

Title: Strong Gravity Lecture - 230307

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Collection: Strong Gravity (2022/2023)

Date: March 07, 2023 - 10:15 AM

URL: <https://pirsa.org/23030043>

(A) Penrose Process

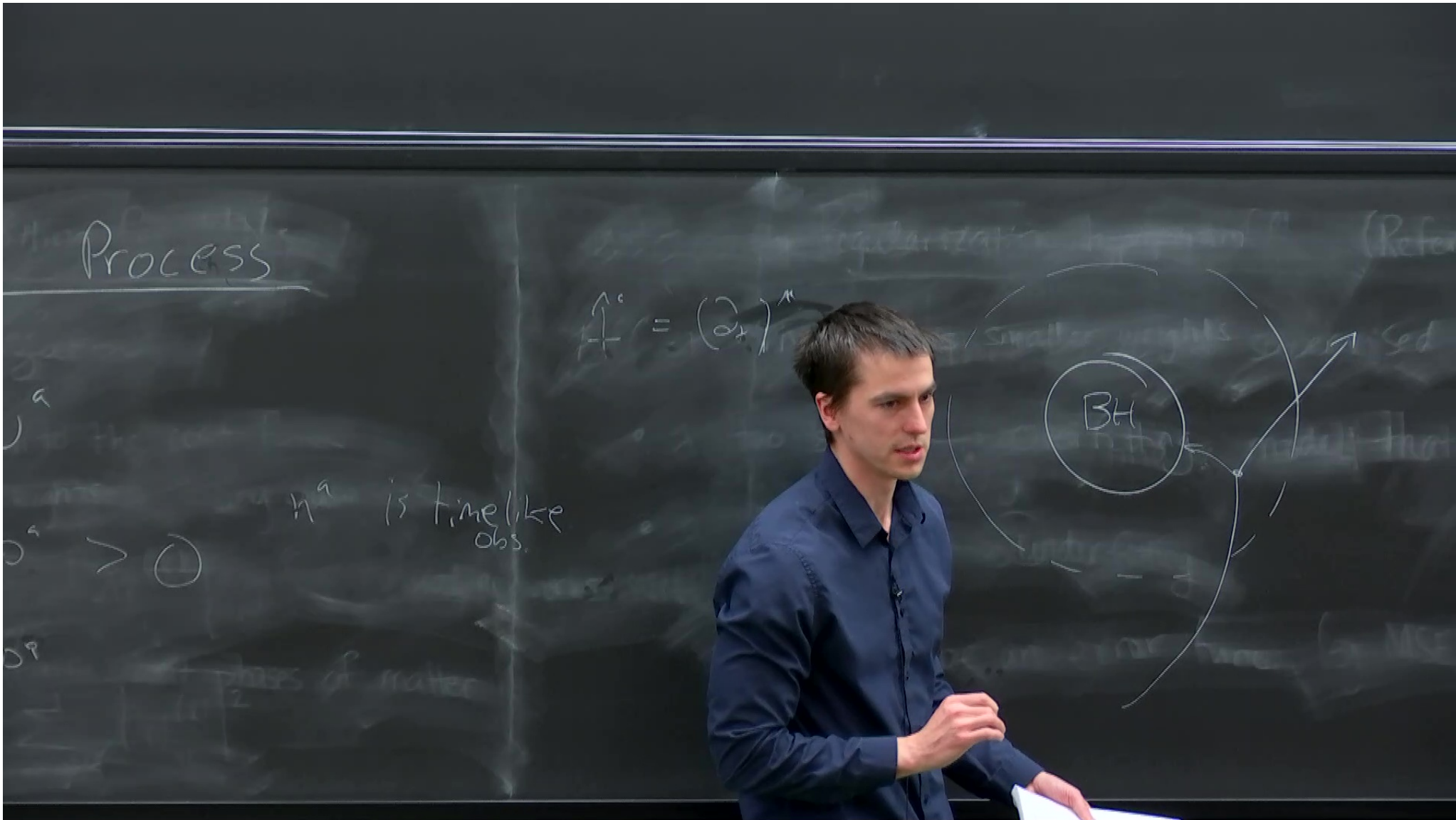
$$p^a = m u^a$$

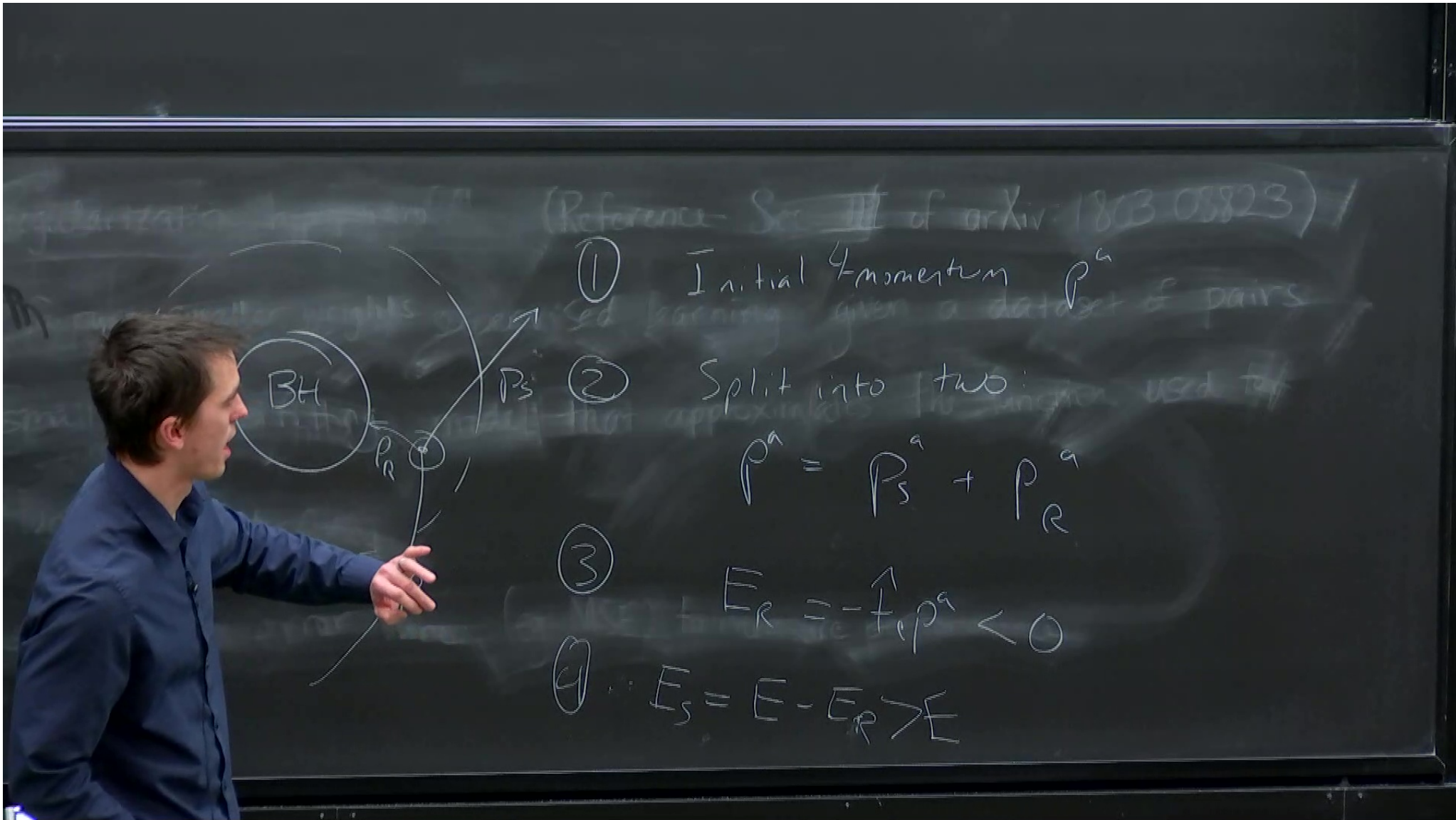
$$E_n = -n_a p^a > 0$$

$$E_{\leftarrow} = -\hat{t}^a p_a$$

n^a is timelike
obs.

$$\hat{t}^a = (2+1)\hat{n}^a$$





$$\frac{\bar{r}}{\bar{r}_0} = 1 \quad \bar{a} \geq 0.91$$

$$p^a = \frac{d+}{d\tau} (1, 0, 0, \Omega) \quad (m=1)$$

$$\Omega := \frac{d\phi/d\tau}{d+ / d\tau}$$

Normalized

Normalization

$$g_{\tau\tau} = -1 = \left(\frac{dt}{d\tau}\right)^2 \left(g_{tt} + g_{\phi\phi} \Omega^2 + 2\Omega g_{t\phi} \right) \quad (**)$$

$$\Rightarrow \frac{dt}{d\tau} = - \left(g_{tt} + g_{t\phi} \Omega \right) \quad (***)$$

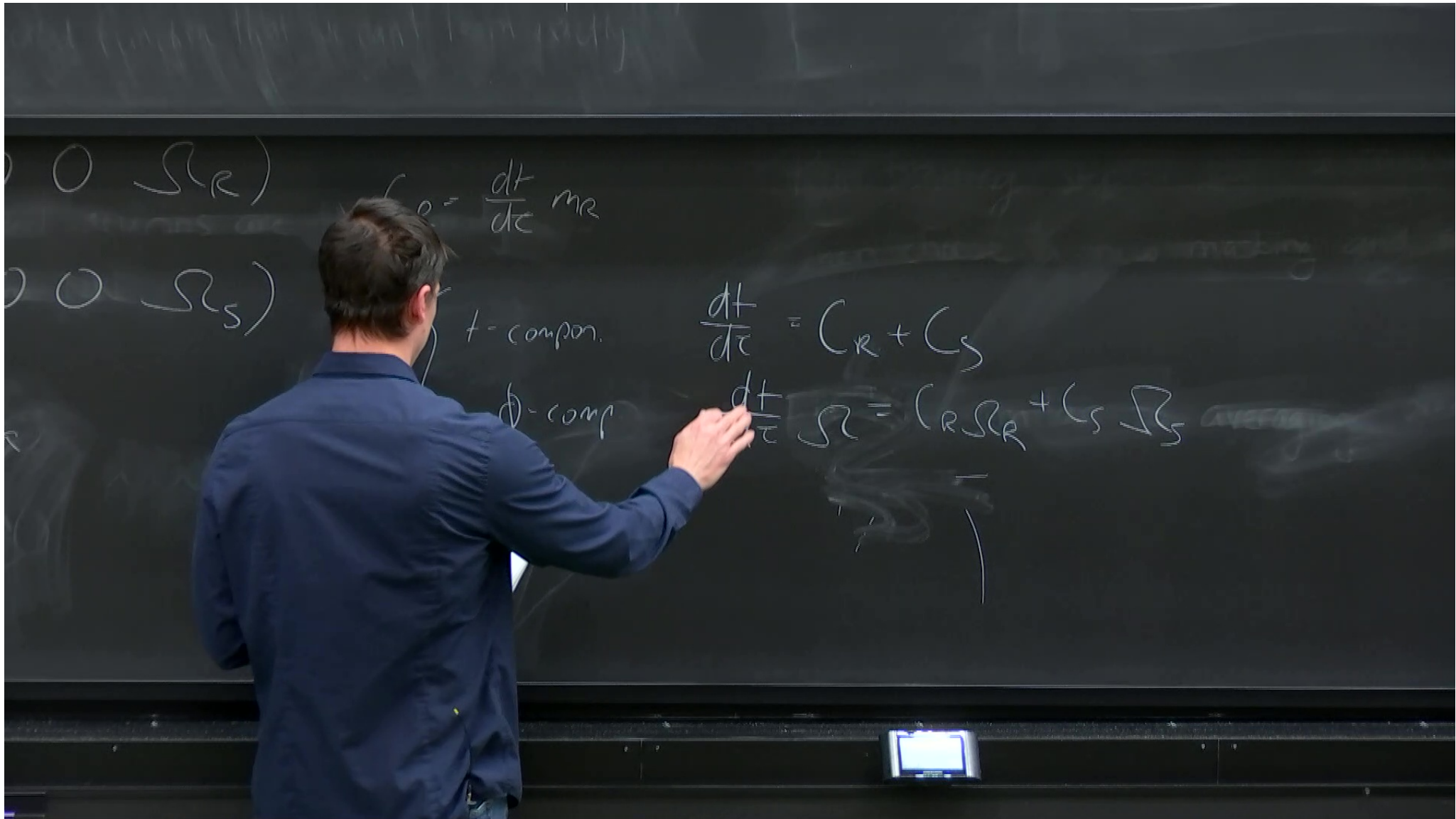
$$\Rightarrow \left(g_{\phi\phi} + g_{t\phi}^2 \right) \Omega^2 + 2g_{t\phi} (1 + g_{tt}) \Omega + g_{tt} (1 + g_{tt}) = 0$$

$$P_R^{\uparrow} = C_R (1 \ 0 \ 0 \ \Omega_R)$$

$$C_R = \frac{dt}{dc} m_R$$

$$P_S^{\uparrow} = C_S (1 \ 0 \ 0 \ \Omega_S)$$

$$P^{\uparrow} = P_R^{\uparrow} + P_S^{\uparrow}$$



$$C_R = \frac{dt}{dc} m_R$$

t-compon.

ϕ -comp

$$\frac{dt}{dc} = C_R + C_S$$

$$\frac{dt}{dc} \Omega = C_R \Omega_R + C_S \Omega_S$$

$$C_S = \frac{\frac{dt}{dc} (\Omega - \Omega_R)}{(\Omega_S - \Omega_R)}$$

$$C_S$$

$$C_R \Omega_R + C_S \Omega_S$$

$$\frac{(\Omega - \Omega_R)}{(\Omega_S - \Omega_R)}$$

$$E_S = -\vec{A}_a \cdot \vec{p}_S^a = -C_S (g_{++} + g_{+p} \Omega_S)$$

$$= \left(\frac{g_{++} + g_{+p} \Omega_S}{g_{++} + g_{+p} \Omega} \right) \left(\frac{\Omega - \Omega_R}{\Omega_S - \Omega_R} \right)$$

$$\eta = \frac{E_S - E}{E} = \frac{E_S}{E} - 1$$

$$\Omega_{\pm} = \frac{a_{\pm} g_{\pm\phi} \pm \sqrt{g_{\pm\phi}^2 - g_{\pm\pm} g_{\phi\phi}}}{g_{\phi\phi}}$$

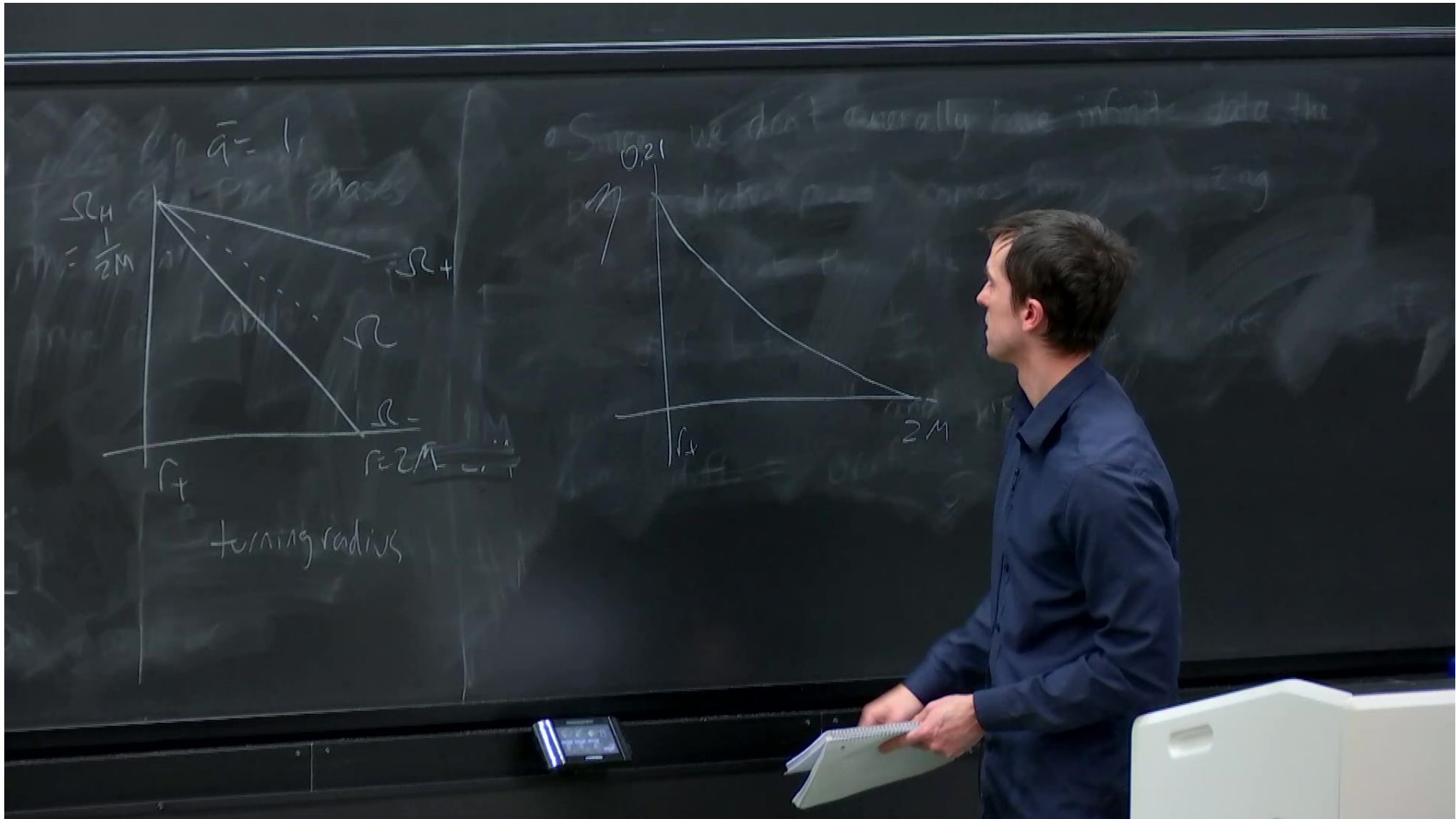


$\bar{a} \rightarrow 1$, $\Omega_3 \rightarrow \Omega_+$
 $\Omega_4 \rightarrow \Omega_-$

$\bar{a} = 1$
 two phases

Labels

\hat{r}_+



Empirical test of data

$$\hat{\lambda}^a = \hat{\lambda}^c + \Omega_{H1} \hat{\phi}^a$$



$$p_a \hat{\lambda}^a < 0$$

$$-SE + \Omega_{H1} \delta J < 0$$

$$\delta J \Omega_{H1} < SE$$



BH Area Thm and Thermodynamics

Hawking (1971)

$$\delta A_{BH} \geq 0$$

ASSUMING cosmic censorship
Null Energy Condition

BH Area

$$A_{BH} = \int \sqrt{|g_{\theta\theta} g_{\phi\phi}|} \Big|_{r=r_+} d\theta d\phi$$

$$= 4\pi (r_+^2 + a^2)$$

$$(r_+ = M + \sqrt{M^2 - a^2})$$

$$A_{BH} = 8\pi [M^2 + \sqrt{M^4 - J^2}]$$

$d\theta d\phi$

angular momentum

$$J = aM$$

$$\delta A_{BH} = \frac{J}{\sqrt{M^2 - J^2}} \left(\left(\frac{2M\sqrt{M^2 - J^2}}{J} + \frac{2M^3}{J} \right) \delta M - \delta J \right)$$

$d\phi$
 angular momentum
 $J = aM$

$$\frac{2r_+}{\bar{a}} = \Omega_H^{-1}$$

$$\frac{\delta A_{BH}}{\delta T_H} = \frac{\bar{a}}{\sqrt{1 - \bar{a}^2}} \left(\Omega_H^{-1} \delta M \right)$$



$$\delta A_{BH} = \frac{J}{\sqrt{M^2 - J^2}} \left(\left(\frac{2M\sqrt{M^2 - J^2}}{J} + \frac{2M^3}{J} \right) \delta M - \delta J \right)$$

$$\frac{2r_+}{\bar{a}} = \Omega_{H}^{-1}$$

$$\frac{\delta A_{BH}}{\delta T} = \frac{\bar{a}}{\sqrt{1 - \bar{a}^2}} \left(\Omega_{H}^{-1} \delta M - \delta J \right)$$

$$\delta A_{BH} \geq 0 \Leftrightarrow \delta J \Omega_{H} < \delta M$$

$$M_{ir} = \sqrt{\dots}$$

$$\delta A_{BH} = \frac{J}{\sqrt{M^2 - J^2}} \left(\left(\frac{2M\sqrt{M^2 - J^2}}{J} + \frac{2M^3}{J} \right) \delta M - \delta J \right)$$

$$\frac{2r_+}{\bar{a}} = \Omega_{H}^{-1}$$

$$\frac{\delta A_{BH}}{\delta T} = \frac{\bar{a}}{\sqrt{1 - \bar{a}^2}} \left(\Omega_{H}^{-1} \delta M - \delta J \right)$$

$\delta A_{BH} \geq 0 \Rightarrow \delta J \Omega_{H}^{-1} < \delta M$

$$M_{ir} = \sqrt{\frac{A_{BH}}{16\pi}}$$

$$\bar{a} = 0 \quad M_{ir} = M_{BH}$$

$$\bar{a} = 1 \quad M_{ir} = \frac{1}{\sqrt{2}} M_{BH}$$

$$E_{rot} = M_{BH} - M_{ir} \hat{=} 0.29 M_{BH} \text{ for } \bar{a} = 1$$