

Title: Cosmology Lecture (230404)

Speakers: Neal Dalal

Collection: Cosmology (2022/2023)

Date: April 04, 2023 - 3:45 PM

URL: <https://pirsa.org/23040054>

Last time: FRW, geodesics, distances

---

Evolution of  $g_{\mu\nu}$ :  $ds^2 = -dt^2 + a^2(t) [dr^2 + D(r)^2 d\Omega^2]$

Last time: FRW, geodesics, distances

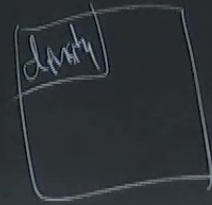
Evolution of  $g_{\mu\nu}$  :  $ds^2 = -dt^2 + a^2(t) [dr^2 + D^2(r) d\Omega^2]$

Last time: FRW, geodesics, distances

Evolution of  $g_{\mu\nu}$ :  $ds^2 = -dt^2 + a^2(t) [dr^2 + D^2(r) d\Omega^2]$

Einstein eqns:  $G_{\mu\nu} = 8\pi G T_{\mu\nu}$

↑ density & flux of  $E, P$



perfect fluid w/  $\rho, P, u^\mu$

$$T^{\mu\nu} = (\rho + P) u^\mu u^\nu + P g^{\mu\nu}$$

frame where  $u^\mu = (1, 0, 0, 0)$

$$T^{\mu}_{\nu} = \begin{pmatrix} -\rho & & & \\ & P & & \\ & & P & \\ & & & P \end{pmatrix}$$

perfect fluid w/  $\rho, P, u^\mu$

$$T^{\mu\nu} = (\rho + P) u^\mu u^\nu + P g^{\mu\nu}$$

frame where  $u^\mu = (1, 0, 0, 0)$

$$T^{\mu}_{\nu} = \begin{pmatrix} -\rho & & & \\ & P & & \\ & & P & \\ & & & P \end{pmatrix}$$

$$\nabla_\nu T^{\mu\nu} = 0$$

$\mu=0$  : continuity eqn

$\mu=i$  : Euler eqn

$$\frac{\partial \rho}{\partial t} + 3 \frac{\dot{a}}{a} [\rho + P] = 0$$

$$P = w\rho \Rightarrow \dot{\rho} + 3\frac{\dot{a}}{a}(1+w)\rho = 0 \quad \boxed{\text{const. } w}$$

$$\int \frac{d\rho}{\rho} = -3(1+w) \int \frac{da}{a} \Rightarrow \rho \propto a^{-3(1+w)}$$

$$\text{NR: } P \ll \rho \Rightarrow w \approx 0 \Rightarrow \rho \propto a^{-3}$$

$$\text{R: } P = \rho/3 \Rightarrow w = \frac{1}{3} \Rightarrow \rho \propto a^{-4}$$

Const.  $w$

$$\propto a^{-3(1+w)}$$

$$p < \frac{1}{a}$$
$$E = \sqrt{p^2 + m^2}$$



$$P = w\rho \Rightarrow \dot{\rho} + 3\frac{\dot{a}}{a}(1+w)\rho = 0 \quad (\text{const. } w)$$

$$\int \frac{d\rho}{\rho} = -3(1+w) \int \frac{da}{a} \Rightarrow \rho \propto a^{-3(1+w)}$$

$$\text{NR: } P \ll \rho \Rightarrow w \approx 0 \Rightarrow \rho \propto a^{-3}$$

$$\text{R: } P = \rho/3 \Rightarrow w = \frac{1}{3} \Rightarrow \rho \propto a^{-4}$$

$$\left. \begin{array}{l} \rho \propto \frac{1}{a} \\ E = \sqrt{p^2 + m^2} \end{array} \right|$$

$$G^0_0 = -3 \left[ \left( \frac{\dot{a}}{a} \right)^2 + \frac{k}{(aR)^2} \right] = 8\pi G T^0_0 = -8\pi G \rho$$

$$G^i_j = - \left[ 2 \frac{\dot{a}}{a} + \left( \frac{\dot{a}}{a} \right)^2 + \frac{k}{(aR)^2} \right] \delta^i_j$$

$$H^2 = \frac{8\pi G}{3} \rho - \frac{k}{(aR)^2}$$

ist  
Friedmann

$$\left[ \frac{\dot{a}}{a} + \left( \frac{\dot{a}}{a} \right)^2 + \left( \frac{k}{aR} \right)^2 \right] = -4\pi G P$$

$$\frac{4\pi G \rho}{3}$$

$$\frac{\dot{a}}{a} = -\frac{4\pi G}{3} \rho - 4\pi G P =$$

$$-\frac{4\pi G}{3} (\rho + 3P) = \frac{\dot{a}}{a}$$

2nd  
Friedmann

example

$$\left[ \frac{\dot{a}}{a} + \left( \frac{\dot{a}}{a} \right)^2 + \left( \frac{k}{aR} \right)^2 \right] = -4\pi G P$$

2nd  
Friedmann

$$\frac{4\pi G \rho}{3}$$

$$\frac{\dot{a}}{a} = -\frac{4\pi G}{3}\rho - 4\pi G P = -\frac{4\pi G}{3}(\rho + 3P) = \frac{\dot{a}}{a}$$

Example:  $k=0$ , const  $w$   
 $R \rightarrow \infty$

$$H^2 = \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3}\rho = \frac{8\pi G\rho_0}{3} a^{-3(1+w)} = H_0^2 a^{-3(1+w)}$$

$$\rho = \rho_0 a^{-3(1+w)}$$

$$\frac{\dot{a}}{a} = H_0 a^{-\frac{3(1+w)}{2}} \Rightarrow \int a^{\frac{3(1+w)}{2}-1} da = \int H_0 dt$$

$$\frac{2}{3(1+w)} a^{\frac{3(1+w)}{2}} = H_0 t$$

$$\text{NR: } w=0 \Rightarrow a \propto t^{2/3}$$

$$\text{rel: } w = \frac{1}{3} \Rightarrow a \propto t^{1/2}$$

$$\text{NR: } w=0 \Rightarrow a \propto t^{2/3}$$

$$\text{rel: } w = \frac{1}{3} \Rightarrow a \propto t^{1/2}$$

$$w = -1: \frac{\dot{a}}{a} = H_0 \Rightarrow \int \frac{da}{a} = H_0 \int dt$$

$$a \propto e^{H_0 t}$$

$w \neq -1$

Multiple component  
critical density  $\rho_{crit}$

$H_0 \int dt$   
 $e H_0 t \Rightarrow$  inflation  
DE



$w \neq -1$

Multiple component

critical density  $\rho_{crit} = \frac{3H^2}{8\pi G}$

$\Omega_k = \frac{-k}{a^2 R^2 H^2}$

$H_0 \int dt$   
 $H_0 t \Rightarrow$  inflation  
DE

$\Omega_i(t) = \frac{\rho_i(t)}{\rho_{crit}(t)}$

$\Omega_m(t) = \rho_m / \rho_{crit}$

$\Omega_r$   
 $\Omega_\Lambda$

WF-1

# Multiple component

critical density  $\rho_{crit} \equiv \frac{3H^2}{8\pi G}$

$$\Omega_k = \frac{-k}{a^2 K^2 H^2}$$

$$\Omega_i(t) = \frac{\rho_i(t)}{\rho_{crit}(t)}$$

$$\Omega_m(t) = \rho_m / \rho_{crit}$$

$$\Omega_r$$

$$\Omega_k$$

$$\Omega_m + \Omega_r + \Omega_k + \Omega_k + \dots = 1$$



$$\left. \begin{aligned} \rho_m &= \rho_{m0} a^{-3} \\ \rho_r &= \rho_{r0} a^{-4} \end{aligned} \right\}$$

$$\frac{H^2}{H_0^2} = \frac{8\pi G}{3H_0^2} \left[ \rho_{m0} a^{-3} + \rho_{r0} a^{-4} \right]$$

$$\frac{H^2}{H_0^2} = \left[ \frac{\rho_{m,0} a^{-3}}{\rho_{crit,0}} + \frac{\rho_{r,0} a^{-4}}{\rho_{crit,0}} + \frac{\rho_{vac,0} a^0}{\rho_{crit,0}} + \dots \right]^{-\frac{k}{2}} \frac{1}{a^2}$$

$$\frac{H^2}{H_0^2} =$$

$$\left[ \Omega_{m0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{\Lambda 0} a^0 + \Omega_{k0} a^{-2} + \dots \right]$$

$$\left. \frac{H^2}{H_0^2} \right\} \left[ \Omega_{m0} a^{-3} + \Omega_{r0} a^{-4} + \Omega_{\Lambda 0} a^0 + \Omega_{k0} a^{-2} + \dots \right]$$

$$\Omega_m > 1, \Omega_r = \Omega_\Lambda = 0$$

$\Downarrow$

$\Omega_k < 0 \Rightarrow$  positive curvature

$$H^2 = H_0^2 \left[ \frac{\Omega_m}{a^3} + \frac{\Omega_k}{a^2} \right]$$

$$\Omega_m > 1, \Omega_r = \Omega_\Lambda = 0$$

$\Downarrow$

$\Omega_k < 0 \Rightarrow$  positive curvature

$$H^2 = H_0^2 \left[ \frac{\Omega_{m0}}{a^3} + \frac{\Omega_{k0}}{a^2} \right] = H_0^2 \left[ \frac{1+\delta_0}{a^3} - \frac{\delta_0}{a^2} \right] \text{ at } a=1$$

$$\Omega_{m0} = 1 + \delta_0 \text{ at } a=1$$

$$\Omega_{k0} = -\delta_0$$



$$\Omega_m > 1, \Omega_r = \Omega_\Lambda = 0$$

⇓

$\Omega_k < 0 \Rightarrow$  positive curvature

$$\left(\frac{\dot{a}}{a}\right)^2 = H^2 = H_0^2 \left[ \frac{\Omega_{m0}}{a^3} + \frac{\Omega_{k0}}{a^2} \right] = H_0^2 \left[ \frac{1+\delta_0}{a^3} - \frac{\delta_0}{a^2} \right]$$

$$a_{\text{max}} = \frac{1+\delta_0}{\delta_0}$$

$$\Omega_{m0} = 1 + \delta_0 \quad \text{at } a=1$$

$$\Omega_{k0} = -\delta_0$$

