

Title: Advanced General Relativity - 240306 (afternoon)

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Collection: Advanced General Relativity (PHYS7840)

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URL: <https://pirsa.org/24030004>

Gauss-Codazzi

$$4R_{\mu\nu\rho\sigma} n^\mu e_a^\nu e_b^\rho e_c^\sigma = D_c K_{ab} - D_b K_{ac}$$

$$4R_{\alpha\rho\sigma\delta} e_a^\alpha e_b^\rho e_c^\sigma e_d^\delta = R_{abcd} - \varepsilon(K_{ac}K_{bd} - K_{ad}K_{bc})$$

$$4G_{\mu\nu} n^\mu n^\nu = -\frac{1}{2}(\varepsilon R + K_{ab}K^{ab} - K^2)$$

$$4G_{\mu\alpha} n^\mu e_a^\alpha = D_b K_a^b - D_a K$$

$$K_{ab} = e_a^\alpha e_b^\beta \nabla_\alpha n_\beta$$

$$= e_a^\alpha e_b^\beta \nabla_\beta n_\alpha$$

$$= \frac{1}{2} e_a^\alpha e_b^\beta (\nabla_\alpha n_\beta + \nabla_\beta n_\alpha)$$

$$= \frac{1}{2} e_a^\alpha e_b^\beta \mathcal{L}_n g_{\alpha\beta}$$

Painlevé-Gullstrand's version of Schwarzschild

$$= -f dt^2 + f^{-1} dr^2 + r^2 d\Omega^2$$

$$= 1 - 2M/r$$

$$\text{surface: } \Phi = t + \int \frac{\sqrt{2M/r}}{f} dr = t + 4M \left\{ \sqrt{r/2M} + \frac{1}{2} \ln \frac{\sqrt{r/2M} - 1}{\sqrt{r/2M} + 1} \right\}$$

$$-\nabla_\alpha \Phi = -\left(1, \frac{\sqrt{2M/r}}{f}, 0, 0\right) \quad g^{\alpha\beta} n_\alpha n_\beta = -1$$

$$\text{coordinates: } \gamma^\alpha = (t, r, \theta, \varphi)$$

$$e_a^\alpha = \partial x^\alpha / \partial y^a$$

$$\text{relations: } t = \text{const} - \int \frac{\sqrt{2M/r}}{f} dr$$

$$e_r^\alpha =$$

$$r = r$$

$$\theta = \theta$$

$$\varphi = \varphi$$

wurzchild

$$\frac{\sqrt{2mr}}{f} dr = + + 4M \left\{ \sqrt{r/2m} + \frac{1}{2} \ln \frac{\sqrt{r/2m} - 1}{\sqrt{r/2m} + 1} \right\}$$

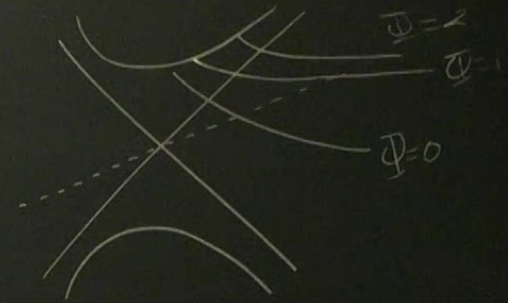
$$\int g^{\alpha\beta} n_{\alpha} n_{\beta} = -1$$

$$e_a^{\alpha} = \partial x^{\alpha} / \partial y^a$$

$$e_r^{\alpha} = \left(-\frac{\sqrt{2mr}}{f}, 1, 0, 0 \right)$$

$$e_{\theta}^{\alpha} = (0, 0, 1, 0)$$

$$e_{\phi}^{\alpha} = (0, 0, 0, 1)$$



$$\begin{aligned} r &= r \\ \theta &= \theta \\ \varphi &= \varphi \end{aligned}$$

$$\begin{aligned} e_{\theta}^{\alpha} &= (0, 1, 0, 1, 0) \\ e_{\varphi}^{\alpha} &= (0, 0, 1, 0, 1) \end{aligned}$$

$$h_{ab} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & r^2 & 0 \\ 0 & 0 & r^2 \sin^2 \theta \end{pmatrix}$$

$$h_{ab} dy^a dy^b = dr^2 + r^2 d\Omega^2 = \text{flat!}$$

$$K_{ab} = \begin{pmatrix} \frac{1}{2} \sqrt{2M/r^3} & 0 & 0 \\ 0 & -\sqrt{2M/r} & 0 \\ 0 & 0 & -\sqrt{2M/r} \sin^2 \theta \end{pmatrix}$$

$$K = -\frac{3}{2} \sqrt{2M/r^3}$$

$$= \frac{1}{2} \vec{e}_a \vec{e}_b \chi_n \text{ etc}$$

GAUSS

2D surfaces in 3D space (flat)



$$\frac{1}{2} R = k_1 k_2 = \frac{1}{r_1 r_2}$$

$$R_{abcd} = \underbrace{k_{ac} k_{bd}}_{\text{intrinsic}} - \underbrace{k_{ad} k_{bc}}_{\text{extrinsic}}$$

↓
R

↓
eigenvalues (k_{ab}) = k₁, k₂

$$k_i = \frac{1}{r_i}, \quad i=1, 2$$

↓
eigenvalues (h_{ab}) = K_1, K_2

$$A = \int \sqrt{h} d^2y$$

$$\delta A = \int K \left(\frac{1}{2} \right) \delta S$$

$$\text{minimal surface: } \delta A = 0 \\ \Rightarrow K = 0$$

INTRINSIC TENSORS

embedding relations

$$h_{ab} =$$

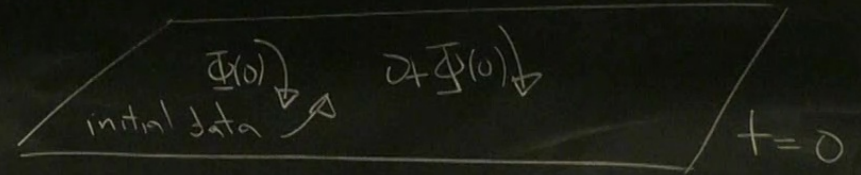
$$K_{ab} =$$

Initial-value problem

flat 4D spacetime,

$$\square \Phi = 0$$

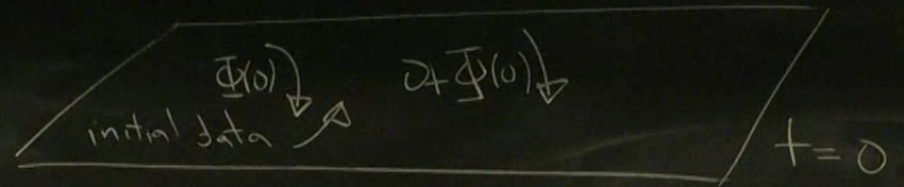
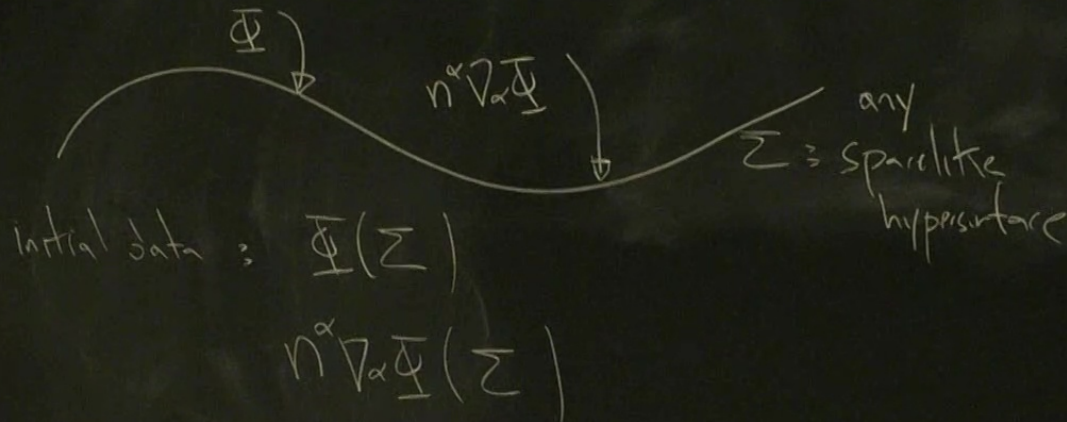
curved spacetime =



Initial-value problem

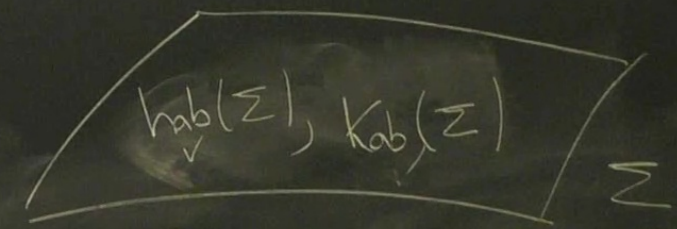
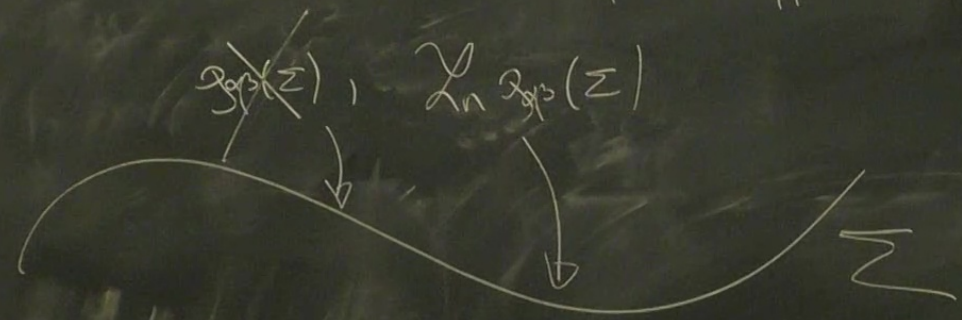
flat 4D spacetime,
 $\square \Phi = 0$

curved spacetime:



GR

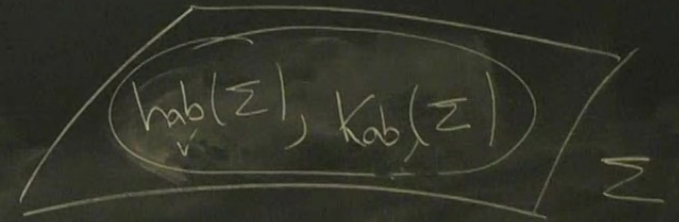
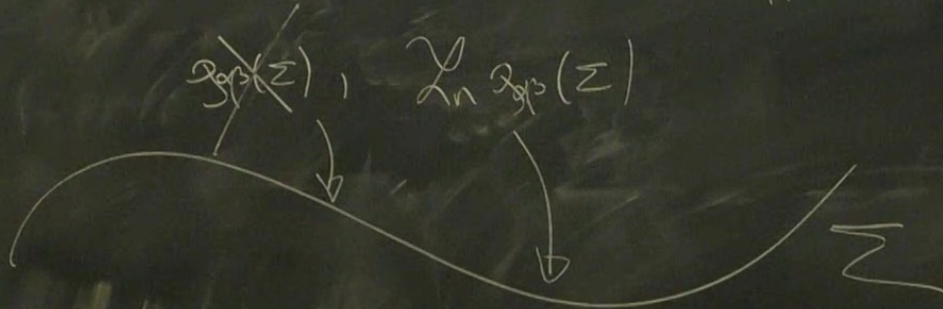
initial moment \equiv any spacelike hypersurface



EFE : $\partial_t h_{ab} = \dots$
 $\partial_t K_{ab} = \dots$

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initial moment \equiv any spacelike hypersurface

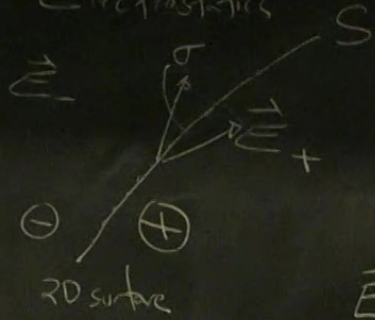


EFE : $\partial_t h_{ab} = \dots$
 $\partial_t K_{ab} = \dots$

Constraints : $R - K_{ab}K^{ab} + K^2 = 8\pi T_{\mu\nu} n^\mu n^\nu \equiv$ energy density

$D_b K^b_a - D_a K = 4\pi T_{\mu\nu} n^\mu e^{\nu}_a \equiv$ matter current

Electrostatics



$$L = \dots = \dots \ominus - \dots \ominus$$

$$[\vec{E} \cdot \hat{n}] = \sigma / \epsilon_0$$

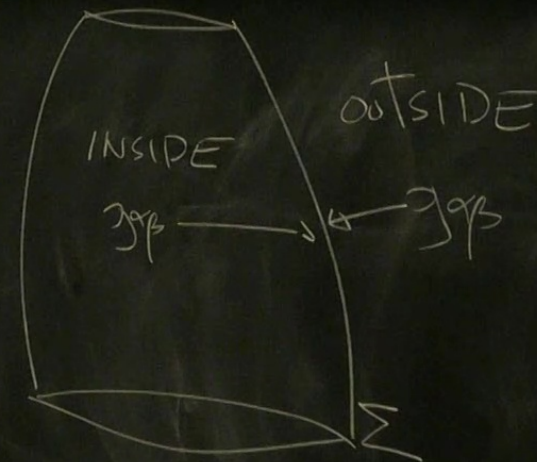
$$[\vec{E} \cdot \vec{e}] = 0$$

$$\vec{E} = -\vec{\nabla}V$$

$$[V] = 0$$
$$[\partial_n V] = -\sigma / \epsilon_0$$

Junction conditions

gravitational collapse



$$L = \int_V (\mathbf{v} \cdot \mathbf{h}) - \int_V \rho \phi$$

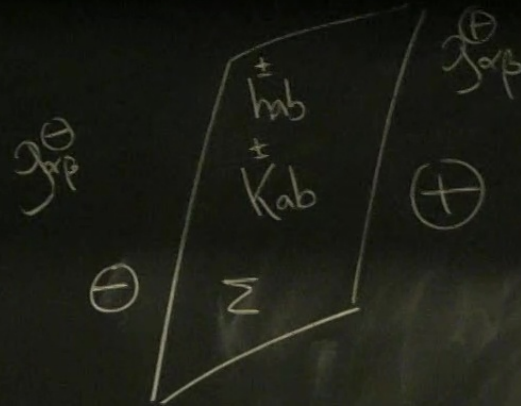
$$[\mathbf{E} \cdot \hat{n}] = \sigma / \epsilon_0$$

$$[\mathbf{E} \cdot \hat{e}] = 0$$

$$\mathbf{E} = -\nabla V$$

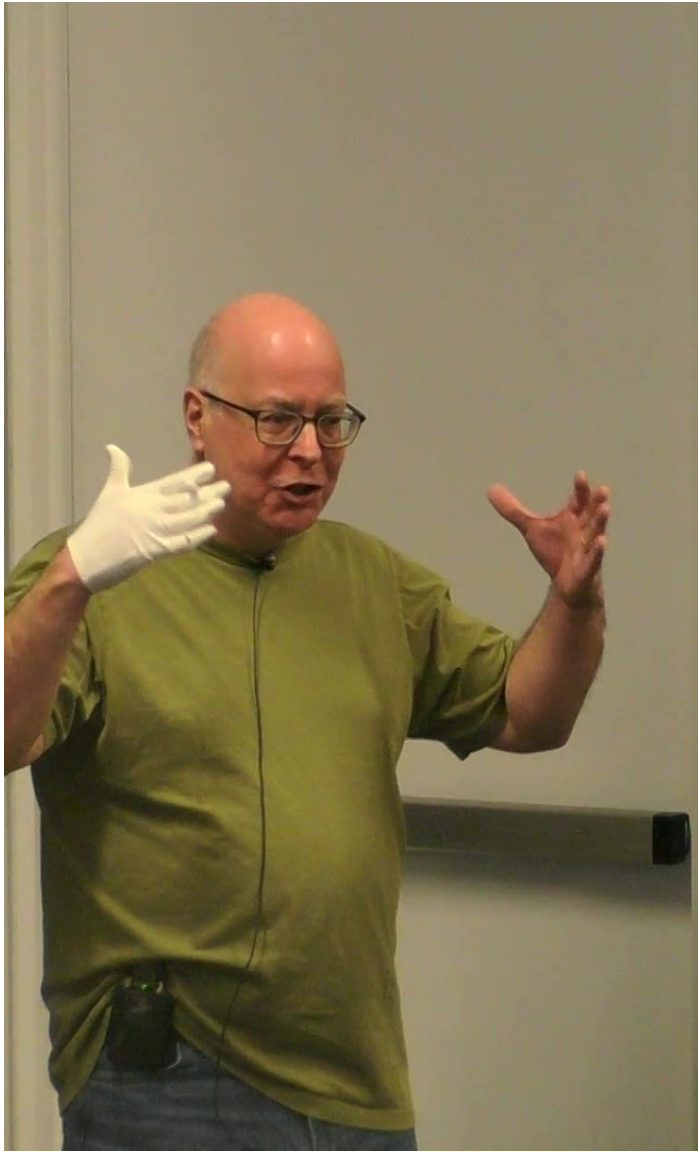
$$[V] = 0$$

$$[\partial_n V] = -\sigma / \epsilon_0$$



$$[h_{ab}] = 0$$

$$[K_{ab}] = -\sigma / \epsilon_0$$



Junction conditions
gravitational collapse

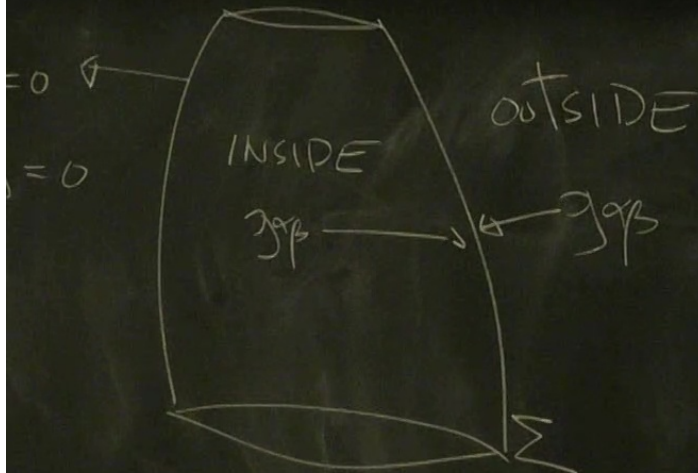
$[h_{ab}] = 0$ ←
 $[K_{ab}] = 0$

INSIDE OUTSIDE

$\int \rho_B$ ← $\int \rho_B$

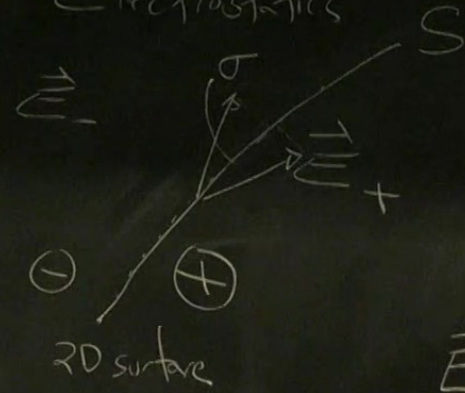
metric conditions

gravitational collapse



SCHWARZ

Electrostatics



$$[E_{\parallel}] = \dots$$

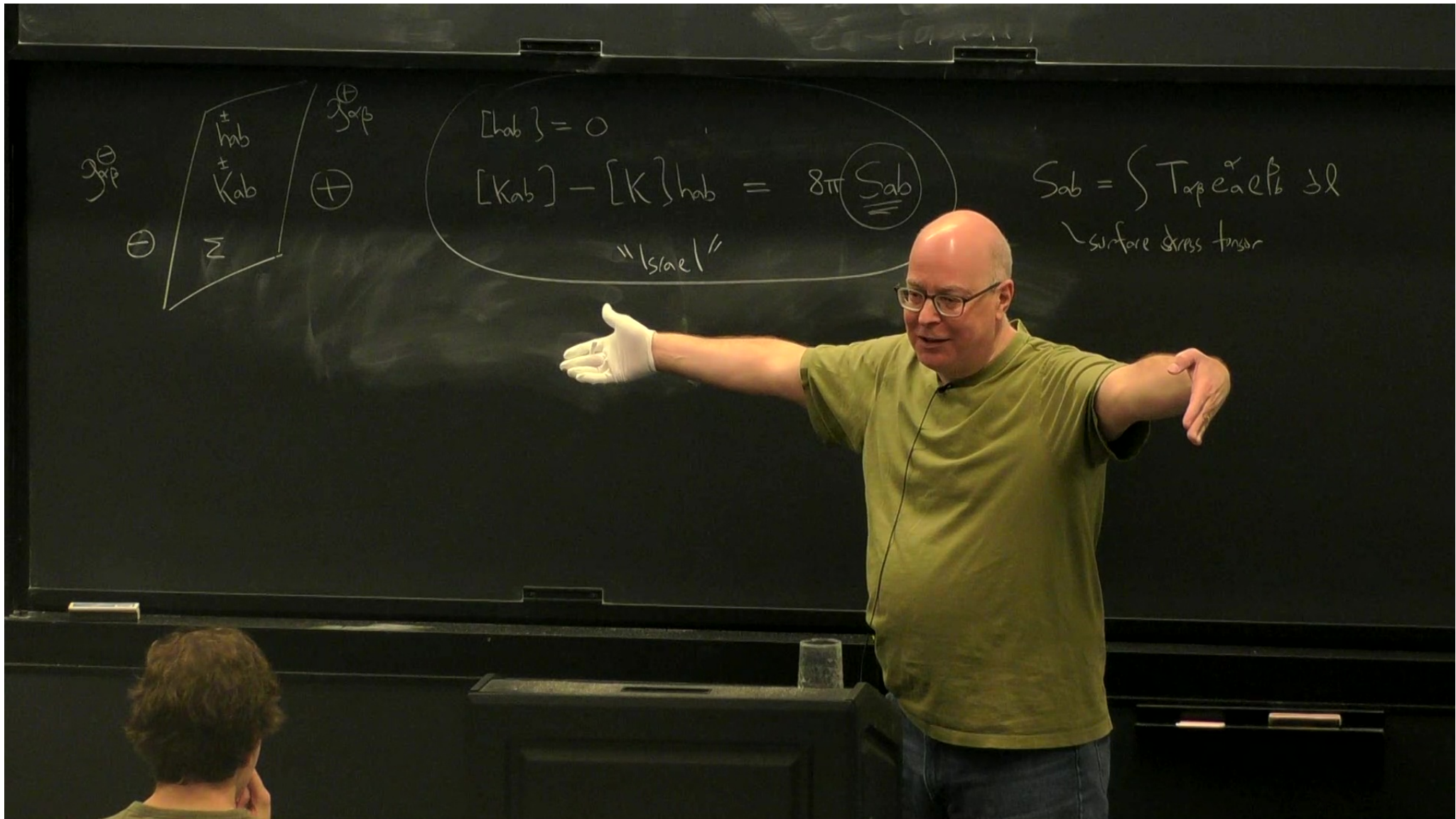
$$[\vec{E} \cdot \hat{n}] = 0$$

$$[\vec{E} \cdot \vec{e}] = 0$$

$$\vec{E} = -\vec{\nabla}V$$

$$[V] = 0$$

$$[\partial_n V] = -\sigma/\epsilon_0$$



$$[h_{ab}] = 0$$

$$[K_{ab}] - [K]h_{ab} = 8\pi \underline{S_{ab}}$$

"Israel"

$$S_{ab} = \int T_{\alpha\beta} e^{\alpha} e^{\beta} \wedge \alpha$$

surface stress tensor