

Title: Research Talk 9 - Black hole cohomologies in N=4 Yang-Mills

Speakers:

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Abstract: I will explain a cohomology problem in the $SU(N)$ maximal super-Yang-Mills theory which we expect to capture BPS black holes in AdS. I also explain the recent progress for the $SU(2)$ theory.

Black hole cohomologies in $\mathcal{N} = 4$ Yang-Mills

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Talk based on collaborations with

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- Eunwoo Lee (Seoul National Univ.)
- Siyul Lee (Univ. of Michigan)
- Jaemo Park (Postech)

“The shape of non-graviton operators for SU(2)” [arXiv:2209.12696](#).

“Towards quantum black hole microstates” [arXiv:2304.10155](#).

See also:

- Chi-Ming Chang, Ying-Hsuan Lin,
“Words to describe a black hole” [arXiv:2209.06728](#).
- Budzik, Gaiotto, Kulp, Williams, Wu, Yu,
“Semi-chiral operators in 4d N=1 gauge theories” [arXiv:2306.01039](#).
- Chang, Feng, Lin, Tao,
“Decoding stringy near-supersymmetric black holes” [arXiv:2306.04673](#).
- Budzik, Murali, Vieira,
“Following black hole states” [arXiv:2306.04693](#).

The problem & motivations

The problem: **1/16-BPS states** in 4d maximal SYM w/ $SU(N)$ gauge group

- Among $16Q + 16S$ supercharges, pick a pair Q & $S = Q^\dagger$ that the BPS states preserve.
- BPS states saturate the bound: $\{Q, Q^\dagger\} \sim E - (R_1 + R_2 + R_3 + J_1 + J_2) \geq 0$
- Nilpotency $Q^2 = 0$: **BPS states** \sim **harmonic forms** $\xleftrightarrow{1 \text{ to } 1}$ **Q-cohomology classes**.
- We shall find & discuss the representatives of new cohomology classes.

Motivations:

- Special case of 1/4-BPS states in $N = 1$ QFT: Information on SUSY dynamics
[Budzik, Gaiotto, Kulp, Williams, Wu, Yu] (2023)
- Our main motivation: **BPS black hole microstates** in $AdS_5 \times S^5$
→ Requires studying **strong coupling** & **large N** .

I will explain recent (perhaps modest) progress in this program.

- Cohomologies at **weak-coupling** (1-loop $\sim O(g_{YM}^2)$)
- Want to eventually study $SU(N \gg 1)$. \leftrightarrow Today, I will report **SU(2)** & a bit of **SU(3)**.
- Qualitative features & rough comparison w/ “gravity dual” (\leftarrow mostly omitted today)

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Historical remarks

2004~ : [Gutowski, Reall] ... Constructions of BPS black holes in $AdS_5 \times S^5$

2005~ : [Romelsberger] [Kinney, Maldacena, Minwalla, Raju] ...

Efficient way to study BPS spectrum → the “superconformal index”

- At that time, it was unclear how to see black holes from this index.
- Direct studies w/ **weak-coupling cohomology**: No new solutions found beyond BPS gravitons.
[Berkooz, Reichmann, Simon] (2006) [Grant, Grassi, SK, Minwalla] (2008) [Chang, Yin] (2013) ...

2018 ~ : [Cabo Bizet, Cassani, Martelli, Murthy] [Choi, J. Kim, SK, Nahmgoong] [Benini, Milan] ...

Realized how to see black holes from the index. (I.e. computed their entropy.)

- This index is coupling-independent → Counts weak-coupling cohomologies.
- Since it counted black holes, we became confident that new cohomologies should exist.

2022 ~ : [Chang, Lin] [Choi, SK, E. Lee, Park] ...

With these recent realizations & promises, revisited the cohomology problem.

- Especially because the cohomologies carry more information than the index.

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Weak-coupling setup

SU(N) maximal SYM on R^4 : all fields in adjoint rep. (written in N=1 language)

3 chiral multiplets: $\phi_m, \bar{\phi}^m$ and $\psi_{m\alpha}, \bar{\psi}^m_{\dot{\alpha}}$ ($m = 1,2,3$)

vector multiplet: $A_\mu \sim A_{\alpha\beta}$ and $\lambda_\alpha, \bar{\lambda}_{\dot{\alpha}}$ ($\mu = 1, \dots, 4$) ($\alpha = \pm, \dot{\alpha} = \dot{\pm}$)

Gauge-invariant local BPS operators: (at $x^\mu = 0$ on R^4)

- Free limit ($g_{YM} \rightarrow 0$): Any gauge-invariants of the **invariant fields** under Q & Q^\dagger :

$\bar{\phi}^m, \psi_{m+}, \bar{\lambda}_{\dot{\alpha}}, f_{++} \equiv F_{1+i2, 3+i4}$ & derivatives $\partial_{1+i2} \equiv \partial_1 - i\partial_2, \partial_{3+i4} \equiv \partial_3 - i\partial_4$ acting on them

- Not all of them are BPS at $g_{YM} \neq 0$: At small $g_{YM} \ll 1$, classical SUSY on them reads

$Q \bar{\phi}^m = 0, Q \psi_{m+} \sim g_{YM} \epsilon_{mnp} [\bar{\phi}^n, \bar{\phi}^p], Q f_{++} \sim g_{YM} [\psi_{m+}, \bar{\phi}^m], Q \bar{\lambda}_{\dot{\alpha}} = 0, [Q, D_{+\dot{\alpha}}] \sim g_{YM} [\bar{\lambda}_{\dot{\alpha}}, \cdot]$

- Classical Q & Q^\dagger at $\frac{1}{2}$ -loop \rightarrow Anomalous dimension $QQ^\dagger + Q^\dagger Q \sim E - E_{BPS}$ at **1-loop**, $O(g_{YM}^2)$.

Goal: Find **free BPS operators which remain BPS at 1-loop level**.

- Remain BPS at strong-coupling? A non-renormalization theorem pursued. [Chang, Lin] (2022)
- But their index captures BH's \rightarrow So empirically, **many of them should remain BPS**.
- So expecting "certain" non-renormalization, we study 1-loop cohomologies.

The strategy

The problem at finite $N = 2, 3, \dots$:

- Construct all cohomologies at given charges, and mod out those from **BPS multi-gravitons**
→ [Chang, Lin] (2022) did it & found the first non-graviton cohomology for SU(2).

What are the “**BPS gravitons at finite N** ” ...?

- Chiral primaries $\text{tr}[\bar{\phi}^{(m_1)} \dots \bar{\phi}^{(m_n)}]$, descendants (1-particle) & their products (multi-particle)
- Finite N subtlety: # of independent operators reduces due to **trace relations**
→ QFT dual of “**stringy exclusion principle**” due to **giant gravitons**
 - [Maldacena, Strominger] (1998) [Susskind, McGreevy, Toumbas] (2000)
 - [Jevicki, Ramgoolam] (1999) [Grisaru, Myers, Tafjord] (2000)
 - [Ho, Ramgoolom, Tatar] (1999) [Hashimoto, Hirano, Itzhaki] (2000)
- So they reflect all the expected BPS graviton physics, including the finite N effects.

Remainders have chances to describe black holes. (→ Call “**black hole operators**”)

- Conservatively, results at low $N = 2, 3, \dots$ are just intermediate steps towards large N .
- Progressively, lessons on black holes in wildly quantum ($G_N \sim 1/N^2$) gravity?

A streamlined strategy

To quickly diagnose which charge sectors host new operators, we use the index:

- Grade operators with charges appearing in the index, say with

$$j \equiv 6(R + J) = 2(R_1 + R_2 + R_3) + 3(J_1 + J_2) \geq 0.$$

- Compute the full index & that over gravitons \rightarrow Subtract. [Choi, SK, E. Lee, S. Lee, Park] (2023)

SU(2) \rightarrow Study the index $Z(t) = \text{Tr}[(-1)^F t^j]$

- In full generality, only have series expansions till certain order:

$$\begin{aligned} Z(t) &= 1 + 6t^4 - 6t^5 - 7t^6 + 18t^7 + 6t^8 - 36t^9 + 6t^{10} + 84t^{11} - 80t^{12} - 132t^{13} + 309t^{14} - 18t^{15} - 567t^{16} \\ &\quad + 516t^{17} + 613t^{18} - 1392t^{19} - 180t^{20} + 2884t^{21} - 1926t^{22} - 4242t^{23} + 7890t^{24} + 792t^{25} - 15876t^{26} \\ &\quad + 13804t^{27} + 15177t^{28} - 37536t^{29} + 7049t^{30} + 57522t^{31} - 58704t^{32} + \dots \\ Z_{\text{grav}}(t) &= 1 + 6t^4 - 6t^5 - 7t^6 + 18t^7 + 6t^8 - 36t^9 + 6t^{10} + 84t^{11} - 80t^{12} - 132t^{13} + 309t^{14} - 18t^{15} - 567t^{16} \\ &\quad + 516t^{17} + 613t^{18} - 1392t^{19} - 180t^{20} + 2884t^{21} - 1926t^{22} - 4242t^{23} + 7891t^{24} + 786t^{25} - 15864t^{26} \\ &\quad + 13804t^{27} + 15138t^{28} - 37476t^{29} + 7048t^{30} + 57414t^{31} - 58566t^{32} + \dots \end{aligned}$$

$$\boxed{Z - Z_{\text{grav}} = -t^{24} + 6t^{25} - 12t^{26} + 0t^{27} + 39t^{28} - 60t^{29} + t^{30} + 108t^{31} - 138t^{32} + \dots}$$

- Computed exactly in “**BMN truncation**”: $Q\bar{\phi}^m = 0$, $Q\psi_{m+} \sim \epsilon_{mnp}[\bar{\phi}^n, \bar{\phi}^p]$, $Qf_{++} \sim [\psi_{m+}, \bar{\phi}^m]$

[Berenstein, Maldacena, Nastase] (2002)

[Nakwoo Kim, Klose, Plefka] (2003)

$$[Z - Z_{\text{grav}}]_{\text{BMN}} = -\frac{t^{24}}{1 - t^{12}} \cdot \frac{1}{(1 - t^8)^3} \cdot (1 - t^2)^3$$

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Constructions

An ∞ -tower of non-graviton cohomologies, whose representatives are: $n = 0, 1, 2, \dots$

$$\begin{aligned}
 O_n = & (f \cdot f)^n \epsilon^{c_1 c_2 c_3} (\phi^a \cdot \psi_{c_1}) (\phi^b \cdot \psi_{c_2}) (\psi_a \cdot \psi_b \times \psi_{c_3}) \\
 & + n (f \cdot f)^{n-1} \epsilon^{b_1 b_2 b_3} \epsilon^{c_1 c_2 c_3} (f \cdot \psi_{b_1}) (\phi^a \cdot \psi_{c_1}) (\psi_{b_2} \cdot \psi_{c_2}) (\psi_a \cdot \psi_{b_3} \times \psi_{c_3}) \\
 & - \left(\frac{n}{72} + \frac{n(n-1)}{108} \right) (f \cdot f)^{n-1} \epsilon^{a_1 a_2 a_3} \epsilon^{b_1 b_2 b_3} \epsilon^{c_1 c_2 c_3} (\psi_{a_1} \cdot \psi_{b_1} \times \psi_{c_1}) (\psi_{a_2} \cdot \psi_{b_2} \times \psi_{c_2}) (\psi_{a_3} \cdot \psi_{b_3} \times \psi_{c_3})
 \end{aligned}$$

- In particular, the “threshold” non-graviton operator at t^{24} order:

$$O_0 = \epsilon^{p_1 p_2 p_3} (\phi^m \cdot \psi_{p_1}) (\phi^n \cdot \psi_{p_2}) (\psi_m \cdot \psi_n \times \psi_{p_3})$$

(Used 3d vector notation for SU(2) adjoints: $A \cdot B \sim \text{tr}(AB)$ and $A \times B \sim [A, B]$)

This tower fully explains the BMN index:

$$[Z - Z_{\text{grav}}]_{\text{BMN}} = -\frac{t^{24}}{1 - t^{12}} \cdot \frac{1}{(1 - t^8)^3} \cdot (1 - t^2)^3$$

“core black hole” primary operators O_n

descendants within BMN

Limited dressings by gravitons $\text{tr}(2\bar{\phi}^m f + \epsilon^{mnp} \psi_n \psi_p)$
(only 3 out of 17 gravitons in BMN sector)

- These operators don't really look like black holes in most senses. (E.g. small entropy)
- But they exhibit subtle properties, qualitatively reminiscent of black holes. (\rightarrow next slide)

A no-hair theorem?

To appreciate the last point, see the BMN results:

- Q satisfies Leibniz rule \rightarrow The product (BH) \times (graviton) is another cohomology.
- There are 17 different species of graviton particles in the BMN sector.
- But “black hole operators” O_n **abhor dressings by all but 3 gravitons**: $\text{tr}(2\bar{\phi}^m f + \epsilon^{mnp} \psi_n \psi_p)$.

$$[Z - Z_{\text{grav}}]_{\text{BMN}} = \boxed{-\frac{t^{24}}{1 - t^{12}}} \cdot \boxed{\frac{1}{(1 - t^8)^3}} \cdot (1 - t^2)^3$$

This feature continues in the general SU(2) index:

- Eliminating $PSU(1,2|3) \subset PSU(2,2|4)$ superconformal descendants of O_0 , the remainder is:

$$\chi_0(t) = -t^{24} + \boxed{6t^{25} - 12t^{26} + 0t^{27} + 39t^{28} - 60t^{29} + t^{30} + 108t^{31}} - 135t^{32} + \dots$$

$$\downarrow$$

$$Z - Z_{\text{grav}} - \chi_0(t) = -3t^{32} + \dots$$

- The “boring” range $25 \leq j \leq 31 \rightarrow$ Many product cohomologies absent in the index.
- Simplest possibility: All Q-exact, i.e. absent \leftarrow Checked explicitly for many (next slide).

Only a partial no-hair theorem:

- Conformal **primaries** of gravitons: **29 of 32 don't dress O_0** (at least invisible in the index).
- Conformal **descendants**...? (more in our paper \rightarrow next page)

Illustration: Q-exactness at low levels

$$\begin{aligned}
 t^{28}: \quad O_0(\bar{\phi}^{(m)} \cdot \bar{\phi}^{(n)}) &= -\frac{1}{14}Q[20\epsilon^{rs(m)}(\bar{\phi}^{(n)} \cdot \psi_{p+})(\bar{\phi}^{(p)} \cdot \psi_{r+})(\bar{\phi}^{(q)} \cdot \psi_{q+})(f_{++} \cdot \psi_{s+}) \\
 &\quad -20\epsilon^{prs}(\bar{\phi}^{(m)} \cdot \psi_{p+})(\bar{\phi}^{(n)} \cdot \psi_{r+})(\bar{\phi}^{(q)} \cdot \psi_{q+})(f_{++} \cdot \psi_{s+}) \\
 &\quad +30\epsilon^{prs}(\bar{\phi}^{(m)} \cdot \psi_{p+})(\bar{\phi}^{(n)} \cdot \psi_{r+})(\bar{\phi}^{(q)} \cdot \psi_{s+})(f_{++} \cdot \psi_{q+}) \\
 &\quad -7\epsilon^{a_1 a_2 p} \epsilon^{b_1 b_2 (m)} (\bar{\phi}^{(n)} \cdot \psi_{p+})(\bar{\phi}^{(q)} \cdot \psi_{q+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+}) \\
 &\quad +18\epsilon^{a_1 a_2 p} \epsilon^{b_1 b_2 (m)} (\bar{\phi}^{(n)} \cdot \psi_{q+})(\bar{\phi}^{(q)} \cdot \psi_{p+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+})]
 \end{aligned}$$

$$\begin{aligned}
 t^{29}: \quad O_0(\bar{\phi}^m \cdot \bar{\lambda}_{\dot{\alpha}}) &= \frac{1}{8}Q[40\epsilon^{mnp}(f_{++} \cdot \psi_{q+})(\bar{\lambda}_{\dot{\alpha}} \cdot \psi_{r+})(\bar{\phi}^{(q)} \cdot \psi_{n+})(\bar{\phi}^r \cdot \psi_{p+}) \\
 &\quad -4\epsilon^{ma_1 a_2} \epsilon^{nb_1 b_2} (\bar{\lambda}_{\dot{\alpha}} \cdot \psi_{n+})(\bar{\phi}^p \cdot \psi_{p+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+}) \\
 &\quad +6\epsilon^{ma_1 a_2} \epsilon^{nb_1 b_2} (\bar{\lambda}_{\dot{\alpha}} \cdot \psi_{p+})(\bar{\phi}^p \cdot \psi_{n+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+}) \\
 &\quad +\epsilon^{na_1 a_2} \epsilon^{pb_1 b_2} (\bar{\lambda}_{\dot{\alpha}} \cdot \psi_{n+})(\bar{\phi}^m \cdot \psi_{p+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+})]
 \end{aligned}$$

$$\begin{aligned}
 t^{30}: \quad O_0(\bar{\phi}^m \cdot \psi_{n+} - \frac{1}{3}\delta_n^m \bar{\phi}^p \cdot \psi_{p+}) \\
 &= \frac{1}{4}Q[\epsilon_{npq}\epsilon^{ra_1 a_2} \epsilon^{qb_1 b_2} \epsilon^{mc_1 c_2} (\bar{\phi}^p \cdot \psi_{r+})(\psi_{a_1+} \cdot \psi_{a_2+})(\psi_{b_1+} \cdot \psi_{b_2+})(\psi_{c_1+} \cdot \psi_{c_2+})]
 \end{aligned}$$

Concluding remarks

“Construction” (in a limited sense) of BPS black hole microstates

- Found new “black hole cohomologies” for SU(2)
- What I presented are **not** actual BPS states, even at 1-loop. Just representatives.

[Chang, Feng, Lin, Tao] (2023) [Budzik, Murali, Vieira] (2023)

- SU(3) in BMN sector [Jae Hyeok Choi, Jehyun Lee, Siyul Lee] (work in progress)

$$Z - Z_{grav} = \boxed{-t^{24}} + 3t^{26} - 3t^{28} - 10t^{30} + 15t^{32} + 24t^{34} - 36t^{36} - 45t^{38} + 39t^{40} + 124t^{42} + O(t^{44})$$

Same threshold level as SU(2)...!

- Higher N? Higher charges? Analytic structures? Insights from emergent structures of holomorphically twisted QFT? ... [Budzik, Gaiotto, Kulp, Williams, Wu, Yu] (2023)

Better picture on hairy (BPS) black holes in AdS₅ x S⁵ ?

[Bhattacharyya, Minwalla, Papadodimas] (2010) [Markeviciute, Santos] [Markeviciute] (2018)

[Niehoff, Santos, Way] (2015) [Chesler] (2021) [SK, Kundu, E. Lee, J. Lee, Minwalla, Patel] (2023) ...

- Studied BPS perturbations of BPS black holes, dual to $(\partial_{++})^{m_1} (\partial_{+-})^{m_2} \text{tr}(X^2 + Y^2 + Z^2)$

Similar “partial no-hair” behavior as SU(2) cohomologies. [Choi, SK, E. Lee, S. Lee, Park] (2023)

Perturbative dressing forbidden:	$m_1 + m_2 < 4q/\ell^2$	q : “size” parameter of BH ℓ : AdS radius
Allowed (large conformal descendants):	$m_1 + m_2 \geq 4q/\ell^2$	

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