

Title: 3 short stories about the quantum <-> gravity interface

Speakers: Etera Livine

Series: Quantum Gravity

Date: October 12, 2023 - 2:30 PM

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Abstract: Building on the fact that quantum uncertainty is a dynamical degree of freedom in itself in quantum mechanics, I will start with the remark that its evolution can provide a notion of intrinsic clock. To illustrate the non-triviality of this idea, we'll check the dynamics of the uncertainty if gravitizing quantum mechanics into a non-linear Schrödinger equation, and will see that it (surprisingly?) follows the same trajectories as planetary orbits. A third thread will come from the black hole mini-superspace in general relativity. We will see that, assuming that the quantization preserves all the symmetries of the theory, universal quantum gravity corrections can also be written as a non-linear Schrödinger equation, leading to a non-trivial evolution of the "classicality" of black hole wave-packets along the radial coordinate.

Zoom link: <https://pitp.zoom.us/j/95808416532?pwd=cHJnVXpzdDd3eXdFbUIMamFsMFV3UT09>

Three Short Stories on Quantum \leftrightarrow Gravity

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ENS Lyon, CNRS

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Oct 2023



Three Short Stories

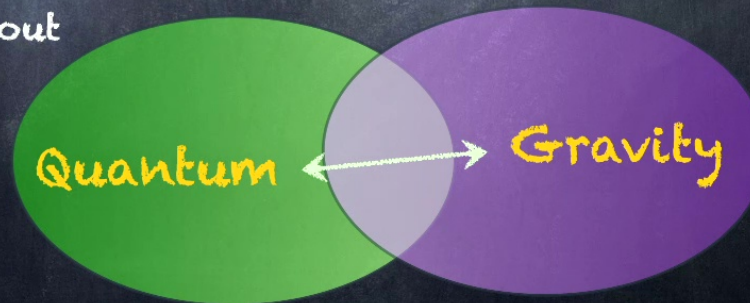
- My first is about the dynamics of quantum uncertainty
- My second is about gravitizing 1d quantum mechanics
- My third is about the quantization of the BH minisuperspace

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Three Short Stories

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↳ My whole is about
the interface



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Interface Quantum \leftrightarrow Gravity

Quantum Gravity

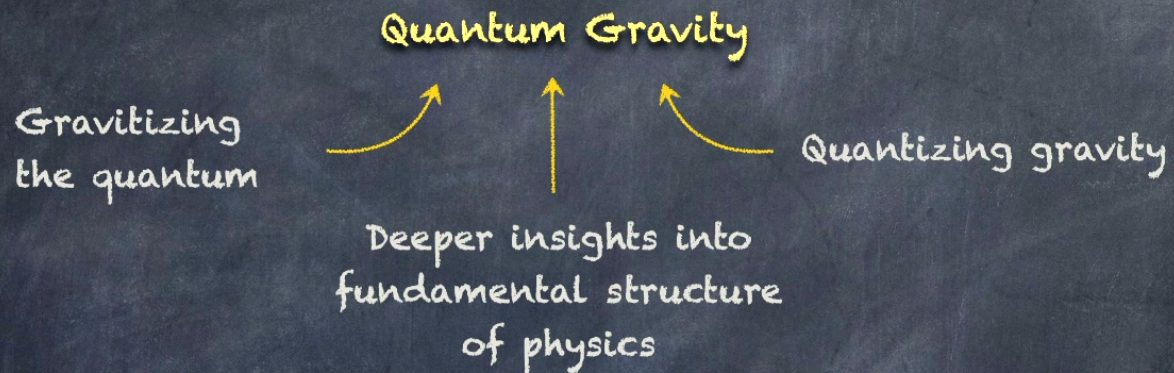
More than just
Quantizing gravity



QG Question has grown beyond
the initial mission of taming
the quantum fluctuation of the
space-time geometry

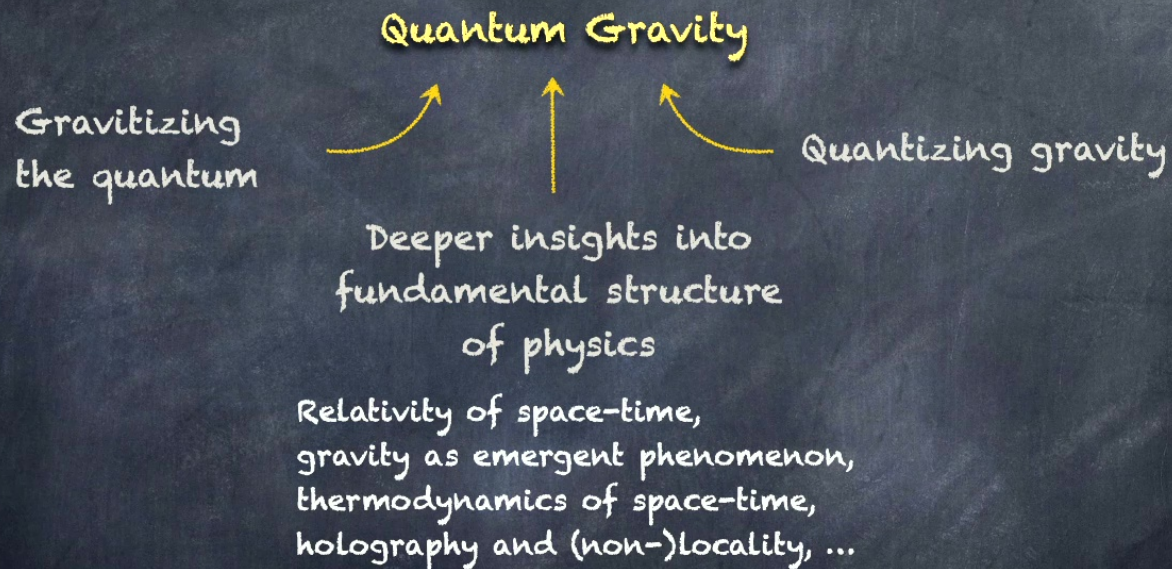
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Interface Quantum \leftrightarrow Gravity



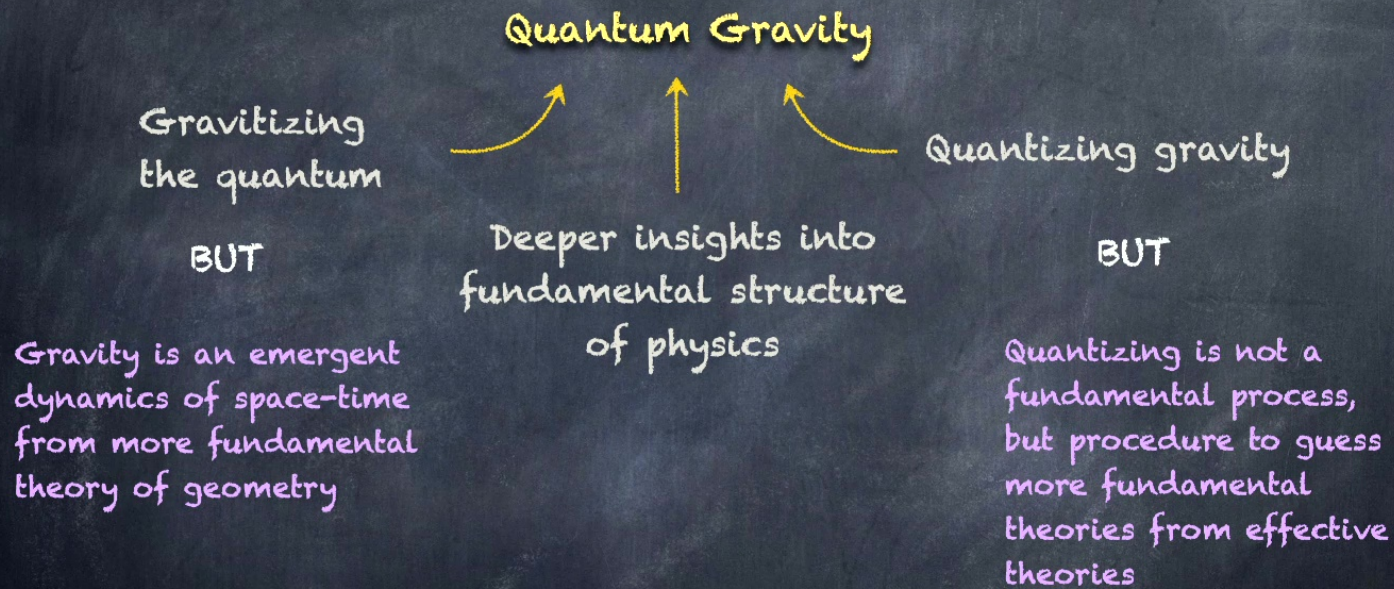
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Interface Quantum \leftrightarrow Gravity



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Interface Quantum \leftrightarrow Gravity



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Interface Quantum \leftrightarrow Gravity

Quantum Gravity

Gravitizing
the quantum

BUT

Gravity is an emergent
dynamics of space-time
from more fundamental
theory of geometry



Quantizing gravity

BUT

Quantizing is not a
fundamental process,
but procedure to guess
more fundamental
theories from effective
theories

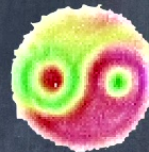
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Interface Quantum \leftrightarrow Gravity

Quantum Gravity



A deepening of BOTH concepts



1. Dynamics of quantum uncertainty
2. Gravitizing 1d quantum mechanics
3. Quantization of the BH minisuperspace



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Dynamics of Quantum Uncertainty

And so, the story begins

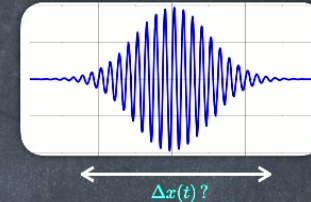
Quantum Uncertainty

A legitimate
degree of freedom

Reflects the quantum
fluctuations

Yes, things can
quantum-fluctuate
independently from
their (semi-)classical
evolution

Experimentally relevant :
Squeezing modes, qu→cl transition
and wave-function collapse



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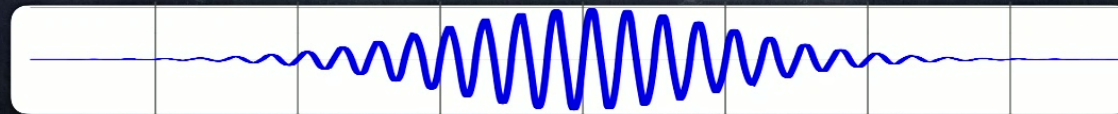
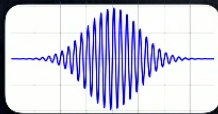
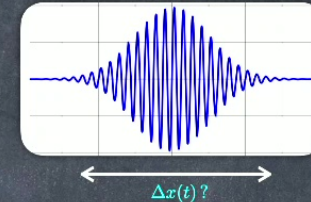
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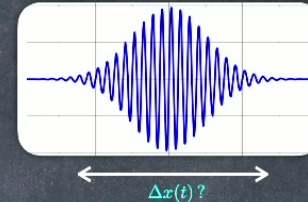
Stretched Quantum uncertainty = Quantum Superposition

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Dynamics of Quantum Uncertainty

And so, the story begins

Quantum Uncertainty



$x(t)$ versus $\Psi(x, t)$

Classical particle
Mechanics

Quantum particle
Mechanics

as field Mechanics

Wave-function carries
an infinity of d.o.f.s

Fourier modes $\Psi_p(t)$

Multipole moments $\langle x^a p^b \rangle$

Gaussian wave-packets

$\langle x \rangle, \langle p \rangle, \langle x^2 \rangle, \langle p^2 \rangle, \langle xp \rangle, \dots, \langle x^n p^{N-n} \rangle, \dots$

"Classical" d.o.f.s

non-Gaussianities

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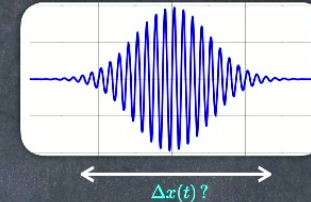
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4

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Dynamics of Quantum Uncertainty

And so, the story begins

Use **Quantum Uncertainty** as physical observable ?

↳ e.g. as an internal clock for (semi-classical) systems ?

If something evolves,
it can be used as clock

Then could it be applied
to the pb of time in
Quantum Gravity ?...

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Goal: Define Relational Observables between a system's main d.o.f.s
and its quantum fluctuation ?

Like $\langle g_{\mu\nu}(x) \delta g_{\alpha\beta}(x) \rangle$

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But here: Illustrate the idea in a much Much MUCH simpler context !

Back to the simplest model in classical mechanics

Solo adventures [2212.09442, 2305.03847, 2307.01061]



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Once upon a time, before (or after) QM

There was classical mechanics:

Consider a classical oscillator with a time-dependent frequency:

$$s_\omega[t, q(t)] = \int dt \left[\frac{1}{2} m \dot{q}^2 - \frac{1}{2} m \omega(t)^2 q^2 \right] \longrightarrow \ddot{q} + \omega^2 q = 0$$

Explicit time dependence opens the door to
explore different choices of clock
and investigate role of time reparametrization

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So let's play the time reparametrization game:

$$\left| \begin{array}{l} t \mapsto \tilde{t} = f(t) \\ q(t) \mapsto \tilde{q}(\tilde{t}) = h(t)^{\frac{1}{2}} q(t) \end{array} \right. \quad h = \dot{f}$$

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$$s_{\tilde{\omega}}[\tilde{t}, \tilde{q}(\tilde{t})] = s_\omega[t, q(t)], \quad \text{with } q = h^{-\frac{1}{2}} \tilde{q} \circ f \quad \text{and} \quad \omega^2 = h^2 (\tilde{\omega} \circ f)^2 + \frac{1}{2} \text{Schw}[f]$$

It is a "symmetry":

Virasoro group action on
Sturm-Liouville operators

$$d_{\tilde{t}}^2 \tilde{q} + \tilde{\omega}(\tilde{t})^2 \tilde{q} = 0 \quad \Leftrightarrow \quad d_t^2 q + \omega(t)^2 q = 0$$

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Standard symmetry for
vanishing Schwarzian:
SL(2, R) Mobius transf !

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Synchronizing clock

We can start with an arbitrary time-dependent frequency and map it to a constant frequency :

$$\omega^2 = h^2 \Omega^2 + \frac{1}{2} \text{Schw}[f] \quad \xrightarrow{\eta \equiv h^{-\frac{1}{2}}} \quad \ddot{\eta} + \omega^2 \eta = \frac{\Omega^2}{\eta^3} \quad \text{Auxiliary NL diff eqn}$$

Then we have a standard harmonic oscillator, we can solve it !

Defines a « synchronizing clock »
for which the oscillator beats
regularly

$$t \mapsto T = f(t) = \int^t h = \int^t \frac{1}{\eta^2}$$

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Can be used to define the Ermakov-Lewis invariant :

$$2I_\eta[q] = (\eta \dot{q} - \dot{\eta} q)^2 + \frac{\Omega^2 q^2}{\eta^2}, \quad d_t I_\eta[q] = 0$$

Looks like some angular momentum ?? Looks weird

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But where does this constant of motion mean ?

Noether charges for sym under Möbius transformations of time (conformal transformations)

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Classical solution from Quantum mechanics

Move up to quantum mechanics with potential:

$$i\hbar\partial_t\psi = -\frac{\hbar^2}{2m}\partial_x^2\psi + \frac{1}{2}m\omega(t)^2x^2\psi$$

Look at semi-classical regime: dynamics of Gaussian wave-packet

$$\psi(t, x) = Ne^{i\gamma} e^{i\frac{xp}{\hbar}} e^{-A(x-q)^2} \quad A = \frac{1}{4\alpha^2} \left(1 - i\frac{2\alpha\beta}{\hbar}\right)$$

$$\langle\hat{x}\rangle = q, \quad \langle\hat{p}\rangle = p, \quad \langle\hat{x}^2\rangle = q^2 + \alpha^2, \quad \langle\hat{x}\hat{p}\rangle_{sym} = pq + \alpha\beta$$

Then effective eqn of motions for position and uncertainty:

$$\dot{q} = \frac{p}{m}, \quad \dot{p} = -m\omega^2q$$

$$\dot{\alpha} = \frac{\beta}{m}, \quad \dot{\beta} = -m\omega^2\alpha + \frac{\hbar^2}{4m\alpha^3}$$

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Proper time reparametrization

Quantum mechanics lift a veil on classical mechanism,
but also lessons for QG ?

- Quantum uncertainty has a non-trivial evolution
- Defines an interesting "internal" clock, for which potential becomes time-independent
- Use pulsating quantum uncertainty (\neq higher moments) as clock ?

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Gravitizing wave-packets

Once upon at the start of time, Gravity should affect the Quantum

To start with the beginning, let's go simple and look at **1d QM**

Schrödinger
action

$$S[\psi] = \int dt dx \left[i\hbar \bar{\psi} \partial_t \psi - \frac{\hbar^2}{2m} \partial_x \bar{\psi} \partial_x \psi \right]$$

A 1+1-d
non-relativistic
field theory

Which **interaction terms** would be coupled by **Newton's constant G** ?

$$[G] = M^{-1} L^3 T^{-2}$$

$$[\hbar] = M L^2 T^{-1}$$

$$[\psi]^2 = L^{-1}$$

Assume real translation-invariant polynomial term, with linear coupling in G (first order) :

$$[P[\psi]] = M^2 L^{-2}$$

m^2

Only 2 possibilities :
 $\bar{\psi} \partial_x \psi$ or $|\psi|^4$

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$$m^2$$

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 $\bar{\psi}\partial_x\psi$ or $|\psi|^4$

$$S_G[\psi] = \int dt dx \left[i\hbar\bar{\psi}\partial_t\psi - \frac{\hbar^2}{2m}\partial_x\bar{\psi}\partial_x\psi + iaGm^2(\bar{\psi}\partial_x\psi - \psi\partial_x\bar{\psi}) + bGm^2|\psi|^4 \right]$$

Drift term: $x \rightarrow x-vt$

Non-linear Schrödinger

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Gravitizing wave-packets

NLSE as gravity-deformed 1d QM : what physics ?

$$S_G[\psi] = \int dt dx \left[i\hbar \bar{\psi} \partial_t \psi - \frac{\hbar^2}{2m} \partial_x \bar{\psi} \partial_x \psi + b G m^2 |\psi|^4 \right]$$

field conjugate field

$$i\hbar \{ \psi(x), \bar{\psi}(y) \} = \delta(x - y)$$

Poisson bracket

Hamiltonian

$$H[\psi] = \int dx \left[\frac{\hbar^2}{2m} \partial_x \bar{\psi} \partial_x \psi - b G m^2 |\psi|^4 \right]$$

Laplacian leads to diffusion of wave-packets = Kinetic term

Self-interaction = Scattering term creates clustering force when $b > 0$

Localisation of wave-packet requires extra-force e.g. coming from potential well

Non-linearity allows for self-localisation

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Gravitizing wave-packets

To understand mechanism, can look at effective dynamics of Gaussian wave-packets

Approximation, for small perturbations

Plug ansatz $\psi_{\text{Gaussian}}(t, x) = N e^{i\gamma} e^{i\frac{xp}{\hbar}} e^{-A(x-q)^2}$

in action
$$S_G[\psi] = \int dt dx \left[i\hbar \bar{\psi} \partial_t \psi - \frac{\hbar^2}{2m} \partial_x \bar{\psi} \partial_x \psi + bGm^2 |\psi|^4 \right]$$

Give effective action for "Gaussian mini-superspace":

$$S_G^{\text{eff}} = \int dt \left[p\dot{q} + \beta\dot{\alpha} - H_{\text{eff}}[p, q, \alpha, \beta] \right] \quad \text{with} \quad H_{\text{eff}} = \frac{p^2}{2m} + \frac{\beta^2}{2m} + \frac{\hbar^2}{8m\alpha^2} - \frac{bGm^2}{2\sqrt{\pi}\alpha}$$

Conformal potential
coming from uncertainty
Like angular momentum

Quartic self-interaction
of wave-function
Like gravitational
potential



Gravity as
analogue model for
non-linear QM

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Gravitizing wave-packets

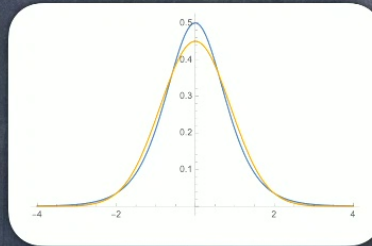
Compare exact solitons and approx Gaussian wave-packet:

$$\psi = \mathcal{A} \frac{e^{-i\omega t} e^{\frac{i}{\hbar} x p}}{\cosh \gamma(x - vt)}$$

$$\frac{1}{\gamma} = \frac{\hbar^2}{b G m^3} \frac{1}{N}$$

$$\Delta x = \sqrt{\langle x^2 \rangle - \langle x \rangle^2} = \frac{\pi}{\gamma \sqrt{12}} = \frac{\pi}{\sqrt{12}} \frac{1}{N} \frac{\hbar^2}{b G m^3}$$

$$\frac{\pi}{\sqrt{12}} \simeq 0.907$$



$$\psi(x, t) = K e^{i\gamma} e^{\frac{i}{\hbar} p(x-q)} e^{-A(x-q)^2}$$

$$\alpha_0 = \frac{\sqrt{\pi}}{2} \frac{\hbar^2}{b G m^3}$$

$$\frac{\sqrt{\pi}}{2} \simeq 0.886$$

Elliptic trajectory should approx exact pulsating solitons

↪ Oscillating quantum uncertainty,
Like beating heart of wave-packet,
a.k.a. internal clock

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Gravitizing wave-packets

Moral of the story ?

Gravity effects on QM

Self-interaction of wave-function

Non-linear Schrödinger

Approximated by gravity in field space ?!

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The BH Mini-Superspace

Let us look at spherically-symmetric static metrics in GR

- ↳ Classical vacuum solution is Schwarzschild metric
- ↳ Describes BH geometry with horizon
- ↳ Proof-of-concept configuration for GR

4

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The BH Mini-Superspace

Let us look at spherically-symmetric static metrics in GR

- ↳ Classical vacuum solution is Schwarzschild metric
- ↳ Describes BH geometry with horizon
- ↳ Proof-of-concept configuration for GR

$$S[g] = \frac{1}{l_P^2} \int_{\mathcal{M}} d^4x [\mathcal{R} - 2\Lambda]$$

$$ds^2 = \epsilon (-N^2(r)dr^2 + \gamma_{tt}(r)dt^2) + \gamma_{\theta\theta}(r)d\Omega^2$$

$$\gamma_{tt} := 2\beta(r)/\alpha(r), \quad \gamma_{\theta\theta} := l_s^2 \alpha(r) \quad c = \frac{l_0 l_s^2}{l_p^3} = \frac{V_{IR}}{V_{UV}}$$

$\epsilon = -$ BH exterior

$\epsilon = +$ BH interior

Evolution along r : $d\tau = \sqrt{\frac{2\beta}{\alpha}} N(r) dr$

$$S_\epsilon[\alpha, \beta] = \epsilon c l_P \int d\tau \left[\frac{\epsilon}{l_s^2} - \frac{\epsilon \alpha}{l_\Lambda^2} + \frac{\beta \dot{\alpha}^2 - 2\alpha \dot{\alpha} \dot{\beta}}{2\alpha^2} \right]$$

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The BH Mini-Superspace

This reduces GR to classical mechanics with finite nb of d.o.f.s :

Metric components $\leftarrow S_\epsilon[\alpha, \beta] = \epsilon c l_P \int d\tau \left[\frac{\epsilon}{\ell_s^2} - \frac{\epsilon \alpha}{\ell_\Lambda^2} + \frac{\beta \dot{\alpha}^2 - 2\alpha \dot{\alpha} \dot{\beta}}{2\alpha^2} \right]$

Proper radial coordinate, with lapse hidden

Energy off-set, fixed by fiducial length scale

Potential, given by cosmological constant

Kinetic term, 2d mechanics

Hamiltonian :

$$\mathcal{H} = \mathbf{H}^{(\Lambda)} - \frac{c l_P}{\ell_s^2} \quad \text{with} \quad \mathbf{H}^{(\Lambda)} = -\frac{1}{\epsilon c l_P} \left[\alpha p_\alpha p_\beta + \frac{1}{2} \beta p_\beta^2 \right] + \frac{c l_P}{\ell_\Lambda^2} \alpha$$

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The BH Mini-Superspace

Let us check the symmetries i.e. constants of motion :

- Free theory: given by conformal Killing vectors of field space metric
- Check how non-vanishing potential affects those observables

Part of a larger programme of finding **HIDDEN symmetries of GR**

↳ Hints that there are **MORE** symmetries than gauge sym under spacetime diffeomorphism

↳ Geroch group,
Janis-Newman
algo, ...

Ladder sym
For QNMs

↳ Edge modes,
Boundary
sym alg

Hidden symmetries, beyond Killing
vectors of field space metric

Work in progress with Ben Achour
[1806.09290, 1904.06149, 1909.13390,
2001.11807, 2004.05841, 2110.01466,
2202.12828, 2207.07312, 2302.07644]

and with Geiller [2010.07059,
2107.03878, 2205.02615, 2308.09200]

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The BH Mini-Superspace

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- Free theory: given by conformal Killing vectors of field space metric
- Check how non-vanishing potential affects those observables

Find maximal symmetry algebra :

Schrödinger alg

$$\mathfrak{sh}(2) = (\mathfrak{sl}(2, \mathbb{R}) \oplus \mathfrak{so}(1, 1)) \oplus_s (\mathbb{R}^2 \oplus \mathbb{R}^2)$$

Translations

$$P_+^{(\Lambda)} = \sqrt{\alpha} p_\alpha + \frac{\beta p_\beta}{2\sqrt{\alpha}} - \epsilon \frac{c^2 \ell_P^2}{\ell_\Lambda^2} \frac{\sqrt{\alpha}}{p_\beta}$$

$$P_-^{(\Lambda)} = \sqrt{\alpha} p_\beta$$

Gallilean Boosts

$$B_+^{(\Lambda)} = \epsilon \frac{\beta}{\sqrt{\alpha}} - \frac{2c^2 \ell_P^2}{3\ell_\Lambda^2} \frac{\sqrt{\alpha}}{p_\beta^2} + \frac{1}{c\ell_P} \tau P_+^{(\Lambda)}$$

$$B_-^{(\Lambda)} = \epsilon 2\sqrt{\alpha} + \frac{1}{c\ell_P} \tau P_-$$

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The BH Mini-Superspace

Let us check the symmetries i.e. constants of motion :

- Free theory: given by conformal Killing vectors of field space metric
- Check how non-vanishing potential affects those observables

Find maximal symmetry algebra :

Schrödinger alg

$$\mathfrak{sh}(2) = (\mathfrak{sl}(2, \mathbb{R}) \oplus \mathfrak{so}(1, 1)) \oplus_s (\mathbb{R}^2 \oplus \mathbb{R}^2)$$

Conformal transf

$$\begin{aligned} Q_+^{(\Lambda)} &= c l_P \mathbf{H}^{(\Lambda)}, \\ Q_0^{(\Lambda)} &= D^{(\Lambda)} - \tau \mathbf{H}^{(\Lambda)}, \\ c l_P Q_-^{(\Lambda)} &= -2\epsilon c l_P \beta^{(\Lambda)} - 2\tau D^{(\Lambda)} + \tau^2 \mathbf{H}^{(\Lambda)} \end{aligned}$$

with $D^{(\Lambda)} = (\alpha p_\alpha + \beta p_\beta) - \epsilon \frac{4c^2 \ell_P^2}{3\ell_\Lambda^2} \frac{\alpha}{p_\beta}$

$$\{Q_0, Q_\pm\} = \pm Q_\pm, \quad \{Q_+, Q_-\} = -2Q_0$$

Translations

$$\begin{aligned} P_+^{(\Lambda)} &= \sqrt{\alpha} p_\alpha + \frac{\beta p_\beta}{2\sqrt{\alpha}} - \epsilon \frac{c^2 \ell_P^2}{\ell_\Lambda^2} \frac{\sqrt{\alpha}}{p_\beta} \\ P_-^{(\Lambda)} &= \sqrt{\alpha} p_\beta \end{aligned}$$

so(1,1) boosts

$$J^{(\Lambda)} = 2\alpha p_\alpha - \epsilon \frac{4c^2 \ell_P^2}{3\ell_\Lambda^2} \frac{\alpha}{p_\beta}$$

Gallilean Boosts

$$\begin{aligned} B_+^{(\Lambda)} &= \epsilon \frac{\beta}{\sqrt{\alpha}} - \frac{2c^2 \ell_P^2}{3\ell_\Lambda^2} \frac{\sqrt{\alpha}}{p_\beta^2} + \frac{1}{c l_P} \tau P_+^{(\Lambda)} \\ B_-^{(\Lambda)} &= \epsilon 2\sqrt{\alpha} + \frac{1}{c l_P} \tau P_- \end{aligned}$$

Central charge

$$\begin{aligned} \{P_-, P_+\} &= 0, & \{B_-, B_+\} &= 0, \\ \{B_\pm, P_\pm\} &= 0, & \{B_\pm, P_\mp\} &= \epsilon, \\ \{J, B_\pm\} &= \pm B_\pm, & \{J, P_\pm\} &= \pm P_\pm, \end{aligned}$$

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The BH Mini-Superspace

BH Mass is a constant of motion : $\ell_M = -\epsilon \frac{2\ell_s^3}{c^2 \ell_P^2} J^{(\Lambda)} P_-^{(\Lambda)}$

But not a Casimir of Schrödinger alg, since BH mini-superspace describe phase space for all spherically-sym metrics with arbitrary mass

Now we quantize the model :

↳ Quantize: wave-function of the metric components $\Psi(\alpha, \beta)$

↳ Free Quantum Mechanics : $S[\Psi] = \int d\tau \left[i\hbar \bar{\Psi} \partial_\tau \Psi + \frac{\hbar^2}{c \ell_P} \partial_\tau \bar{\Psi} \partial_\tau \Psi \right]$

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Universal QG Corrections

But, is QG meant to be a mere quantization of GR ...

... or a deformed quantization of GR ?

All approaches always have (at least) another ingredient !

e.g. Area gap, string tension, $1/N$ corrections, sprinkling density, ...

Always about
some discreteness ... !



So ...

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The BH Mini-Superspace

BH Mass is a constant of motion : $l_M = -\epsilon \frac{2l_s^3}{c^2 l_P^2} J^{(\Lambda)} P_-^{(\Lambda)}$

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Universal QG correction preserving Schrodinger symmetry

$$S[\Psi] = \int d\tau \left[i\hbar \bar{\Psi} \partial_\tau \Psi + \frac{\hbar^2}{c l_P} \partial_\tau \bar{\Psi} \partial_\tau \Psi + \kappa^2 |\Psi|^4 \right]$$

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The BH Mini-Superspace

Universal QG correction preserving
Schrodinger symmetry

$$S[\Psi] = \int d\tau \left[i\hbar\bar{\Psi}\partial_\tau\Psi + \frac{\hbar^2}{c\ell_P} \partial_\tau\bar{\Psi}\partial_\tau\Psi + \kappa^2|\Psi|^4 \right]$$

What kind of phenomenology ?

- ↳ Self-interaction of BH wave-function
- ↳ New parameter controls attraction/repulsion of probability peaks !
- ↳ Modulates "evolution" of metric superpositions along radial direction
- ↳ Explores qu → cl transition in QG
- ↳ New type of analogue QG models for BH

Ben Achour, L, Oriti, [2302.07644]

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At the interface Quantum \leftrightarrow Gravity

What's next ?

More work !



The Quantum
of
Gravity ?



BH-WH
oscillation
accelerant

Non-linear corner
inducer of soft
graviton EPR pairs

Holographic
dissolvant of
spacetime bulk

LQG spin oil
to reduce
information
viscosity

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At the interface Quantum \leftrightarrow Gravity

What's next ?

Questions !



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