

Title: Gravitational Physics Lecture

Speakers: Ruth Gregory

Collection: Gravitational Physics

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L4: Black Holes & Branes

Last time:

$$ds^2 = A^2(r) dt^2 - B^2(r) dr^2 - C^2(r) d\Omega_{D-2}^2$$

$$R^t_t = \frac{1}{B^2} \left(\frac{A''}{A} - \frac{A'B'}{AB} + (D-2) \frac{A'C'}{AC} \right)$$

$$R^r_r = \frac{1}{B^2} \left(\frac{A''}{A} + (D-2) \frac{C''}{C} - \frac{B'}{B} \left(\frac{A'}{A} + (D-2) \frac{C'}{C} \right) \right)$$

$$R^d_\beta = \delta^\alpha_\beta \left[\frac{1}{B^2} \left(\frac{C''}{C} - \frac{C'B'}{CB} + \frac{A'C'}{AC} + (D-3) \frac{C'^2}{C^2} \right) - \frac{1}{C^2} R^{\alpha\beta} \right]$$

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For a subspace of curvature κ ,

$$\frac{R_{\alpha\beta}}{n} \sim \frac{(D-3)\kappa}{n-1} \delta^{\alpha}_{\beta}$$

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Add extra building blocks

$$ds^2 = A^2 dt^2 - B^2 dr^2 - c^2 d\Omega_{n-1}^2 - \overset{\text{FLAT}}{\downarrow} D^2 dy_p^2$$

Clear that y & Ω do not mix at connection level.

$$\underline{\Theta}^\alpha_\beta = \underline{\Theta}^\alpha_\rho \underline{f}_\rho, \quad \underline{\Theta}^\alpha_r = \frac{c'}{c} \underline{\omega}^\alpha$$
$$\underline{\Theta}^i_r = \frac{D'}{D} \underline{\omega}^i, \quad \underline{\Theta}^t_r = \frac{A'}{A} \underline{\omega}^t$$

$$\underline{\Theta}^\alpha_\beta = \underline{\Theta}^\alpha_\alpha f, \quad \underline{\Theta}^\alpha_r = \frac{c'}{c} \underline{\omega}^\alpha$$

$$\underline{\Theta}^i_r = \frac{D'}{D} \underline{\omega}^i, \quad \underline{\Theta}^t_r = \frac{A'}{A} \underline{\omega}^t$$

2-forms are same or similar,
get extra cross terms:

$$\text{e.g. } \underline{R}^t_i = \underline{\Theta}^t_r \wedge \underline{\Theta}^r_i = \frac{A'D'}{AD} \underline{\omega}^t \wedge \underline{\omega}^i$$

↑ ignored with η .

$$\underline{\Theta}^\alpha_\beta = \underline{\Theta}^\alpha_{\alpha\beta}, \quad \underline{\Theta}^\alpha_r = \frac{C'}{C} \underline{\omega}^\alpha$$

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↑
ignored with η .

$$R^{tr}_{tr} = \frac{1}{B^2} \left(\frac{A''}{A} - \frac{A'D'}{AD} \right) \quad R^{ti}_{tj} = \frac{A'D'}{AB^2D} \delta^i_j$$

$$R^{ir}_{jr} = \frac{1}{B^2} \left(\frac{D''}{D} - \frac{B'D'}{BD} \right)$$

$$R^{i\alpha}_{j\beta} = \frac{C'D'}{B^2D} \delta^i_j \delta^\alpha_\beta$$

$$R^{ij}_{kl} = \frac{D''}{D^2} (\delta^i_k \delta^j_l - \delta^i_l \delta^j_k)$$

plus the
same
 R^{td}_{tr}
 $R^{r\beta}_{r\alpha}$
 $R^{r\alpha}_{rp}$
from Friday

$$\underline{\Theta}^\alpha_\beta = \underline{\Theta}^\alpha_{\alpha\beta}, \quad \underline{\Theta}^\alpha_r = \frac{c'}{c} \underline{\omega}^\alpha$$

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(R = d\Theta + \Theta \wedge \Theta)

↑
ignored with \eta.

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plus the
same
 $R^{t\alpha}_{t\beta}$
 $R^{\alpha\beta}_{\gamma\delta}$
 $R^{\alpha\alpha}_{\beta\beta}$
from Friday

$$\underline{\Theta}^\alpha_\beta = \underline{\Theta}^\alpha_\rho \underline{\omega}^\rho_\beta, \quad \underline{\Theta}^\alpha_r = \frac{c'}{c} \underline{\omega}^\alpha_r$$

$$\underline{\Theta}^i_r = \frac{D'}{D} \underline{\omega}^i_r, \quad \underline{\Theta}^t_r = \frac{A'}{A} \underline{\omega}^t_r$$

2-forms are same or similar,
get extra cross terms:

$$\text{e.g. } \underline{R}^t_i = \underline{\Theta}^t_r \wedge \underline{\Theta}^r_i = \frac{A'D'}{AD} \underline{\omega}^t_r \wedge \underline{\omega}^r_i$$

↑
lowered with η .

(R = d Θ + $\Theta \wedge \Theta$)

$$R^{tr}_{tr} = \frac{1}{B^2} \begin{pmatrix} A'' & -A'D' \\ A & AD \end{pmatrix} \quad R^{ti}_{tj} = \frac{A'D'}{AB^2} \delta^i_j$$

$$R^{ir}_{jr} = \frac{1}{B^2} \begin{pmatrix} D'' & -B'D' \\ D & BD \end{pmatrix}$$

$$R^{i\alpha}_{j\beta} = \frac{C'D'}{B^2CD} \delta^i_j \delta^\alpha_\beta$$

$$R^{ij}_{kl} = \frac{D''}{D^2} (\delta^i_k \delta^j_l - \delta^i_l \delta^j_k)$$

plus the same

$$R^{k\alpha}_{l\beta} \delta^\alpha_\gamma \delta^\beta_\delta$$

$$R^{\alpha\beta}_{\gamma\delta}$$

$$R^{\alpha\beta}_{\gamma\delta}$$

from Friday

$$R^{\alpha\beta}_{\gamma\delta} = -\frac{1}{C^2} R^{\alpha\beta}_{\rho\sigma} \delta^\rho_\gamma \delta^\sigma_\delta + \frac{C''}{C^2 D^2} (\delta^\alpha_\gamma \delta^\beta_\delta - \delta^\alpha_\delta \delta^\beta_\gamma)$$

Hence $R_t^t = \frac{1}{B^2} \left(\frac{A''}{A} - \frac{A'B'}{AB} + p \frac{A'D'}{AD} + n \frac{A'C'}{AC} \right)$

$$R_{ij}^i = \frac{\delta_{ij}}{B^2} \left(\frac{D''}{D} - \frac{D'B'}{DB} + \frac{A'D'}{AD} + n \frac{C'D'}{CD} + (p-1) \frac{D'^2}{D^2} \right)$$

$$R_{\beta}^{\alpha} = \frac{1}{B^2} \left(\frac{C''}{C} - \frac{C'B'}{CB} + \frac{A'C'}{AC} + p \frac{C'D'}{CD} + (n-1) \frac{C'^2}{C^2} \right) - \frac{1}{C^2} R_{\alpha\beta}^{\alpha}$$

$$R_r^r = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C''}{C} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right)$$

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$$R_r^r = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C''}{C} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right)$$

Curvature
of dt^2

$$\frac{R_{\alpha\beta}}{n} \sim \frac{(D-3)\kappa}{n-1} \delta^{\alpha\beta}$$

Add extra building blocks

$$ds^2 = A^2 dt^2 - B^2 dr^2 - c^2 \underbrace{d\Omega_{n-1}^2}_{n\text{-dim}} - \underbrace{D^2 dy_p^2}_{p\text{-dim}}$$

FLAT, $i,j,k \dots$

2-form
get ext
e.g

that y & Ω do not mix at connection level.

$$R_{tt} = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{c''}{c} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{c'}{c} + p \frac{D'}{D} \right) \right)$$

curvature
of $d\Omega^2$

Examples $p=0$, $n=2$ i.e. $D=4$ Take $C=r$ (areal)

$$C'=1, C''=0.$$

$$G_1^t = \frac{1}{r^2} - \frac{1}{B^2} \left(-\frac{2B'}{Br} + \frac{1}{r^2} \right) = \frac{1}{r^2} - \frac{1}{r^2} \left(\frac{r}{B^2} \right)' = 8\pi\epsilon T_0.$$

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$$\Rightarrow \frac{1}{B^2} = 1 - \frac{2G}{r} \int 4\pi r^2 T^0_0 dr.$$

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$$R^r_r = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C''}{C} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right) \quad \text{of } dt^2$$

$$B^{-2} = 1 - \frac{2GM(r)}{r}$$

$$R_r^r = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C''}{C} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right) \quad \text{of } d\Omega^2$$

$$B^{-2} = 1 - \frac{2GM(r)}{r}$$

Next, $R_t^t - R_r^r = \frac{2}{rB^2} \left(\frac{A'}{A} + \frac{B'}{B} \right)$

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If $T_o = T_r$, then $A \propto 1/B$

$$\left(\frac{A'}{A} + \frac{nc'}{c} + \frac{pD'}{D} \right)$$

covariant
of dt^2

$$\frac{1}{B^2} = 1 - \frac{2GM}{r}$$

$$\frac{1}{r}$$

"energy"

In vacuum, we get

$$A^2 = \frac{1}{B^2} = 1 - \frac{2GM}{r}$$

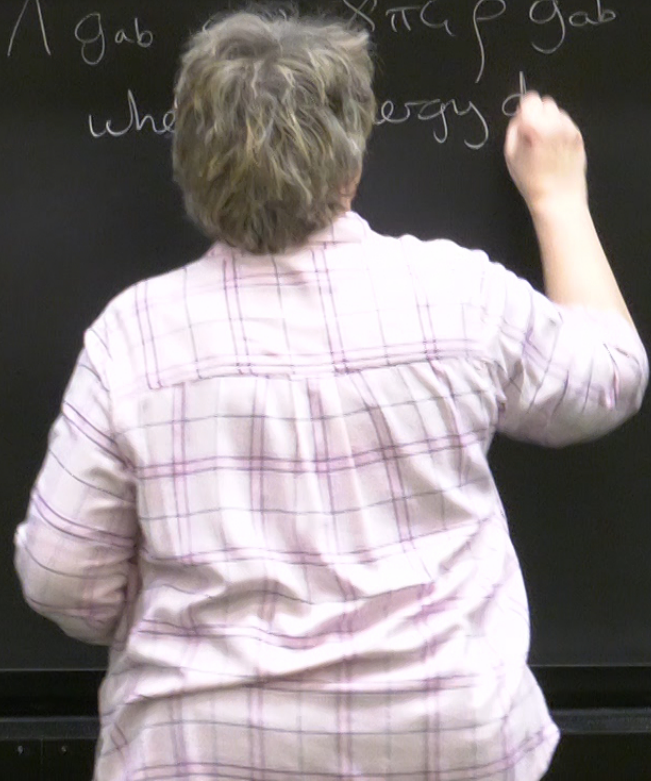
What about Λ

Cosmological constant

$$G_{ab} = 8\pi G T_{ab} + \Lambda g_{ab}$$

$$\Lambda g_{ab} \quad 8\pi G \rho g_{ab}$$

where energy density



$$\Rightarrow \frac{1}{B^2} = 1 - \frac{2M}{r} \quad \left. \begin{array}{l} 4\pi r^2 \rho dr \\ \text{"energy" between } r \text{ \& } r+dr. \end{array} \right\}$$

$$\Lambda g_{ab} \leftrightarrow 8\pi G \rho g_{ab}$$

where energy density = - (pressure)

g_{ab}

$$\Rightarrow \frac{1}{B^2} = 1 - \frac{4\pi r}{r} \int_{4\pi r}^{r+dr} \text{ "energy" between } r \text{ \& } r+dr.$$

$$\Lambda_{gab} \leftrightarrow 8\pi G \rho_{gab}$$

where energy density = -(pressure) = Tension

gab

$$\Rightarrow \frac{1}{B^2} = 1 - \frac{2M}{r} \quad \left. \begin{array}{l} 4\pi r^2 \rho dr \\ \text{"energy" between } r \text{ \& } r+dr. \end{array} \right\}$$

$$\Lambda g_{ab} \leftrightarrow 8\pi G \rho g_{ab}$$

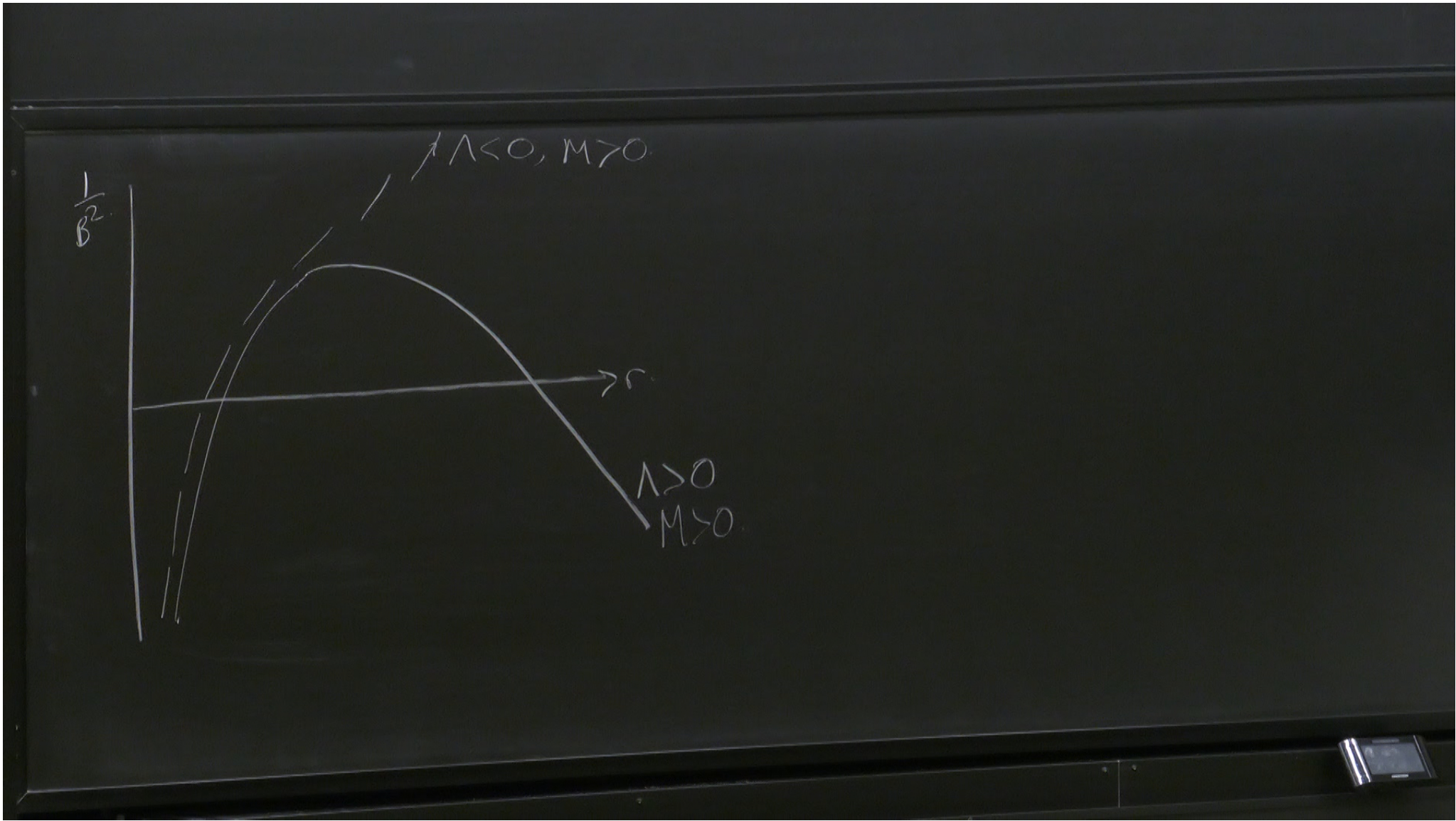
where energy density = - (pressure) = Tension

With Λ

$$\frac{1}{B^2} = 1 - \frac{2M}{r} - \int 4\pi r^2 \frac{\Lambda}{8\pi G} dr$$

$$= 1 - \frac{\Lambda r^2}{3} \quad \left(-\frac{2GM}{r} \right)$$

g_{ab}



"Energy" between r & $r+dr$.

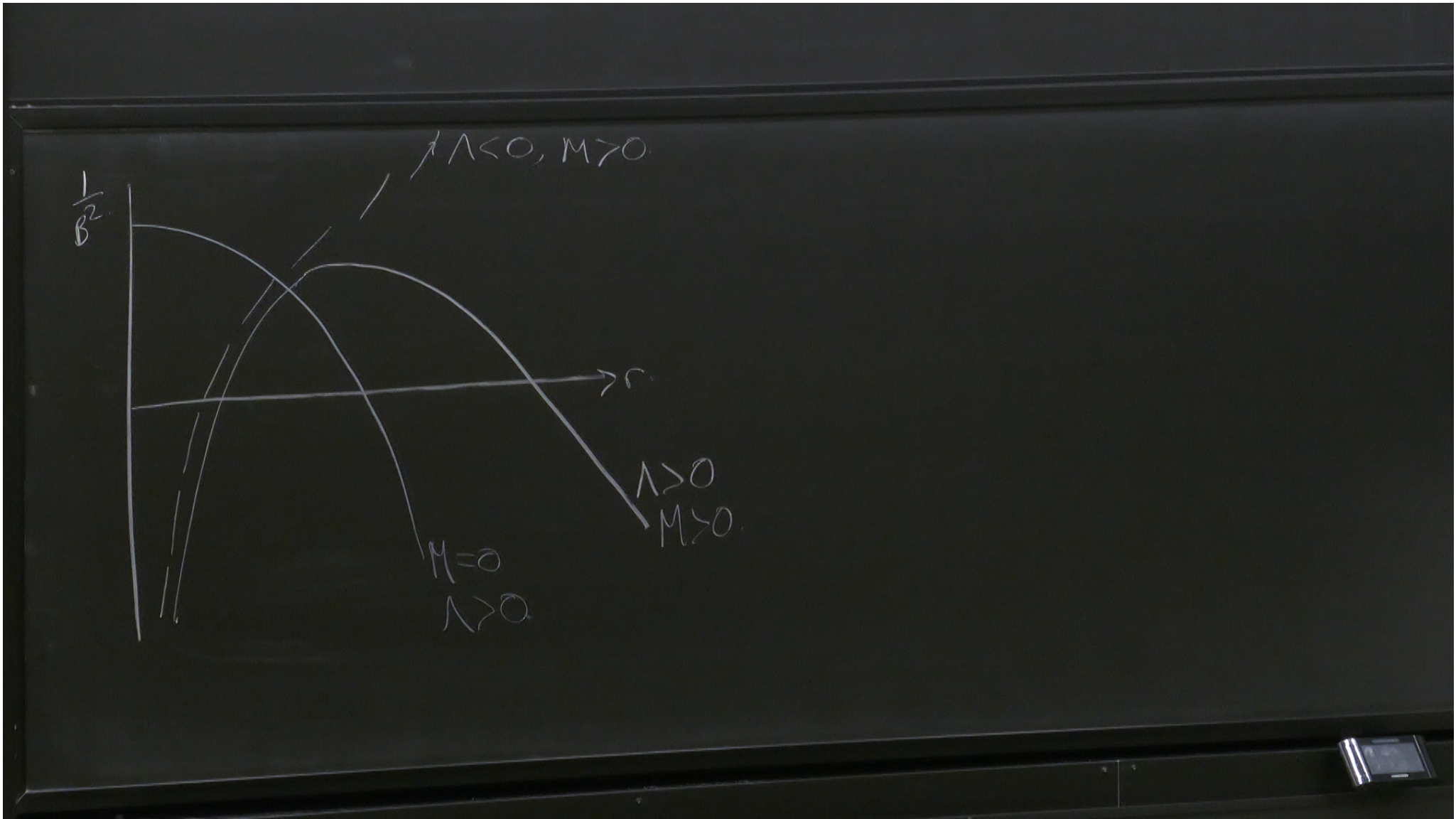
$$\Lambda g_{ab} \leftrightarrow 8\pi G \rho g_{ab}$$

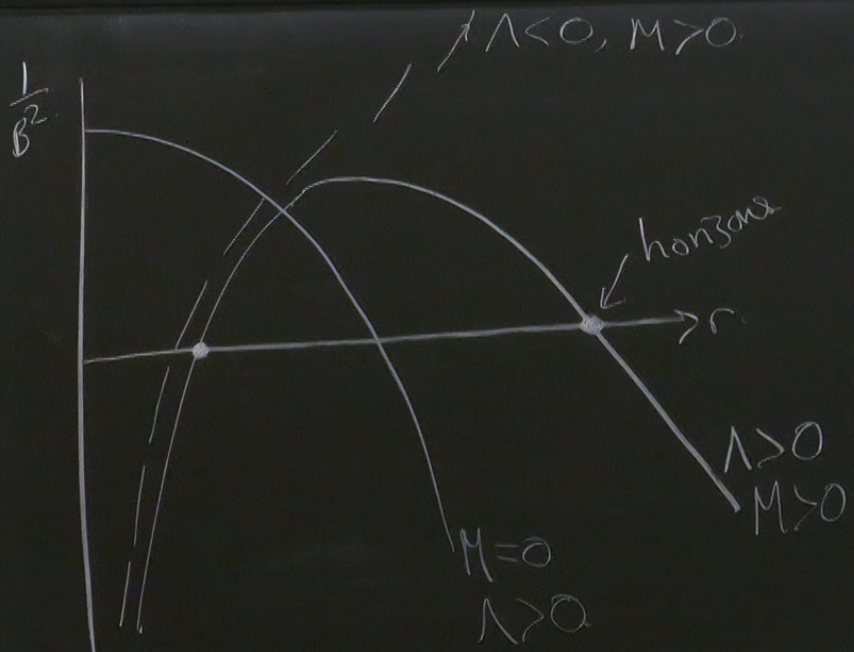
where energy density = - (pressure) = Tension

With Λ

$$\frac{1}{B^2} = 1 - \frac{2GM}{r} - \int 4\pi r^2 \frac{\Lambda}{8\pi G} dr$$
$$= 1 - \frac{\Lambda r^2}{3} \quad \left(-\frac{2GM}{r} \right)$$

$A \propto 1/B$
still.





$\Lambda > 0$ is de Sitter spacetime (dS)
 $\Lambda < 0$ is anti-de Sitter (AdS)

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(M=0) $ds^2 = \left(1 - \frac{r^2}{L^2}\right) dt^2 - \frac{dr^2}{\left(1 - \frac{r^2}{L^2}\right)} - r^2 d\Omega^2$ where $\Lambda = \frac{3}{L^2}$

$\Lambda > 0$
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$g_{tt} \rightarrow 0$ at $r = L = \sqrt{\frac{3}{\Lambda}}$

- A cosmological horizon

$\Lambda > 0$ is de Sitter spacetime (dS)

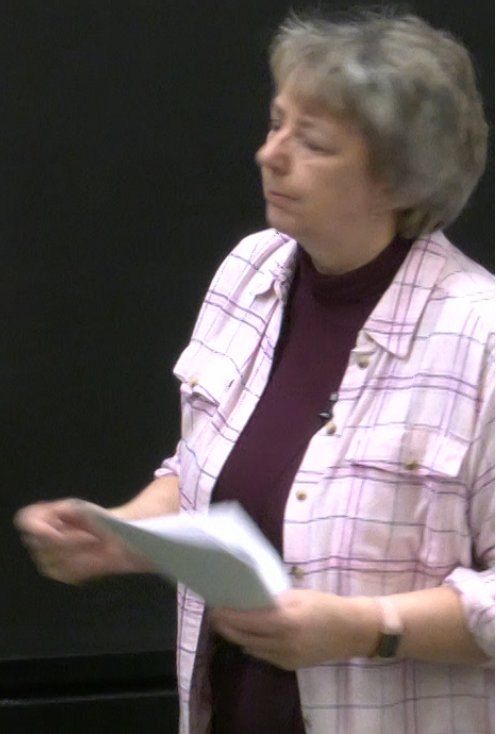
$\Lambda < 0$ is anti-de Sitter (AdS)

(M=0) $ds^2 = \left(1 - \frac{r^2}{L^2}\right) dt^2 - \frac{dr^2}{\left(1 - \frac{r^2}{L^2}\right)} - r^2 d\Omega^2$ ← where $\Lambda = \frac{3}{L^2}$

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STATIC
PATCH



de Sitter spacetime (dS)

Sitter (AdS)

$$ds^2 = \frac{dt^2}{(1 - \frac{r^2}{L^2})} - r^2 d\Omega^2$$

where $\Lambda = \frac{3}{L^2}$

STATIC
PATCH

$$L = \sqrt{\frac{3}{\Lambda}}$$

1 horizon

dS has many coord systems.

$$ds^2 = dt^2 - L^2 \cosh^2 \frac{t}{L} d\Omega^2 \quad \text{GLOBAL}$$

$$= dT^2 - e^{2T/L} dx^2 \quad \text{FLAT FRW.}$$

de Sitter spacetime (dS)

Sitter (AdS)

$$ds^2 = \frac{dt^2}{(1 - \frac{r^2}{L^2})} - r^2 d\Omega^2$$

where $\Lambda = \frac{3}{L^2}$

STATIC
PATCH

$$L = \sqrt{\frac{3}{\Lambda}}$$

1 horizon

dS has many coord systems

$$ds^2 = d\tau^2 - L^2 \cosh^2 \frac{\tau}{L} d\Omega^2 \quad \text{GLOBAL}$$

$$= dT^2 - e^{2T/L} d\underline{x}^2 \quad \text{- FLAT FRW}$$

Can be represented as a hyperboloid
embedded in (D+1) dim
Minkowski



de Sitter spacetime (dS)

Sitter (AdS)

$$ds^2 = \frac{dt^2}{(1 - \frac{r^2}{L^2})} - r^2 d\Omega^2$$

where $\Lambda = \frac{3}{L^2}$

STATIC
PATCH

$$L = \sqrt{\frac{3}{\Lambda}}$$

horizon

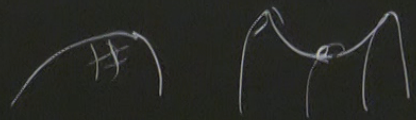
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AdS is hyperbolic - negatively curved)

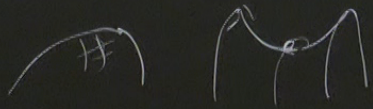


If $\kappa \neq 1$ (ie $d\Omega^2$ not a sphere)

$$A^2 = B^{-2} = \kappa - \frac{\Lambda}{3} r^2 - \frac{2GM}{r}$$

need $\Lambda < 0$, $\kappa = 0, -1$ possible

AdS is hyperbolic - negatively curved) (ii) $ds^2 = A^2 g_{\mu\nu} dx^\mu dx^\nu - dt^2$
($p=0$, analytically continue $d\Omega^2$, $A=0$)



If $\kappa \neq 1$ (ie $d\Omega^2$ not a sphere)

$$A^2 = B^{-2} = \kappa - \frac{\Lambda}{3} r^2 - \frac{2GM}{r}$$

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g curved) (ii) $ds^2 = A^2 g_{\mu\nu} dx^\mu dx^\nu - dt^2$
($p=0$, analytically continue dt^2 , $A=0$)

$$R_r^r = (0-1) \frac{A''}{A}$$

$$R_{\nu}^{\mu} = \left(\frac{A''}{A} + (0-2) \frac{A'^2}{A^2} - \frac{(0-2)\kappa}{A^2 L} \right) \delta_{\nu}^{\mu}$$

Take $g_{\mu\nu}$ constant curvature, length scale L .

In vacuum $A''=0$

$$A'^2 = \kappa/L^2$$

$\kappa=1$ for nontrivial soln

$$A = A_0 \pm r/L$$
$$= 1 \pm r/L$$

Consider

$$A = 1 - |r|/L$$

$$g_{\mu\nu} dx^\mu dx^\nu - dt^2$$

(infinite dL^2 , $A=0$)

$$\left((D-2) \frac{A'^2}{A^2} - \frac{(D-2)\kappa}{A^2 L^2} \right) \delta_{\mu\nu}^{CM}$$

of curvature, length scale L .

In vacuum $A''=0$

$$A'^2 = \kappa/L^2$$

$\kappa=1$ for nontrivial soln

$$A = A_0 \pm r/L \\ = 1 \pm r/L$$

Consider

$$A = 1 - |r|/L$$

$$A' = \mp 1/L \quad r \geq 0$$

$$A'' = -\frac{2}{L} \delta(r)$$

need $\Lambda < 0$, $K=0, -1$ possible

take $g_{\mu\nu}$ constant

Reconstruct energy-momentum

$$G^r_r = 0 \quad G^M_\nu = \frac{2(p-2)}{L} \delta(r) \delta^M_\nu$$

Concentrated source at $r=0$,

Mirror image under $r \leftrightarrow -r$.

T^M_ν like a Λ in μ - ν space

$$ds^2 = ($$

Take g_{μν} constant curvature, length scale L.

$$A' = \mp 1/L \quad r \geq 0$$
$$A'' = -\frac{2}{L} \delta(r).$$

$$ds^2 = \left(1 - \frac{|r|}{L}\right)^2 \left[dt^2 - L^2 \cosh^2\left(\frac{t}{L}\right) dl^2 \right] - dr^2$$

← FLAT SPACETIME

Looks like horizon at L.

Define $\rho = (L-r) \cosh t/L$ ←

$$\tau = (L-r) \sinh t/L$$

$$\rho^2 - \tau^2 = (L-r)^2; \quad d\tau^2 - d\rho^2 = \frac{(L-r)^2}{L^2} dt^2 - dr^2$$

Take $g_{\mu\nu}$ constant curvature, length scale L .

$$A' = \mp 1/L \quad r \gtrless 0$$

$$A'' = -\frac{2}{L} \delta(r).$$

$$ds^2 = \left(1 - \frac{|r|}{L}\right)^2 \left[dt^2 - L^2 \cosh^2\left(\frac{t}{L}\right) dl^2 \right] - dr^2$$

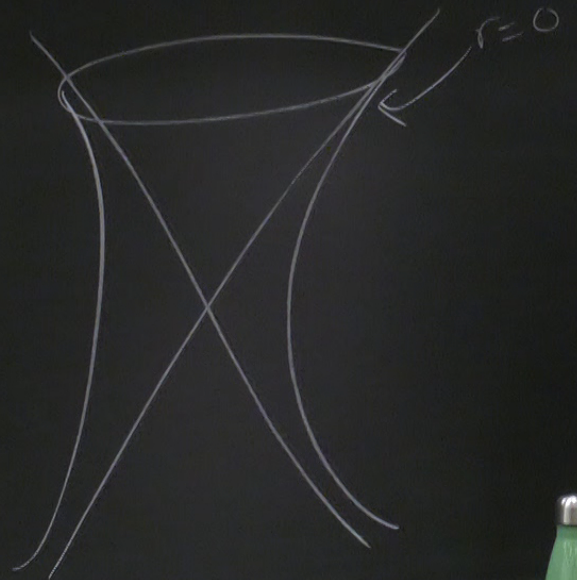
Looks like horizon at L .

Define $\rho = (L-r) \cosh t/L$ ←

$\tau = (L-r) \sinh t/L$

$$\rho^2 - \tau^2 = (L-r)^2; \quad d\tau^2 - d\rho^2 = \left(\frac{1-r}{L}\right)^2 dt^2 - dr^2$$

← FLAT SPACETIME



$$r=0 \quad p^2 - \tau^2 = L^2$$
$$r=L \quad t=0 \quad \leftarrow \quad p = \tau = 0$$

$$K'_{\tau} = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C}{C} + p \frac{D}{D} - \frac{B}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right) \quad \text{of class}$$

$\tau = L \iff$ LIGHTCONE OF ORIGIN.

$$R'' = \frac{1}{B^2} \left(\frac{A''}{A} + n \frac{C''}{C} + p \frac{D''}{D} - \frac{B'}{B} \left(\frac{A'}{A} + n \frac{C'}{C} + p \frac{D'}{D} \right) \right)$$

of dr^2

$r = L \iff$ LIGHTCONE OF ORIGIN. \rightarrow DOMAIN WALL

$r < 0$ identical

2 copies of interior of hyperboloid

"glued" at hyperboloid.