

Title: Cosmology Lecture

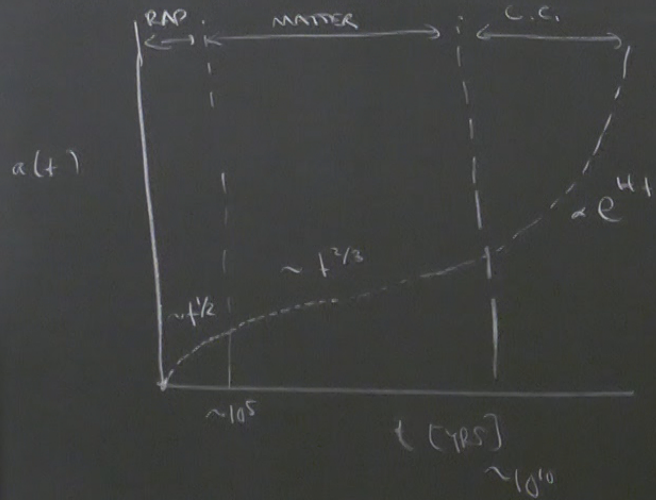
Speakers: Kendrick Smith

Collection: Cosmology 2023/24

Date: February 29, 2024 - 1:00 PM

URL: <https://pirsa.org/24020036>

$$ds^2 = -dt^2 + a(t)^2 dx^2$$



$$H^2 = \frac{8\pi G}{3} \rho$$

$$H \equiv \frac{d \log a}{dt}$$

$$\dot{\rho} = -3H(\rho + p)$$

$$T_{\mu\nu} = \begin{pmatrix} \rho(t) & & & \\ & & & \\ & & & \\ & & & \end{pmatrix}$$

$$a(t)^2 p(t) \delta_{ij}$$

$$\Leftrightarrow \frac{d\rho}{d \log a} = -3(\rho + p)$$

CONSTANT - w

$$w = \frac{p(t)}{\rho(t)} = \text{const} \quad \text{"EQUATION OF STATE"}$$

$$= \begin{cases} 1/3 & \text{RADIATION} \\ 0 & \text{MATTER} \\ -1 & \text{COSMOLOGICAL CONSTANT} \end{cases}$$

$$\frac{d\rho}{d \log a} = -3(\rho + p) = -3(1+w)\rho$$

$$\rho = \rho_0 a^{-3(1+w)}$$

$(\rho)_0 = \text{"TODAY"}$ (AT $a=1$)

$$= \begin{cases} a^{-3} & \text{MATTER } [w=0] \end{cases}$$

$$= \begin{cases} a^{-4} & \text{RADIATION } [w=\frac{1}{3}] \end{cases}$$

$$= \begin{cases} a^0 & \text{COSMOLOGICAL CONSTANT } [w=0] \end{cases}$$

$$H^2 = \frac{8\pi G}{3} \rho_0 a^{-3(1+w)}$$

$$H = H_0 a^{-3(1+w)/2}$$

$$H_0 = \left[\frac{8\pi G}{3} \rho_0 \right]^{1/2}$$

$$\frac{da}{dt} = H_0 a^{-(1+3w)/2}$$

$$\Rightarrow a(t) = \begin{cases} \left[\frac{3(1+w)}{2} H_0 t \right]^{\frac{2}{3+3w}} \\ e^{H_0 t} \end{cases}$$

$$w > -1$$

$$w = -1$$

$$\propto \begin{cases} t^{1/2} \\ t^{2/3} \\ e^{H_0 t} \end{cases}$$

$$w = \frac{1}{3} \quad [\text{RADIATION}]$$

$$w = 0 \quad [\text{MATTER}]$$

$$w = -1 \quad [\text{COSM. CONSTANT}]$$

Λ CDM

$$\text{CONST. } w \Rightarrow \begin{cases} \rho = \rho_0 a^{-3(1+w)} \\ p = w\rho_0 a^{-3(1+w)} \end{cases}$$

$$\Lambda\text{CDM} \Rightarrow \begin{cases} \rho = \rho_{\text{r0}} a^{-4} + \rho_{\text{m0}} a^{-3} + \Lambda \\ p = \frac{1}{3}\rho_{\text{r0}} a^{-4} + 0 - \Lambda \end{cases}$$

$$G_{\mu\nu} = 8\pi G T_{\mu\nu}$$

$$H = \left[\frac{8\pi G}{3} (\rho_{r0} a^{-4} + \rho_{m0} a^{-3} + \Lambda) \right]^{1/2}$$

=

$$\Omega_r = \frac{\rho_{r0}}{\rho_{tot,0}} \quad \Omega_m = \frac{\rho_{m0}}{\rho_{tot,0}}$$

$$\Omega_r =$$

1/2

$$\Omega_r = \frac{P_{r0}}{P_{tot,0}}$$

$$\Omega_m = \frac{P_{m0}}{P_{tot,0}}$$

$$\Omega_\Lambda = \frac{\Lambda}{P_{tot,0}}$$

$$\Omega_r = (8.5 \pm 0.7) \times 10^{-5}$$

$$\Omega_m = 0.311 \pm 0.006$$

$$\Omega_\Lambda = 0.689 \pm 0.006$$

$$\left\{ \begin{array}{l} \Omega_b = 0.05 \\ \Omega_c = 0.25 \end{array} \right.$$

$$H = - \left[\frac{8\pi G}{3} (\rho_{r0} a^{-4} + \rho_{m0} a^{-3} + \Lambda) \right]^{1/2}$$

$$= \left[\frac{8\pi G}{3} \rho_{tot} (\Omega_r a^{-4} + \Omega_m a^{-3} + \Omega_\Lambda) \right]^{1/2}$$

$$\Omega_r =$$

$$\Omega_r$$

$$\Omega_m$$

$$\Omega_\Lambda$$

$$H$$

$$\begin{aligned}
 t(a) &= - \left[\frac{8\pi G}{3} (\rho_{r0} a^{-4} + \rho_{m0} a^{-3} + \Lambda) \right]^{1/2} \\
 &= - \left[\frac{8\pi G}{3} \rho_{tot,0} (\Omega_r a^{-4} + \Omega_m a^{-3} + \Omega_\Lambda) \right]^{1/2} \\
 &= - H_0 \left[\Omega_r a^{-4} + \Omega_m a^{-3} + \Omega_\Lambda \right]^{1/2}
 \end{aligned}$$

$$\Omega_r + \Omega_m + \Omega_\Lambda = 1 \quad \text{BY DEFINITION}$$

$$\Omega_r = \frac{\rho_{r0}}{\rho_{tot,0}} \quad \Omega_m = \frac{\rho_{m0}}{\rho_{tot,0}} \quad \Omega_\Lambda = \frac{\Lambda}{\rho_{tot,0}}$$

$$\Omega_r = (8.5 \pm 0.7) \times 10^{-5}$$

$$\Omega_m = 0.311 \pm 0.006$$

$$\Omega_\Lambda = 0.689 \pm 0.006$$

$$\left\{ \begin{array}{l} \Omega_b = 0.05 \\ \Omega_c = 0.25 \end{array} \right.$$

$$\begin{aligned}
 H_0 &= \left[\frac{8\pi G}{3} \rho_{tot,0} \right]^{1/2} = (70 \pm 3) \text{ (km s}^{-1}\text{) Mpc}^{-1} \\
 &= (0.072 \pm 0.003) \text{ Gyr}^{-1}
 \end{aligned}$$

$$\frac{1}{a} \frac{da}{dt} = H(a)$$

$$\Rightarrow dt = \frac{da}{a H(a)}$$

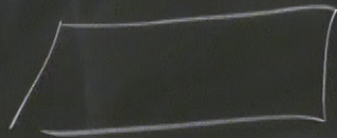
$$t = \int_0^a \frac{da'}{a' H(a')} = H_0^{-1} \int_0^a \frac{da'}{a'} \left[\Omega_r a'^{-4} + \Omega_m a'^{-3} + \Omega_\Lambda \right]$$

$$\frac{1}{a} \frac{da}{dt} = H(a)$$

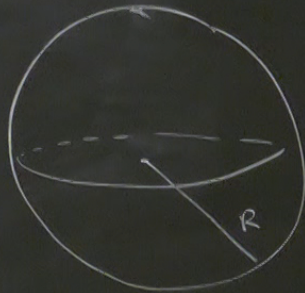
$$\Rightarrow dt = \frac{da}{a H(a)}$$

$$t = \int_0^a \frac{da'}{a' H(a')} = H_0^{-1} \int_0^a \frac{da'}{a'} \left[\Omega_r a'^{-4} + \Omega_m a'^{-3} + \Omega_\Lambda \right]^{-1/2}$$

• SPATIAL CURVATURE $K \sim [\text{COMOVING LENGTH}]^{-2}$

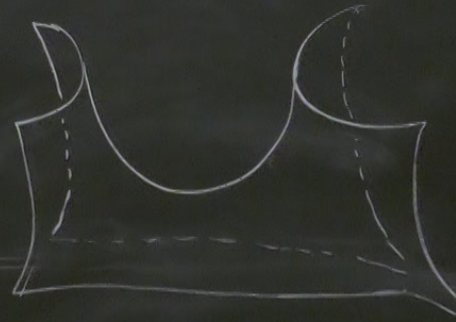


$$K=0$$



$$K > 0$$

$$= \frac{1}{R^2}$$



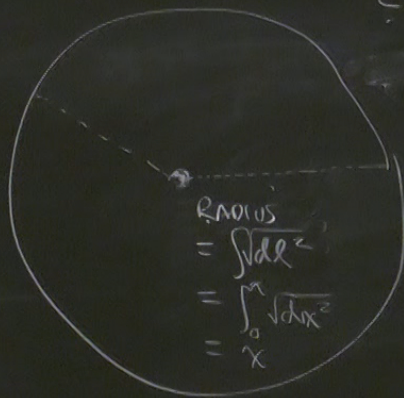
$$K < 0$$

$$ds^2 = -d$$

$$ds^2 = -dt^2 + a(t)^2 dl^2$$

$$dl^2 = (dx)^2 + S_k(x)^2 \left[(d\theta)^2 + (\sin \theta)^2 d\phi^2 \right]$$

$$S_k(x) = \begin{cases} x & \text{IF } k=0 \\ \frac{1}{\sqrt{k}} \sin(\sqrt{k}x) & \text{IF } k > 0 \\ \frac{1}{\sqrt{-k}} \sinh(\sqrt{-k}x) & \text{IF } k < 0 \end{cases}$$

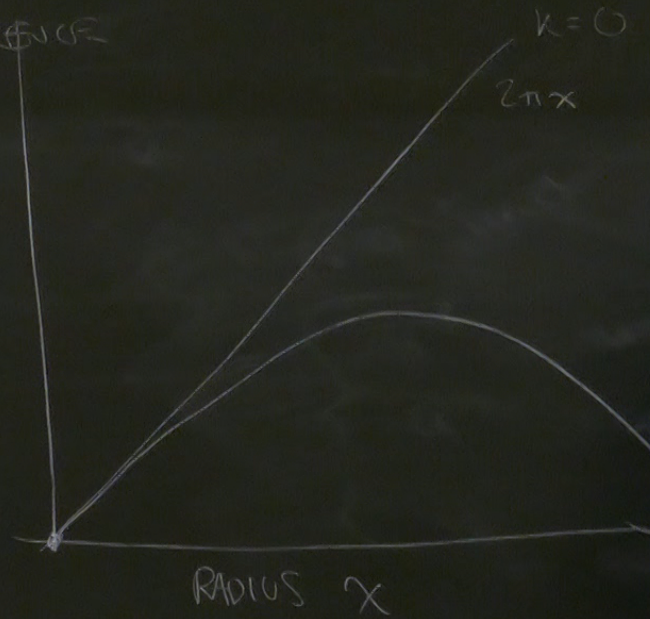


CIRCUMFERENCE

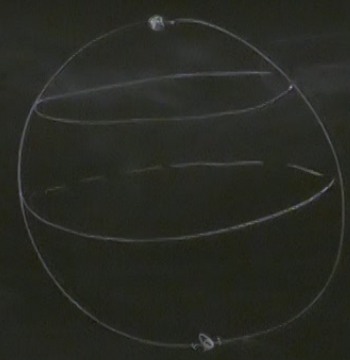
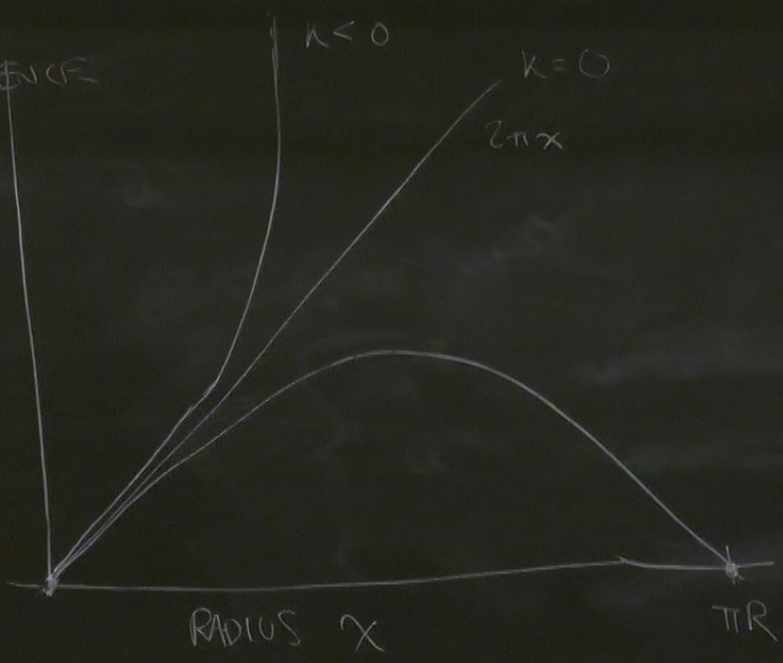
$$\begin{aligned}
 &= \int dl \\
 &= \int_0^{2\pi} \sqrt{S_k(r)^2 d\varphi^2} \\
 &= \int_0^{2\pi} S_k(r) d\varphi \\
 &= 2\pi S_k(r)
 \end{aligned}$$

CIRCUMFERENCE

$$2\pi S_k(r)$$



CIRCUMFERENCE
 $2\pi S_k(x)$



$$x \leftrightarrow r \quad r = S_K(x)$$

$$dl^2 = \frac{(dr)^2}{1 - Kr^2} + r^2 \left[(d\theta)^2 + (\sin\theta)^2 d\phi^2 \right]$$

$$\text{CIRCUMFERENCE} = 2\pi r$$

$$\text{RADIUS } x = \begin{cases} r & K=0 \\ \frac{1}{\sqrt{K}} \sin^{-1}(\sqrt{Kr}) & K > 0 \\ \frac{1}{\sqrt{-K}} \sinh^{-1}(\sqrt{-Kr}) & K < 0 \end{cases}$$

OS X

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$$\underbrace{H^2 + \frac{K}{a^2}}_{\Lambda_{\text{eff}}/3} = \underbrace{\frac{8\pi G}{3} \rho}_{8\pi G T_{\text{eff}}/3}$$

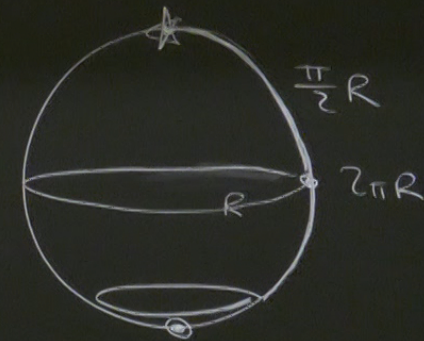
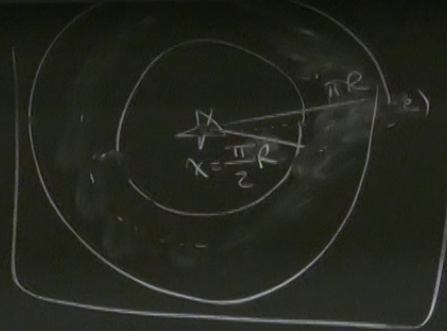
$$\dot{\rho} = -3H(\rho + p)$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p)$$

$\Lambda = 0$
 $\Lambda > 0$
 $\Lambda < 0$

πR

$$\underbrace{H^2 + \frac{K}{a^2}}_{G_{00}/3} = \underbrace{\frac{8\pi G}{3} \rho}_{8\pi G T_{00}/3}$$



$$\ddot{\rho} = -3H(\rho + \dot{\rho})$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p)$$