Title: Emergent classical gauge symmetry from quantum entanglement

Speakers: Joshua Kirklin

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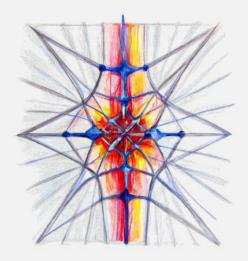
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Abstract: Inspired by the emergence of bulk diffeomorphism invariance in holography, I will give an explicit description how entanglement between quantum subsystems can lead to emergent gauge symmetry in a classical limit. Along the way, I will provide a precise characterisation of when it is consistent to treat a quantum subsystem classically in such a limit, and show that this gives strong constraints on the entanglement structure of classical states. I will explain how this generically leads to emergent fundamentally non-local classical degrees of freedom, which may nevertheless be accounted for in a kinematically local way if one employs an appropriately redundant description. The mechanism I describe is general and elementary, but for concreteness I will exhibit a toy example involving three entangled spins at high angular momentum, and I will also describe a significant generalisation of this toy example based on coadjoint orbits. If there is time, I will discuss evidence for the role this phenomenon plays in gravity. This talk is based on arXiv:2209.03979.

Zoom link: https://pitp.zoom.us/j/92066956880?pwd=OTRySTIOVGgvM3RCRmkzWHFVSUF3Zz09

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# Emergent classical gauge symmetry from quantum entanglement



Josh Kirklin

Okinawa Institute of Science and Technology 沖縄科学技術大学院大学

Seminar at Perimeter Institute, 15th December 2022

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Any theory of quantum gravity must in some way account for bulk diffeomorphism invariance. There are basically two options:

- 1. Diffeomorphism invariance is fundamentally part of the quantum theory.
- 2. Diffeomorphism invariance is only *emergent* in the semiclassical regime.

Evidence to take the latter seriously: AdS/CFT. The bulk spacetime itself is emergent at large N. So diffeomorphisms of that bulk are also emergent.

Diffeomorphism invariance is a kind of gauge symmetry. What does it mean for a gauge symmetry to be emergent?

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Gauge symmetries are *redundancies* in our description of reality, whose purpose is to allow for a better conceptual and computational grasp of the underlying physics.

In differing regimes, we may use different descriptions of reality, which are redundant in different ways.

A gauge symmetry is *emergent* if our description of physics in a more fundamental regime is less redundant than our description in a less fundamental regime.

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quantum theory \xrightarrow{classical\ limit} classical theory more redundant description (emergent\ gauge\ symmetry)
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In field theory and gravity, gauge symmetries allow us to use mathematically local structures to describe non-local degrees of freedom.

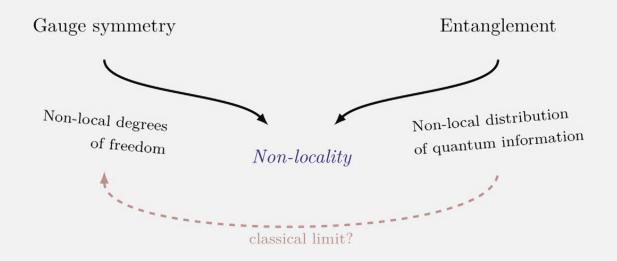
So, instead of emergent gauge symmetry, you may also think of this talk as being about emergent non-local degrees of freedom.

[Rovelli, 2013, "Why gauge?"] [Witten, 2016, "Symmetry and Emergence"]

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In holography, the emergence of the bulk appears to be largely quantum information theoretic in nature.

I will focus on one aspect of this story: the link between *gauge symmetry* and *entanglement*. Motivation comes from gravity, but I won't restrict to the gravitational setting.



Quantum entanglement would give rise to emergent classical gauge symmetry.

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This talk: Identify the mechanism for this to happen.

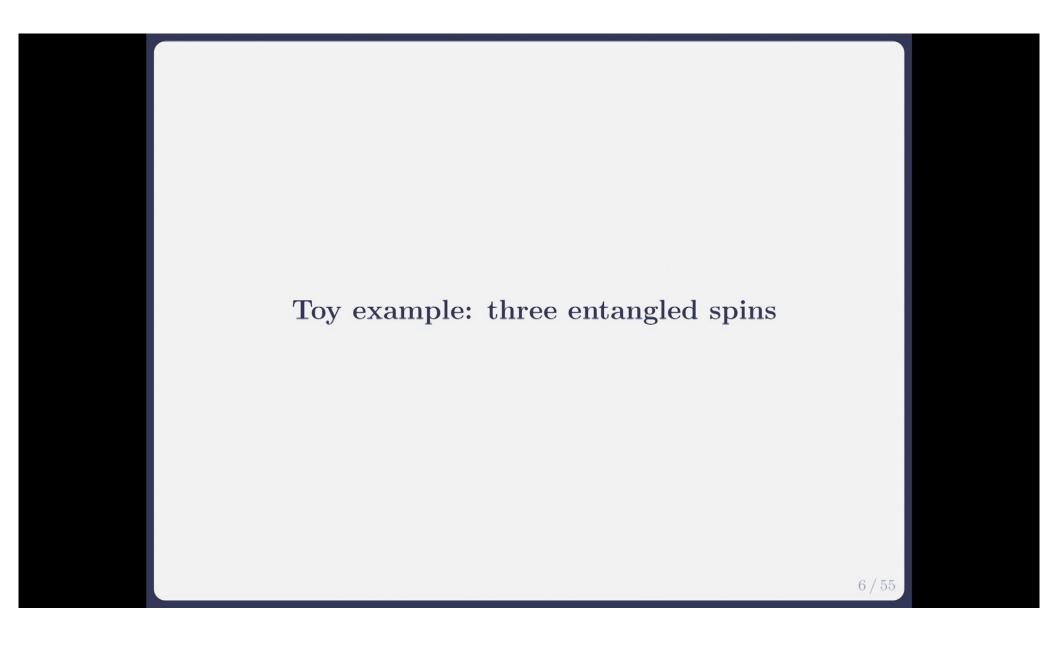
Precise and elementary, but also very general (not model-specific).

#### Takeaways:

- You don't need something like traditional constraint quantisation to quantise a theory with gauge symmetry you can use entanglement instead.
- The structure of multipartite entanglement simplifies significantly in the classical limit you can describe it with gauge symmetry.

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Consider a spin j with Hilbert space  $\mathcal{H}$ . Let  $\hat{\mathbf{J}}$  be the angular momentum operator, and let  $|j,m\rangle$  be usual eigenbasis.

Spin coherent states are an overcomplete basis of  $\mathcal{H}$ :

$$|\mathbf{n}\rangle = \hat{D}(\mathbf{n}) |j, -j\rangle, \text{ where } \hat{D}(\mathbf{n}) = \exp\left(i\frac{\theta}{\sin(\theta)}(\mathbf{n}_i \times \mathbf{e}^3) \cdot \hat{\mathbf{J}}\right),$$

where **n** is a unit 3-vector of angle  $\theta$  from  $\mathbf{e}^3 = (0, 0, 1)$ .

These states provide us with a notion of a classical limit at large j.

$$\lim_{j \to \infty} |\langle \mathbf{n} | \mathbf{n}' \rangle|^2 = \begin{cases} 1 & \text{if } \mathbf{n} = \mathbf{n}', \\ 0 & \text{otherwise.} \end{cases}$$

The classical state space is  $S^2$ , and any classical observable  $A:S^2\to\mathbb{C}$  has an operator representation

$$\hat{A} = \int_{S^2} \frac{\mathrm{d}^2 \mathbf{n}}{4\pi} (2j+1) |\mathbf{n}\rangle \langle \mathbf{n}| A(\mathbf{n}),$$

such that

$$\hat{A} | \mathbf{n} \rangle \approx A(\mathbf{n}) | \mathbf{n} \rangle$$
.

## Classical limit of unentangled spins

Suppose we have three spins  $j_1, j_2, j_3$ .

We can get a classical limit by defining

$$|\mathbf{n}_1, \mathbf{n}_2, \mathbf{n}_3\rangle = |\mathbf{n}_1\rangle \otimes |\mathbf{n}_2\rangle \otimes |\mathbf{n}_2\rangle$$

These states satisfy

$$\lim_{j_1,j_2,j_3\to\infty} \left| \langle \mathbf{n}_1,\mathbf{n}_2,\mathbf{n}_3|\mathbf{n}_1',\mathbf{n}_2',\mathbf{n}_3' \rangle \right|^2 = \begin{cases} 1 & \text{if } \mathbf{n}_i = \mathbf{n}_i', \quad i = 1,2,3, \\ 0 & \text{otherwise.} \end{cases}$$

The classical state space is  $S^2 \times S^2 \times S^2$ .

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## Classical limit of entangled spins

Let  $|0,0\rangle \in \mathcal{H} = \mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$  be the unique state of zero angular momentum.

This is a highly entangled state:

$$|0,0\rangle = \sum_{m_1=-j_1}^{j_1} \sum_{m_1=-j_2}^{j_2} \sum_{m_3=-j_3}^{j_3} \underbrace{\begin{pmatrix} j_1 & j_2 & j_3 \\ m_1 & m_2 & m_3 \end{pmatrix}}_{\text{Wigner } 3j\text{-symbol}} |j_1,m_1\rangle \otimes |j_2,m_2\rangle \otimes |j_3,m_3\rangle.$$

Act with SU(2) representations of the first two spins:

$$|\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2\rangle = (U_1(\psi_1, \mathbf{m}_1) \otimes U_2(\psi_2, \mathbf{m}_2) \otimes \mathbb{1}_3) |0, 0\rangle.$$

where  $U_i(\psi_i, \mathbf{m}_i) = \exp(2i\psi_i \,\mathbf{m} \cdot \hat{\mathbf{J}}^i)$ .

Here  $\psi_1, \psi_2 \in [0, \pi)$  and  $\mathbf{m}_1, \mathbf{m}_2$  are unit 3-vectors. Each of  $(\psi_1, \mathbf{m}_1)$  and  $(\psi_2, \mathbf{m}_2)$  parametrise points on  $S^3$ , the group manifold of SU(2).

Thus, we have an  $S^3 \times S^3$  of such states.

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## Classical limit of entangled spins

These states form an overcomplete basis for  $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$ . Moreover:

$$\lim_{j_1,j_2,j_3\to\infty} \left| \langle \psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2 | \psi_1', \psi_2'; \mathbf{m}_1', \mathbf{m}_2' \rangle \right|^2 = \begin{cases} 1 & \text{if } \psi_i = \psi_i' \text{ and } \mathbf{m}_i = \mathbf{m}_i', \\ 0 & \text{otherwise.} \end{cases}$$

Thus, these states give a classical limit, with classical state space  $S^3 \times S^3$ .

Classical observables  $A:S^3\times S^3\to \mathbb{C}$  may be represented as operators:

$$\hat{A} = \int_{S^3 \times S^3} d\mu \ A(\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2) \ |\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2\rangle \langle \psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2|,$$

(for some measure  $\mu$ ) which obey

$$\hat{A} | \psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2 \rangle \approx A(\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2) | \psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2 \rangle$$
.

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## Classical limit of entangled spins

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Thus, we have an  $S^3 \times S^3$  of such states.

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Two classical limits of the same quantum system:

states:  $|\mathbf{n}_1, \mathbf{n}_2, \mathbf{n}_3\rangle$   $|\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2\rangle$ 

regime:  $j_1, j_2, j_3 \to \infty$   $j_1, j_2, j_3 \to \infty$ 

entanglement: separable highly entangled

Hilbert space:  $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$   $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$ 

classical state space:  $S^2 \times S^2 \times S^2$   $S^3 \times S^3$ 

local structure: preserved ?

What happened to the local structure (i.e. decomposition into three spins)?

To answer this, consider

$$\rho_{i}(\psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}) = \operatorname{tr}_{\overline{i}} |\psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}\rangle \langle \psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}|,$$

$$\rho_{ij}(\psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}) = \operatorname{tr}_{\overline{i}\overline{j}} |\psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}\rangle \langle \psi_{1}, \psi_{2}; \mathbf{m}_{1}, \mathbf{m}_{2}|,$$

i.e. the reduced states of each spin i and pair of spins ij.

One may show:

$$\rho_i(\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2) = \frac{\mathbb{1}_i}{2j_i + 1},$$

so there is actually *only one* possible classical state for each individual spin.

Similarly (w.l.o.g. consider ij = 23):

$$\rho_{23}(\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2) \approx \frac{\hat{\pi}_{23}(\psi_1, \mathbf{m}_1)}{2j_1 + 1},$$

where

$$\hat{\pi}_{23}(\psi, \mathbf{m})\hat{\pi}_{23}(\psi', \mathbf{m}') \approx \delta_{\psi\psi'}\delta_{\mathbf{m}\mathbf{m}'}\hat{\pi}_{23}(\psi, \mathbf{m})$$

are a set of approximately mutually orthogonal projection operators, labelled by points in  $S^3$ . Thus, there are  $S^3$  possible classical states for each pair of spins.

Aside: 'classically resolvable'. See later.

To recover the local structure we can use gauge symmetry.

For each spin define a space of kinematical states  $\mathcal{N}_i^{\text{kin.}} = \text{SU}(2) = S^3$ , and introduce a gauge group G = SU(2) that acts simultaneously (i.e. diagonally) from the right.

Then the space of *physical* states for the full system is

$$\frac{\mathcal{N}_1^{\text{kin.}} \times \mathcal{N}_2^{\text{kin.}} \times \mathcal{N}_3^{\text{kin.}}}{G} = \frac{\text{SU}(2) \times \text{SU}(2) \times \text{SU}(2)}{\text{SU}(2)} = \text{SU}(2) \times \text{SU}(2) = S^3 \times S^3.$$

Similarly, the space of *physical* states of each individual spin is a singleton:

$$\frac{\mathcal{N}_i^{\text{kin.}}}{G} = \frac{\text{SU}(2)}{\text{SU}(2)},$$

while the space of *physical* states for each pair of spins is an  $S^3$ :

$$\frac{\mathcal{N}_i^{\text{kin.}} \times \mathcal{N}_j^{\text{kin.}}}{G} = \frac{\text{SU}(2) \times \text{SU}(2)}{\text{SU}(2)} = \text{SU}(2) = S^3,$$

as required. This  $S^3$  parametrises bilocal degrees of freedom: the kinematical state of one spin relative to the kinematical state of another.

Two classical limits of the same quantum system:

states:  $|\mathbf{n}_1, \mathbf{n}_2, \mathbf{n}_3\rangle$   $|\psi_1, \psi_2; \mathbf{m}_1, \mathbf{m}_2\rangle$ 

regime:  $j_1, j_2, j_3 \to \infty$   $j_1, j_2, j_3 \to \infty$ 

entanglement: separable highly entangled

Hilbert space:  $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$   $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$   $\leftarrow$  no constraints

classical state space:  $S^2 \times S^2 \times S^2$   $S^3 \times S^3$ 

local structure: preserved preserved

gauge symmetry: none SU(2)

Thus, quantum entanglement has led to emergent classical gauge symmetry.

This was a toy example. For most of the rest of the talk, I will describe the general mechanism underlying this phenomenon.

To recover the local structure we can use gauge symmetry.

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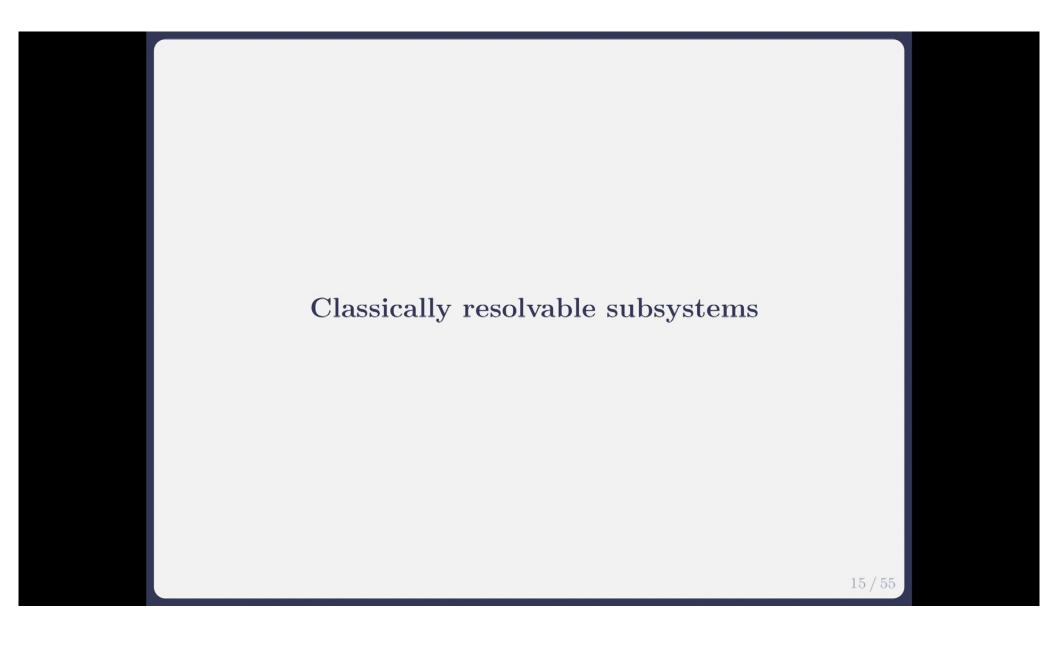
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## Locality and the classical limit

Physical systems have 'local structure' if they are *composite*, i.e. divisible into subsystems (e.g. the spins, subregion in QFT, ...).

Each subsystem s has a set of observables  $\mathcal{O}_s$  'local to' that subsystem.

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If  $A \in \mathcal{O}_{s \cup s'}$  cannot be formed as a combination of observables in  $\mathcal{O}_s$  and  $\mathcal{O}_{s'}$ , then A measures non-local degrees of freedom (e.g. A is a Wilson line).

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To understand *emergent gauge symmetry*, we need to know what happens to *quantum* subsystems in a *classical limit*. Actually not all quantum subsystems will be well-behaved in this limit (e.g. very small subregion in QFT).

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## What is a quantum subsystem?

A quantum subsystem is a von Neumann algebra  $A_s \subset L(\mathcal{H})$ , where  $\mathcal{H}$  is the Hilbert space of the full system.

In this talk: assume  $\dim(\mathcal{H})$  finite. Will also assume no pre-existing quantum gauge symmetry. This implies  $\mathcal{A}_s$  is a *Type I factor*. Can then write:

$$\mathcal{A}_s = \mathcal{B}(\mathcal{H}_s) \otimes \mathbb{1}_{\bar{s}}, \quad \text{where} \quad \mathcal{H} = \overbrace{\mathcal{H}_s}^{\text{Hilbert space of } s} \otimes \underbrace{\mathcal{H}_{\bar{s}}}_{\text{Hilbert space of complement of } s}$$

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Hilbert space of s

Hilbert space of complement of s

More generally:

$$\mathcal{A}_i = \mathbb{1}_1 \otimes \cdots \otimes \mathbb{1}_{i-1} \otimes \mathcal{B}(\mathcal{H}_i) \otimes \mathbb{1}_{i+1} \otimes \cdots \otimes \mathbb{1}_n,$$

where

$$\mathcal{H} = \mathcal{H}_1 \otimes \cdots \otimes \mathcal{H}_i \otimes \cdots \otimes \mathcal{H}_n$$
.

is a tensor factorisation into subsystem Hilbert spaces  $\mathcal{H}_i$ .

#### What is a classical limit?

Let  $\mathcal{N}$  be space of classical states,  $C(\mathcal{N})$  be set of functions  $\mathcal{N} \to \mathbb{C}$ , and pick a set of operators  $\mathcal{C} \subset L(\mathcal{H})$  that is 'approximately isomorphic' to  $C(\mathcal{N})$ :

$$C(\mathcal{N}) \to \mathcal{C}, \quad A \mapsto \hat{A},$$

i.e.

$$\widehat{A^*} \approx \widehat{A}^{\dagger}, \quad \widehat{AB} \approx \widehat{A}\widehat{B}, \quad \widehat{(\alpha A + \beta B)} \approx \alpha^* \widehat{A}^{\dagger} + \beta^* \widehat{B}^{\dagger}, \quad \text{etc...}$$

(" $\approx$ " denotes equality in the classical limit  $\chi \to 0$  where  $\chi = \hbar, G, 1/N^2, \ldots$ )

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(" $\approx$ " denotes equality in the classical limit  $\chi \to 0$  where  $\chi = \hbar, G, 1/N^2, \ldots$ )

At leading order in  $\chi$ , the map  $A(x) \mapsto \hat{A}$  is implemented by

$$\hat{A} \approx \int_{\mathcal{N}} d\mu(x) \frac{N}{N(x)} \hat{\pi}(x) A(x).$$

for some measure  $\mu$  on  $\mathcal{N}$  and projection operators  $\hat{\pi}(x)$  obeying

$$\hat{\pi}(x)\hat{\pi}(y) \approx \delta_{xy}\hat{\pi}(x)$$
, for all  $x, y \in \mathcal{N}$ ,

Here,  $N = \dim(\mathcal{H})$  and  $N(x) = \operatorname{rank}(\hat{\pi}(x))$ .

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## 'Complete' classical limit

By measuring  $\hat{\pi}(x)$  we can determine with high precision if the state of the classical degrees of freedom is x.

If  $N(x) = \operatorname{rank}(\hat{\pi}(x)) > 1$ , then there is more than one quantum state consistent with a given classical state x. This indicates that there are still some 'left over' quantum degrees of freedom in the classical limit.

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## 'Complete' classical limit

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If  $N(x) = \operatorname{rank}(\hat{\pi}(x)) > 1$ , then there is more than one quantum state consistent with a given classical state x. This indicates that there are still some 'left over' quantum degrees of freedom in the classical limit.

A 'complete' classical limit is one for which N(x) = 1, so that classical degrees of freedom suffice to determine the full state. Then we may write

$$\hat{\pi}(x) = |x\rangle \langle x|,$$

and

$$\hat{A} \approx \int_{\mathcal{N}} d\mu(x) N |x\rangle \langle x| A(x),$$

and we have

$$\hat{A}|x\rangle \approx A(x)|x\rangle$$
 for all  $\hat{A} \in \mathcal{C}, x \in \mathcal{N}$ .

## What is the classical limit of a subsystem?

Two sets of operators:

- $A_s = \mathcal{B}(\mathcal{H}_s) \otimes \mathbb{1}_{\bar{s}}$  defining quantum subsystem s.
- $\mathcal{C}$  defining (complete) classical limit.

To understand the classical limit of the subsystem, consider the intersection:

$$C_s = A_s \cap C$$
.

This consists of operators measuring classical degrees of freedom in s.

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This consists of operators measuring classical degrees of freedom in s.

Gelfand-Naimark theorem implies  $C_s$  is approximately isomorphic to an algebra  $C(\mathcal{N}_s)$  of functions on a space  $\mathcal{N}_s$  of classical subsystem states.

Explicitly:

$$\mathcal{N}_s = \left\{ x_s \mid x_s : \hat{A} \mapsto \langle x | \hat{A} | x \rangle, x \in \mathcal{N} \right\} \subset \mathcal{C}_s^*.$$

The classical observable  $A_s \in C(\mathcal{N}_s)$  corresponding to  $\hat{A} \in \mathcal{C}_s$  is defined via  $A_s(x_s) = x_s(\hat{A})$ .

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This construction of  $\mathcal{N}_s$  works for any quantum subsystem. But in general  $\mathcal{N}_s$  doesn't fully account for physics in s.

It is not guaranteed that knowledge of classical degrees of freedom in s suffices to determine complete state of s.

So a complete classical limit for the full system does not necessarily imply its quantum subsystems behave in a completely classical way.

Consider classical subsystem operators:

$$C_{s} = \left\{ \hat{A} \mid \hat{A} = \hat{A}_{s} \otimes \mathbb{1}_{\bar{s}} \approx \int_{\mathcal{N}} d\mu(x) N |x\rangle \langle x| A(x) \right\}$$

In an extreme case: the only  $\hat{A}_s$  satisfying this condition is  $\hat{A}_s \propto \mathbb{1}_s$ . Then there is only one element in  $\mathcal{N}_s$ , i.e. no classical degrees of freedom.

We would need to describe the subsystem in a completely quantum way.

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## Classical resolvability

We are interested in the opposite case, where classical degrees of freedom suffice to describe subsystem.

Suppose we know full system is in some classical state in  $\mathcal{N}$  (but we don't know which one).

We will say subsystem s is classically resolvable if knowledge of  $x_s$  suffices to determine its quantum state  $\rho_s$  (to a high degree of accuracy in the classical limit).

More precisely, suppose  $x, y \in \mathcal{N}$  correspond to  $x_s, y_s \in \mathcal{N}_s$ . We already know:

$$\rho_s(x) = \rho_s(y) \implies x_s = y_s$$

(by definition of  $x_s, y_s$ ). Subsystem s is classically resolvable if reverse is true:

$$x_s = y_s \implies \rho_s(x) = \rho_s(y).$$

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## States of classically resolvable subsystem

Suppose s classically resolvable,  $\hat{A} \in \mathcal{C}_s$ .

$$\hat{A} = \hat{A}_s \otimes \mathbb{1}_{\bar{s}} \approx \int_{\mathcal{N}} d\mu(x) \, N \, |x\rangle \, \langle x| \, A(x). \tag{*}$$

 $\frac{1}{N_{\bar{s}}} \operatorname{tr}_{\bar{s}}$  both sides, where  $N_{\bar{s}} = \dim(\mathcal{H}_{\bar{s}})$ :

$$\hat{A}_s \approx \int_{\mathcal{N}} d\mu(x) N_s \rho_s(x) A_s(x_s),$$

where  $N_s = \dim(\mathcal{H}_s)$ , and  $A_s(x_s) = x_s(\hat{A}_s) = \langle x | (\hat{A}_s \otimes \mathbb{1}_{\bar{s}}) | x \rangle = A(x)$ .

Classical resolvability means we can set

$$A_s(x_s) \propto \begin{cases} 1 & \text{if } \rho_s(x) \approx \rho_s(y), \\ 0 & \text{otherwise.} \end{cases}$$

Then  $\hat{A}_s$  approximately proportional to  $\rho_s(x)$ .

But (\*) implies  $\hat{A}_s$  is approximately proportional to a projection operator.

## States of classically resolvable subsystem

Thus, in a classically resolvable subsystem s, the reduced density matrix  $\rho_s(x)$  is always approximately proportional to a projection operator, for any  $x \in \mathcal{N}$ :

$$\rho_s(x) \approx \frac{\hat{\pi}_s(x_s)}{N_s(x_s)} \quad \text{where} \quad N_s(x_s) = \text{rank}(\hat{\pi}_s(x_s)).$$

(c.f. toy model)

The projection operator  $\hat{\pi}_s(x_s)$  is a classical operator. It measures whether  $x_s$  is the classical state of subsystem s.

The classical degrees of freedom in subsystem s can't be in more than one state:

$$\hat{\pi}_s(x_s)\hat{\pi}_s(y_s) \approx \delta_{x_s y_s}\hat{\pi}_s(x_s).$$

 $\rho_s(x)$  determines the way in which s is entangled with other subsystems. So these are strong constraints on entanglement.

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## Classically resolvable subsystem $\Rightarrow$ complete classical limit for subsystem

General classical subsystem operator for classically resolvable subsystem:

$$\hat{A}_s \approx \int_{\mathcal{N}_s} d\mu_s(x_s) \, \frac{N_s}{N_s(x_s)} \, \hat{\pi}_s(x_s) A_s(x_s),$$

 $\mu_s$  the pushforward of  $\mu$  to  $\mathcal{N}$ .

So subsystem s may be treated with a self-contained classical limit of the kind previously described.

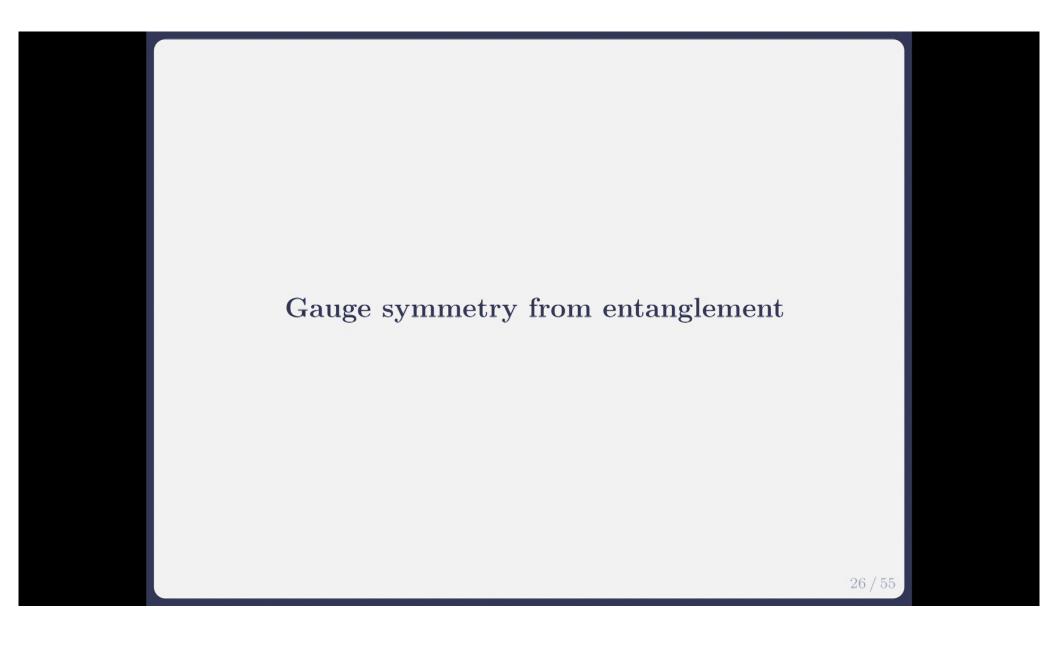
Note: in general  $\hat{\pi}_s(x_s)$  has rank greater than 1, so this is not a complete classical limit.

But a classically resolvable system can be described using only classical degrees of freedom...

This is consistent because we are assuming s is part of a larger completely classical system. This is extra information compared to before, where we only considered classical limits of isolated systems.

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#### Gauge symmetry from entanglement

Take a completely classical limit of a quantum system. Classical states  $x \in \mathcal{N}$  corresponding to quantum states  $|x\rangle \in \mathcal{H}$ .

Assume the quantum system has a classically resolvable 'local structure', i.e. division into classically resolvable subsystems (and unions of subsystems also classically resolvable). Then we decompose

$$\mathcal{H} = \mathcal{H}_1 \otimes \cdots \otimes \mathcal{H}_i \otimes \cdots \otimes \mathcal{H}_n$$
.

For each subsystem  $s_i$  there is a classical space  $\mathcal{N}_i$  of subsystem states, and a map from  $x \in \mathcal{N}$  to corresponding  $x_i \in \mathcal{N}_i$ . Reduced states are approximately proportional to projection operators acting on  $\mathcal{H}_i$ :

$$\rho_i(x) \approx \frac{\hat{\pi}_i(x_i)}{N_i(x_i)} \quad \text{where} \quad N_i(x_i) = \text{rank}(\hat{\pi}_i(x_i)).$$

Classical subsystem operators may be written

$$\hat{A}_i \approx \int_{\mathcal{N}_i} d\mu_i(x_i) \frac{N_i}{N_i(x_i)} \hat{\pi}_i(x_i) A_i(x_i)$$
 where  $N_i = \dim(\mathcal{H}_i)$ .

Projection operators are mutually approximately orthogonal:

$$\hat{\pi}_i(x_i)\hat{\pi}_i(y_i) \approx \delta_{x_iy_i}\hat{\pi}_i(x_i).$$

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# Gauge symmetry from entanglement

 $\rho_i(x)$  tells us how  $s_i$  is entangled with the other subsystems in the state x.

In this case,  $\rho_i(x) \propto \hat{\pi}_i(x_i)$  means some part of  $s_i$  is maximally entangled with some part of its complement.

Roughly speaking:  $s_i$  and its complement share  $\log_2(N_i(x_i))$  maximally entangled qubits / Bell pairs.

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#### The unentangled case

Sanity check: *separable subsystems*. Should be no gauge symmetry. Not hard to confirm this.

Unentangled subsystems implies  $\rho_i(x)$  is rank 1, so  $\rho_i(x) = |x_i\rangle \langle x_i|$  for some  $|x_i\rangle \in \mathcal{H}_i$ . We have  $\langle x_i|y_i\rangle \approx \delta_{x_iy_i}$ . Classical subsystem operators may be written

$$\hat{A}_i \approx \int_{\mathcal{N}_i} d\mu_i(x_i) N_i |x_i\rangle \langle x_i| A_i(x_i).$$

So separable subsystems undergo a *complete classical limit*, unlike entangled case.

Set of subsystem states  $x_i$  determines full system state via

$$|x\rangle\langle x| = |x_1\rangle\langle x_1|\otimes\cdots\otimes|x_i\rangle\langle x_i|\otimes\cdots\otimes|x_n\rangle\langle x_n|$$
.

x also determines  $x_1, \ldots, x_n$ , so we have a bijection

$$\mathcal{N} \longleftrightarrow \mathcal{N}_1 \times \cdots \times \mathcal{N}_i \times \cdots \times \mathcal{N}_n$$
.

No non-local degrees of freedom, so no gauge symmetry.

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### The entangled case

Return to the entangled case. Classical subsystem operators:

$$C_i = \left\{ \hat{A}_i \approx \int_{\mathcal{N}_i} d\mu_i(x) \, \frac{N_i}{N_i(x_i)} \hat{\pi}_i(x_i) A_i(x_i) \mid A_i : \mathcal{N}_i \to \mathbb{C} \right\}.$$

Define

$$C_{\text{local}} = C_1 \otimes \cdots \otimes C_i \otimes \cdots \otimes C_n.$$

This is the set of operators which only depend on local degrees of freedom.

Most general  $\hat{A}_{local} \in \mathcal{C}_{local}$ :

$$\hat{A}_{local} \approx \int_{\mathcal{N}_1} d\mu_1(x_1) \cdots \int_{\mathcal{N}_n} d\mu_n(x_n) \ N \,\hat{\pi}_{local}(x_1, \dots, x_n) \, A_{local}(x_1, \dots, x_n),$$

where

$$\hat{\pi}_{local}(x_1,\ldots,x_n) = \hat{\pi}_1(x_1) \otimes \cdots \otimes \hat{\pi}_n(x_n)$$

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## Non-local degrees of freedom

Define  $C_{\text{non-local}} = C \setminus C_{\text{local}}$ .

If  $\mathcal{C}_{\text{non-local}}$  is non-empty, then there are non-local classical degrees of freedom.

Suppose  $|x\rangle$  is entangled, and  $\hat{A}_{local} \in \mathcal{C}_{local}$  satisfies  $\hat{A}_{local} |x\rangle = 1$ . Then

$$\operatorname{rank}(\hat{A}_{\operatorname{local}}) \ge \operatorname{rank}(\hat{\pi}_{\operatorname{local}}(x_1, \dots, x_n)) > 1.$$

On the other hand,  $|x\rangle\langle x|$  is rank 1 and  $|x\rangle\langle x|\in\mathcal{C}$ .

Therefore  $C_{\text{non-local}}$  is non-empty: it contains  $|x\rangle \langle x|$ .

 $|x\rangle\langle x|$  measures if the classical state is x. Thus, to know if the classical state is x, we have to measure non-local degrees of freedom.

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#### Non-local degrees of freedom

N.B. presence of non-local degrees of freedom depends on structure of entanglement between subsystems in the state  $|x\rangle$ .

Separable implies no non-local degrees of freedom.

Subsystems could be *entangled* in some states, but *separable* in others. Moreover, when entangled, can be entangled in different ways.

One can show that the number of non-local degrees of freedom shared by two subsystems  $s_i, s_j$  is counted by their mutual information

$$I_{i:j}(x) = \operatorname{tr}_{ij}(\rho_{ij}(x)\log\rho_{ij}(x)) - \operatorname{tr}_{i}(\rho_{i}(x)\log\rho_{i}(x)) - \operatorname{tr}_{j}(\rho_{j}(x)\log\rho_{j}(x)).$$

They share non-local degrees of freedom if and only if the mutual information is non-vanishing (in the classical limit).

Can be interpreted as a variable 'bulk topology'.

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#### Kinematical states

Have shown that entanglement between classically resolvable subsystems leads to classically emergent non-local degrees of freedom.

Now: account for these non-local degrees of freedom using gauge symmetry.

First step is to construct a space of kinematical states for each subsystem. A kinematical state will include a purification of  $\rho_i(x)$ :

$$|\psi_i\rangle \in \mathcal{H}_i \otimes \widetilde{\mathcal{H}}_i(x_i)$$
 such that  $\widetilde{\operatorname{tr}}_i(|\psi_i\rangle \langle \psi_i|) = \rho_i(x)$ ,

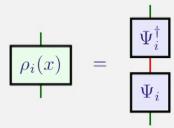
where  $\widetilde{\mathcal{H}}_i(x_i)$  is an auxiliary Hilbert space of sufficiently high dimension.

For simplicity we can set  $\widetilde{\mathcal{H}}_i(x_i) = \mathbb{C}^{N_i(x_i)}$ .

We are adding  $\log_2 N_i(x_i)$  qubits to the subsystem — these are edge modes.

More notationally convenient to view the purification as a map  $\Psi: \mathcal{H}_i \to \widetilde{\mathcal{H}}_i(x_i)^*$ .

The purification condition may then be written  $\rho_i(x) = \Psi_i^{\dagger} \Psi_i$ .



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#### Kinematical states

Kinematical subsystem state space:

$$\mathcal{N}_i^{\text{kin.}} = \{ (x_i, \Psi_i) \mid x_i \in \mathcal{N}_i, \, \Psi_i \in \mathcal{P}_i(x_i) \},$$

where  $\mathcal{P}_i(x_i)$  is the space of purifications of  $\rho_i(x)$ .

Full kinematical state space:

$$\mathcal{N}^{\text{kin.}} = \mathcal{N}_1^{\text{kin.}} \times \cdots \times \mathcal{N}_i^{\text{kin.}} \times \cdots \times \mathcal{N}_n^{\text{kin.}}$$

A general kinematical state may be written

$$((x_1, \Psi_1), \ldots, (x_i, \Psi_i), \ldots, (x_n, \Psi_n)) \in \mathcal{N}^{\text{kin.}}$$

How do we project to physical states? Two steps: impose constraints, then gauge reduce.

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Let  $X: x \mapsto (x_1, \dots, x_n)$  map full system states  $x \in \mathcal{N}$  to corresponding subsystem states  $x_i \in \mathcal{N}_i$ , and let

$$image(X) = \overline{\mathcal{N}_1 \times \cdots \times \mathcal{N}_n}.$$

First constraint is almost trivial:

$$(x_1,\ldots,x_n)\in\overline{\mathcal{N}_1\times\cdots\times\mathcal{N}_n}.$$

So  $(x_1, \ldots, x_n)$  can come from at least one state of the full system.

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Let us introduce some additional reference structures. Analogous to e.g. coordinate systems in gravity.

1. A section Y of X, i.e. a map

$$Y: \overline{\mathcal{N}_1 \times \cdots \times \mathcal{N}_1} \to \mathcal{N}$$

such that  $X \circ Y$  is the identity.

2. A choice of purification  $\Phi_i(x_i) \in \mathcal{P}_i(x_i)$  for each subsystem  $s_i$  and each  $x_i \in \mathcal{N}_i$ .

From this, define gluing states for  $(x_1, \ldots, x_n) \in \overline{\mathcal{N}_1 \times \cdots \times \mathcal{N}_n}$ :

$$|\sigma(x_1,\ldots,x_n)\rangle = N_1(x_1)\ldots N_n(x_n) \left(\Phi_1(x_1)\otimes\cdots\otimes\Phi_n(x_n)\right) |Y(x_1,\ldots,x_n)\rangle.$$

$$|\sigma(x_1, \dots, x_n)\rangle = N_1(x_1) \dots N_n(x_n)$$

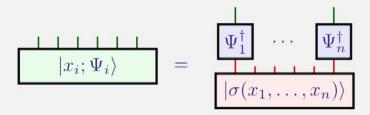
$$|Y(x_1, \dots, x_n)\rangle$$

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We can use these states to glue together the kinematical subsystem states.

If 
$$(x_1, \ldots, x_n) \in \overline{\mathcal{N}_1 \times \cdots \times \mathcal{N}_n}$$
, we define  $|x_i; \Psi_i\rangle \in \mathcal{H}$ , by

$$|x_i; \Psi_i\rangle = (\Psi_1^{\dagger} \otimes \cdots \otimes \Psi_n^{\dagger}) |\sigma(x_1, \dots, x_n)\rangle.$$



Essentially, take tensor product of purifications for each subsystem, then project auxiliary degrees of freedom onto  $|\sigma(x_1,\ldots,x_n)\rangle$ .

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One can show: we can obtain any physical state by gluing in this way.

But not all kinematical states when glued together will yield sensible classical states. Restricting to those which do:

$$\overline{\mathcal{N}^{\text{kin.}}} = \left\{ \left( (x_1, \Psi_1), \dots, (x_n, \Psi_n) \right) \in \mathcal{N}^{\text{kin.}} \mid (x_1, \dots, x_n) \in \overline{\mathcal{N}_1 \times \dots \times \mathcal{N}_n} \right.$$
and  $|x_i; \Psi_i\rangle = |y\rangle$  for some  $y \in \mathcal{N} \right\}.$ 

This is the 'constraint surface'  $\overline{\mathcal{N}^{\text{kin.}}} \subset \mathcal{N}^{\text{kin.}}$ .

Gauge reduction map is

$$R: \overline{\mathcal{N}^{\text{kin.}}} \to \mathcal{N}, \quad ((x_1, \Psi_1), \dots, (x_n, \Psi_n)) \mapsto |x_i; \Psi_i\rangle.$$

Any physical observable may be written as a function of the kinematical subsystem states. In particular, *non-local* physical observables can be decomposed into *local* kinematical observables.

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We can also glue together any proper subset of the full set of subsystems.

E.g. the physical states of a union of subsystems  $s_i \cup s_j$  are in one-to-one correspondence with the reduced states  $\rho_{ij}(x)$ , which can be obtained by gluing the kinematical states  $(x_i, \Psi_i)$  and  $(x_j, \Psi_j)$ :

$$\rho_{ij}(x) = (\Psi_i \otimes \Psi_j) \sigma_{ij}(x_i, x_j) (\Psi_i \otimes \Psi_j)^{\dagger},$$

where

$$\sigma_{ij}(x_i, x_j) = \widetilde{\operatorname{tr}}_{\overline{ij}} |\sigma(x_1, \dots, x_n)\rangle \langle \sigma(x_1, \dots, x_n)|.$$

Classical resolvability of  $s_i$ ,  $s_j$  and  $s_i \cup s_j$  ensure:  $\sigma_{ij}$  only depends on  $x_i, x_j$ .

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## Gauge transformations

A local gauge transformation is a change of purifications  $\Psi_i \to U_i^{\dagger} \Psi_i$ , with  $U_i \in U(N_i(x_i))$ . This changes the kinematical state of the subsystem, but not its physical state.

Thus, the *local* gauge group in subsystem  $s_i$  is  $U(N_i(x_i))$ .

In order to leave the physical state  $|x_i; \Psi_i\rangle$  of the full system invariant,  $U_i$  must obey

$$|\sigma(x_1,\ldots,x_n)\rangle = (U_1 \otimes \cdots \otimes U_n) |\sigma(x_1,\ldots,x_n)\rangle.$$

This defines the *global* gauge group:

$$G(x) = \operatorname{Stab}_{U(N_1(x_1)) \times \cdots \times U(N_n(x_n))} (|\sigma(x_1, \dots, x_n)\rangle).$$

N.B. these are state dependent gauge groups. (Same is true in gravity.)

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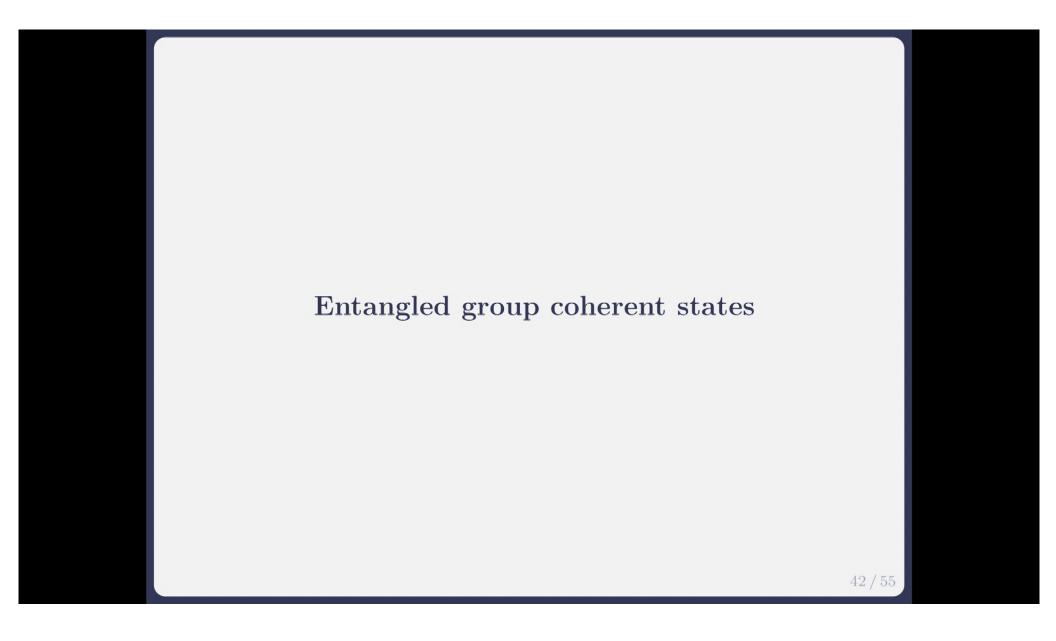
### Summary

- For subsystems to have good classical limits, they must be *classically* resolvable, meaning the subsystem state can be determined by classical measurements alone.
- Classical resolvability implies that the reduced state in any subsystem is approximately proportional to one of a set of mutually orthogonal projection operators,  $\rho_i(x) \propto \hat{\pi}_i(x_i)$ .
- If  $rank(\hat{\pi}_i(x_i)) > 1$ , then there are emergent non-local degrees of freedom.
- These can be accounted for by introducing a gauge symmetry where the kinematical states are *purifications* of  $\rho_i(x)$ . In other words,  $\log_2 N_i(x_i)$  auxiliary qubits as 'edge modes'.
- We can glue together arbitrary collections of subsystems by projecting auxiliary qubits onto entangled gluing states  $|\sigma(x_1,\ldots,x_n)\rangle$ .
- Any classical physical observable may be decomposed into a combination of local kinematical observables.

This provides a precise general picture of how entanglement leads to classically emergent gauge symmetry.

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### Entangled group coherent states

The toy model of three spins is a special case of a larger family based on unitary representations of Lie groups.

Let  $\mathcal{H} = \mathcal{H}_1 \otimes \cdots \otimes \mathcal{H}_n$ , and let  $G_i$  act unitarily and irreducibly on  $\mathcal{H}_i$ .

In separable case, there is a well-known construction of classical limits, with classical space of states  $\mathcal{N}_i$  for each subsystem being a *coadjoint orbit* of G.

Construct coherent states for full system by taking the tensor product of coherent states for the subsystems.

Then space of states for full system is  $\mathcal{N} = \mathcal{N}_1 \times \cdots \times \mathcal{N}_n$ , which is a coadjoint orbit of  $G = G_1 \times \cdots \times G_n$ .

There is no emergent gauge symmetry (consistent with no entanglement).

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## Entangled group coherent states

In the paper: a modification of this construction.

Involves the choice of a subgroup

$$H \subseteq G = G_1 \times \cdots \times G_n$$

with special properties. I explain how to construct a set of coherent states for full system, with entanglement determined by H.

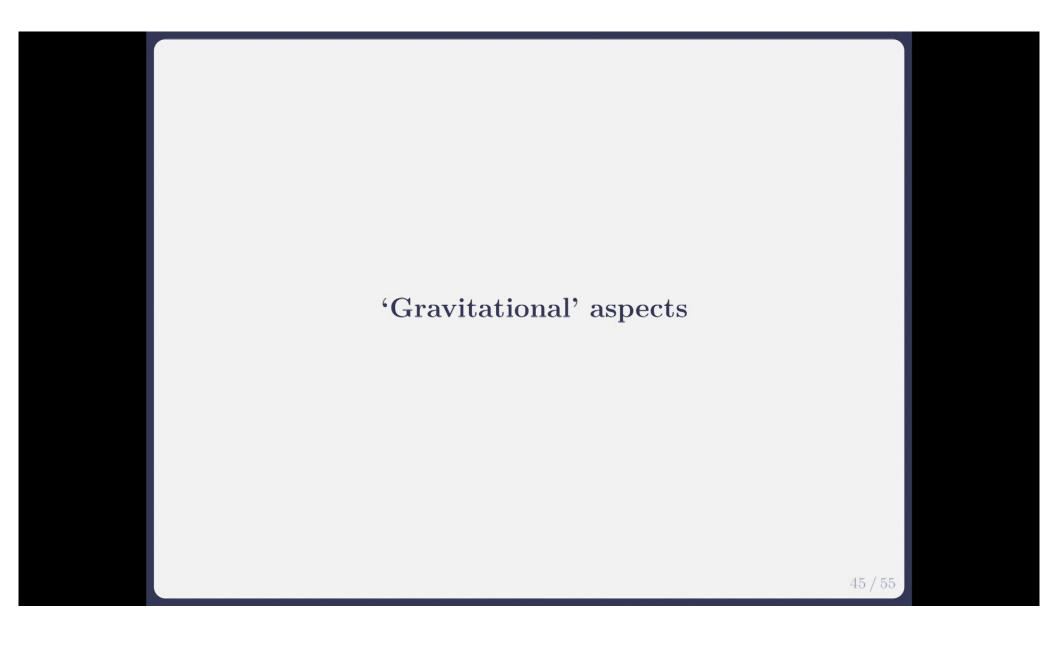
I show: this yields a good classical limit, and all subsystems are classically resolvable. Also: the entanglement leads to an emergent gauge symmetry, with a certain gauge group K satisfying

$$H \subseteq K \subseteq G$$
.

For the three spin toy model:  $G_i = SU(2)$ , H is the diagonal subgroup, and K = H.

This gives a very large and varied class of classical limits with gauge symmetry emerging from entanglement.

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## Gravity from entanglement

There is a general expectation that spacetime can emerge from structure of quantum entanglement.

So diffeomorphism invariance of that spacetime should also emerge from entanglement.

The mechanism described here is a very generic (model-independent) way for this to happen.

Not too much of a stretch to suggest that it is general enough to include the gravitational case.

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#### Classical resolvability of subregions in gravity

Reduced density matrix of spacetime subregion:

$$\rho = \frac{\exp(-\hat{A}/4G + \dots)}{\operatorname{tr}(\exp(-\hat{A}/4G + \dots))},$$

where  $\hat{A}$  is an area operator. [Jafferis, Lewkowycz, Maldecena, Suh, 2015]

In classical  $G \to 0$  limit, this is proportional to a projection operator onto minimal area states.

Can show fidelity of reduced density matrices  $\rho, \rho'$  for two different subsystem states obeys [Kirklin, 2019]

$$\operatorname{tr}\left(\sqrt{\sqrt{\rho}\rho'\sqrt{\rho}}\right) = \exp\left(-\mathcal{O}(1/G)\right) \to 0,$$

which implies projection operators for  $\rho, \rho'$  are approximately mutually orthogonal.

Thus, gravitational subsystems are *classically resolvable*, consistent with present work.

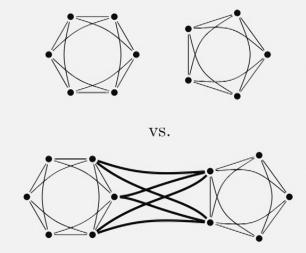
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# Variable 'bulk topology'

Different states have different kinds of entanglement, so different kinds of non-local degrees of freedom. Interpret this as different 'bulk topologies'.

For example:



Represents entanglement structure in two different states.

Dots are subsystems, lines are present when mutual information is non-vanishing.

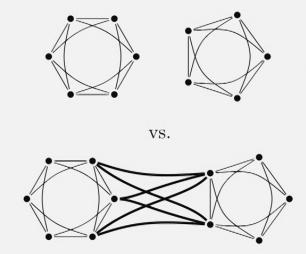
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### Modular symmetries

The emergent gauge transformations I have described are approximate *modular* symmetries of each subsystem — i.e. transformations which do not change the reduced density matrix.

In gravity, subregion modular symmetries are transformations of geometric edge modes. [Czech, de Boer, Ge, Lamprou, 2019]

So this is consistent. (Suggests a direct entanglement interpretation of geometric edge modes...)

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#### Quantum error correction

QEC has had important conceptual implications in gravity.

QEC also plays a role here. Suppose we take a classical limit of a system with Hilbert space  $\mathcal{H}$  and obtain a classical state space  $\mathcal{N}$  with emergent gauge symmetry. So

$$\mathcal{N}^{\text{kin.}} \supset \overline{\mathcal{N}^{\text{kin.}}} \to \mathcal{N}.$$

Suppose we do constrained quantisation of  $\mathcal{N}$ . So quantise  $\mathcal{N}^{\text{kin.}}$  to  $\mathcal{H}^{\text{kin.}}$ , and identify physical states

$$\mathcal{H}^{\mathrm{phys.}} \subset \mathcal{H}^{\mathrm{kin.}}$$

by imposing some operator constraints.

There is then a sense in which original Hilbert space  $\mathcal{H}$  is embedded as a code subspace of  $\mathcal{H}^{\text{kin.}}$ .

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#### Conclusion

- A general mechanism for the emergence of classical gauge symmetry from quantum entanglement.
- This proceeded from an understanding of classical resolvability.
- Toy model of three entangled spins, and a group-theoretic generalisation.
- Evidence that the mechanism could be responsible (in part) for diffeomorphism invariance in gravity.

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#### Future directions

- Phase space structure (Berry curvature  $\rightarrow$  symplectic form).
- Semi-classical limit (some degrees of freedom remain quantum... to what extent do they respect emergent gauge symmery?).
- $\dim(\mathcal{H}) = \infty$ , Type II and Type III subsystems, compatibility with pre-existing quantum gauge symmetry.
- Specific gravitational applications: coadjoint orbits of the *corner* symmetry group, relationship with spin network states, etc.
- Other implications of classical resolvability (e.g. entropy cone...).
- Dynamics... chaos, decoherence, etc.

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