

Title: Quantum Analysis of the Bianchi IX model: Exploring Chaos

Speakers: Sara Fernandez Uria

Series: Quantum Gravity

Date: December 07, 2023 - 2:30 PM

URL: <https://pirsa.org/23120035>

Abstract: According to the Belinski-Khalatnikov-Lifshitz conjecture, the Bianchi IX spacetime describes the evolution of each spatial point close to a generic spacelike singularity. However, near the singularity, quantum effects are expected to be relevant. Therefore, in this work a quantum analysis of the model is performed, mainly focusing on its chaotic nature. Considering some minimal approximations, it is possible to encode all the information of the quantum degrees of freedom in certain canonical variables, expanding thus the classical phase space. In this way, we can apply the usual methods of dynamical systems for studying chaos. In particular, two techniques are considered. On the one hand, an analytical study is carried out, which provides an isomorphism between the quantum dynamics of Bianchi IX and the geodesic flow on a Riemannian manifold. On the other hand, by means of numerical simulations, the fractal dimension of the boundary between points with different outcome in the space of initial data is studied. The main conclusion is that, although the quantum system is chaotic, the quantum effects considerably reduce this behavior as compared to its classical counterpart.

Zoom link <https://pitp.zoom.us/j/91559466008?pwd=UzQvTGRkR3VQWm9MWDlaaVAyNi9EQT09>

Quantum Analysis of the Bianchi IX model: Exploring Chaos

Sara F. Uria

Department of Physics and EHU Quantum Center, University of the Basque Country, Bilbao, Spain

Work in collaboration with D. Brizuela (UPV/EHU), M. Bojowald (PSU), and P. Calizaya-Cabrera (LSU)

Based on: [arXiv:2307.00063](https://arxiv.org/abs/2307.00063), [arXiv:2307.13040](https://arxiv.org/abs/2307.13040)

Perimeter Institute for Theoretical Physics, December 7, 2023

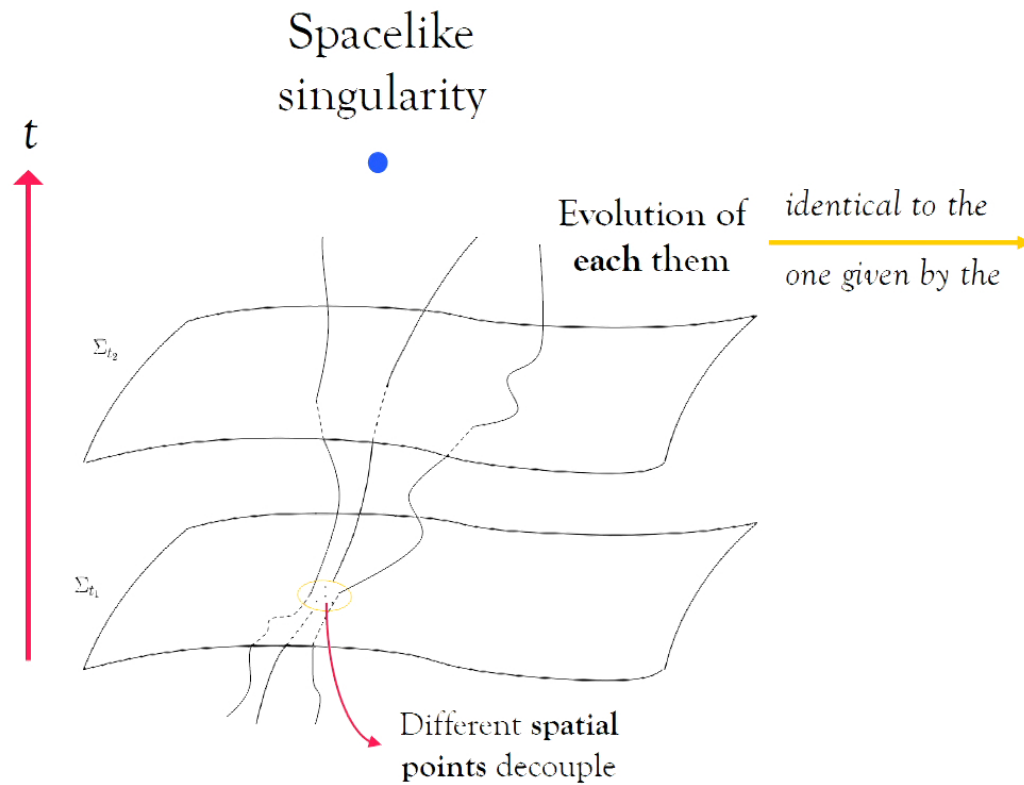


Introduction

BKL conjecture

(1970)

Supported by many numerical studies



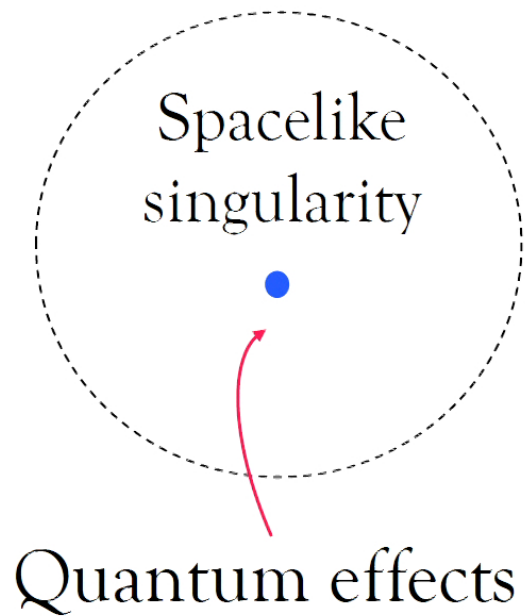
(vacuum Bianchi IX)

Mixmaster
model

- Dominated by pure gravity
- Oscillatory
- Chaotic

Introduction

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*Main goal of
my research*



Mixmaster model

- Dominated by pure gravity
- Oscillatory
- Chaotic

Classical model: Mixmaster

Metric

$$ds^2 = -N^2 dt^2 + \sum_{i=1}^3 a_i(t)^2 \sigma_i^2$$

Lapse function

Scale factors

1-forms on the
3-sphere

In terms of the Euler angles (θ, ϕ, ψ)

$$\begin{aligned}\sigma_1 &:= \sin \psi d\theta - \cos \psi \sin \theta d\phi, \\ \sigma_2 &:= \cos \psi d\theta + \sin \psi \sin \theta d\phi, \\ \sigma_3 &:= -d\psi - \cos \theta d\phi\end{aligned}$$

- Spatially homogenous and anisotropic
 - Solution of the vacuum Einstein field equations
 - Spacelike singularity at $t = 0$, where $a_i \rightarrow 0$
 - Describe the evolution of each spatial point close to a generic spacelike singularity (**BKL conjecture**)
- } vacuum **Bianchi IX** spacetime

Classical model: Mixmaster

- Follow the ADM formalism

Misner
variables

$$\beta_+ := -\frac{1}{2} \ln \left[\frac{a_3}{(a_1 a_2 a_3)^{1/3}} \right], \quad \beta_- := \frac{1}{2\sqrt{3}} \ln \left(\frac{a_1}{a_2} \right), \quad \alpha = \frac{1}{3} \ln(a_1 a_2 a_3)$$

Shape parameters

Spatial volume

Singularity $\alpha \rightarrow -\infty$

Hamiltonian
constraint

$$C = \frac{1}{2} e^{-3\alpha} (-p_\alpha^2 + p_+^2 + p_-^2) + e^\alpha V(\beta_+, \beta_-) = 0$$

Potential

- p_α, p_+, p_- conjugate momenta: $\{\alpha, p_\alpha\} = \{\beta_+, p_+\} = \{\beta_-, p_-\} = 1$

Classical model: Mixmaster

The dynamics?

- Freedom to choose any time parametrization (General Relativity)
- Choose an internal variable as the time variable
 - α is a monotonic function: choose it, gauge $\alpha = t$

**Hamiltonian
constraint**

$$C = \frac{1}{2} e^{-3\alpha} (-p_\alpha^2 + p_+^2 + p_-^2) + e^\alpha V(\beta_+, \beta_-) = 0$$

Solve it for p_α

$$H := -p_\alpha = [p_+^2 + p_-^2 + 2e^{4\alpha} V(\beta_+, \beta_-)]^{1/2}$$

**Physical
Hamiltonian**

Classical model: Mixmaster

Physical Hamiltonian

$$H := -p_\alpha = [p_+^2 + p_-^2 + 2e^{4\alpha} V(\beta_+, \beta_-)]^{1/2}$$

Equations
of motion

$$\frac{d\beta_\pm}{d\alpha} = \{\beta_\pm, H\} = \frac{\partial H}{\partial p_\pm} = \frac{p_\pm}{H}, \quad \frac{dp_\pm}{d\alpha} = \{p_\pm, H\} = -\frac{\partial H}{\partial \beta_\pm} = -\frac{e^{4\alpha}}{H} \frac{\partial V}{\partial \beta_\pm}$$

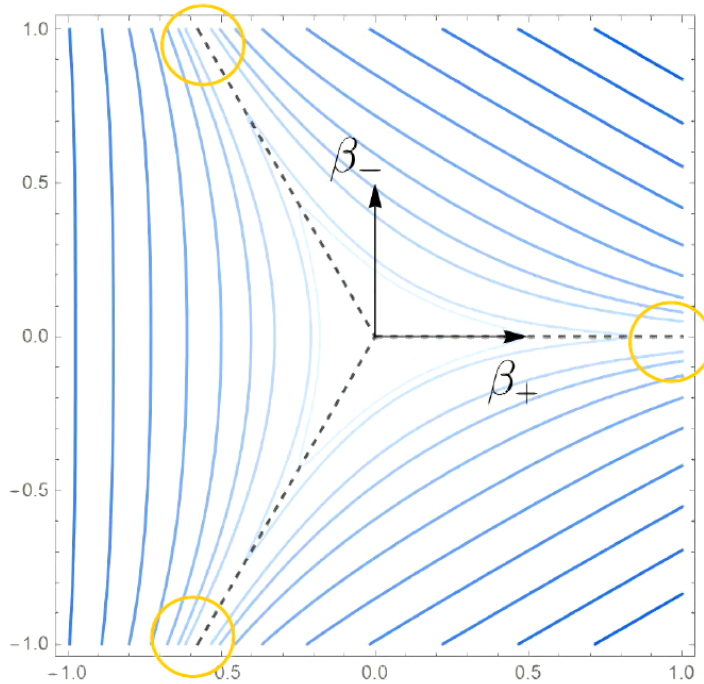
Qualitative dynamics close to the singularity ($\alpha \rightarrow -\infty$) ?



Analyze the potential $V(\beta_+, \beta_-)$

Classical model: Mixmaster

$$V(\beta_+, \beta_-) := \frac{1}{6} [e^{-8\beta_+} + 2e^{4\beta_+} (\cosh 4\sqrt{3}\beta_- - 1) - 4e^{-2\beta_+} \cosh 2\sqrt{3}\beta_-]$$



Equipotential lines

- $V(\beta_+, \beta_-)$ is negligible:

$$H = [p_+^2 + p_-^2 + 2e^{4\alpha} V(\beta_+, \beta_-)]^{1/2} \approx (p_+^2 + p_-^2)^{1/2}$$

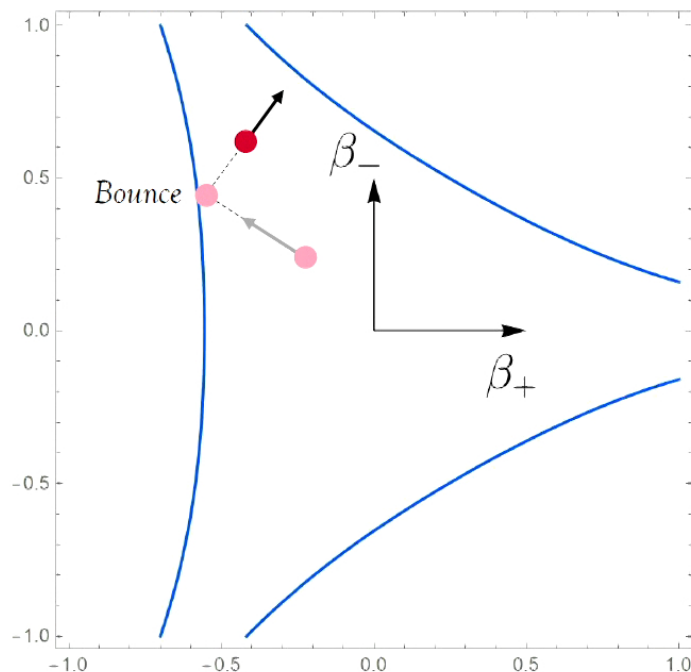
**Free
dynamics**

- $V(\beta_+, \beta_-)$ is not negligible:

Infinite potential walls

- ★ Along the 3 symmetry semiaxes, no potential wall is encountered: **3 exits**

Classical model: Mixmaster



Analogy

The system is a point particle with coordinates (β_+, β_-) moving in a **potential well with 3 exits**

~ similar to the billiard picture

$$\star H = [p_+^2 + p_-^2 + 2e^{4\alpha}V(\beta_+, \beta_-)]^{1/2}$$

Potential walls move away as $\alpha \rightarrow -\infty$

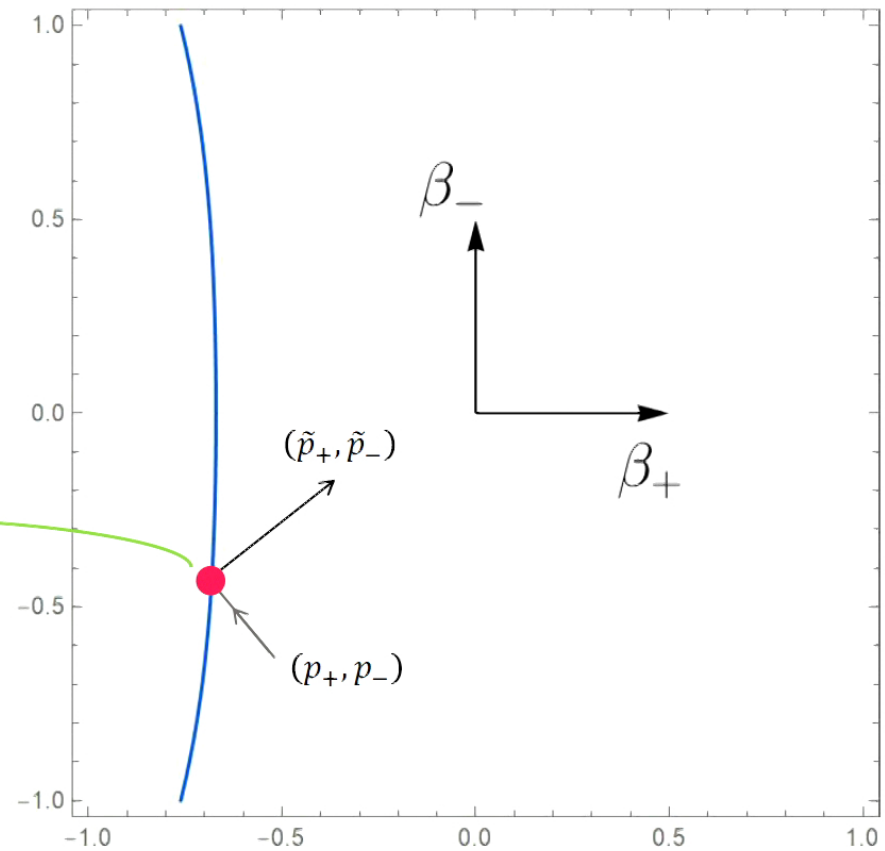
Classical model: Mixmaster

Bounces

Specific transition laws

$$\tilde{p}_- = p_-$$

$$\tilde{p}_+ = \frac{1}{3} [4(p_+^2 + p_-^2)^{1/2} - 5p_+]$$



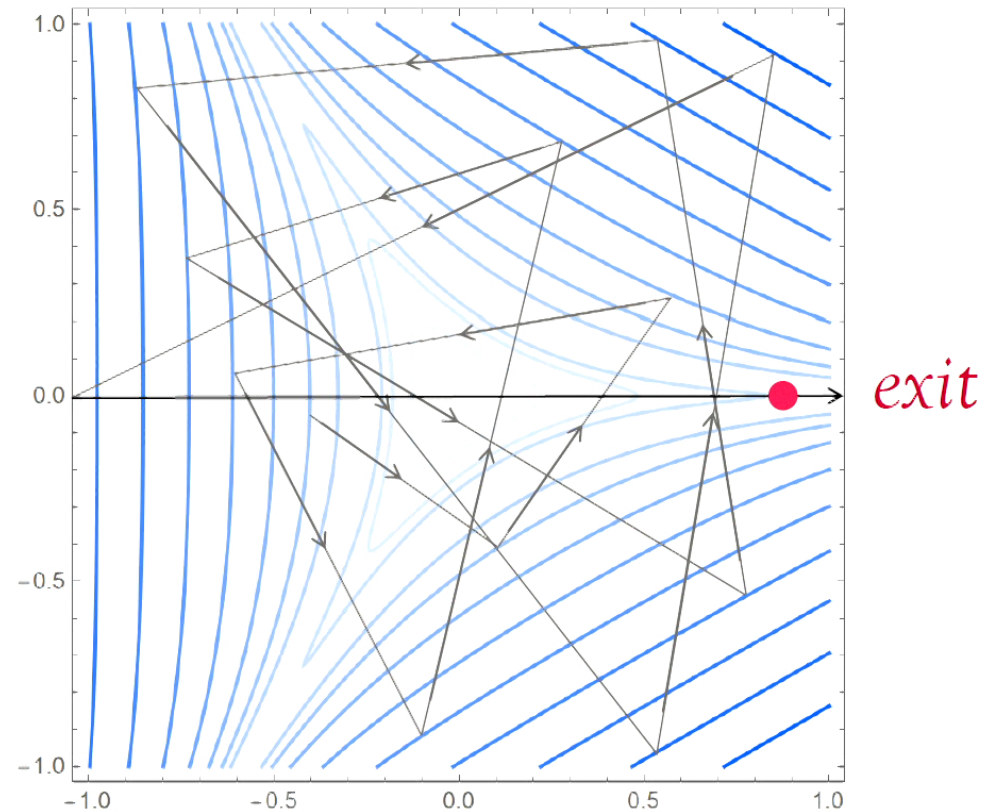
Classical model: Mixmaster

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Constantly
bouncing
against the
potential walls
(billiard)

Final state
 $\alpha \rightarrow +\infty$

(most points)
“escape”
through one
of the 3 exits



Classical model: Mixmaster

Chaotic motion

Slight difference in the initial conditions: **not escape** or escape from **another exit**

How to prove this?

- Nature of general relativity: **observer independent methods** \longrightarrow ~~Usual dynamical techniques~~

1. Fractal methods:

N. J. Cornish and J. J. Levin, *The Mixmaster universe: A Chaotic Farey Tale*, Phys. Rev. D **55**, 7489 (1997)

2. Lyapunov exponent:

G. P. Imponente and G. Montani, *On the Covariance of the Mixmaster Chaoticity*, Phys. Rev. D **63**, 103501 (2001)

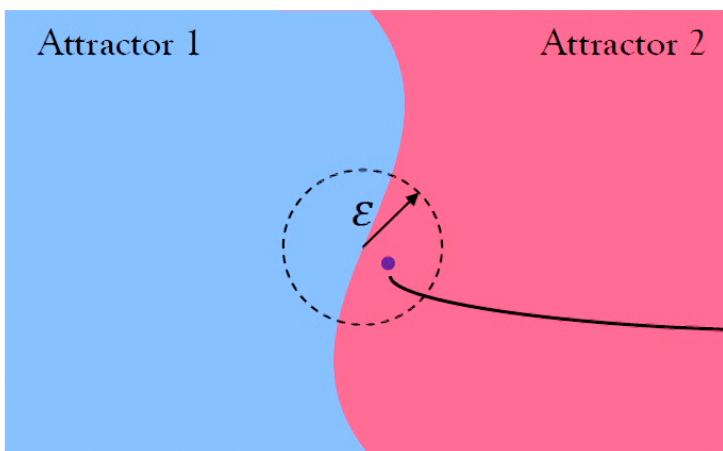
Classical model: Mixmaster

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1. Fractal methods: N. J. Cornish and J. J. Levin, *The Mixmaster universe: A Chaotic Farey Tale*, Phys. Rev. D **55**, 7489 (1997)

Main ideas:

- Systems with more than one attractor (repeller), exit or final outcome
 - Divide the space of initial conditions accordingly, in terms of the attractor/exit they end up (basins of attraction)



- * $\varepsilon \equiv$ radius of the uncertainty in the initial conditions
- * $f(\varepsilon) \equiv$ fraction of the space of initial conditions with uncertain outcomes for each ε

Uncertain outcome: **Attractor 1** or **Attractor 2**?

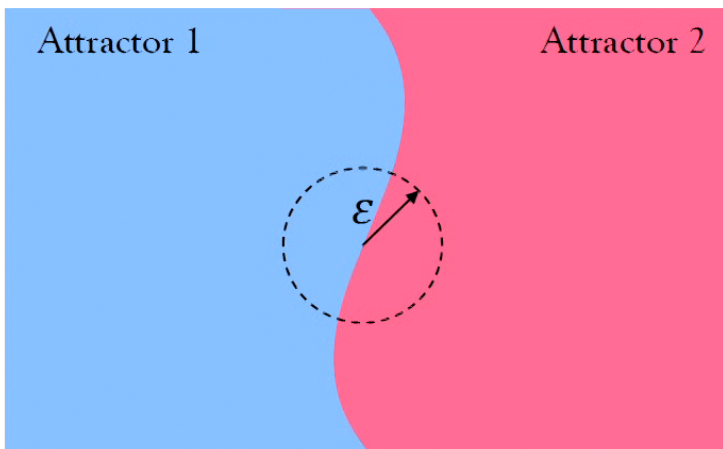
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- **Not chaotic:** no amplification of the initial error

$$f(\varepsilon) \propto \varepsilon$$

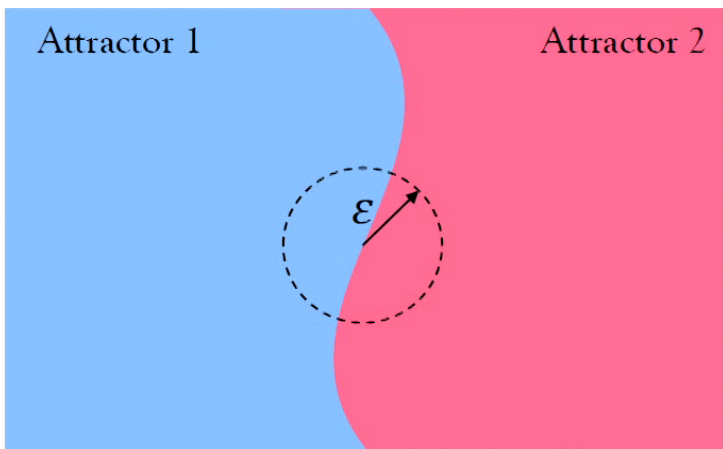
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- **Chaotic:** amplification of the initial error

$$f(\varepsilon) \propto \varepsilon^\delta \quad (0 \leq \delta < 1)$$

Classical model: Mixmaster

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1. **Fractal methods:** N. J. Cornish and J. J. Levin, *The Mixmaster universe: A Chaotic Farey Tale*, Phys. Rev. D **55**, 7489 (1997)

(Uncertainty exponent) $\delta = D - D_b$ $\left. \begin{array}{l} \\ \end{array} \right\} = 1, \text{ smooth boundary}$

Dimension of the space of initial conditions \leftarrow

Dimension of the boundary \leftarrow



- **Not chaotic:** no amplification of the initial error

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Classical model: Mixmaster

13

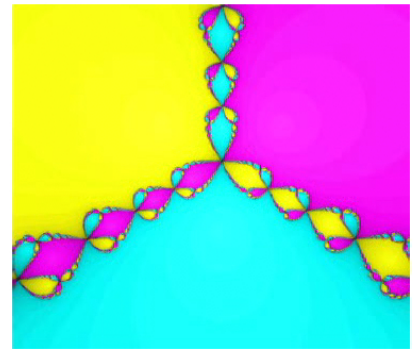
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Dimension of the space of initial conditions \leftarrow

\leftarrow Dimension of the boundary



- **Not chaotic:** no amplification of the initial error

$$f(\varepsilon) \propto \varepsilon$$

- **Chaotic:** amplification of the initial error

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Classical model: Mixmaster

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Dimension of the space of initial conditions \leftarrow D
 Dimension of the boundary \leftarrow D_b

- **Not chaotic:** no amplification of the initial error

$$f(\varepsilon) \propto \varepsilon \quad \text{Smooth boundary}$$

- **Chaotic:** amplification of the initial error

$$f(\varepsilon) \propto \varepsilon^\delta \quad (0 \leq \delta < 1) \quad \text{Fractal boundary}$$

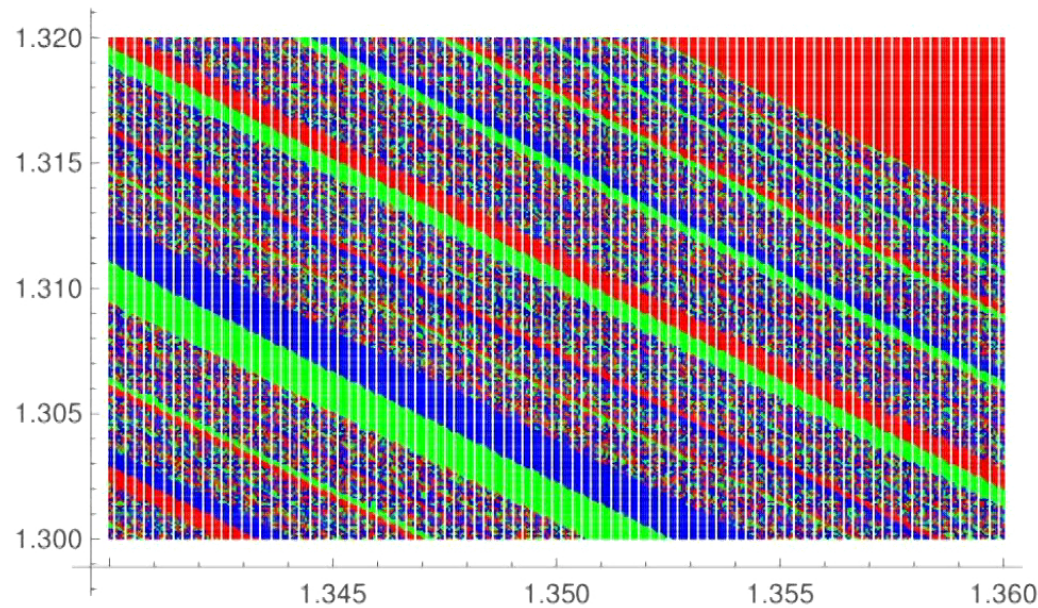
Compute δ \longrightarrow Quantify the level of chaos!

Classical model: Mixmaster

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1. **Fractal methods:** N. J. Cornish and J. J. Levin, *The Mixmaster universe: A Chaotic Farey Tale*, Phys. Rev. D **55**, 7489 (1997)

Classify each point of the space of initial data in terms of its final state (exit)



● Exit 1 ● Exit 2 ● Exit 3

Fractal structure ↔ Chaos

Quantify it: $\delta = 0.14$

Classical model: Mixmaster

2. Lyapunov exponent:

G. P. Imponente and G. Montani, *On the Covariance of the Mixmaster Chaoticity*, Phys. Rev. D **63**, 103501 (2001)

Main ideas:

- The **value** of the Lyapunov exponent is *not invariant*: depends on the time parametrization
 - The **sign** of the Lyapunov exponent is **invariant**: $\lambda > 0 \rightarrow$ chaotic
- The Lyapunov exponent is an **accurate** way of representing chaos \longleftrightarrow some **specific conditions** are satisfied

With the **usual variables** to describe the Mixmaster model they are *not satisfied* !

Change of variables

Misner-Chitre variables:

$$\alpha = -e^{\Gamma\xi}, \quad \beta_+ := e^{\Gamma\sqrt{\xi^2 - 1}}\cos\theta, \quad \beta_- := e^{\Gamma\sqrt{\xi^2 - 1}}\sin\theta, \quad (\theta \in (0, \pi), \xi \geq 1, \Gamma \in \mathbb{R})$$

Classical model: Mixmaster

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2. Lyapunov exponent:

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- Each trajectory on the phase space is **isomorphic** to a **geodesic** on a Riemannian manifold with metric:

$$ds^2 = E^2 \left[\frac{d\xi^2}{\xi^2 - 1} + (\xi^2 - 1)d\theta^2 \right]$$



Necessary conditions
to compute the
Lyapunov exponent

Chaotic geodesic flow \longleftrightarrow The Mixmaster model is **chaotic**

- The geodesic deviation equation shows that **initially close** geodesics **exponentially diverge** \rightarrow Lyapunov exponent $\lambda > 0$ \longrightarrow The system is **chaotic**

Complete picture?

Quantum effects?

Quantum model: formalism

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Canonical quantization: $\beta_{\pm} \rightarrow \hat{\beta}_{\pm}, \quad p_{\pm} \rightarrow \hat{p}_{\pm}, \quad [\hat{\beta}_m, \hat{p}_n] = i\hbar\delta_{mn} \quad (m, n \in \{+, -\})$

Instead of solving the **Schrödinger** equation $\hat{H}\psi(\beta_{\pm}, \alpha) = i\hbar \frac{\partial}{\partial \alpha} \psi(\beta_{\pm}, \alpha)$, we will study how the **quantum moments** of the wave function **evolve**, based on this equivalence:

Information
of the
quantum
state

Wave function $\psi(\beta_{\pm}, \alpha)$

equivalent

$$\left\{ \Delta(\beta_+^m p_+^n \beta_-^k p_-^l) := \left\langle (\hat{\beta}_+ - \beta_+)^m (\hat{p}_+ - p_+)^n (\hat{\beta}_- - \beta_-)^k (\hat{p}_- - p_-)^l \right\rangle \right\}_{m,n,k,l=0}^{+\infty}$$

Infinite set of quantum moments

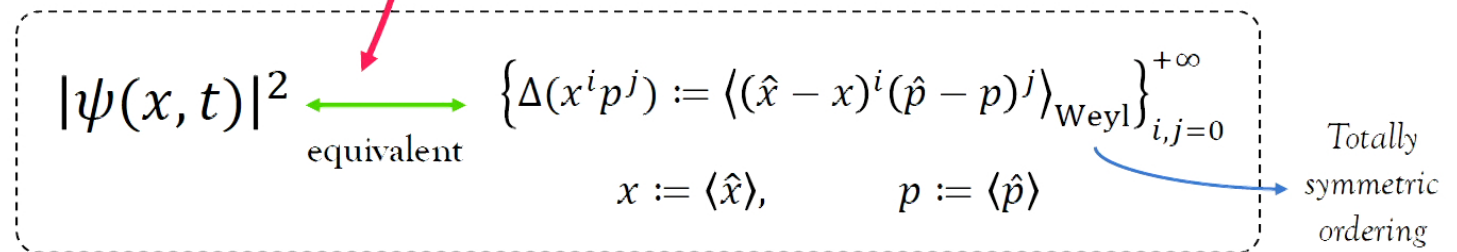
$$\beta_{\pm} := \langle \hat{\beta}_{\pm} \rangle, \quad p_{\pm} := \langle \hat{p}_{\pm} \rangle$$

Quantum model: formalism

Motivation?

- Probability distribution $P(x)$ \longleftrightarrow equivalent \longleftrightarrow Moment generating function
- $$M(s) := \int_E e^{sx} P(x) dx \quad \left(\langle x^n \rangle = \frac{d^n M(s)}{ds^n} \Big|_{s=0} \right)$$

Information of the quantum state



Dimensions

$$[\Delta(x^i p^j)] = [\hbar]^{(i+j)/2}$$

$\rightarrow i + j = \text{order}$

Central and completely symmetric quantum moments

e.g. $\text{Var}(x) = \Delta(x^2)$

Quantum model: formalism

The dynamics of the quantum moments? Governed by the following Hamiltonian:

Hamiltonian operator

Taylor expansion around the expectation values $\beta_{\pm} := \langle \hat{\beta}_{\pm} \rangle$, $p_{\pm} := \langle \hat{p}_{\pm} \rangle$

$$H_Q := \langle \hat{H}(\hat{\beta}_+, \hat{p}_+, \hat{\beta}_-, \hat{p}_-, \alpha) \rangle = H + \sum_{m+n+k+l=2}^{+\infty} \frac{1}{m! n! k! l!} \frac{\partial^{m+n+k+l} H(\beta_+, p_+, \beta_-, p_-, \alpha)}{\partial \beta_+^m \partial p_+^n \partial \beta_-^k \partial p_-^l} \Delta(\beta_+^m p_+^n \beta_-^k p_-^l)$$

Classical Hamiltonian

Equations of motion

$$\{\langle \hat{f} \rangle, \langle \hat{g} \rangle\} = \frac{1}{i\hbar} \langle [\hat{f}, \hat{g}] \rangle$$

$$\frac{d\beta_{\pm}}{d\alpha} := \{\beta_{\pm}, H_Q\} = \frac{\partial H_Q}{\partial p_{\pm}}, \quad \frac{dp_{\pm}}{d\alpha} := \{p_{\pm}, H_Q\} = -\frac{\partial H_Q}{\partial \beta_{\pm}}, \quad \frac{d\Delta(\beta_+^m p_+^n \beta_-^k p_-^l)}{d\alpha} := \{\Delta(\beta_+^m p_+^n \beta_-^k p_-^l), H_Q\}$$

Completely **equivalent** to the evolution given by the **Schrödinger equation**

Quantum model: formalism

Remarks:

- It can be applied to **any** quantum system [Bojowald, Skirzewski (2006)]
- Known for a very long time: thanks to computers “rediscovered” in the last ~15 years
- Usually H_Q cannot be written in a closed-form (infinite series)
 - Particular exception: **polynomial** Hamiltonians
 - Practical purposes: apply a **cut-off**

$$H_Q := \langle \hat{H}(\hat{x}, \hat{p}, t) \rangle = H + \sum_{i+j=2}^N \frac{1}{i!j!} \frac{\partial^{i+j} H(x, p, t)}{\partial x^i \partial p^j} \Delta(x^i p^j)$$

~ Semiclassical study

Quantum model: formalism

Advantages:

- Deal directly with **measurable quantities** (expectation values) instead of with the wave function
- Clearer to **interpret**: work in the **phase space** directly
- Closer to the classical description: classical e.o.m. + quantum corrections
 - Immediate to obtain the **classical limit**
 - Very appropriate for **semiclassical** studies
- No need to define the **Hilbert space**: “assume” that there is an internal product
- Equations with **total derivatives** instead of partial (Schrödinger)

Quantum model: set-up

Our particular model: **Mixmaster**

H_Q cannot be written in a closed-form

$$H_Q := \langle \hat{H}(\hat{\beta}_+, \hat{p}_+, \hat{\beta}_-, \hat{p}_-, \alpha) \rangle = H + \sum_{i+j+k+l=2}^{+\infty} \frac{1}{i! j! k! l!} \frac{\partial^{i+j+k+l} H(\beta_+, p_+, \beta_-, p_-, \alpha)}{\partial \beta_+^i \partial p_+^j \partial \beta_-^k \partial p_-^l} \Delta(\beta_+^i p_+^j \beta_-^k p_-^l)$$

(classical Hamiltonian) $H = [p_+^2 + p_-^2 + 2e^{4\alpha} V(\beta_+, \beta_-)]^{1/2}$

How to deal with this?

- Apply a *cut-off*: D. Brizuela and S. F. Uria (2022)
- Consider other *approximations*: M. Bojowald, D. Brizuela, P. Calizaya Cabrera, and S. F. Uria (2023)

Quantum model: set-up

Consider two assumptions: **(Semiclassical)**

$$1. \quad H_Q := \langle \hat{H} \rangle \approx \sqrt{\langle \hat{H}^2 \rangle} \longrightarrow H_Q^2 \approx \langle \hat{H}^2(\hat{\beta}_+, \hat{p}_+, \hat{\beta}_-, \hat{p}_-, \alpha) \rangle = \langle \hat{p}_+^2 + \hat{p}_-^2 + 2e^{4\alpha} V(\hat{\beta}_+, \hat{\beta}_-) \rangle$$

$$= p_+^2 + p_-^2 + \Delta(p_+^2) + \Delta(p_-^2) + 2e^{4\alpha} \sum_{i+j=2}^{+\infty} \frac{1}{i!j!} \frac{\partial^{i+j} V(\beta_+, \beta_-)}{\partial \beta_+^i \partial \beta_-^j} \Delta(\beta_+^i \beta_-^j)$$

2. **Gaussian-like** state all along the evolution:

$$\Delta(\beta_+^{2n} \beta_-^{2m}) = \frac{s_1^{2n} s_2^{2m} (2n)! (2m)!}{2^n 2^m n! m!}, \quad \Delta(\beta_+^n \beta_-^m) = 0 \text{ otherwise } (n, m \in \mathbb{N}, s_1, s_2 \in \mathbb{R}^+)$$

$$\sum_{i+j=2}^{+\infty} \frac{1}{i!j!} \frac{\partial^{i+j} V(\beta_+, \beta_-)}{\partial \beta_+^i \partial \beta_-^j} \Delta(\beta_+^i \beta_-^j) = \frac{1}{6} \left[e^{-8\beta_+ + 32s_1^2} + 2e^{4\beta_+ + 8s_1^2} \left(e^{24s_2^2} \cosh(4\sqrt{3}\beta_-) - 1 \right) - 4e^{-2\beta_+ + 2s_1^2 + 6s_2^2} \cosh(2\sqrt{3}\beta_-) \right]$$

Quantum potential $V_Q(\beta_+, \beta_-, s_1, s_2)$

Quantum model: set-up

$$\longrightarrow H_Q \approx \left[p_+^2 + p_-^2 + \Delta(p_+^2) + \Delta(p_-^2) + 2e^{4\alpha} V_Q(\beta_+, \beta_-, s_1, s_2) \right]^{1/2}$$

Closed-form
effective
Hamiltonian

Some **variables** are *not* **canonically conjugate** !



Transformation to **canonical variables** $\{s_i, p_{s_i}\} = \delta_{ij}$ ($i, j = 1, 2$)

Encode the
quantum effects

$$\Delta(p_+^2) = p_{s_1}^2 + \frac{U_1}{s_1^2}, \quad \Delta(p_-^2) = p_{s_2}^2 + \frac{U_2}{s_2^2}, \quad \text{with } U_1, U_2 \text{ constants}$$

Thus \longrightarrow

$$H_Q \approx \left[p_+^2 + p_-^2 + p_{s_1}^2 + \frac{U_1}{s_1^2} + p_{s_2}^2 + \frac{U_2}{s_2^2} + 2e^{4\alpha} V_Q(\beta_+, \beta_-, s_1, s_2) \right]^{1/2}$$

M. Bojowald, J. Phys. A: Math. Theor. 55, 504006 (2022)

Quantum model: set-up

Advantage of this approximation:

Work on an **extended** (finite) **phase space**
(2 classical + 2 quantum degrees of freedom)



Use the **previous techniques** to study **chaos**



Compare with the classical results

*Is the quantum system
more or less chaotic?*

Quantum model: chaos

1. Lyapunov exponent: *Is it chaotic?*

Problem: in the present coordinate system $(\beta_+, \beta_-, s_1, s_2, p_+, p_-, p_{s_1}, p_{s_2})$ the Lyapunov exponent cannot be properly computed

→ Extend the Misner-Chitre canonical transformation: $(\Gamma, \xi, \theta) \rightarrow (\Gamma, \xi, \theta, \sigma, \phi)$

Classical
Misner-Chitre
transformation

$$\alpha = -e^\Gamma \xi, \quad \beta_+ := e^\Gamma \sqrt{\xi^2 - 1} \cos \theta, \quad \beta_- := e^\Gamma \sqrt{\xi^2 - 1} \sin \theta \cos \sigma,$$

$$s_1 = e^\Gamma \sqrt{\xi^2 - 1} \sin \theta \sin \sigma f_+(\phi, p_\phi), \quad s_2 = e^\Gamma \sqrt{\xi^2 - 1} \sin \theta \sin \sigma f_-(\phi, p_\phi)$$

$$\text{with } f_\pm(\phi, p_\phi) := \frac{1}{\sqrt{2}|p_\phi|} [p_\phi^2 \pm (U_1 - U_2 - h(p_\phi) \sin 2\phi)]^{1/2},$$

$$h(p_\phi) := \sqrt{(p_\phi^2 - (U_1 + U_2))^2 - 4U_1U_2}, \quad \xi \geq 1, \quad \Gamma \in \mathbb{R}, \quad \theta, \sigma \in (0, \pi), \quad \phi \in [0, 2\pi)$$

Quantum model: chaos

1. Lyapunov exponent:

- As in the classical case, **each trajectory** on the phase space is **isomorphic** to a **geodesic** on a Riemannian manifold with metric

$$ds^2 = E^2 \left[\frac{d\xi^2}{\xi^2 - 1} + (\xi^2 - 1)d\theta^2 + (\xi^2 - 1) \sin^2 \theta d\sigma^2 + (\xi^2 - 1) \sin^2 \theta \sin^2 \sigma d\phi^2 \right]$$

✓ Necessary conditions to compute the Lyapunov exponent

Quantum model: chaos

1. Lyapunov exponent:

- As in the classical case, **each trajectory** on the phase space is **isomorphic** to a **geodesic** on a Riemannian manifold with metric

$$ds^2 = E^2 \left[\frac{d\xi^2}{\xi^2 - 1} + (\xi^2 - 1)d\theta^2 + (\xi^2 - 1) \sin^2 \theta d\sigma^2 + (\xi^2 - 1) \sin^2 \theta \sin^2 \sigma d\phi^2 \right]$$

Chaotic geodesic flow \longleftrightarrow The quantum Mixmaster model is chaotic

- The geodesic deviation equation shows that **initially close** geodesics **exponentially diverge** \rightarrow Lyapunov exponent $\lambda > 0$ \longleftrightarrow The quantum model is **still chaotic**

Quantum model: chaos

2. Fractal methods: *More or less chaotic than the classical system?*

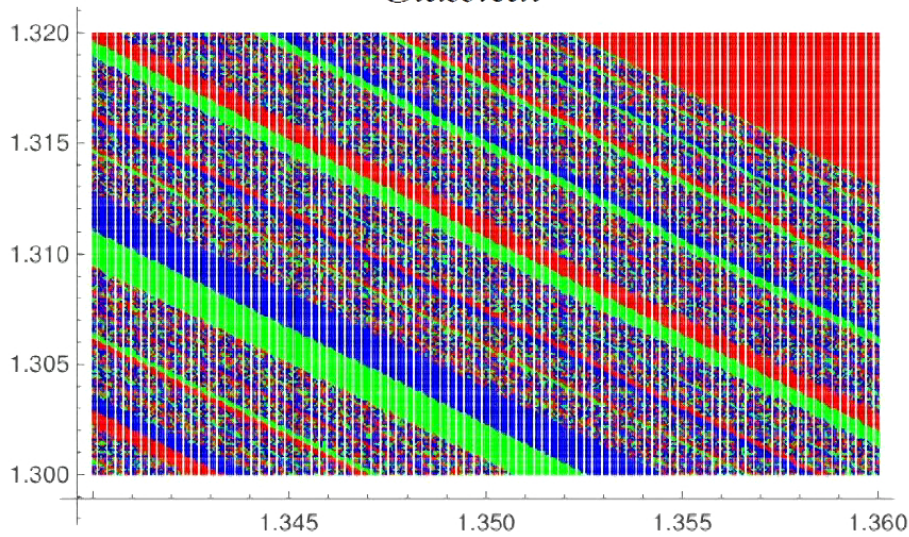
- Quantum dynamics follow the classical picture: **free motion** + **bounces**
 - There are **3 exits** ~ classical exits [D. Brizuela, S. F. Uria (2022)]
- Divide the **space of initial data** as classically: **exit** through which the system **escapes**
- For **different cross-sections** (initial values of the quantum moments) **measure** the **uncertainty exponent** \longleftrightarrow the **level of chaos**
- Compare the results with the **classical** one

Quantum effects reduce the level of chaos!

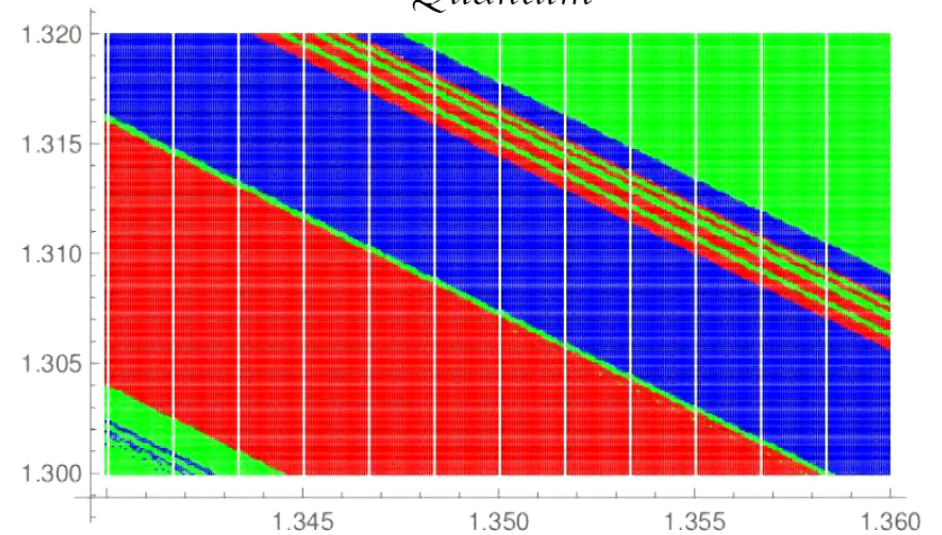
Quantum model: chaos

2. Fractal methods:

Classical



Quantum



Less fractal \longleftrightarrow Less chaotic

Quantum model: chaos

Interesting result: **implications?** (early universe...)

More approaches: **(Ongoing and future work)**

1. Apply iteratively the **discrete map** of the **quantum transition laws** obtained in [D. Brizuela, S. F. Uria (2022)]
2. Study **chaos far** from the **singularity**, where General Relativity prevails:
 - ✦ If the system is *still* chaotic, *no quantum theory* can fully **eliminate** the model's **chaotic nature**.
3. Consider a different quantization: **polymeric**

Conclusions

- Present the classical Mixmaster model: **chaotic** character
 - Numerically: **fractal methods**
 - Analytically: **Lyapunov exponent**
- Remark that **quantum effects** must be taken into account (close to the singularity)
- Build a **quantum Mixmaster model**
 - Perform a **decomposition** of the wave function into its **quantum moments**
 - Consider two minimal approximations:
 - **Semiclassical regime**
 - **General enough** for different theories of quantum gravity
 - Describe the system with **canonical variables** and **finite phase space**
- Study **quantum chaos** with the previous methods (numerical + analytical):

Quantum effects reduce the level of chaos!



Thanks for your
attention