

Title: Advanced General Relativity - 240327 (afternoon)

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Collection: Advanced General Relativity (PHYS7840)

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## Kerr event horizon

Intrinsic coordinates  $y^a = (v, \theta, \chi)$

Embedding relations:  $X^\alpha = X^\alpha(y^a)$  :

$$V = v$$

$$r = r_+$$

$$\theta = \theta$$

$$\psi = \Omega_+ v + \chi$$

Tangent vectors:

$$K^\alpha = \frac{\partial X^\alpha}{\partial v} = (1, 0, 0, \Omega_+)$$

$$e_\theta^\alpha = \frac{\partial X^\alpha}{\partial \theta} = (0, 1, 0, 0)$$

$$e_\chi^\alpha = \frac{\partial X^\alpha}{\partial \chi} = (0, 0, 0, 1)$$

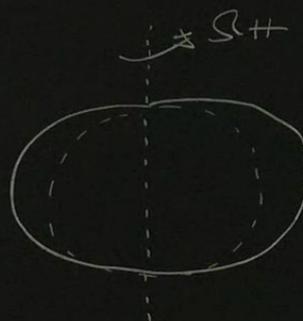
$$\Omega_+ = \frac{a}{r_+^2 + a^2}$$

$$r_+ = \frac{r_+ - M}{r_+^2 + a^2}$$

$$r_+ = M + \sqrt{M^2 - a^2}$$

Induced metric:  $\Omega_{AB} = g_{\alpha\beta} e^{\alpha} e^{\beta}$

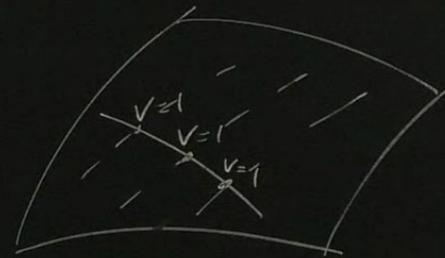
$$\begin{cases} \Omega_{\theta\theta} = r_+^2 + a^2 \cos^2 \theta \\ \Omega_{\phi\phi} = \frac{(r_+^2 + a^2)^2}{r_+^2 + a^2 \cos^2 \theta} \sin^2 \theta \end{cases}$$



$$\sqrt{\Omega} = (r_+^2 + a^2) \sin \theta$$

Cross-sectional area:

$$\begin{aligned} A &= \int \sqrt{\Omega} d^2 \theta \\ &= (r_+^2 + a^2) \int \sin \theta d\theta d\phi \\ &= 4\pi (r_+^2 + a^2) \\ &= 8\pi M r_+ \end{aligned}$$



$\partial x$ 

$$r_+ = M + \sqrt{M^2 - a^2}$$

$$= (r_+ + a)$$

 $4\pi$  $= 8\pi$ 

$$A = 8\pi M \left( M + \sqrt{M^2 - (J/M)^2} \right)$$

$$A = A(M, J)$$

$$\text{BH \# 1 : } (M, J)$$

$$\text{BH \# 2 : } (M + \delta M, J + \delta J)$$

Differential law

$$\delta A = \left( \frac{\partial A}{\partial M} \right) \delta M + \left( \frac{\partial A}{\partial J} \right) \delta J$$

$$\frac{\kappa}{8\pi} \delta A = \delta M - \Omega_H \delta J$$

$$\rightarrow \delta M = \frac{\kappa}{8\pi} \delta A + \Omega_H \delta J$$

$$4\pi = 8\pi r^2 +$$

$$\sqrt{M^2 - (J/M)^2}$$

$$A = A(M, J)$$

BH #1 : (M, J)

BH #2 ; (M + \delta M, J + \delta J)

$$\delta A = \left( \frac{\partial A}{\partial M} \right) \delta M + \left( \frac{\partial A}{\partial J} \right) \delta J$$

$$\delta M - \Omega_H \delta J$$

$$\delta M = \frac{\kappa}{8\pi} \delta A + \Omega_H \delta J$$

$$M(A, J)$$

$$\frac{\kappa}{8\pi} = \left( \frac{\partial M}{\partial A} \right)_J$$

$$\Omega_H = \left( \frac{\partial M}{\partial J} \right)_A$$

$$+ \sqrt{M^2 - (J/M)^2}$$

$$A = A(M, J)$$

$$\text{BH \# 1} = (M, J)$$

$$\text{BH \# 2} = (M + \delta M, J + \delta J)$$

$$\delta A = \left( \frac{\partial A}{\partial M} \right) \delta M + \left( \frac{\partial A}{\partial J} \right) \delta J$$

$$\delta M - \Omega_H \delta J$$

$$\rightarrow \delta M = \frac{\kappa}{8\pi} \delta A + \Omega_H \delta J$$

"TSS"

$$M(A, J)$$

$$\frac{\kappa}{8\pi} = \left( \frac{\partial M}{\partial A} \right) \Big|_J$$

$$\Omega_H = \left( \frac{\partial M}{\partial J} \right) \Big|_A$$

## Integral law

Euler thm on homogeneous functions:

$$f(x^a) : f(\lambda x^a) = \lambda^K f(x)$$
$$\Rightarrow Kf = \sum_a x^a \frac{\partial f}{\partial x^a}$$

$$[M] = L^2$$
$$[A] = L^2$$
$$[J] = L^2$$

$$M = M(A, J)$$

$$M(\lambda A, \lambda J) = \lambda^{1/2} M(A, J) \rightarrow M \text{ is homo. func. of deg}$$

$$\text{Euler: } \frac{1}{2} M = A \frac{\partial M}{\partial A} + J \frac{\partial M}{\partial J}$$

## Integral law

Euler thm on homogeneous functions:

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$$M(\lambda A, \lambda J) = \lambda^{1/2} M(A, J) \rightarrow$$

M is homo. func. of deg

$$\text{Euler: } \frac{1}{2}M = A \frac{\partial M}{\partial A} + J \frac{\partial M}{\partial J} = \frac{K}{8\pi} A + \Omega$$

$$M = \frac{K}{4\pi} A + 2\Omega$$

homogeneous functions:

$$f(x^a) : f(\lambda x^a) = \lambda^K f(x^a)$$

homogeneous function of degree  $K$ .

$$\Rightarrow Kf = \sum_a x^a \frac{\partial f}{\partial x^a}$$

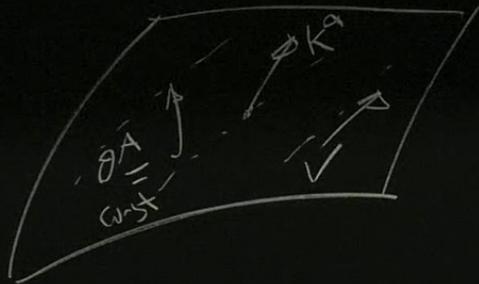
$$A, J \quad (A, \lambda J) = \lambda^{1/2} M(A, J) \rightarrow M \text{ is homo. func. of degree } \frac{1}{2}.$$

$$\frac{1}{2}M = A \frac{\partial M}{\partial A} + J \frac{\partial M}{\partial J} = \frac{K}{8\pi} A + 2\Omega_H J$$

$$M = \frac{K}{4\pi} A + 2\Omega_H J$$

(Smarr's formula)

Null hypersurfaces



$$y^a = (v, \theta^A)$$

$$k^\alpha = \frac{\partial X^\alpha}{\partial v}, \quad e_A^\alpha = \frac{\partial X^\alpha}{\partial \theta^A}, \quad N_\alpha$$

$$V^A = (V, \theta^A)$$

$$K^\alpha = \frac{\partial X^\alpha}{\partial V}, \quad e_A^\alpha = \frac{\partial X^\alpha}{\partial \theta^A} \quad , \quad N_\alpha$$

$$\Omega_{AB} = \partial_\alpha p e_A^\alpha e_B^\alpha$$

$$K^\rho \nabla_\rho K^\alpha = \kappa K^\alpha$$

$$K^\rho \nabla_\rho e_A^\alpha = \omega_A K^\alpha + B_A^B e_B^\alpha$$

$$K_\alpha K^\alpha = 0$$

$$K_\alpha e_A^\alpha = 0$$

$$N_\alpha N^\alpha = 0, \quad N_\alpha K^\alpha = -1, \quad N_\alpha e_A^\alpha = 0$$

$$B_{AB} = e_A^\alpha e_B^\beta \nabla_\alpha K_\beta = \frac{1}{2} \textcircled{H} \Omega_{AB} + \mathcal{T}_{AB}$$

$$= \partial_V \Omega_{AB}$$

$$\textcircled{H} = \frac{1}{\sqrt{\Omega}} \partial_V \sqrt{\Omega}$$

$$\frac{\delta A}{\delta V} = \int_S \textcircled{H} \delta S$$

$$\delta S = \sqrt{\Omega} \delta^2 \Theta$$

Integration :

$$d\Sigma_\alpha = -K_\alpha \partial V dS$$

$$dS_{\text{opt}} = 2 K[\alpha N/\beta] dS$$

Gauss-Codazzi eqns :

$$R_{\mu\nu} K^\mu K^\nu = -\partial_\nu \mathbb{H} + \mathcal{L}(\mathbb{H}) - \frac{1}{2} \mathbb{H}^2 - \sigma_{AB} \sigma^{AB} \quad (1)$$

$$R_{\mu\nu} K^\mu e_A^\nu = \partial_\nu W_A - \mathcal{D}_A K$$

$\partial x$

$$r_+ = M + \sqrt{M^2 - a^2}$$

$$= (r_+ + a)$$

$4\pi$

$$= 8\pi M$$

Integration:

$$\partial \bar{\Sigma}_\alpha = -K_\alpha \partial V \partial S$$

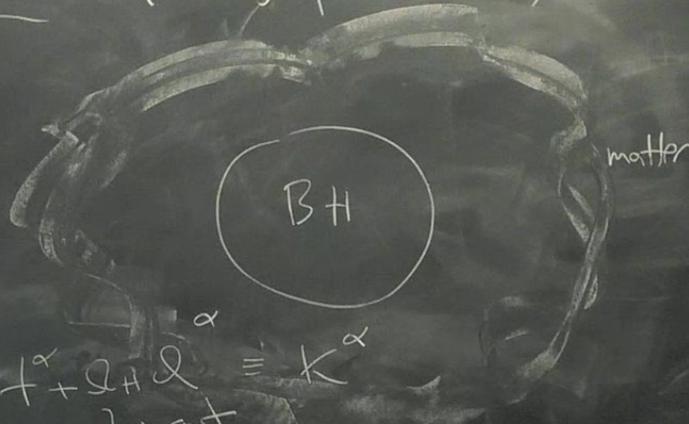
$$\partial S_{\text{app}} = 2 K_{[\alpha} N_{\beta]} \partial S$$

Mass-Codazzi eqns:

$$\begin{cases} R_{\mu\nu} K^\mu K^\nu = -\partial_V \Theta + K \Theta - \frac{1}{2} \Theta^2 - \sigma_{AB} \sigma^{AB} & (\text{Raychaudhuri}) \\ R_{\mu\nu} K^\mu e_A^\nu = \partial_V W_A - \partial_A K - \frac{1}{2} \partial_A \Theta + D_B \sigma_A^B + \Theta W_A \end{cases}$$

Stationary BH (nothing depends on time)

NOT KERR!



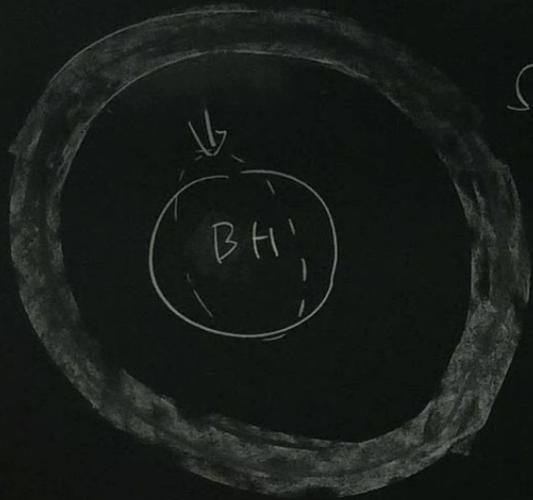
Hawking (1972)

— must be either static  
( $t^\alpha$  hypersurface orthogonal  
 $\rightarrow \omega_{\alpha\beta} = \nabla_{[\alpha} t_{\beta]} = 0$ )

or axisymmetric

$t^\alpha$  not hypersurface orthogonal  
 $\phi^\alpha$  axial killing vector.

Why stationary  $\rightarrow$  axisymmetry?



$$\Omega_H = \text{const}$$



Tidal forces  $\rightarrow \Omega \downarrow$   
 $\hookrightarrow$  torques

$$\rightarrow K^\alpha = t^\alpha + \Omega_H \ell^\alpha = \text{null on event horizon.}$$

$\hookrightarrow$  tangent to null generators ;

$$K^\beta \nabla_\beta K^\alpha = \kappa K^\alpha$$

Stationary  $\rightarrow$  axisymmetry  $\rightarrow K^\alpha = t^\alpha + \Omega_H \partial^\alpha$   
 $\hookrightarrow \partial_V(\text{anything}) = 0$   
 $\hookrightarrow$  tangent to null g

$$\boxed{\textcircled{H} = 0}$$

$$\boxed{\partial_V \textcircled{H} = 0}$$

Raychaudhuri =

$$\sigma_{AB} \sigma^{AB} + R_{\alpha\beta} K^\alpha K^\beta = 0$$

→ tangent to null generators ;

$$K^\alpha v^\beta K^\alpha K^\beta$$

0

$$\textcircled{H} = 0$$

Null energy condition :

$$T_{\alpha\beta} K^\alpha K^\beta \geq 0$$

$$\rightarrow R_{\alpha\beta} K^\alpha K^\beta \geq 0$$

$$\underbrace{\sigma_{AB}}_{\neq 0} + \underbrace{R_{\alpha\beta} K^\alpha K^\beta}_{\neq 0} = 0$$

$$\Rightarrow \begin{cases} \sigma_{AB} = 0 \\ R_{\alpha\beta} K^\alpha K^\beta = 0 = T_{\alpha\beta} K^\alpha K^\beta \end{cases}$$

= WA K ...

$$\textcircled{H} = \frac{1}{\sqrt{\Omega}} \omega \sqrt{\Omega}$$

$$\frac{\partial \textcircled{H}}{\partial V} = \dots$$
  
$$\delta S = \sqrt{\Omega} \delta^2 \Theta$$

$+ R_{\alpha\beta} K^\alpha K^\beta = 0$

$$\Rightarrow \begin{cases} \sigma_{AB} = 0 \\ R_{\alpha\beta} K^\alpha K^\beta = 0 = T_{\alpha\beta} K^\alpha K^\beta \end{cases}$$

$\frac{K}{\sqrt{\dots}}, e_A = \frac{U_X}{\partial X^A} \rightarrow N_\alpha \quad N_\alpha N^\alpha = 0, N_\alpha K^\alpha = -1, N_\alpha e_A^\alpha = 0$

$e_A^\alpha e_B^\beta$

$$B_{AB} = e_A^\alpha e_B^\beta \nabla_\alpha K_\beta = \frac{1}{2} \textcircled{+} \Omega_{AB} + \check{\sigma}_{AB}$$

$$= \omega \Omega_{AB}$$

$A K^\alpha + B_A^B e_A^\alpha$

$$\textcircled{+} = \frac{1}{\sqrt{\Omega}} \omega \sqrt{\Omega}$$

$$\frac{\delta A}{\delta V} = \int_S \textcircled{+} \delta S$$

$$\delta S = \sqrt{\Omega} \delta^2 \Theta$$