

Title: General Relativity for Cosmology Lecture - 100523

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Collection: General Relativity for Cosmology

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Abstract: Zoom: <https://pitp.zoom.us/j/91640855624?pwd=dWVWV2doSnBhUS9JUkhjQVBwY0h0dz09>

(source (for essay): see also my paper 1510.02751)

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Then, new:

- Specified $\nabla \Rightarrow$ specified parallel transport in M
 \Rightarrow implicitly specified "shape" of M

Question:How does ∇ determine "shape"? Through:Torsion & Curvature!In General Relativity: one assumes torsionless ∇ , i.e.: $T = 0$.Idea: "(Extended) equivalence principle":

Christoffel Γ will express gravitation and pseudo forces.
 Therefore, we require that around each $p \in M$ there exists a chart so that $\Gamma(p) = 0$ (i.e. no such forces in free fall).

This rules out the existence of torsion:Why? The torsion is a tensor. \Rightarrow It transforms linearly with invertible Jacobian matrices

$$\bar{T}_{jk}^i(\bar{x}) = \frac{\partial \bar{x}^i}{\partial x^a} \frac{\partial x^b}{\partial \bar{x}^j} \frac{\partial x^c}{\partial \bar{x}^k} T_{bc}^a(x)$$

 \Rightarrow If T_{ij} vanishes in one cds, it vanishes in all cds.

$$\begin{aligned} t_{ijk} &:= 2(\Gamma_{ik} - \Gamma_{jk}) \\ C_{ijk} &:= \frac{1}{2}(\Gamma_{ik}^e - \Gamma_{jk}^e) \end{aligned} \quad \left\{ \begin{array}{l} \Gamma_{ij}^k = \Gamma_{(ijk)}^{k\text{ sym}} + \Gamma_{(ijk)}^{k\text{ anti}} \\ \Rightarrow \bar{\Gamma}_{(ijk)}^{k\text{ anti}} = \frac{\partial \bar{x}^e}{\partial x^i} \frac{\partial x^j}{\partial \bar{x}^k} \frac{\partial x^k}{\partial \bar{x}^l} \Gamma_{(ijk)}^{l\text{ sym}} \end{array} \right. !$$

Definition: $T_{ijk}^l := 2 \Gamma_{(ijk)}^{l\text{ sym}}$ is the "Torsion tensor"(Notice: Since Γ is not a tensor, but Γ_{ijk}^l is, T_{ijk}^l is not a tensor)Proposition:Vice versa, if $T_{ijk}^l(x) = 0 \forall x \in M$,then there is for every $p \in M$ a chart with $\Gamma_{ijk}^l(p) = 0$.Recall: ξ is autoparallel to a path $\gamma: t \rightarrow x(t)$ if $\stackrel{M}{T}(M)$

$$\nabla_{\dot{\gamma}} \xi = 0$$

Meaning: ξ is parallel transported along the path γ in M .

$$\text{Explicitly: } \frac{d\xi^k}{dt} + \Gamma_{ij}^k \frac{dx^i}{dt} \xi^j = 0$$

Geodesics:A curve $\gamma: t \rightarrow x(t)$ is called a geodesic if $\dot{\gamma}$ is autoparallel along γ , i.e. if

$$\nabla_{\dot{\gamma}} \dot{\gamma} = 0$$

Meaning: The path γ in M is such that the path's tangent vectors are parallel translates of each other.

Christoffel Γ will express gravitational and pseudo forces.

Therefore, we require that around each pEM there exists a chart so that $\Gamma(p) = 0$ (i.e. no such forces in freefall).

This rules out the existence of torsion:

Why? The torsion is a tensor.

\Rightarrow It transforms linearly with invertible Jacobian matrices

$$\bar{T}_{ijk}(\bar{x}) = \frac{\partial \bar{x}^i}{\partial x^a} \frac{\partial x^b}{\partial \bar{x}^j} \frac{\partial x^c}{\partial \bar{x}^k} T_{abc}(x)$$

\Rightarrow If T_{ij} vanishes in one cds, it vanishes in all cds.

⇒ In charts, geodesics $x^i(t)$ obey:

$$\frac{d^2 x^i}{dt^2} + \Gamma(x)^i{}_{jk} \frac{dx^j}{dt} \frac{dx^k}{dt} = 0 \quad (*)$$

Theory of ordinary differential equations:

Given $p = y(0)$, each initial condition $\xi = y(0)$ belongs to a unique geodesic y_ξ of nonzero length.

Subscript indicates initial condition vector

Notice: If $y_\xi(t)$ solves $(*)$ then $y_\xi(\lambda t)$

also solves $(*)$ and for $\lambda \in \mathbb{R}$:

$$y_\xi(t) = y_\xi(\lambda t) \quad (G)$$

Recall:

ξ is autoparallel to a path $\gamma: t \rightarrow x(t)$ if

$$T'(m)$$

$$\nabla_\gamma \xi = 0$$

Meaning: ξ is parallel transported along the path γ in M .

$$\text{Explicitly: } \frac{d\xi^k}{dt} + \Gamma^k{}_{ij} \frac{dx^i}{dt} \xi^j = 0$$

Geodesics:

A curve $\gamma: t \rightarrow x(t)$ is called a geodesic if ξ is autoparallel along γ , i.e. if

$$\nabla_\gamma \xi = 0$$

Meaning: The path γ in M is such that the path's tangent vectors are parallel translates of each other.

"Exponential map"

Consider a fixed point $p \in M$.

The exponential map is defined through:

$$\exp_p: T_p(M) \rightarrow M \quad \begin{smallmatrix} \text{(rally from a neighborhood} \\ \text{of 0 in } T_p(M) \text{ to a neighborhood} \\ \text{of } p \in M) \end{smallmatrix}$$

$$\exp_p: \xi \rightarrow \exp_p(\xi) := y_\xi(1)$$

where y is the geodesic with $y_\xi(0) = p, y'_\xi(0) = \xi$.

Observe:

From (G) we obtain:

$$\nu(\lambda) = \nu(1) = \exp_p(\lambda \nu) \quad (E)$$

\Rightarrow Given $p = \gamma(0)$, each initial condition $\xi = \dot{\gamma}(0)$ belongs to a unique geodesic γ_ξ of nonzero length.

script indicates initial condition vector

□ Notice: If $\gamma_\xi(t)$ solves (*) then $\gamma_\xi(\lambda t)$

also solves (*) and for $\lambda \in \mathbb{R}$:

$$\gamma_\xi(t) = \gamma_\xi(\lambda t) \quad (G)$$

(Exercise: verify)

"Geodesic" or "Riemann normal" coordinates:

□ \exp_p is a diffeomorphism from a neighborhood of $0 \in T_p(M) \cong \mathbb{R}^n$ into a neighborhood of the point $p \in M$.

$\Rightarrow \exp_p$ provides a chart around p :

□ Choose a basis, say e_1, e_2, \dots, e_n of $T_p(M)$, then:

$$\xi = \xi^i e_i$$

□ Through \exp_p , the ξ^i become the coordinates of points in a neighborhood of $p \in M$:

$$(\xi^1, \dots, \xi^n) \rightarrow \exp_p(\xi^i e_i) \in M$$

□ These $\{\xi^i\}$ are called "normal" or "geodesic coordinates."

$$\exp_p : T_p(M) \rightarrow M \quad (\forall p \in M)$$

$$\exp_p : \xi \rightarrow \exp_p(\xi) := \gamma_\xi(1)$$

where γ is the geodesic with $\gamma_\xi(0) = p, \dot{\gamma}_\xi(0) = \xi$.

□ Observe:

From (G) we obtain:

$$\gamma_\xi(\lambda) = \gamma_\xi(1) = \exp_p(\lambda \xi) \quad (E)$$

\Rightarrow Geodesics, γ , through p are straight lines in a normal cds about p !

□ Recall (E):

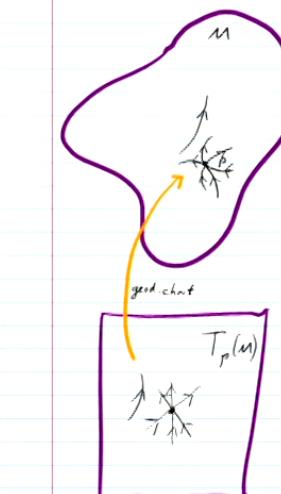
$$\gamma_\xi(\lambda) = \exp_p(\lambda \xi)$$

for varying λ one moves along the geodesic in M

for varying λ one moves along a straight line in the coordinate system of the ξ^i !

□ Thus: In geodesic cds, geodesics through p are straight lines of constant velocity ξ .

□ Does this mean $\Gamma_{i,j}^k(p) = 0$? No!



□ Choose a basis, say e_1, e_2, \dots, e_n of $T_p(M)$, then:

$$\xi = \xi^i e_i$$

□ Through exp., the ξ^i become the coordinates of points in a neighborhood of $p \in M$:

$$(\xi^1, \dots, \xi^n) \rightarrow \exp_p(\xi^i e_i) \in M$$

□ These $\{\xi^i\}$ are called "normal" or "geodesic coordinates."

Geodesic eqn. at p:

$$\frac{d^2 x^k}{dt^2} + \Gamma^k_{ij}(p) \dot{x}^i \dot{x}^j = 0$$

Thus: $(\Gamma^k_{1j} \xi_1(p) + \Gamma^k_{2j} \xi_2(p)) \frac{dx^i}{dt} \frac{dx^j}{dt} = 0$

$\xi = 0$ because of type (antisymmetric), Γ^k_{ij} (symmetric)¹³

$$\Rightarrow \Gamma^k_{ij} \xi_j(p) = 0 \text{ in geodesic cds.}$$

⇒ Indeed: If the torsion vanishes, $\Gamma^k_{ij} \xi_j(p) = \frac{1}{2} T^k_{ij}(p) = 0$ then for each $p \in M$ there exists a chart in which the entire gravity and pseudo force field vanishes at p :

$$\Gamma^k_{ij} \xi_j(p) = 0$$

Note:
Quantum fluctuations
may induce torsion!
So, let's nevertheless ask:

What would torsion mean, geometrically?



along the geodesic
in M

a straight line
in the coordinate
system of the chart!

□ Thus: In geodesic cds, geodesics through p are straight lines of constant velocity ξ .

□ Does this mean $\Gamma^k_{ij}(p) = 0$? No!

Abstract definition of Torsion:

□ Assume ξ_1 and ξ_2 are tangent vectors at $p \in M$:

Then, the Torsion map is defined as:

$$\mathcal{T}: T_p(M) \times T_p(M) \rightarrow T_p(M)$$

The will be the amount by which an infinitesimal parallelogram spanned by ξ_1 and ξ_2 does not close.

$$\mathcal{T}: \xi_1, \xi_2 \rightarrow \mathcal{T}(\xi_1, \xi_2) := \underbrace{\nabla_{\xi_1} \xi_2 - \nabla_{\xi_2} \xi_1}_{\text{we could also write: } = \omega(\mathcal{T}(\xi_1, \xi_2))} - [\xi_1, \xi_2]$$

□ It is used to define the Torsion tensor, J ,

$$J \in T_{p,2}(M)$$

through:

looking for a vector
such that
a (1,2) brace yields a number

$$\rightarrow J(\omega, \xi_1, \xi_2) := \underbrace{\langle \omega, \mathcal{T}(\xi_1, \xi_2) \rangle}_{\text{contraction yields a number}} \in \mathbb{R}$$

$$\Rightarrow \Gamma_{ij}^k(p) = 0 \text{ in geodesic cds.}$$

\Rightarrow Indeed: If the torsion vanishes, $\Gamma_{ij}^k(p) = \frac{1}{2} T_{ij}^k(p) = 0$ then for each $p \in M$ there exists a chart in which the entire gravity and pseudo force field vanishes at p :

$$\Gamma_{ij}^k(p) = 0$$

Note:
Quantum fluctuations
may induce torsion!
So, let's nonetheless ask:

What would torsion mean, geometrically?

Compare with prior definition:

□ Choose canonical bases $w := dx^k$, $\xi_1 := \frac{\partial}{\partial x^1}$, $\xi_2 := \frac{\partial}{\partial x^2}$:

$$\begin{aligned} \square \quad T_{ij}^k &:= dx^k(T(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j})) \\ &= \langle dx^k, T(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}) \rangle \quad (\text{more convenient notation}) \\ &= \langle dx^k, \underbrace{\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j}}_{\text{Recall: } \frac{\partial^2}{\partial x^i \partial x^j} = \Gamma_{ij}^k} - \underbrace{\nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^i}}_{\frac{\partial^2}{\partial x^j \partial x^i} f = 0 \quad \forall f} - [\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}] \rangle \\ &= \langle dx^k, \Gamma_{ij}^k \frac{\partial}{\partial x^k} - \Gamma_{ji}^k \frac{\partial}{\partial x^k} \rangle = \Gamma_{ij}^k \delta_{ki} - \Gamma_{ji}^k \delta_{ki} \end{aligned}$$

$$\Rightarrow T_{ij}^k = \Gamma_{ij}^k - \Gamma_{ji}^k$$

$$S: T_p(M) \times T_p(M) \rightarrow T_p(M) \quad \text{parallelogram spanned by } \xi_1 \text{ and } \xi_2 \text{ does not close.}$$

$$T: \xi_1, \xi_2 \rightarrow T(\xi_1, \xi_2) := \nabla_{\xi_1} \xi_2 - \nabla_{\xi_2} \xi_1 - [\xi_1, \xi_2]$$

□ It is used to define the Torsion tensor, T ,

$$T \in T_{p,2}(M)$$

through:

feeding 1 covariant vector
to a (1,1) bimod yields a number

$$T(w, \xi_1, \xi_2) := \underbrace{\langle w, T(\xi_1, \xi_2) \rangle}_{\substack{\in T_p(M) \\ \in T_p(M)}} \in \mathbb{R}$$

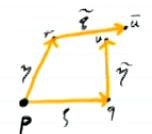
for prof it's a tensor, i.e. Strain tensor

we could also write: $= w(T(\xi_1, \xi_2))$

contraction yields
a number

Geometric meaning of torsion? Parallelograms would not close!

Travel from p infinitesimally in ξ and then γ direction, and compare with the reverse. (In flat space: $x' + \gamma' + \xi' = x' + \xi' + \gamma'$)



$$\begin{aligned} \xi, \gamma &\in T_p \\ \tilde{\xi} &\in T_p \\ \tilde{\gamma} &\in T_q \end{aligned}$$

Result parallel transport: $\nabla_{\tilde{\xi}} v = 0$

$$\frac{dv^k}{dt} + \Gamma_{ij}^k \frac{dx^i}{dt} v^j = 0$$

$$\tilde{\xi}(r) = ?$$

$$\begin{aligned} \tilde{\xi}^k(x + \gamma) &\approx \xi^k(x) + \frac{d\xi^k}{dt}(x) \\ &= \xi^k(x) - \Gamma(x)^k_{ij} \gamma^i \xi^j \end{aligned}$$

$$\Rightarrow (\text{ds. of } \bar{u}: x^i + \gamma^i + \xi^i - \Gamma(x)^i_{ij} \gamma^j \xi^i)$$



$$\square J_{ij} := dx^k \left(J \left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right) \right)$$

$$= \langle dx^k, J \left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right) \rangle \quad (\text{more convenient notation})$$

$$= \langle dx^k, \underbrace{\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} - \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^i}}_{\text{Recall: } \frac{\partial^2}{\partial x^i \partial x^j} = \Gamma_{ij}^k \frac{\partial}{\partial x^k}} - [\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}] \rangle$$

$$\left(\frac{\partial^2}{\partial x^i \partial x^j} - \frac{\partial^2}{\partial x^j \partial x^i} \right) f = 0 \quad \forall f$$

$$= \langle dx^k, \Gamma_{ij}^k \frac{\partial}{\partial x^i} - \Gamma_{ji}^k \frac{\partial}{\partial x^j} \rangle = \Gamma_{ij}^k \delta_{ji}^k - \Gamma_{ji}^k \delta_{ij}^k$$

$$\square \Rightarrow J^k_{ij} = \Gamma_{ij}^k - \Gamma_{ji}^k$$

Analogously obtain: (ds. of u : $x^i + \xi^i + \gamma^a - \Gamma(x)^k_{ij} \xi^j \gamma^i$)

Torsion!

$$\Rightarrow \text{Cd. distance from } u \text{ to } \bar{u} \text{ is: } (\Gamma(u)^k_{ji} - \Gamma(u)^k_{ij}) \gamma^i \xi^j = T_{ji}^k \xi^j$$

Comment: We had:

$$\hat{\xi}^k(x + \gamma) \approx \xi^k(x) + \frac{d\xi^k}{dt}(t) = \xi^k(x) - \Gamma(x)^k_{ij} \gamma^i \xi^j$$

$$\text{this is also:} \quad = \xi^k(x) - (\gamma^i \xi^k_{,i} + \Gamma(x)^k_{ij} \gamma^i \xi^j) + \gamma^i \xi^j_{,i}$$

$$= \xi^k(x) - \gamma^i \xi^k_{,i} + \gamma^i \xi^j_{,j}$$

Thus: cd distance from u to \bar{u} is:

$$(x^i + \xi^i + \gamma^a - \gamma^i \xi^k_{,i} + \gamma^i \xi^j_{,j}) - (x^i - \xi^i - \gamma^i + \xi^i \gamma^j_{,j} - \gamma^i \xi^j_{,j}) = T_{ji}^k \xi^j$$

Recall that indeed: $J: \gamma, \xi \rightarrow J(\gamma, \xi) = \nabla_\gamma \xi - \nabla_\xi \gamma - [\gamma, \xi]$



$$\beta, \gamma \in T_p^1 \\ \tilde{\xi} \in T_p^1 \\ \tilde{\gamma} \in T_q^1$$

Recall parallel transport: $\nabla_\gamma v = 0$

$$\frac{dv^k}{dt} + \Gamma_{ij}^k \frac{dx^i}{dt} v^j = 0$$

$$\tilde{\xi}(r) = ?$$

$$\tilde{\xi}^k(x + \gamma) \approx \xi^k(x) + \frac{d\xi^k}{dt}(t) \quad \text{Now use } v = \xi, \frac{dx}{dt} = \gamma^i.$$

$$= \xi^k(x) - \Gamma(x)^k_{ij} \gamma^i \xi^j$$

$$\Rightarrow \text{(ds. of } \bar{u}: x^i + \xi^i + \gamma^a - \Gamma(x)^k_{ij} \gamma^i \xi^j \text{)}$$

Curvature:

Assume ξ_1, ξ_2 and ξ_3 are tangent vectors at $p \in M$.

\square The curvature map, R , is defined through:

$$R: \xi_1, \xi_2, \xi_3 \rightarrow R(\xi_1, \xi_2) \xi_3 = \underbrace{(\nabla_{\xi_1} \nabla_{\xi_2} - \nabla_{\xi_2} \nabla_{\xi_1} - [\nabla_{\xi_1}, \nabla_{\xi_2}]) \xi_3}_{\text{an operator, or map, acting on } \xi_3} \in T_p^1(M)$$

\square It defines the curvature tensor, R ,

$$R \in T_p^1(M) \quad \text{can be fed one vector and 3 vectors to yield a number}$$

through:

$$R(\omega, \xi_1, \xi_2, \xi_3) := \underbrace{\langle \omega, R(\xi_1, \xi_2) \xi_3 \rangle}_{\omega(R(\xi_1, \xi_2) \xi_3)} \in \mathbb{R}$$

$$\hat{\xi}^k(x+\gamma) \approx \xi^k(x) + \frac{d\xi^k}{dt}(x) = \xi^k(x) - \Gamma(x)^k_{ij} \gamma^i \xi^j$$

this is also:

$$\begin{aligned} &= \xi^k(x) - (\gamma^i \xi^k_{,i} + \Gamma(x)^k_{ij} \gamma^i \xi^j) + \gamma^i \xi^j_{,i} \\ &= \xi^k(x) - \gamma^i \xi^k_{,i} + \gamma^i \xi^j_{,j} \end{aligned}$$

Thus: cd distance from x to \tilde{x} is:

$$(x^a + \gamma^a + g^a - \gamma^i \xi^k_{,i} + \gamma^i \xi^k) - (x^a - g^a - \gamma^a + \xi^i \gamma^k_{,i} - \gamma^i \xi^k_{,i}) = \Gamma^a_{ijk} \gamma^i \xi^j$$

Recall that indeed: $T: \gamma, \xi \rightarrow T(\gamma, \xi) = \nabla_\gamma \xi - \nabla_\xi \gamma - [\gamma, \xi]$

In a chart:

$$\begin{aligned} R^i_{jkl} &= \langle dx^i, R\left(\frac{\partial}{\partial x^j}, \frac{\partial}{\partial x^l}\right)\frac{\partial}{\partial x^k} \rangle \\ &= \langle dx^i, \left(\underbrace{\frac{\partial^2}{\partial x^j \partial x^l} - \frac{\partial^2}{\partial x^l \partial x^j}}_{=0} - \frac{\partial^2}{\partial x^k \partial x^l} \right) \frac{\partial}{\partial x^i} \rangle \\ &= \langle dx^i, \underbrace{\frac{\partial^2}{\partial x^j \partial x^k} \frac{\partial}{\partial x^l}}_{=0} - \underbrace{\frac{\partial^2}{\partial x^k \partial x^l} \frac{\partial}{\partial x^j}}_{=0} \rangle \\ &= \langle dx^i, \left(\underbrace{\Gamma^s_{jik} + \Gamma^s_{kj} \Gamma^i_{ls} - \Gamma^s_{sik} - \Gamma^s_{ks} \Gamma^i_{ls}}_{=0} \right) \frac{\partial}{\partial x^j} \rangle \\ &= \Gamma^i_{jik} - \Gamma^i_{kjl} + \underbrace{\Gamma^s_{kj} \Gamma^i_{ls} - \Gamma^s_{ks} \Gamma^i_{ls}}_{(\text{at origin of geodesic cds they vanish})} \end{aligned}$$

$$R: \xi_1, \xi_2, \xi_3 \rightarrow R(\xi_1, \xi_2) \xi_3 = \underbrace{(\nabla_{\xi_1} \nabla_{\xi_2} - \nabla_{\xi_2} \nabla_{\xi_1} - [\nabla_{\xi_1}, \nabla_{\xi_2}]) \xi_3}_{\text{an operator, or map, acting on } \xi_3}$$

It defines the curvature tensor, R ,

\hookrightarrow can be fed one vector and 3 vectors to yield a number
 $R \in T^1_3(M)$

through:

$$R(\omega, \xi_1, \xi_2, \xi_3) := \underbrace{\langle \omega, R(\xi_1, \xi_2) \xi_3 \rangle}_{\omega(R(\xi_1, \xi_2) \xi_3)} \in \mathbb{R}$$

Curvature tensor's meaning?

Intuition:

- Contains derivations of Γ \Rightarrow
- expresses variation in gravitational forces \Rightarrow
- expresses the strength and direction of "tidal forces".

Geometry:

- Curvature expresses noncommutativity of two parallel transports, namely:

$$\begin{aligned}
 &= \left\langle dx^i, \underbrace{\nabla_{x^j} \Gamma_{ij}^k}_{=0} \frac{\partial}{\partial x^k} - \nabla_{x^j} \Gamma_{ik}^l \frac{\partial}{\partial x^l} \right\rangle \\
 &= \left\langle dx^i, \left(\Gamma_{ij,k}^l + \Gamma_{ij}^m \Gamma_{mk}^l - \Gamma_{ik,j}^l - \Gamma_{kj,i}^l \right) \frac{\partial}{\partial x^l} \right\rangle \\
 &= \Gamma_{ij,k}^l - \Gamma_{ik,j}^l + \underbrace{\Gamma_{ij}^m \Gamma_{mk}^l - \Gamma_{kj,i}^l}_{(\text{at origin of geodesic } \dot{c} \text{ do they vanish})}
 \end{aligned}$$

Proposition: (Ricci Identity)

Assume the torsion vanishes and that ξ is a vector field. Then:

$$\xi^a_{;cd} - \xi^a_{;dc} = R^a_{bcd} \xi^b$$

(here: $\xi^a_{;cd} := \xi^a_{;c;d}$ etc.)

Remark: (absl. necessary to derive because need Taylor expansion)
See, e.g., but by Stewarts Einstein

It implies that for parallel transport along infinitesimal parallelogram:

$$(\hat{s}-s)^a \approx \eta^a \nu^c R^a_{bcd} \xi^d$$



expresses the strength and direction of "tidal forces".

geometry:

Curvature expresses noncommutativity of two parallel transports, namely:

Proof of Ricci identity:

Assume ξ, η, ν are vector fields.

Then, $R(\xi, \eta)\nu := \nabla_\xi (\nabla_\eta \nu) - \nabla_\eta (\nabla_\xi \nu) - \nabla_{[\xi, \eta]} \nu$ reads

use: $\nabla_\eta \nu = \eta^i \nu_{,i} = \eta^i \nu_{,i} \left(\frac{\partial^2 \nu}{\partial x^i \partial x^j} \right) = \eta^i \nu_{,i} \left(\frac{\partial^2 \nu}{\partial x^j \partial x^i} \right) = \eta^i \nu_{,i} \left(\frac{\partial^2 \nu}{\partial x^j \partial x^i} \right) \dots$

$$\begin{aligned}
 R^a_{bcd} \xi^b \eta^c \nu^d &= (\nu_{,d} \eta^c)_{,b} \xi^b - (\nu_{,b} \eta^c)_{,d} \xi^b \\
 &\quad - \nu_{,d} (\eta^c \xi^b - \xi^c \eta^b)
 \end{aligned}$$

used Torsion = $T(\xi, \eta) = \eta^i \xi_{,i} - \xi^i \eta_{,i} - [\xi, \eta] = 0$
i.e. $[\xi, \eta] = \nabla_\xi \eta - \nabla_\eta \xi$

Terms cancel:

$$\Rightarrow R^a_{bcd} \xi^b \eta^c \nu^d = (\nu_{,d} \eta^c - \nu_{,b} \eta^d) \xi^b \eta^c$$

True $\forall \xi, \eta \Rightarrow R^a_{bcd} \nu^d = \nu^a_{,b;c} - \nu^a_{,c;b} \quad \checkmark$

$$\xi^a_{;bcd} - \xi^a_{;cad} = R^a_{bcd} \xi^b$$

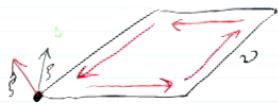
(here: $\xi^a_{;cd} := \xi^a_{;c;d}$ etc.)

Remark:

(had to derive because need Taylor expansion, i.e., e.g., but by Stewart or Einstein)

It implies that for parallel transport along infinitesimal parallelogram:

$$(\hat{\xi} - \xi)^a \approx \gamma^b \nabla^c R^a_{bcd} \xi^b$$



The "Bianchi Identities":

- They are automatic relations among torsion and curvature, by construction.

Preparation: \triangleright for maps!

Consider an arbitrary $T_p(M)$ -linear map:

$$K: \underbrace{\xi_1 \times \xi_2 \times \dots \times \xi_r}_{\text{tangent vector}} \rightarrow \underbrace{K(\xi_1, \dots, \xi_r)}_{\text{tangent vector}} \quad (\text{e.g.: Torsion or curvature map})$$

i.e. at each $p \in M$:

$$K: T_p(M)^r \rightarrow T_p(M)^1$$

$$\text{in basis: } R^a_{bcd} \xi^b \gamma^c v^d = (\nu^a_{id} \gamma^d)_{;cd} - (\nu^a_{id} \xi^d)_{;cd} - \nu^a_{id} (\gamma^c \xi^d - \xi^c \gamma^d)$$

used $Torsion = T(\xi_i, \xi_j) = \nu_{ij} \xi_k - \nu_{ji} \xi_k - [\xi_i, \xi_j] = 0$
i.e. $[\xi_i, \xi_j] = \nabla_{\xi_i} \xi_j - \nabla_{\xi_j} \xi_i$

Terms cancel:

$$\Rightarrow R^a_{bcd} \xi^b \gamma^c v^d = (\nu^a_{id} \xi^d)_{;cd} - \nu^a_{id} \xi^d$$

True $\forall \xi, \gamma \Rightarrow R^a_{bcd} v^d = \nu^a_{id} \xi^d - \nu^a_{id} \xi^d \checkmark$

□ We can view K as a tensor $\tilde{K} \in T_p(M)^1$,

namely:

$$\tilde{K}(w, \xi_1, \dots, \xi_r) := \langle w, K(\xi_1, \dots, \xi_r) \rangle$$

□ Now let the usual derivative of the tensor R define the derivative of the map K :

$$\langle w, (\nabla_{\xi} K)(\xi_1, \dots, \xi_r) \rangle := \underbrace{\nabla_{\xi} \tilde{K}(w, \xi_1, \dots, \xi_r)}_{\text{new concept: covariant derivative of a map } K: T_p(M)^r \rightarrow T_p(M)^1}$$

usual cov. derivative of a $(1, r)$ tensor when fed one covariant vector

Using \triangleright for map:

Preparation: ∇ for maps!

Consider an arbitrary $\mathbb{F}(M)$ -linear map:

$$K: \underbrace{\mathfrak{g}_1 \times \mathfrak{g}_2 \times \dots \times \mathfrak{g}_r}_{\text{tangent vectors}} \xrightarrow{\substack{\text{(e.g., Torsion or curvature map)} \\ \text{ }} } \underbrace{K(\mathfrak{g}_1, \dots, \mathfrak{g}_r)}_{\text{tangent vector}}$$

i.e. at each $p \in M$:

$$K: T_p(M)^r \rightarrow T_p(M)^1$$

1st Bianchi Identity:

$$\sum_{\text{cyclic}} R(\mathfrak{g}, \gamma) v = \sum_{\text{cyclic}} (\mathcal{T}(\mathcal{T}(\mathfrak{g}, \gamma), v) + \nabla_{\mathfrak{g}} \mathcal{T})(\gamma, v)$$

2nd Bianchi Identity:

$$\sum_{\text{cyclic}} ((\nabla_{\mathfrak{g}} R)(\gamma, v) + R(\mathcal{T}(\mathfrak{g}, \gamma), v)) = 0$$

with obvious simplification in case $\mathcal{T} = 0$.

Note: They are automatically obeyed equations, just like any set of lin. operators obeys the Jacobi identity with respect to $[,]$. Indeed that's why:

Now let the usual derivative of the tensor R define the derivative of the map K :

$$\langle w, (\nabla_{\mathfrak{g}} K)(\mathfrak{g}_1, \dots, \mathfrak{g}_r) \rangle := \nabla_{\mathfrak{g}} \tilde{K}(w, \mathfrak{g}_1, \dots, \mathfrak{g}_r)$$

new concept:
covariant derivative
of a map $K: T_p(M)^r \rightarrow T_p(M)^1$

usual cov. derivative
of a $(1, r)$ tensor
when fed one covector & r vectors

Using ∇ for map:

Proof of 1st Bianchi: (assuming no torsion)

$$\sum_{\text{cyclic}} R(\mathfrak{g}, \gamma) v = 0$$

Indeed:

$$(\nabla_{\mathfrak{g}} \nabla_{\gamma} - \nabla_{\gamma} \nabla_{\mathfrak{g}}) v - \nabla_{[\mathfrak{g}, \gamma]} v + \text{cyclic}$$

↓ skip by 1 cyclically ↓ step by 1 cyclically

$$= \nabla_{\mathfrak{g}} (\nabla_{\gamma} v - \nabla_{\gamma} v) - \nabla_{[\mathfrak{g}, \gamma]} v + \text{cyclic}$$

Exercise: Prove that: $\nabla_{\gamma} v - \nabla_{\gamma} v = [\gamma, v]$ (easy!) without torsion

$$= \nabla_{\mathfrak{g}} [\gamma, v] - \nabla_{[\mathfrak{g}, \gamma]} v + \text{cyclic}$$

↓ because again $a b - b a = [a, b]$

$$= [\mathfrak{g}, [\gamma, v]] + \text{cyclic}$$

2nd Bianchi Identity:

$$\sum_{\text{cyclic}} \left((\nabla_\gamma R)(\eta, v) + R(\tau(\delta, \eta), v) \right) = 0$$

with obvious simplification in case $\tau = 0$.
↙ or like the homogeneous Maxwell equations

Note: They are automatically obeyed equations, just like any set of lin. operators obeys the Jacobi identity with respect to $[,]$. Indeed that's why:

$= 0$ by Jacobi identity for all lin. maps.

Recall:

Assume A, B, C are linear maps $V \rightarrow V$

$$\text{Then: } [A, [B, C]] + [C, [A, B]] + [B, [C, A]] = 0$$

i.e., the Jacobi identity holds.

Proof: Simply spell out the commutators.

Remark: This means that, e.g., in quantum mechanics, all objects that

This is why the job that need to be representable as operators on the Hilbert space identity is one of the axioms of Lie Algebras. must obey the Jacobi identity, e.g., generators of symmetries.

$$\begin{aligned}
 & (\nabla_\delta \nabla_\gamma - \nabla_\gamma \nabla_\delta) v - \nabla_{[\delta, \gamma]} v + \text{cyclic} \\
 & \quad \downarrow \text{skip by 1 cyclically} \quad \downarrow \text{skip by 1 cyclically} \\
 & = \nabla_\delta (\nabla_\gamma v - \nabla_\gamma \gamma) - \nabla_{[\gamma, v]} \delta + \text{cyclic} \\
 & \text{Exercise: Prove that: } \nabla_\gamma v - \nabla_v \gamma = [\gamma, v] \quad \text{[easy!]} \\
 & = \underbrace{\nabla_\delta [\gamma, v] - \nabla_{[\gamma, v]} \delta}_{\text{!! because again } \nabla_b - \nabla_a b = [a, b]} + \text{cyclic} \\
 & = [\delta, [\gamma, v]] + \text{cyclic}
 \end{aligned}$$