

Title: Quantum Field Theory in Curved Spacetime (PM) - 2023-03-24

Speakers: Sergey Sibiryakov

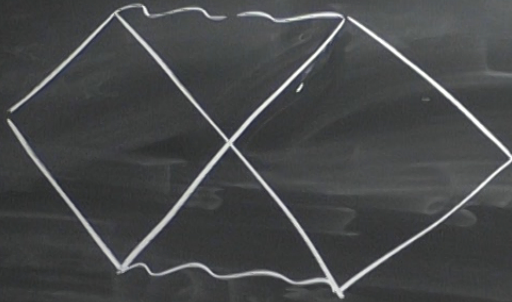
Collection: Quantum Field Theory in Curved Spacetime

Date: March 24, 2023 - 2:00 PM

URL: <https://pirsa.org/23030097>

Abstract: <https://pitp.zoom.us/j/95867663528?pwd=dmd5Y3FaVHJ0SnJWSIFscEY4cmhYUT09>

I) Eternal BH



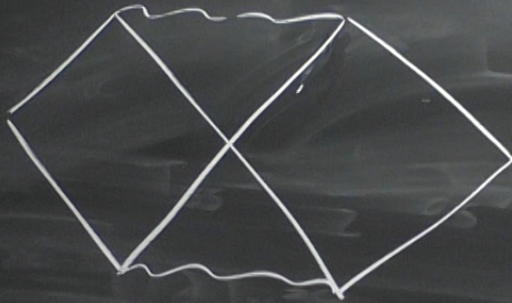
$$ds^2 = -\left(1 - \frac{2M}{r}\right) du dv$$

$$= -\frac{2M}{r} e^{-r/2M} d\bar{u} d\bar{v}$$

$$\bar{u} = -4M e^{-u/4M}$$

$$\bar{v} = 4M e^{v/4M}$$

I) Eternal BH



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$$a) \psi_{\omega, R}^B = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega u} ; \psi_{\omega, L}^B = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega v}$$

$$a_{\omega}^B |0\rangle_B = 0 \leftarrow \text{Boulware vacuum}$$

$$b) \quad \varphi_{\omega, R}^H = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{u}} \quad , \quad \varphi_{\omega, L}^H = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{v}}$$

$$b) \psi_{\omega, R}^H = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{u}}$$

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This state is not empty

Ex. 1

$$\beta_{\tilde{\omega}\omega, RR} = - \langle \psi_{\tilde{\omega}R}^B, (\psi_{\omega R}^H)^* \rangle = \frac{-i}{2\pi\sqrt{\omega\tilde{\omega}}} \left(\frac{i}{4M\tilde{\omega}} \right)^{4M\tilde{\omega}} \cdot \Gamma(1+i4M\tilde{\omega})$$

||
 $\beta_{\tilde{\omega}\omega, LL}$

$$a_{\omega}^H |0\rangle_H = 0 \leftarrow \text{Hartle-Hawking vacuum}$$

This state is not empty

$$\langle \Psi_{\tilde{\omega}, R}, (\Psi_{\omega, R}) \rangle = 2\pi \sqrt{\omega \tilde{\omega}} (4M\omega)$$

$$\beta_{\tilde{\omega}, LL}$$

$$\Gamma(1 + i4M\tilde{\omega})$$

$$G^+(t, r, t', r') = -\frac{1}{4\pi} \ln[(\bar{u} - \bar{u}' - i\epsilon)(\bar{v} - \bar{v}' - i\epsilon)]$$

$$\int_0^\infty \frac{d\omega}{4\pi\omega} \left(e^{-i\omega(\bar{u} - \bar{u}')} + e^{-i\omega(\bar{v} - \bar{v}')} \right)$$

$\partial_{\bar{u}} G^+$

$$\frac{1}{4\pi} e^{i\pi} = -\frac{1}{4\pi} (\bar{u} - \bar{u}' - i\epsilon)$$

1. $\Gamma = \Gamma'$ - fixed

$$\bar{u} - \bar{u}' = -4M \left(e^{-\frac{u}{4M}} - e^{-\frac{u'}{4M}} \right)$$

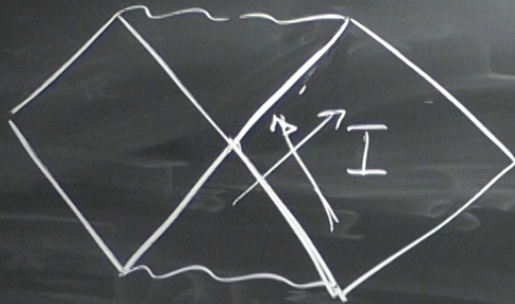
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This state is not empty

$\beta \tilde{\omega}_{LL}$

$\Gamma(1 + i4M\tilde{\omega})$

$$G^+(t, r, t', r) = -\frac{1}{4\pi} \ln \left[\text{ch} \left(\frac{t-t'-i\epsilon}{4M} \right) - 1 \right] + \tilde{G}(r)$$
$$= -\frac{1}{4\pi} \ln \left[\text{ch} \left(\frac{\bar{t}-\bar{t}'-i\epsilon}{4M\sqrt{1-\frac{2M}{r}}} \right) - 1 \right]$$



$$u = t - r_*$$

$$v = t + r_*$$

$$ds^2 = -(1 - \frac{2M}{r}) dt^2 + \frac{dr^2}{1 - \frac{2M}{r}} \rightarrow T = \sqrt{1 - \frac{2M}{r}} t$$

$$= -\frac{2M}{r} e^{-r/2M} d\bar{u} d\bar{v}$$

$$\bar{u} = -4M e^{-u/4M}$$

$$\bar{v} = 4M e^{v/4M}$$

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$$\left. \begin{array}{l} \int_0^{\tilde{\omega}} \\ \int_0^{\tilde{\omega}} \end{array} \right\} \Delta \tilde{\tau} = i8\pi M \sqrt{1-\frac{2M}{r}} \left| = -\frac{1}{4\pi} \ln \left[\text{ch} \left(\frac{\tilde{\tau}-\tilde{\tau}'-i\epsilon}{4M\sqrt{1-\frac{2M}{r}}} \right) - 1 \right] + \tilde{G}(r) \right.$$

$$u = t - r_*$$

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$$a) \psi_{\omega, R}^B = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega u} ; \psi_{\omega, L}^B = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega v}$$

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$$a_{\omega}^B |0\rangle_B = 0 \leftarrow \text{Boulware vacuum}$$

$$\rightarrow F_B(\Delta E) = 0$$

$$\int_0^{\infty} \frac{d\omega}{4\pi\omega} \left(e^{-i\omega(\bar{u}-\bar{u}')} + e^{-i\omega(\bar{\sigma}-\bar{\sigma}')} \right)$$

$$\partial_{\bar{u}} G^+ = (-i) \int_0^{\infty} \frac{d\omega}{4\pi} e^{-i\omega(\bar{u}-\bar{u}')} = -\frac{i}{4\pi(\bar{u}-\bar{u}'-i\epsilon)}$$

$$G_B^+ = -\frac{1}{4\pi} \ln \left[(\bar{u}-\bar{u}'-i\epsilon)(\bar{\sigma}-\bar{\sigma}'-i\epsilon) \right]$$

$$b) \psi_{\omega, R}^H = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{u}}, \quad \psi_{\omega, L}^H = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{v}}$$

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Free falling through horizon

$$\bar{r} = \bar{r}'$$

$$\bar{v} = \bar{u}$$

$$S_H(t, \bar{r}, \bar{t}', \bar{r}') = -\frac{1}{2\pi} \ln(\bar{v} - \bar{u} - i\epsilon)$$

$$u = t - r_*$$

$$v = t + r_*$$

$$\bar{u} = -4M e^{-u/4M}$$

$$\bar{v} = 4M e^{v/4M}$$

2. Free falling through horizon

$$\bar{r} = \bar{r}'$$

$$\bar{t} = \frac{\bar{v} - \bar{u}}{2}$$

$$G_H^+ (\bar{t}, \bar{r}, \bar{t}', \bar{r}') = -\frac{1}{2\pi} \ln (\bar{t} - \bar{t}' - i\epsilon)$$

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2. Free falling through horizon

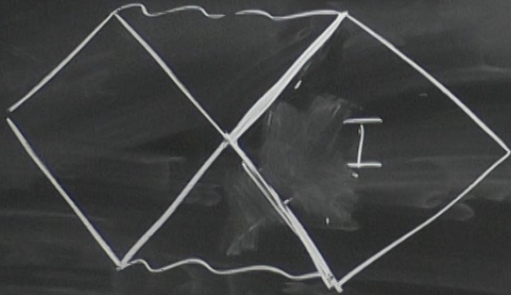
$$\bar{r} = \bar{r}'$$

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$$B(\Delta E) = 0$$

I) Eternal BH



$$u = t - r_*$$

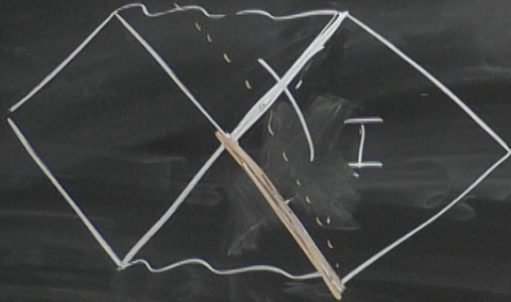
$$v = t + r_*$$

$$\psi_{\omega, R}^u = \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{u}}$$

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I) Eternal BH



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$$\beta(\Delta E) = 0$$

II) HR in a bigger picture

a) GR + horizons + QFT \Rightarrow thermality

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 \Leftarrow ?

thermal + horizons + QFT \Rightarrow GR

T. Jacobson, PRL, 75, 1260 (1995)

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II) HR in a bigger picture

a) GR + horizons + QFT \Rightarrow thermality
 \Leftarrow ?

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T. Jacobson, PRL, 75, 1260 (1995)

$$\beta \tilde{\omega}_{LL}$$

$$2 \Gamma \sqrt{\omega \tilde{\omega}} (4M\tilde{\omega})$$

$$\Gamma (1 + i 4M\tilde{\omega})$$

b) Trans-Planckian problem

BT, U or HH \rightarrow B

$$t, \omega \propto e^{t/2M}$$

$$\beta \tilde{\omega}_{LL}$$

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Unruh, Schutzhold, PRD, 71, 024028 (2005)

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BT, U or HH \rightarrow B

\uparrow $t, \omega \propto e^{t/2M}$

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III) Information paradox

Hawking, PRD, 14, 2460 (1976)

Page, PRL, 71, 1291 (1993)
—————, —————, 3743 ————

Susskind, Thorlacius, Uglum, PRD, 48, 3743 (1993)

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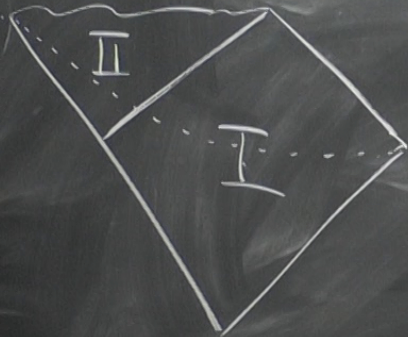
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Hawking, PRD, 14, 2460 (1976)

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————— 3743 ———

Susskind, Thorlacius, Uglum, PRD, 48, 3743 (1993)

1) First formulation

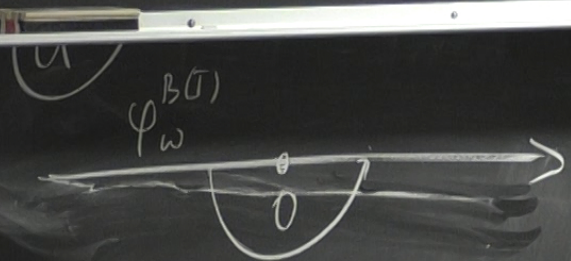


$$\begin{aligned}
 |0\rangle_a &\leftrightarrow \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega\bar{a}} \\
 |0\rangle_B &\leftrightarrow \frac{1}{\sqrt{4\pi\omega}} e^{-i\omega a}
 \end{aligned}
 \left. \vphantom{\begin{aligned} |0\rangle_a \\ |0\rangle_B \end{aligned}} \right\} \begin{aligned} &\frac{1}{\sqrt{4\pi\omega}} e^{-i\omega s} \\ &\text{--- " ---} \end{aligned}$$

$$\psi_{\omega}^{B(I)} = \begin{cases} \bar{u} < 0 : \frac{1}{\sqrt{4\pi\omega}} e^{i4M\omega \ln(-\bar{u}/4M)} \\ \bar{u} > 0 : 0 \end{cases}$$

$$\psi_{\omega}^{B(II)} = \begin{cases} \bar{u} < 0 : 0 \\ \bar{u} > 0 : \frac{1}{\sqrt{4\pi\omega}} e^{-i4M\omega \ln(\bar{u}/4M)} \end{cases}$$

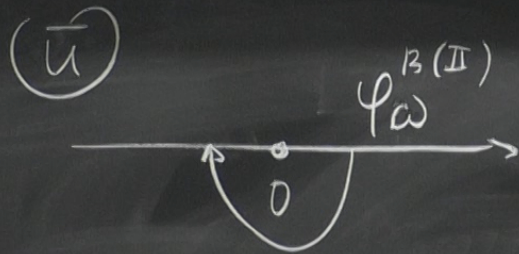
Susskind, Thorlacius, Uglum, PRL, 48, 3743 (1993)



$$\Rightarrow \tilde{\varphi}_\omega^{u(I)} = \sqrt{4\pi\omega} \sqrt{1 - e^{-8\pi M\omega}}$$

$$\times \begin{cases} e^{i\omega 4M \ln(-\bar{u}/4M)}, & \bar{u} < 0 \\ e^{-4\pi M\omega} e^{i\omega 4M \ln(\bar{u}/4M)}, & \bar{u} > 0 \end{cases}$$

$$(\bar{u} > 0 : \frac{1}{\sqrt{4\pi\omega}} e^{-\omega 4M l_n(\bar{u}/4M)})$$



$$\tilde{\varphi}_{\omega}^{u(II)} = \frac{1}{\sqrt{4\pi\omega} \sqrt{1 - e^{-8\pi M\omega}}}$$

$$\left\{ \begin{array}{l} e^{-4\pi M\omega} e^{-\omega 4M l_n(-\bar{u}/4M)}, \bar{u} < 0 \\ e^{-\omega 4M l_n(\bar{u}/4M)}, \bar{u} > 0 \end{array} \right.$$

$$\tilde{\varphi}_\omega^{u(I, II)} \rightarrow \tilde{a}_\omega^{u(I, II)}$$

$$|\beta_\omega^{I II}|^2 = e^{\frac{1}{8\pi M\omega}}$$

$$\tilde{\varphi}_\omega^{u(I)} = \frac{1}{\sqrt{1 - e^{-8\pi M\omega}}} \varphi_\omega^{B(I)} + \frac{e^{-4\pi M\omega}}{\sqrt{1 - e^{-8\pi M\omega}}} (\varphi_\omega^{B(II)})^*$$

$$\tilde{\varphi}_\omega^{u(II)} = \frac{e^{-4\pi M\omega}}{\sqrt{1 - e^{-8\pi M\omega}}} (\varphi_\omega^{B(I)})^* + \frac{1}{\sqrt{1 - e^{-8\pi M\omega}}} \varphi_\omega^{B(II)}$$

$$\tilde{a}_w^{u(I)} = \frac{1}{\sqrt{-}} a_w^{B(I)} - \frac{e^{-4\pi M\omega}}{\sqrt{-}} (a_w^{B(II)})^\dagger$$

$$\tilde{a}_w^{u(II)} = \frac{1}{\sqrt{-}} a_w^{B(II)} - \frac{e^{-4\pi M\omega}}{\sqrt{-}} (a_w^{B(I)})^\dagger$$

Ex 3 $K(a^B, (a^B)^\dagger)$:

$$\tilde{a}_w^{u(I)} = e^{-iK} a_w^{B(I)} e^{iK}$$

$$\tilde{a}_w^{u(II)} = -e^{-iK} a_w^{B(II)} e^{iK}$$

Ex 3 $K(a^B, (a^B)^\dagger) :$

$$\tilde{a}_\omega^{u(I)} = e^{-iK} a_\omega^{B(I)} e^{iK}$$

$$\tilde{a}_\omega^{u(II)} = -e^{-iK} a_\omega^{B(II)} e^{iK}$$

$$0 = \tilde{a}_\omega^u |0\rangle_u = \pm e^{-iK} a_\omega^B e^{iK} |0\rangle_u$$

$$\rightarrow a_\omega^B (e^{iK} |0\rangle_u) = 0$$

$|0\rangle_B$

Susskind, Horowitz, Uglum, PRD, 48, 3743 (1993)

$$|0\rangle_u = e^{-iK} |0\rangle_B = \sum_{\omega_1, \omega_2, \omega'_1, \omega'_2} K_{\omega_1, \omega_2, \omega'_1, \omega'_2} (a_{\omega_1}^{B(I)})^\dagger \dots (a_{\omega_2}^{B(I)})^\dagger$$

$$(a_{\omega'_1}^{B(II)})^\dagger \dots (a_{\omega'_2}^{B(II)})^\dagger |0\rangle_B$$

not accessible

$$\rho = \text{Tr}_{\text{II}} |0\rangle_{\text{u}} \langle 0|_{\text{u}} = \sum_{\omega'} K_{\omega_1 \dots \omega_k; \omega'_1 \dots \omega'_n} K_{\Omega_1 \dots \Omega_\ell; \omega'_1 \dots \omega'_n}^* \cdot | \omega_1 \dots \omega_k \rangle_{\text{B}} \langle \Omega_1 \dots \Omega_\ell |_{\text{B}}$$

Ex 4

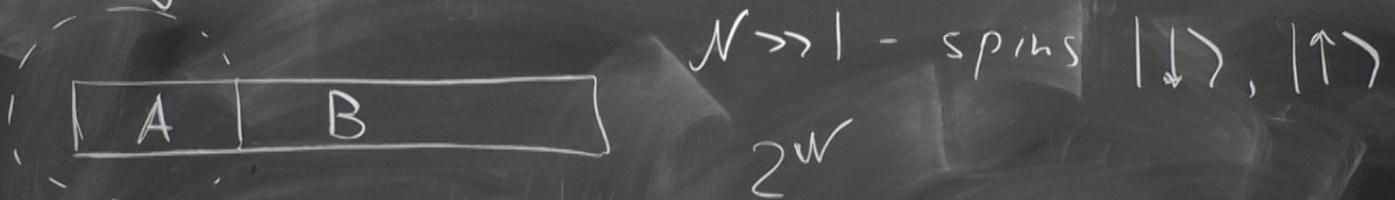
$$\rho = \frac{1}{Z} \exp \left[-\frac{1}{T_{\text{H}}} \int d\omega \omega (a_{\omega}^{\text{R(I)}})^{\dagger} a_{\omega}^{\text{R(I)}} \right]$$

$$\rho = \text{Tr}_{\text{II}} |0\rangle_{\text{u}} \langle 0|_{\text{u}} = \sum_{\omega'} K_{\omega_1 \dots \omega_k; \omega'_1 \dots \omega'_n} K_{\Omega_1 \dots \Omega_\ell; \omega'_1 \dots \omega'_n}^* \cdot | \omega_1 \dots \omega_k \rangle_{\text{B}} \langle \Omega_1 \dots \Omega_\ell |_{\text{B}}$$

Ex 4

$$\rho = \frac{1}{Z} \exp \left[-\frac{1}{T_{\text{H}}} \int d\omega \omega \left(a_{\omega}^{\text{R}(\text{I})} \right)^{\dagger} a_{\omega}^{\text{R}(\text{I})} \right]$$

2) Page's scenario



$$|\Psi\rangle = \sum_{AB} U_{AB} |\psi_A\rangle \otimes |\psi_B\rangle$$

$$\rho_A = \text{Tr}_B |\Psi\rangle\langle\Psi| = \sum_B U_{AB} U_{A'B}^* |\psi_A\rangle\langle\psi_{A'}|$$

$$S_A = -\text{Tr}(\rho_A \ln \rho_A)$$

Ex 5.

1) If ρ_A is pure state $\Rightarrow S_A = 0$

2) $S_A \leq N_A \ln 2$

3) If ρ_A is thermal $\Rightarrow S_A = S_{\text{thermodynamic}}$

4)

$$S_A = -\text{Tr}(\rho_A \ln \rho_A)$$

Ex 5.

1) If ρ_A is pure state $\Rightarrow S_A = 0$

2) $S_A \leq N_A \ln 2$

3) If ρ_A is thermal $\Rightarrow S_A = S_{\text{thermodynamic}}$

4) $S_A = S_B$

Susskind, Thorlacius, Aglum, PRD, 48, 3743 (1993)

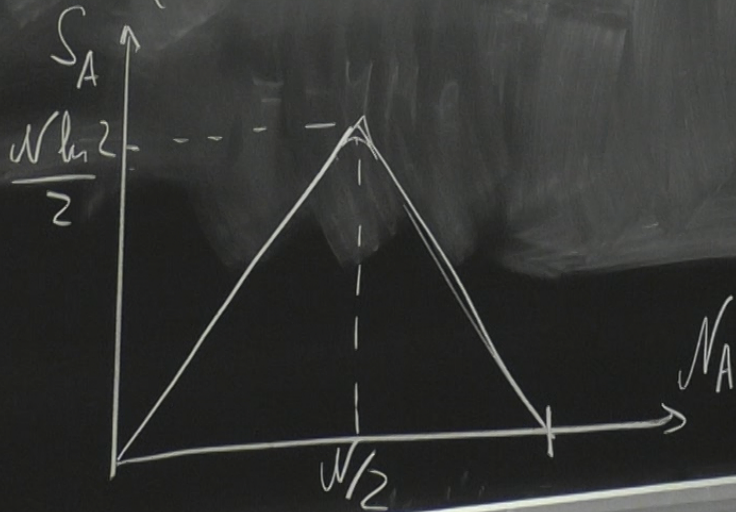
$$\bar{S}_A \begin{cases} N_A \ln 2 + O(2^{N_A - N_B}) & \text{if } N_A < N_B \\ \ln 2 + O(2^{N_B - N_A}) & \text{if } N_B < N_A \end{cases}$$

Susskind, Thorlacius, Aglum, PRD, 48, 3743 (1993)

$$\bar{S}_A = \begin{cases} N_A \ln 2 + O(2^{N_A - N_B}) & \text{if } N_A < N_B \\ N_B \ln 2 + O(2^{N_B - N_A}) & \text{if } N_B < N_A \end{cases}$$

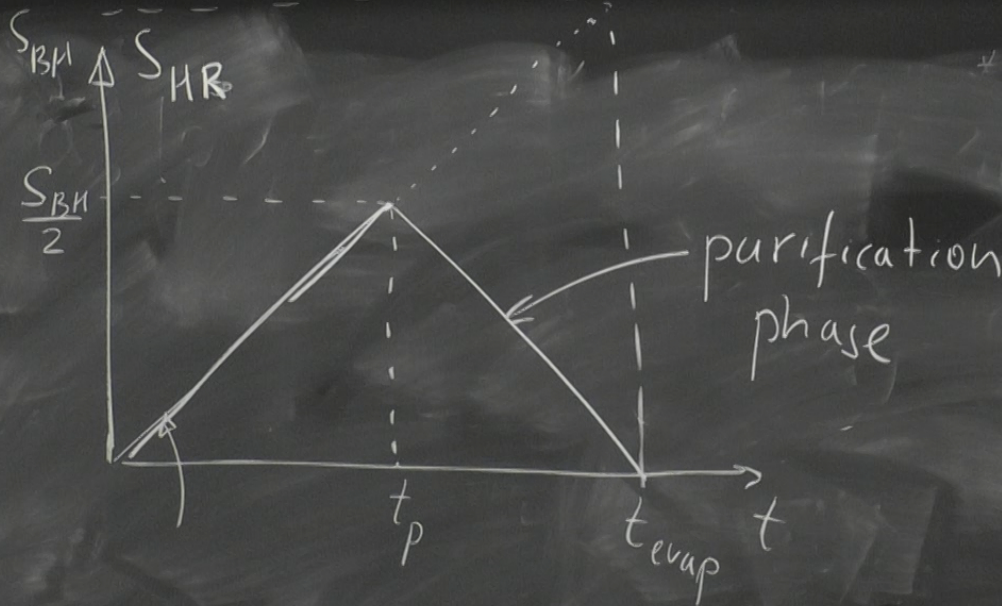
Susskind, Thorlacius, Uglum, PRL, 48, 3743 (1993)

$$\bar{S}_A = \begin{cases} N_A \ln 2 + O(2^{N_A - N_B}) & \text{if } N_A < N_B \\ N_B \ln 2 + O(2^{N_B - N_A}) & \text{if } N_B < N_A \end{cases}$$



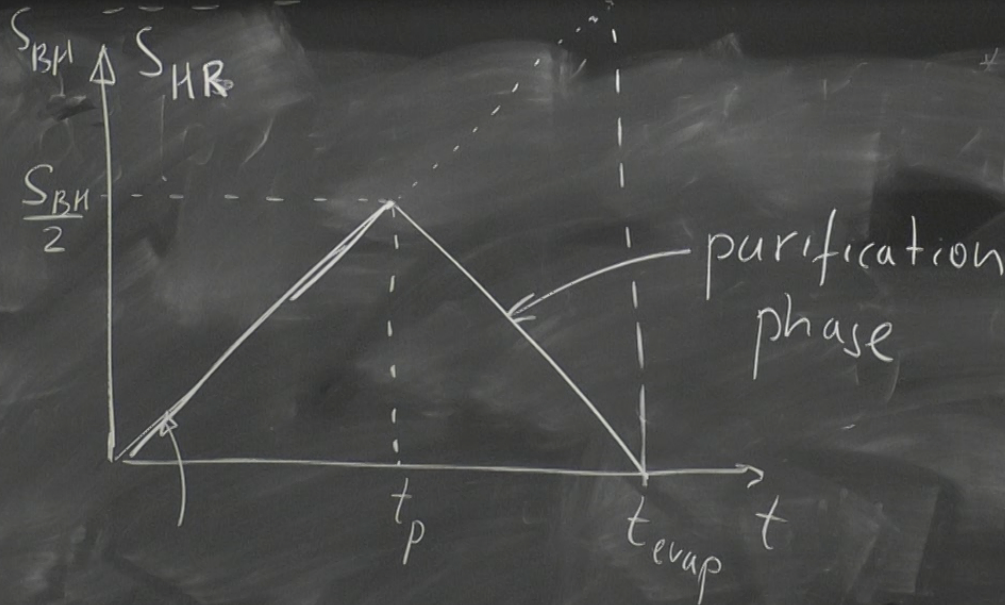
3) If p_A is thermal $\Rightarrow S_A = S_{\text{thermodynamic}}$

$$4) S_A = S_B$$



3) If ρ_A is thermal $\Rightarrow \tilde{S}_A = S_{\text{thermodynamic}}$

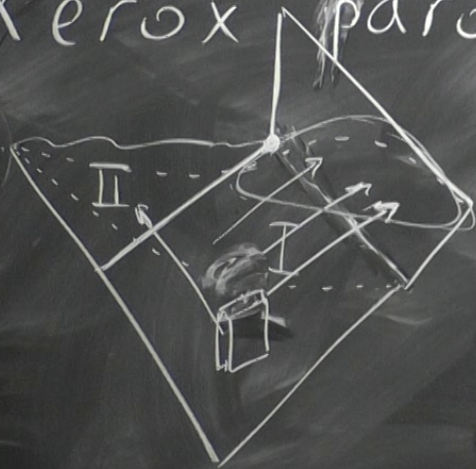
$$4) S_A = S_B$$



Pennington,
JHEP 09, 002
(2020)

2) Page's scenario

B) Xerox paradox

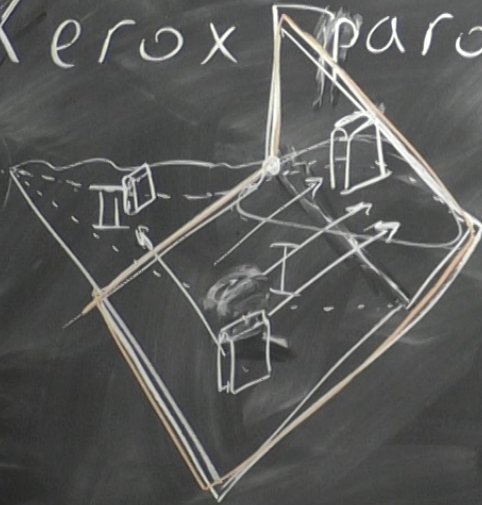


"nice slices" \Rightarrow a) bounded curvature

b) contain almost all HR

2) Page's scenario

3) Xerox paradox



"nice slices" \Rightarrow a) bounded curvature

b) contain almost all HR

$$|\Psi\rangle \rightarrow |\Psi\rangle_{\text{out}} \otimes |\Psi\rangle_{\text{in}}$$