

Title: Grad Student Seminar Series - TBA

Speakers: Conner Dailey

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Zoom link

Scattering Amplitudes in High Energy Limit of Projectable Hořava Gravity

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Outline

1. Introduction and Motivation
2. Methods and Results
3. Projectable Hořava Gravity
4. Generalized Ward Identities
5. Calculating the Amplitudes
6. The Limit $\lambda \rightarrow \infty$
7. Conclusions and Outlook

Introduction and Motivation

- Hořava Gravity is a perturbatively renormalisable metric QFT of gravity* (2301.13580, 2307.13039)
- The Action:

$$S = \frac{1}{2G} \int dt \left(T[\partial_t \text{metric}; \lambda] - V[\text{metric; other couplings}] \right) \quad (1)$$

- The RG flow of the *projectable* version (pHG) at $\lambda = \infty$ has asymptotically free UV fixed points
- The pHG **does not** have a stable perturbative Minkowski vacuum (if the scalar gets strongly coupled it might give something similar to GR at low energies)
- The *non-projectable* version can reproduce the low energy phenomenology but it is a more complicated theory

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Methods and Results

Methods

- We perform the BRST quantization of the theory to build the Hilbert space of the scattering states
- We find the amplitudes using symbolic computer algebra

Results

- The BRST analysis of the modes is generalised to the case of non-relativistic theories and the gauge invariance of the amplitudes is confirmed
- The full set of tree-level 2 to 2 scattering amplitudes is computed
- The limit $\lambda \rightarrow \infty$ is shown to exist

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Formulating the Theory

- The renormalisability is achieved by separating spacetime into space *and* time: the theory is taken to be invariant under anisotropic scaling (like in crystals)

$$x \rightarrow b^{-1} x, \quad t \rightarrow b^{-d} t \tag{2}$$

- The theory can no longer be diffeomorphism invariant, however it can be symmetric under foliation preserving diffeomorphisms (FDiffs)

$$x \rightarrow \tilde{x} = x(\mathbf{x}, t), \quad t \rightarrow \tilde{t}(t), \quad \tilde{t}(t) - \text{monotonic function}$$

- The spacetime metric can therefore be written in terms of ADM-like variables

$$ds^2 = -N^2 dt^2 + \gamma_{ij}(dx^i + N^i dt)(dx^j + N^j dt) \tag{3}$$

- The fields γ , N and N^i transform according to

$$N \rightarrow N \frac{dt}{d\tilde{t}}, \quad N^i \rightarrow \left(N^j \frac{\partial \tilde{x}^i}{\partial x^j} - \frac{\partial \tilde{x}}{\partial t} \right) \frac{dt}{d\tilde{t}}, \quad \gamma_{ij} \rightarrow \gamma_{kl} \frac{\partial x^k}{\partial \tilde{x}^i} \frac{\partial x^l}{\partial \tilde{x}^j}.$$

Formulating the Theory

- The projectability condition:

$$N = N(t) \quad \text{consistent with transformation law} \quad (4)$$

- The Action for projectable model

$$S = \frac{1}{2G} \int d^3x dt \sqrt{\gamma} (K_{ij} K_{ij} - \lambda K^2 - \mathcal{V}) \quad (5)$$

where

$$K_{ij} = \frac{\dot{\gamma}_{ij} - \nabla_i N_j - \nabla_j N_i}{2}, \quad K \equiv \gamma^{ij} K_{ij}$$

is the extrinsic curvature (of the leafs) and

$$\begin{aligned} \mathcal{V}_{3d} &= 2\Lambda - \eta R + \mu_1 R^2 + \mu_2 R_{ij} R_{ij} \\ &+ \nu_1 R^3 + \nu_2 R R_{ij} R^{ij} + \nu_3 R_{ij} R_{jk} R_{ki} + \nu_4 \nabla_i R \nabla_i R + \nu_5 \nabla_i R_{jk} \nabla_i R_{jk} \end{aligned} \quad (6)$$

is the most general potential consisting only of relevant and marginal operators w. r. t. scaling (2)

BRST Quantization

- Ghosts and BRST transformation (Weinberg QFT Vol. II)

$$sh_{ij} = \partial_i c_j + \partial_j c_i + \partial_i c^k h_{jk} + \partial_j c^k h_{ik} + c^k \partial_k h_{ij} , \quad sN^i = \dot{c}^i - N^j \partial_j c^i + c^j \partial_j N^i , \quad (7a)$$

$$sc^i = c^j \partial_j c^i , \quad s\bar{c}_i = b_i , \quad sb_i = 0 . \quad (7b)$$

- The quantum tree-level BRST-invariant action with gauge-fixing fermion Ψ ,

$$S_q = S + \frac{1}{2G} \int d^3x dt s\Psi , \quad \Psi = 2\bar{c}_i F^i - \bar{c}_i O^{ij} b_j , \quad sS_q = 0 \quad (8)$$

where F^i are the gauge-fixing functions and O^{ij} is a non-degenerate operator:

$$F^i = \dot{N}^i + \frac{1}{2} O^{ij} (\partial_k h_j^k - \lambda \partial_j h) , \quad O^{ij} = -\frac{1}{\sigma} (\delta^{ij} \Delta^2 + \xi \partial^i \Delta \partial^j) , \quad (9)$$

- Integrate out the non-dynamical Nakanishi–Lautrup field, and the action takes the form,

$$S_q = S + \int d^3x dt \left(\frac{1}{2G} F^i O_{ij}^{-1} F^j - \frac{1}{G} \bar{c}_i sF^i \right) . \quad (10)$$

BRST Quantization

- Scatter on flat background

$$\gamma_{ij} = \delta_{ij} + h_{ij}, \quad N_i = 0 + N_i, \quad \Lambda = 0$$

- Spectrum

$$\omega_{tt}^2 = \nu_5 k^6, \quad \omega^2 = \nu_s k^6, \quad (11)$$

$$\omega_1^2 = \frac{k^6}{2\sigma}, \quad \omega_0^2 = \frac{(1-\lambda)(1+\xi)}{\sigma} k^6 \quad (12)$$

- Quantum symmetry

$$[\hat{\Phi}_{\text{as}}, Q^{(2)}] = i(s\hat{\Phi})_{\text{linearized}} \quad (13)$$

- Realization on the spectrum of in and out states

$$Q^{(2)} |\psi\rangle = 0, \quad |\psi_1\rangle \sim |\psi_2\rangle \quad \leftrightarrow \quad |\psi_1\rangle = |\psi_2\rangle + Q^{(2)} |\chi\rangle. \quad (14)$$

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General Considerations

- S-matrix

$$\langle q', \text{out} | q, \text{in} \rangle = \langle q', \text{in} | \mathcal{S} | q, \text{in} \rangle , \quad (15)$$

- Symmetry

$$[Q^{(2)}, \mathcal{S}] = 0 \implies \langle \psi' | \mathcal{S} Q^{(2)} | \chi \rangle = 0 \quad \text{for} \quad Q^{(2)} | \psi' \rangle = 0 \quad (16)$$

- In particular, for $|\chi\rangle$ we can take a physical state with added an anti-ghost,

$$|\chi\rangle = \bar{c}_{\mathbf{k}\alpha}^+ |\psi\rangle . \quad (17)$$

- BRST transformation for the anti-ghosts:

$$i[Q^{(2)}, \bar{c}_{\mathbf{k}\alpha}^+]_+ = i \sum_a \mathcal{C}_a \Phi_{\mathbf{k}\alpha}^{a+} , \quad (18)$$

- Substituting this into the symmetry condition we get

$$\sum_a \mathcal{C}_a \langle \psi' | \mathcal{S} \Phi_{\mathbf{k}\alpha}^{a+} | \psi \rangle = 0 . \quad (19)$$

Examples

- Relativistic Yang-Mills in 3+1 dimensions in Feynman gauge

$$\mathcal{L}_q^{\text{YM}} = -\frac{1}{4}F_{\mu\nu}^a F_{\mu\nu}^a - \frac{1}{2}(\partial_\mu A_\mu^a)^2 + \bar{c}^a \partial_\mu D_\mu c^a , \quad (20)$$

where

$$F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g f^{abc} A_\mu^b A_\nu^c , \quad D_\mu c^a = \partial_\mu c^a + g f^{abc} A_\mu^b c^c , \quad (21)$$

g is the coupling constant, and f^{abc} are the structure constants

- The theory is straightforwardly quantized

$$A_\mu^a(x) = \int \frac{d^3 k}{(2\pi)^3 2\omega_k} A_{\mu k}^a e^{-i\omega_k t + i\mathbf{k}\cdot\mathbf{x}} + \text{h.c.} , \quad (22a)$$

$$[A_{\mu k}^a, A_{\nu k'}^{b+}] = 2\omega_k \eta_{\mu\nu} \delta^{ab} (2\pi)^3 \delta(\mathbf{k} - \mathbf{k}') , \quad (22b)$$

where $\omega_k = k$.

Examples

- The BRST transformation of the anti-ghost coincides, up to a sign, with the gauge-fixing function,

$$i[Q^{(2)}, \bar{c}^a]_+ = s\bar{c}^a = -\partial_\mu A_\mu^a , \quad (23)$$

whence we read off

$$i[Q^{(2)}, \bar{c}_k^{a+}]_+ = -ik_\mu A_{\mu k}^{a+} . \quad (24)$$

Substituting this expression into Eqs. (19) we find

$$k_\mu \langle \psi' | \mathcal{S} A_{\mu k}^{a+} | \psi \rangle = k_\mu \langle \psi' | A_{\mu k}^a \mathcal{S} | \psi \rangle = 0 . \quad (25)$$

On the other hand, the scattering amplitudes involving a physical gluon with helicity ± 1 in the initial or final state are given by

$$e_\mu^{(\pm 1)} \langle \psi' | \mathcal{S} A_{\mu k}^{a+} | \psi \rangle , \quad e_\mu^{(\pm 1)*} \langle \psi' | A_{\mu k}^a \mathcal{S} | \psi \rangle \quad (26)$$

Examples

- Non-relativistic Yang–Mills theory (in 4 + 1) with the Lagrangian

$$\mathcal{L}^{\text{YM}} = \frac{1}{2} F_{i0}^a F_{i0}^a - \frac{\kappa_1}{4} D_i F_{jk}^a D_i F_{jk}^a - \frac{\kappa_2}{2} D_i F_{ik}^a D_j F_{jk}^a - g \frac{\kappa_3}{3} f^{abc} F_{ij}^a F_{jk}^b F_{ki}^c \quad (27)$$

- Gauge fixing:

$$S_{g.f.} \sim \int F^a O_{ab}^{-1} F^b, \quad F^a = \dot{\mathcal{A}}^a + \xi \Delta \partial_i A_i^a, \quad O_{ab}^{-1} = \frac{\delta_{ab}}{\xi \Delta} \quad (28)$$

- The dispersion relations are different for the transverse and longitudinal modes, as expected in theories without Lorentz invariance:

$$\omega_{\mathbf{k}1}^2 = (\kappa_1 + \kappa_2) k^4, \quad \omega_{\mathbf{k}0}^2 = \xi k^4 \quad (29)$$

Examples

- The BRST transformation of the anti-ghost is

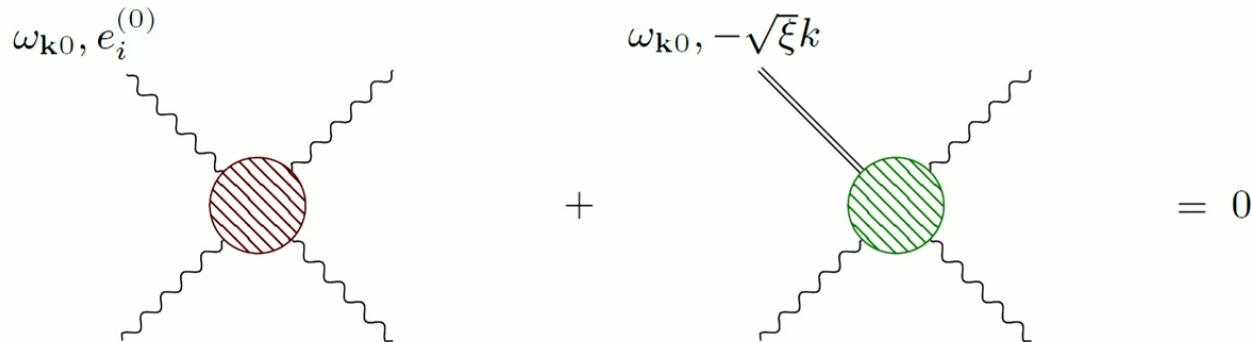
$$i[Q^{(2)}, \bar{c}^a]_+ = s\bar{c}^a = \frac{1}{\xi\Delta} (\dot{\mathcal{A}}^a + \xi\Delta\partial_i A_i^a) . \quad (30)$$

For creation-annihilation operators

$$i[Q^{(2)}, \bar{c}_k^{a+}]_+ = ik (\mathcal{A}_k^{a+} + A_{k0}^{a+}) . \quad (31)$$

Incidentally, this has the same form as in the relativistic case, cf. Eq. (24)! Hence, the constraint (19) becomes

$$\langle \psi' | \mathcal{S} A_{k0}^{a+} | \psi \rangle + \langle \psi' | \mathcal{S} \mathcal{A}_k^{a+} | \psi \rangle = 0 , \quad (32)$$

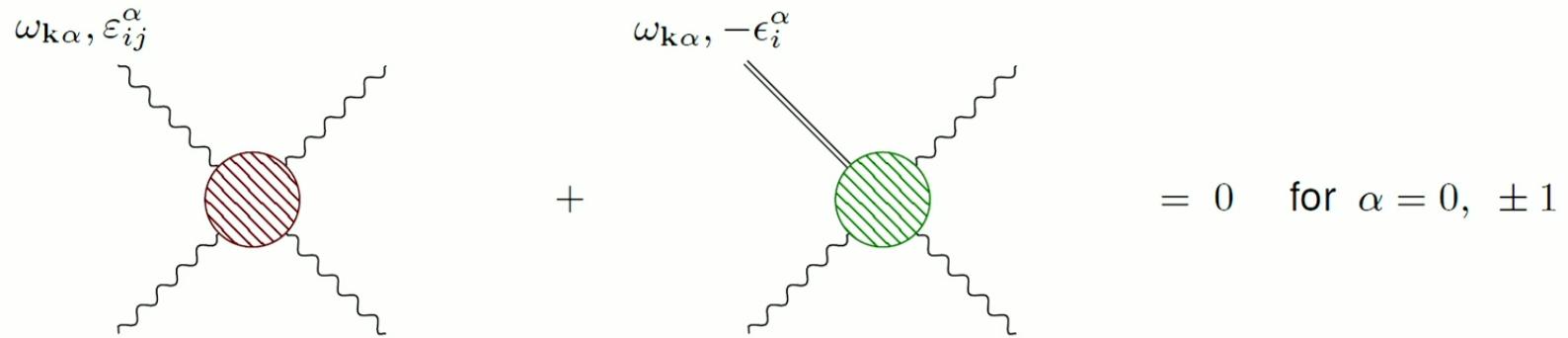


Hořava Gravity

- BRST transformation (for the third time)

$$i[Q^{(2)}, \bar{c}_i]_+ = -\frac{\sigma}{\Delta^2} \dot{N}_i + \frac{\sigma\xi}{(1+\xi)\Delta^3} \partial_i \partial_j \dot{N}_j + \frac{1}{2} (\partial_j h_{ij} - \lambda \partial_i h) \quad (33)$$

- Pictorially (α denotes various polarizations)

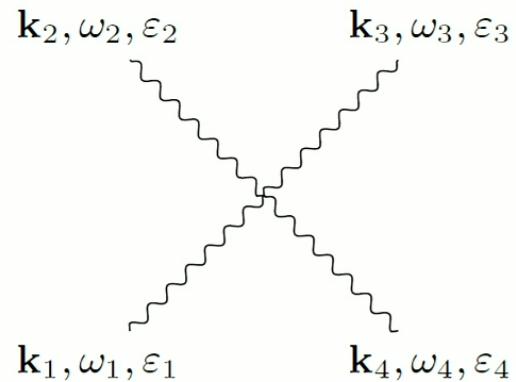


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Feynman Diagrams

$$i\mathcal{M}(\mathbf{k}_I, \omega_I, \alpha_I) =$$



$$+ \quad \begin{array}{c} \text{wavy lines from } \mathbf{k}_1, \omega_1, \varepsilon_1 \text{ to } \mathbf{k}_2, \omega_2, \varepsilon_2 \text{ and } \mathbf{k}_4, \omega_4, \varepsilon_4 \\ h \end{array} + \quad \begin{array}{c} \text{wavy lines from } \mathbf{k}_1, \omega_1, \varepsilon_1 \text{ to } \mathbf{k}_2, \omega_2, \varepsilon_2 \text{ and } \mathbf{k}_3, \omega_3, \varepsilon_3 \\ N \end{array} + (t, u)$$

Overview

- Use standard Feynman rules (and Mathematica) to compute the amplitudes
- Similar to GR

$$\mathcal{M} = \sum \text{contractions of } \varepsilon_i \text{ and } \mathbf{k}_i \quad (34)$$

- In HG the structure is richer due to the presence of higher powers of momenta (higher spatial derivatives) in the vertices. Extra terms:

$$(\mathbf{k}_3 \varepsilon_1 \varepsilon_2 \mathbf{k}_4)(\mathbf{k}_1 \varepsilon_3 \mathbf{k}_1)(\mathbf{k}_2 \varepsilon_4 \mathbf{k}_2), \quad (\mathbf{k}_2 \varepsilon_1 \mathbf{k}_2)(\mathbf{k}_1 \varepsilon_2 \mathbf{k}_1)(\mathbf{k}_4 \varepsilon_3 \mathbf{k}_4)(\mathbf{k}_3 \varepsilon_4 \mathbf{k}_3) \quad (35)$$

- Independent of gauge parameters if on-shell

Head-On Amplitudes

- The helicity amplitudes have the form

$$\mathcal{M}_{\alpha_1 \alpha_2, \alpha_3 \alpha_4} = GE^2 f_{\alpha_1 \alpha_2, \alpha_3 \alpha_4}(\cos \theta; u_s, v_a, \lambda) \quad (36)$$

- Scaling $\sim E^2$ is compatible with unitarity
- Cross section

$$\frac{d\sigma_{\alpha_1 \alpha_2, \alpha_3 \alpha_4}}{\sin \theta d\theta} = \frac{G^2}{72\pi \nu_5 k^2} |f_{\alpha_1 \alpha_2, \alpha_3 \alpha_4}|^2 \propto \lambda_{\text{de Broglie}}^2. \quad (37)$$

diverges at small angles signaling the necessity of an infrared regulator

Head-On Amplitudes

- Higher degree of collinear divergence
- Partially compensated by the orbital wave functions from the angular momentum conservation
- Poles at non-zero angles due to the decay
- Decay makes final states unstable - problem with \mathcal{S} -matrix definition

Head-On Amplitudes

- General Relativity ($x = \cos \theta$)

$$f_{++,++} = f_{--,--} = \frac{1}{1-x^2} \quad (38)$$

- Hořava Gravity:

$$\begin{aligned} f_{++,++} &= f_{--,--} = \\ &= \frac{1}{512\hat{u}_s^2(1-x^2)^3} \left[x^8 \left(-161 - 320v_2^2 + v_2(464 - 720v_3) + 39\hat{u}_s^2 - 9v_3^2(45 - 11\hat{u}_s^2) \right. \right. \\ &\quad \left. \left. + 6v_3(87 - 85\hat{u}_s^2) \right) + 4x^6 \left(231 + 443\hat{u}_s^2 - 72v_3^2\hat{u}_s^2 - 16v_2(21 - 8\hat{u}_s^2) \right. \right. \\ &\quad \left. \left. + 6v_3(63 - 53\hat{u}_s^2) \right) + 2x^4 \left(-287 + 448v_2^2 - 4783\hat{u}_s^2 - 16v_2(49 - 63v_3 + 48\hat{u}_s^2) \right. \right. \\ &\quad \left. \left. + 63v_3^2(9 + \hat{u}_s^2) - 6v_3(147 + 295\hat{u}_s^2) \right) - 4x^2 \left(581 + 128v_2^2 - 6343\hat{u}_s^2 \right. \right. \\ &\quad \left. \left. - 16v_2(35 - 18v_3 + 24\hat{u}_s^2) + 54v_3^2(3 - \hat{u}_s^2) - 6v_3(105 + 269\hat{u}_s^2) \right) - 169 - 64v_2^2 \right. \\ &\quad \left. - 19921\hat{u}_s^2 - 9v_3^2(9 + 17\hat{u}_s^2) + 16v_2(13 - 9v_3 + 32\hat{u}_s^2) + 6v_3(39 - 613\hat{u}_s^2) \right] \end{aligned} \quad (39)$$

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The Limit $\lambda \rightarrow \infty$

- Explicit cancellation of potentially divergent terms
- Manifest regularization of the theory by integrating in an auxiliary non-dynamical scalar field χ :

$$-\frac{\lambda}{2G}\sqrt{\gamma}K^2 \quad \xrightarrow{\lambda \rightarrow \infty} \quad \frac{\sqrt{\gamma}}{G} \left[-\chi K + \frac{\chi^2}{2\lambda} \right] \quad (40)$$

Take the limit (with other couplings fixed) and get for the action of HG,

$$S \xrightarrow[\lambda \rightarrow \infty]{} S' = \frac{1}{2G} \int d^3x dt \sqrt{\gamma} (K_{ij}K^{ij} - 2\chi K - \mathcal{V}) \quad (41)$$

- The field χ takes the role of a Lagrange multiplier constraining the extrinsic curvature to be traceless, $K = 0$
- Propagators in the modified theory are regular: theory is well-defined (albeit inconvenient to use)

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Conclusions and Outlook

- Obtained generalised Ward identities for non-relativistic gauge theories
- Computed the full set of tree level 2 to 2 scattering amplitudes
- We checked explicitly the finiteness of the amplitudes in the limit $\lambda \rightarrow \infty$ and reformulated theory to make it manifestly regular in this limit. This establishes the $\lambda \rightarrow \infty$ as a viable location for asymptotically free UV fixed points
- What about the ground state of pHG?
- Proof of the renormalizability of the non-projectable version