

Title: Lecture - IR S-matrix a

Speakers:

Collection: Celestial Holography Summer School 2024

Date: July 24, 2024 - 10:00 AM

URL: <https://pirsa.org/24070013>

Scattering theory & Infrared Divergences

References:

- Scattering Theory, Taylor
- Scattering Theory of Waves and Particles, Newton
- Lecture notes by D. Tong
damp.cornell.edu/~user/tong/agm/agmten.pdf
- Lecture notes by S. Mizera
2306.05395.



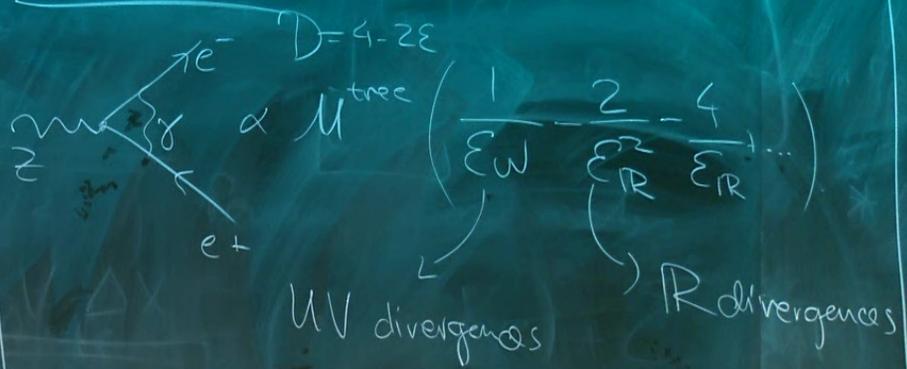
PLAN:

(I) IR divergences in Quantum Mechanics

(II) ——— in QFT.

Goal: Understand the physics.

Motivation:



Heliosophy / Effects of Renormalization

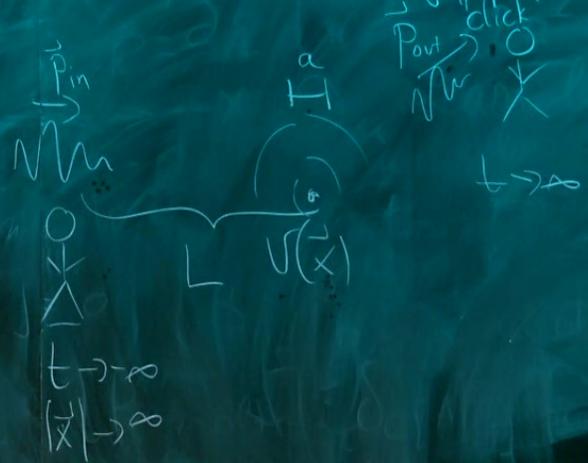
-2ϵ

$$\left(\begin{array}{c|c|c|c} 1 & 2 & 4 & \dots \\ \hline EW & ER & ER & \dots \end{array} \right)$$

divergences IR divergences

S-matrix:
 What is the probability amplitude for measuring \vec{p}_{out} if we sent in \vec{p}_{in} ?

⊕ What is a scattering experiment?



How do we formalize this?

If $V(\vec{x}, t) = V(\vec{x})$ then

$$\Psi(\vec{x}, t) = e^{-iEt} \psi(\vec{x})$$

$$\left(-\frac{\nabla^2}{2m} + V(\vec{x}) \right) \psi(\vec{x}) = E \psi(\vec{x})$$

If the potential $V(\vec{x})$ is short ranged
then asymptotically $|\vec{x}| \rightarrow \infty$:

$$\psi(\vec{x}) \sim e^{\pm i \vec{q} \cdot \vec{x}}, \quad \vec{q} \in \mathbb{R}^3$$

$$E = \frac{\hbar^2 q^2}{2m}$$

• Bound states:

$$\psi(\vec{x}) \sim e^{-\kappa r}$$

$\kappa > 0$

• Scattering states:

$$\psi(\vec{x}) \sim e^{i \vec{p} \cdot \vec{x}}$$

$\kappa = 0$

$$E = \frac{\hbar^2 p^2}{2m}$$

$$p = \hbar q$$

If the potential $V(\vec{x})$ is short ranged
 then asymptotically $|\vec{x}| \rightarrow \infty$

$$\psi(\vec{x}) \sim e^{\pm i \vec{q} \cdot \vec{x}}, \quad \vec{q} \in \mathbb{R}^3$$

$$E = \frac{\hbar^2 q^2}{2m} = \frac{\hbar^2 |\vec{q}|^2}{2m}$$

Bound states:

$$\psi(\vec{x}) \sim \frac{e^{-\kappa r}}{r}, \quad \kappa > 0$$

Scattering states:

$$\psi(\vec{x}) \sim \begin{cases} e^{i \vec{p} \cdot \vec{x}} \\ e^{i \vec{p} \cdot \vec{x}} + \frac{e^{i p r}}{r} \end{cases} < 0$$

$$E = \frac{\hbar^2 p^2}{2m} > 0, \quad p = |\vec{p}|$$

Define:

"In" wavefunctions:

$$\psi_{in} \xrightarrow{r=|\vec{x}| \rightarrow \infty} e^{i \vec{p} \cdot \vec{x}} + f(p, \hat{p} \cdot \hat{x}) \frac{e^{i p r}}{r}$$

"Out" - " - " - " - "

$$\psi_{out} \xrightarrow{r \rightarrow \infty} e^{i \vec{p} \cdot \vec{x}} + \tilde{f}(p, \hat{p} \cdot \hat{x}) \frac{e^{-i p r}}{r}$$

S-matrix:

What is the
 amplitude
 \tilde{p}_{out} if we

Scattering

1st order perturbation

ions:

$$e^{i\vec{p}\cdot\vec{x}} + f(p, \hat{p}\cdot\hat{x}) \frac{e^{ipr}}{r}$$

$$\Rightarrow e^{i\vec{p}\cdot\vec{x}} + \tilde{f}(p, \hat{p}\cdot\hat{x}) \frac{e^{-ipr}}{r}$$

$$\tilde{f} = \int \frac{d^3p}{(2\pi)^3} g(\vec{p}-\vec{p}_0) e^{-i\frac{p^2}{2m}t} \times \left[\underbrace{e^{ipr\cos\theta}}_1 + \underbrace{f(p, \hat{p}\cdot\hat{x}) \frac{e^{ipr}}{r}}_2 \right]$$

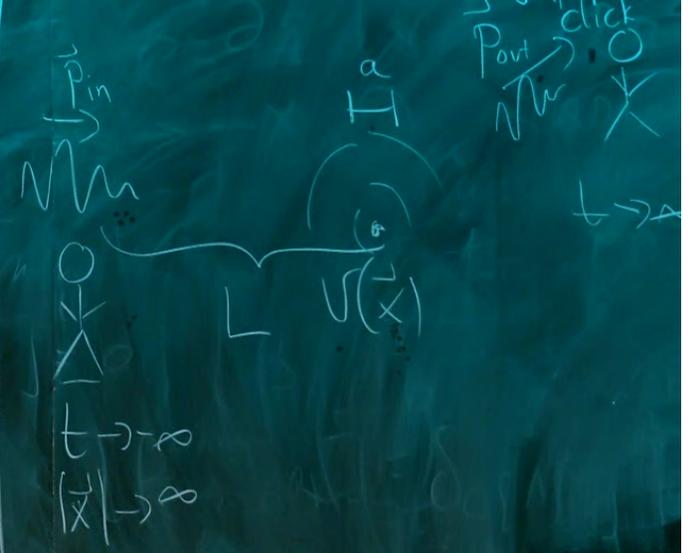
Saddle point analysis:

$$0 = \frac{\partial}{\partial p} \left[-i\frac{p^2}{2m}t + ipr\cos\theta \right]$$

$$\Rightarrow t_1 = \frac{mr\cos\theta}{p}$$

$$t_2 = \frac{mr}{p}$$

⊕ What is a scattering experiment



Logography

Extrapolation $x^2 = (m \cdot \lambda)^2$

$$\vec{p} \cdot \vec{x} + f(p, \hat{p}, \hat{x}) \frac{e^{ipr}}{r}$$

$$\vec{p} \cdot \vec{x} + f(p, \hat{p}, \hat{x}) \frac{e^{-ipr}}{r}$$

$$\tilde{\Psi}^m = \int \frac{d^3p}{(2\pi)^3} g(\vec{p}-\vec{p}_0) e^{-i\frac{p^2}{2m}t} \times \left[\underbrace{e^{ipr \cos \theta}}_1 + \underbrace{f(p, \hat{p}, \hat{x}) \frac{e^{ipr}}{r}}_2 \right]$$

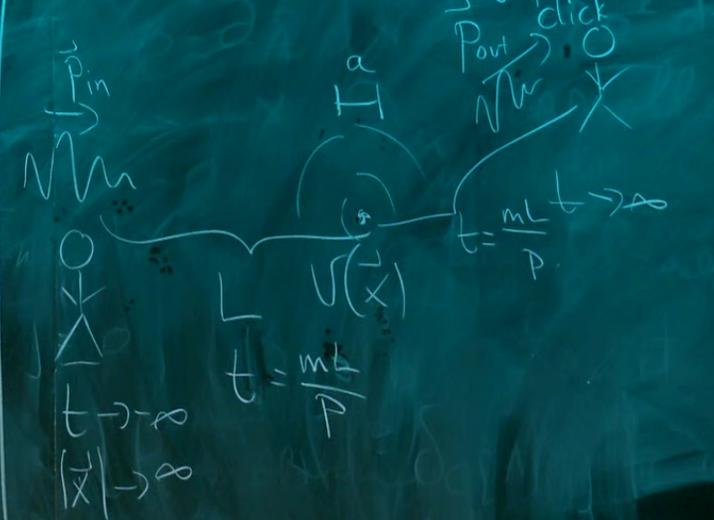
Saddle point analysis:

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$$\Rightarrow t_1 = \frac{mr \cos \theta}{p}$$

$$t_2 = \frac{mr}{p}$$

⊕ What is a scattering experiment?



$$S_{\vec{p}_{in}, \vec{p}_{out}} = \langle \psi^{out} | \psi^{in} \rangle$$

What if we have a Coulomb potential?

Define

• "In" wavefunctions:

$$\psi^{in} \xrightarrow{r=|\vec{x}|\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + f(p, \hat{p}\cdot\hat{x})$$

• "Out" - " " - " "

$$\psi^{out} \xrightarrow{r\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + \tilde{f}(p, \hat{p}\cdot\hat{x})$$

$$S_{\vec{p}_{in}, \vec{p}_{out}} = \langle \psi^{out} | \psi^{in} \rangle$$

What if we have a Coulomb potential?

$$\left(-\frac{\nabla^2}{2m} + \frac{Ze^2}{4\pi r} \right) \psi(\vec{x}) = E \psi(\vec{x})$$

$$\Rightarrow \left(\nabla^2 + 2mE - \frac{2Z}{r} \right) \psi(\vec{x}) = 0$$

Define

• "In" wavefunctions:

$$\psi^{in} \xrightarrow{r=|\vec{x}| \rightarrow \infty} e^{i\vec{p} \cdot \vec{x}} + f(p, \hat{p} \cdot \hat{x})$$

• "Out" - " " -

$$\psi^{out} \xrightarrow{r \rightarrow \infty} e^{i\vec{p} \cdot \vec{x}} + \tilde{f}(p, \hat{p} \cdot \hat{x})$$

Define

"In" wavefunctions:

$$\psi_{in} \xrightarrow{r=|\vec{x}|\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + f(p, \hat{p}\cdot\hat{x}) \frac{e^{ipr}}{r}$$

"Out" - " " - " "

$$\psi_{out} \xrightarrow{r\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + \tilde{f}(p, \hat{p}\cdot\hat{x}) \frac{e^{-ipr}}{r}$$

Asymptotically:

$$\psi(\vec{x}) \sim \frac{e^{\pm ipr + g(r)}}{r}$$

Schr eq $\Rightarrow \frac{dg}{dr} + \left(\frac{dg}{dr}\right)^2 + 2ip \frac{dg}{dr} = \frac{2\epsilon}{r}$

$$\Rightarrow g(r) \sim \mp i \frac{\pi}{p} \log(pr)$$

$$\psi(\vec{x}) \sim \frac{e^{\pm ipr \mp i \frac{\pi}{p} \log(pr)}}{r}$$

[Full solution to the flat space helicity]

Full solution.

$$\psi^{in}(\vec{x}) = e^{i\vec{p}\cdot\vec{x}} e^{-\frac{\pi}{2} \frac{x}{p}} \Gamma\left(1 + \frac{i\delta}{p}\right) F_1\left(-i\delta, 1, i(p\vec{r} - \vec{p}\cdot\vec{x})\right)$$

Define:

• "In" wavefunctions:

$$\psi^{in} \xrightarrow{r=|\vec{x}|\rightarrow\infty} e^{i\vec{p}\cdot\vec{x} + i\frac{\delta}{p}\log(p)} + f(p, \hat{p}\cdot\hat{x}) \frac{e^{ipr - \frac{\delta}{p}\log(p)}}{r}$$

• "Out" - " " - " "

$$\psi^{out} \xrightarrow{r\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + \tilde{f}(p, \hat{p}\cdot\hat{x}) \frac{e^{-ipr}}{r}$$

Scattering in 2D (logarithm) - Effect of Dimensionality

Define

"In" wavefunctions:

$$\psi_{in} \xrightarrow{r=|\vec{x}|\rightarrow\infty} e^{i\vec{p}\cdot\vec{x} + i\frac{\pi}{2}\log(p)} + f(p, \hat{p}\cdot\hat{x}) \frac{e^{ipr}}{r}$$

"Out" - " - "

$$\psi_{out} \xrightarrow{r\rightarrow\infty} e^{i\vec{p}\cdot\vec{x}} + f(p, \hat{p}\cdot\hat{x}) \frac{e^{-ipr}}{r}$$

$$\langle \psi^{out} | \psi^{in} \rangle = -\frac{\chi}{p^2} \frac{\Gamma(1 + \frac{i\chi}{p})}{\Gamma(1 - \frac{i\chi}{p})} \frac{1}{2\sin^2 \frac{\theta}{2}} \times e^{-i\frac{\chi}{p} \log(\sin^2 \frac{\theta}{2})}$$

Key point:

We can envision two types of experiments

probe long-range effects

agnostic to long-range effects
e.g. total cross section $\propto |f|$

⊕ What is



Plan physics and holography

AdS/CFT duality

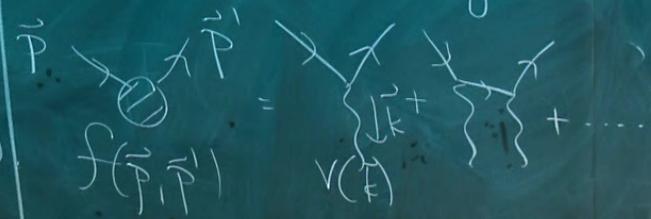
Time delay:

$$0 = \frac{\partial}{\partial p} \left[i p r - \frac{i x}{p} \log [p r (1 - \cos \theta) - i \frac{p^2}{2m} t] \right]$$

$$\Rightarrow t = \frac{m}{p} r + \frac{m x}{p^3} \log [p r (1 - \cos \theta)]$$

What if we don't have the exact solution?

Perturbation theory:



$$V(\vec{k}) = \frac{4\pi g^2/m}{k^2}$$

$$G_T(p, k) = \frac{1}{p^2 - k^2 + i\epsilon}$$

Define:

"In" wavefunction

$$\psi_{in} \quad r = |\vec{x}| \rightarrow \infty$$

"Out" wavefunction

$$\psi_{out} \quad r \rightarrow \infty$$

$$\langle \psi_{out} | \psi_{in} \rangle$$

Born series

$$f^{LO} \propto \frac{-2\gamma}{(\vec{p} - \vec{p}')^2} = \frac{-\gamma}{p^2(1 - \cos\theta)}$$

$$f^{NLO} \propto \int \frac{d^3k}{(2\pi)^3} \frac{1}{[\vec{k} - \vec{p}]^2 [\vec{p} - \vec{k} + i\epsilon]^2 [\vec{p}' - \vec{k}]^2}$$

$\vec{k} \sim \vec{p} \quad = \infty$

$$\psi(\vec{x}) \sim e^{i\vec{p} \cdot \vec{x} + \frac{i\gamma}{p} \log(pr)} + \hat{f}_p \frac{e^{i\vec{p}' \cdot \vec{x} + \frac{i\gamma}{p'} \log(p'r)}}{r}$$

If we blindly expand in γ

$$f(p, \cos\theta) \approx \frac{-\gamma}{p^2(1 - \cos\theta)}$$

$$+ \frac{i\gamma^2}{p^3(1 - \cos\theta)} \left(\dots + \log(pr) \right)$$

⊕ What is a so



Class physics and holography

AdS/CFT correspondence

We have to regulate!

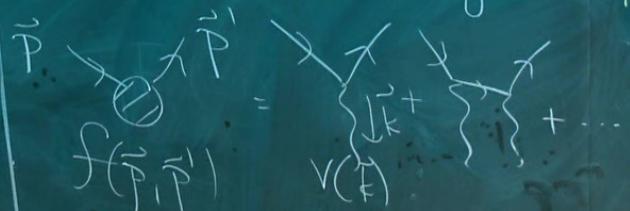
0) $\frac{i\delta^2}{p^2(1-\cos\theta)} \log(pr)$ as $r \rightarrow \infty$

1)

2)

What if we don't have the exact solution?

Perturbation theory: $V(r) \propto \frac{\gamma}{r} e^{-mr}$



$V(\vec{k}) = \frac{4\pi\gamma/m}{k^2 + m^2}$

$G_T(p, k) = \frac{1}{p^2 - k^2 + i\epsilon}$

Born series

$f^{LO} \propto -2\gamma$

$f^{NLO} \propto$

to start [cosmology]

Effect of [uncertainty]

Born series

$(r) \propto \frac{\gamma}{r} e^{-mr}$

$f^{LO} \propto \frac{-2\gamma}{(\vec{p} - \vec{p}')^2} = \frac{-\gamma}{p^2(1-\cos\theta)}$

$f^{NLO} \propto \int \frac{d^3k}{(2\pi)^3} \frac{1}{[(\vec{k} - \vec{p})^2][p^2 - k^2 + i\epsilon][(\vec{p} - \vec{k})^2]}$
 $\vec{k} \sim \vec{p} = \infty$

$\psi(\vec{x}) \sim e^{i\vec{p}\vec{x} + \frac{i\gamma}{p} \log(pr)}$
 $+ \hat{f}(p) \frac{e^{i\vec{p}\vec{x} + \frac{i\gamma}{p} \log(pr)}}{r}$

If we blindly expand in γ

$f(p, \cos\theta) \stackrel{?}{=} \frac{-\gamma}{p^2(1-\cos\theta)}$
 $+ \frac{i\gamma^2}{p^3(1-\cos\theta)} (\dots + \log(pr))$
 $f^{NLO} + O(\gamma^2)$

Ⓡ (What is



$t \rightarrow -\infty$
 $|\vec{x}| \rightarrow \infty$

collan physics and holography
 collan field theory

What if we don't have the exact solution?

We have to regulate!

0) $f^{NLO} = \frac{i\gamma^2}{p^2(1-\cos\theta)} \log(pr)$ as $r \rightarrow \infty$

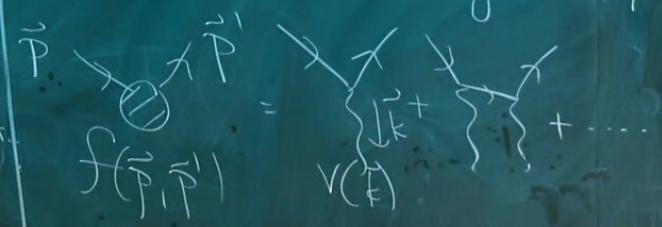
1) Yukawa/mass $f^{NLO} = \frac{i\gamma^2}{p^2(1-\cos\theta)} \log\left(\frac{p}{m}\right)$

2) Dimensional regularization

$f^{NLO} = \frac{i\gamma^2}{p^2(1-\cos\theta)} \left(\frac{1}{\epsilon}\right)$

What if we don't have the exact solution?

Perturbation theory: $V(r) \propto \frac{\gamma}{r} e^{-mr}$



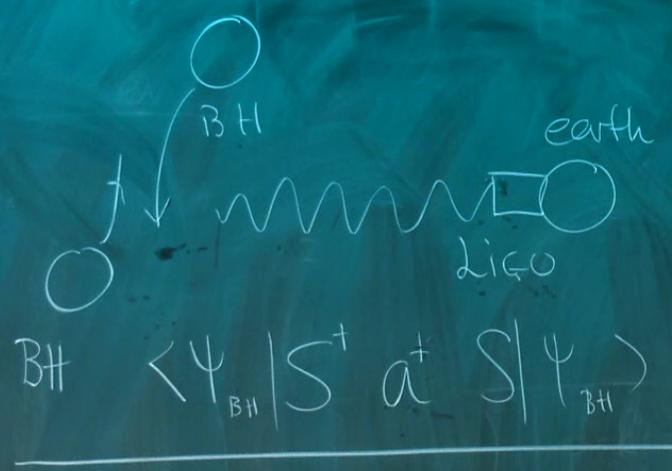
$V(\vec{k}) = \frac{4\pi\gamma/m}{k^2 + m^2}$
 $G(p, k) = \frac{1}{p^2 - k^2 + i\epsilon}$

Born se

$f^{LO} \propto$
 $f^{NLO} \propto$

late part of cosmology / Effects of gravitation

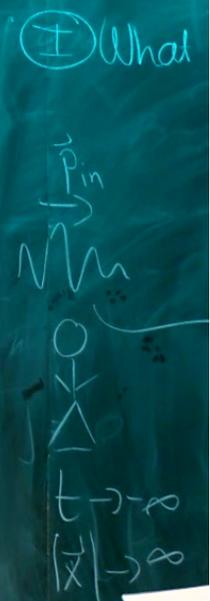
$$V(r) \propto \frac{1}{r} e^{-mr}$$



Two sources of Shapiro time delay:

- ① Graviton escaping the BH potential

$$\Delta t_1 = \frac{1}{c} \log \left(\frac{r_{obs}}{b} \right)$$



don't have the
equation?

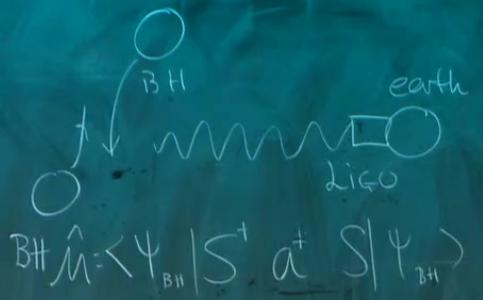
theory: $V(r) \propto \frac{1}{r} e^{-mr}$

$\sum_k \dots$
 $V(r)$

$$k) = \frac{4\pi \gamma m}{K^2 + m^2}$$

$$P(k) = \frac{1}{p^2 k^2 + \epsilon}$$

$\Delta H + V \Delta X$
Lagrange



$$BH \hat{M} = \langle \Psi_{BH} | S^{\dagger} a^{\dagger} S | \Psi_{BH} \rangle$$

$$\hat{M}_{IR} = M^{tree} \exp \left[(\#_1 + \#_2) \frac{1}{\epsilon_{IR}} \right]$$

Two sources of
Shapiro time delay:

- ① Graviton escaping the BH potential
 $\Delta t_1 = \#_1 \log \left(\frac{r_{obs}}{b} \right)$
- ② Delay/advance from approaching each other
 $\Delta t_2 = \#_2 \log \left(\frac{r_{in}}{b} \right)$

