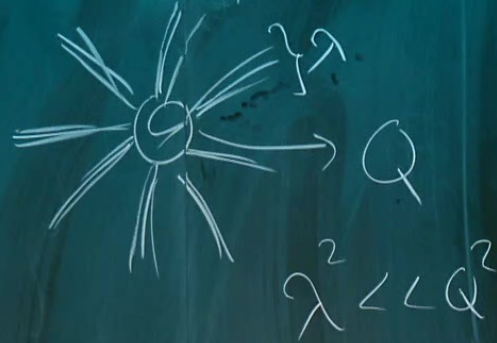
$$A = \langle 0 | W | W_n(0) \rangle \times A^{\text{hard}}$$

2012 04208

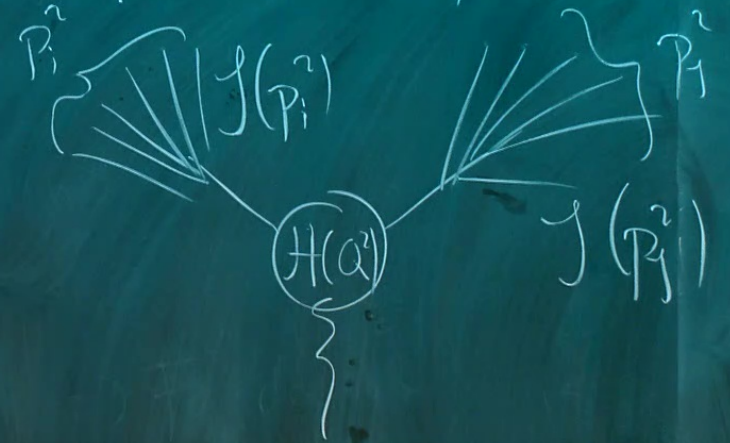
I. QCD.

[Faint handwritten notes]

$M_{n \rightarrow m}$



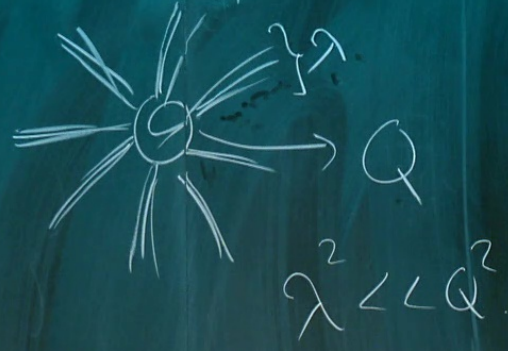
Separate Soft, collinear, hard



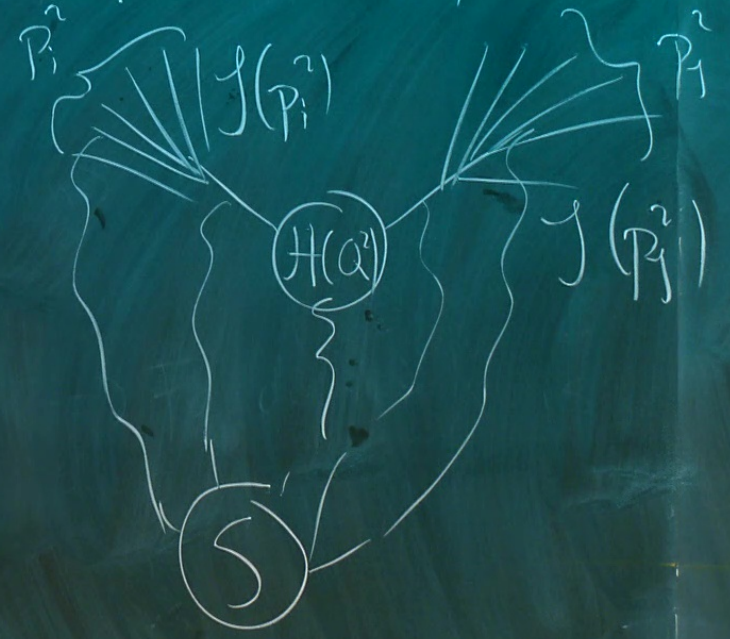
I. QCD.

soft

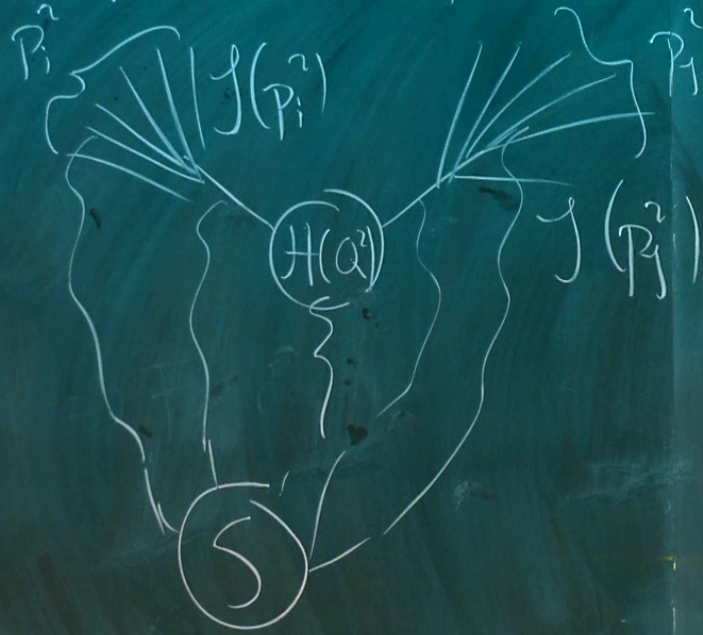
$M_{n \rightarrow m}$



Separate Soft, collinear, hard

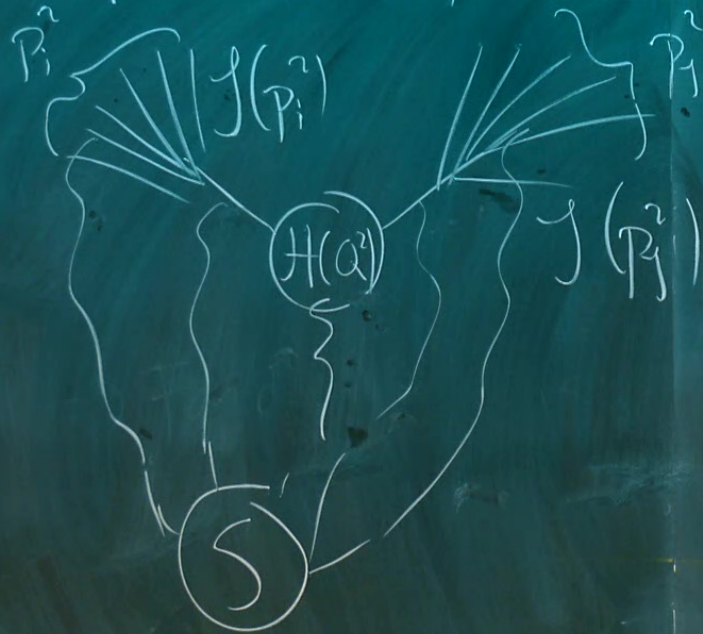


Separate Soft, collinear, hard



Framework: Soft-collinear effective theory
1410.1892.

Separate Soft, collinear, hard



Framework: Soft-collinear effective theory 1410.1892.

$$\mathcal{M}^{\text{QCD}} = \text{SOET}$$

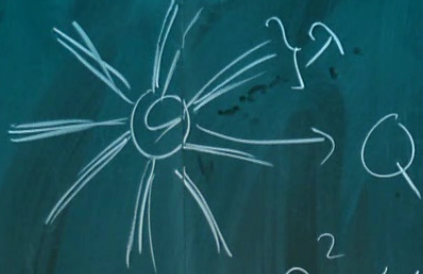
↑
IR divergent

↑
Wilson coefficient

↑
UV divergent

$$\frac{1}{\epsilon_{\text{IR}}} = \frac{1}{\epsilon_{\text{UV}}} + \left(\frac{1}{\epsilon_{\text{IR}}} - \frac{1}{\epsilon_{\text{UV}}} \right)$$

$M_{n \rightarrow m}$



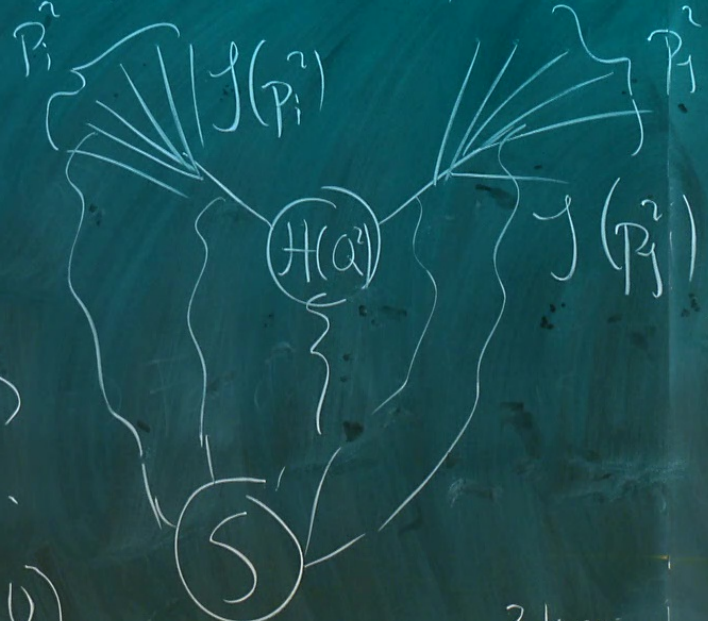
$Q^2 \ll \Lambda^2$

$e^{-i p_{last} \cdot x} e^{-i p_{tot} \cdot x} |f\rangle \langle f| S |i\rangle$

$H \times S^2$

$\langle 0 | W_1 \dots W_n | 0 \rangle$

Separate Soft, collinear, hard



In QED:

M^{3-loop}

M^{4-loop}

IR divergences

$\frac{1}{\epsilon_{IR}}$

Framework:

M^{QCD}

↑
IR divergences