Title: Newton's Cradle Spectra

Speakers: Barbara Soda

Series: Perimeter Institute Quantum Discussions

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Abstract: We present broadly applicable nonperturbative results on the behavior of eigenvalues and eigenvectors under the addition of self-adjoint operators and under the multiplication of unitary operators, in finite-dimensional Hilbert spaces. To this end, we decompose these operations into elementary 1-parameter processes in which the eigenvalues move similarly to the spheres in Newton's cradle. As special cases, we recover level repulsion and Cauchy interlacing. We discuss two examples of applications. Applied to adiabatic quantum computing, we obtain new tools to relate algorithmic complexity to computational slowdown through gap narrowing. Applied to information theory, we obtain a generalization of Shannon sampling theory, the theory that establishes the equivalence of continuous and discrete representations of information. The new generalization of Shannon sampling applies to signals of varying information density and finite length.

Zoom link: https://pitp.zoom.us/j/94120657832?pwd=SmpsWFhhVCtyeXM3a0pVQU9lMGFLdz09

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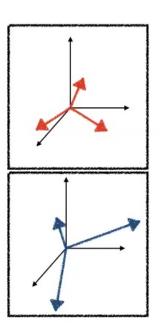
What happens when we add two Hamiltonians?

• Suppose we know the eigenvalues and eigenvectors of a Hamiltonian H_0 .

$$H_0 = \sum_i e_i \ket{e_i} \bra{e_i}$$

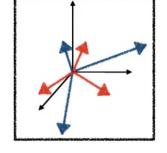
• We also know the eigenvectors and eigenvalues of H'.

$$H' = \sum_{i} f_{i} \ket{f_{i}} \bra{f_{i}}$$



What happens when we add two Hamiltonians?

- Typically, we approach this problem perturbatively.
- Nonperturbative results:
 - Wigner-von Neumann on level repulsion
 - Weyl's inequalities
 - ...too few!



New nonperturbative result: Newton's cradle spectra

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Simpler problem: addition of a projector

What happens to the eigenvalues when we add a rank-1 projector?

$$S(\mu) = S_0 + \mu |v\rangle \langle v|$$

If we know the answer, we can iteratively add a full interaction Hamiltonian:

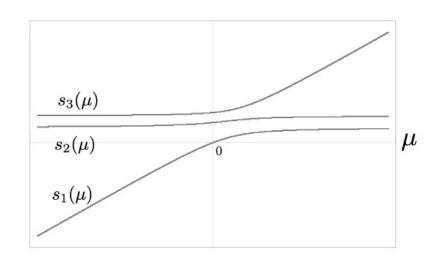
$$S(\mu) = S_0 + \sum_{j} \mu_j |v_j\rangle \langle v_j|$$

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Nonperturbative:

Newton cradle spectra

$$\frac{d\mu(s)}{ds} = \left(\sum_{m=1}^{N} \frac{|v_{m}|^{2}}{s - s_{m}}\right)^{-2} \sum_{r=1}^{N} \frac{|v_{r}|^{2}}{(s - s_{r})^{2}}$$
positive



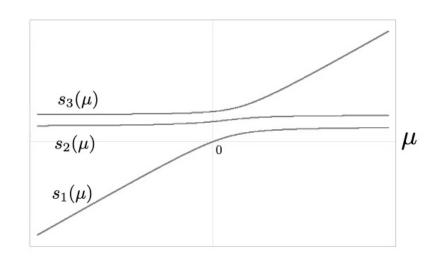
• Generically: $\langle s_i | v \rangle \neq 0, \forall i$ —> "Newton's cradle" behaviour

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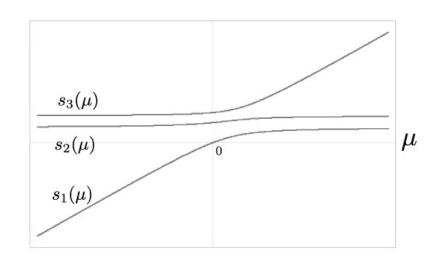
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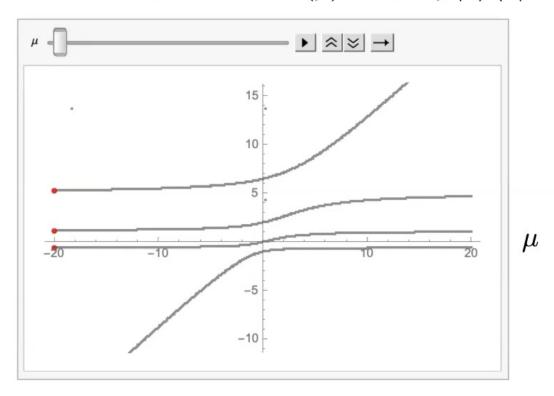


• Generically: $\langle s_i | v \rangle \neq 0, \forall i$ —> "Newton's cradle" behaviour

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Example with 4 eigenvalues.

Animation of eigenvalues of $S(\mu) = S_0 + \mu \ket{v} \bra{v}$



Newton cradle-like motion as a function of the coupling constant μ .

The spectra cover the entire real line exactly once:

$$\bigcup_{n=1,\dots,N}\bigcup_{\mu\in\mathbb{R}\cup\infty}\{s_n(\mu)\}=\mathbb{R}\cup\infty$$

• Therefore, for any $s \in \mathbb{R}$, there is a μ for which s is an eigenvalue of $S(\mu)$:

$$S(\mu)|s\rangle = s|s\rangle$$

• We can calculate the μ exactly using this formula:

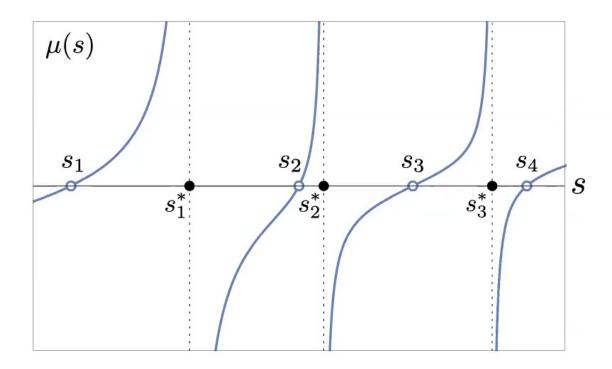
$$\mu(s) = \left(\sum_{i=1}^{N} \frac{|v_n|^2}{s - s_n}\right)^{-1}$$

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Coupling constant μ as a function of an eigenvalue s:

$$\mu(s) = \left(\sum_{i=1}^{N} \frac{|v_n|^2}{s - s_n}\right)^{-1}$$

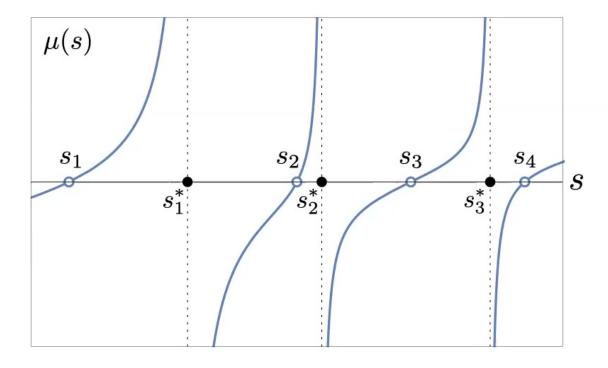
Plot of the dependence:



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$$\mu(s) = \left(\sum_{i=1}^{N} \frac{|v_n|^2}{s - s_n}\right)^{-1}$$

Plot of the dependence:



Special case: Cauchy's interlacing theorem

Theorem paraphrased:

Start with a NxN self-adjoint matrix M, which has eigenvalues $s_1, ..., s_N$.

E.g.
$$M = \begin{pmatrix} a_{11} & a_{12} & a_{13} & a_{14} \\ a_{21} & a_{22} & a_{23} & a_{24} \\ a_{31} & a_{32} & a_{33} & a_{34} \\ a_{41} & a_{42} & a_{43} & a_{44} \end{pmatrix} \text{ has eigenvalues } s_1, s_2, s_3, s_4.$$

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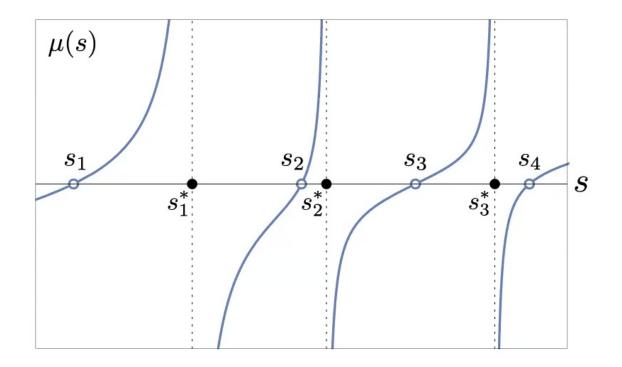
If we cross out a row and a column, the new eigenvalues interlace the old ones.

$$M' = \begin{pmatrix} a_{11} & a_{13} & a_{14} \\ a_{31} & a_{33} & a_{34} \\ a_{41} & a_{43} & a_{44} \end{pmatrix} \text{ has eigenvalues: } s_1', s_2', s_3'$$

$$s_1 \leq s_1' \leq s_2 \leq s_2' \leq s_3 \leq s_3' \leq s_4$$

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$$\mu(s) = \left(\sum_{i=1}^{N} \frac{|v_n|^2}{s - s_n}\right)^{-1} \xrightarrow{\mu \to \infty} \left[\sum_{m=1}^{N} \frac{|v_m|^2}{s_n^* - s_m} = 0 \right]$$



Evolution of eigenvectors

• Assume $|s\rangle$ is an eigenvector corresponding to the eigenvalue s:

$$S(\mu)|s\rangle = s|s\rangle$$

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Evolution of eigenvectors

• Assume $|s\rangle$ is an eigenvector corresponding to the eigenvalue s:

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• The new eigenvector $|s\rangle$ expressed in the original eigenbasis (of S_0):

$$|s\rangle = \frac{e^{i\varphi(s)}}{\mathcal{N}} \sum_{n=1}^{N} \frac{v_n}{|s - s_n|} |s_n\rangle$$

Normalization:
$$\mathcal{N} = \sqrt{\sum_{k} \frac{\left|v_{k}\right|^{2}}{\left(s - s_{k}\right)^{2}}}$$
 Phase arbitrary: $\varphi\left(s\right)$ Simple: $\left|\left\langle s \middle| s_{n} \right\rangle\right| \propto \frac{\left|v_{n}\right|}{\left|s - s_{n}\right|}$

Simple:
$$|\langle s|s_n \rangle| \propto \frac{|v_n|}{|s-s_n|}$$

Newton's cradle of unitary operators

- We start with a unitary matrix U_0 .
- We know its eigenvalues and vectors:

$$U_0 = \sum_i u_i \ket{u_i} \bra{u_i}$$

e.g. if it acts on a 3-dim Hilbert space H,

3 eigenvalues u_1, u_2, u_3 .

We get a behaviour analogous to Newton's cradle when we act with a U(1)

operator family:

$$U(lpha):=\left(\mathbb{1}+\left(e^{ilpha}-1
ight)\left|w
ight>\left< w
ight|\right. U_{0}$$
 , $lpha\in\left[0,2\pi
ight)$

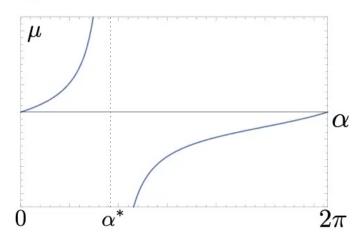
The connection between two Newton cradles

 The self-adjoint and the unitary Newton cradles are related by Cayley transforms:

$$U(\alpha) = (S(\mu) - i\mathbb{1})(S(\mu) + i\mathbb{1})^{-1}, S(\mu) = -i(U(\alpha) + \mathbb{1})(U(\alpha) - \mathbb{1})^{-1}$$

- Moebius transform for the eigenvalues: $u_j = \frac{s_j i}{s_j + i}$
- Relationship between μ and α :

$$\mu(\alpha) = \left(\sum_{m=1}^{N} \frac{|v_m|^2}{s_m^2 + 1} \cot\left(\frac{\alpha}{2}\right) - \sum_{k=1}^{N} \frac{|v_k|^2 s_k}{s_k^2 + 1}\right)^{-1}$$



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Addition of self-adjoint operators translates nicely to multiplication of unitaries,
 but not the exponentiated ones, instead the Cayley transformed unitaries.

$$H \rightarrow e^{iHt}$$

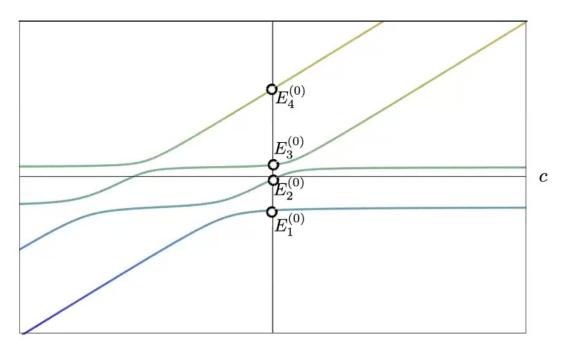
$$H \to (H - i\mathbb{1})(H + i\mathbb{1})^{-1}$$



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Special case: Level repulsion

Eigenvalues of A + cB, where c is the coupling strength, as a function of c:



Plot of eigenvalues of A+cB as a function of c .

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Addition of self-adjoint operators translates nicely to multiplication of unitaries,
 but not the exponentiated ones, instead the Cayley transformed unitaries.

$$H \rightarrow e^{iHt}$$

$$H \to (H - i1)(H + i1)^{-1}$$



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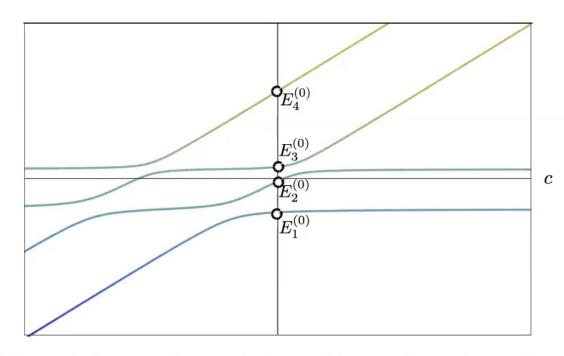
Special case: Level repulsion

Eigenvalues of A + cB, where c is the coupling strength, as a function of c:

Decompose B:

$$B = \sum_{i} b_{i} |v_{i}\rangle \langle v_{i}|$$

Add projectors $\left.b_{i}\left|v_{i}\right\rangle\left\langle v_{i}\right|\right.$ one by one.



Plot of eigenvalues of A+cB as a function of c.

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Special case: Level repulsion

Eigenvalues of A + cB, where c is the coupling strength, as a function of c:

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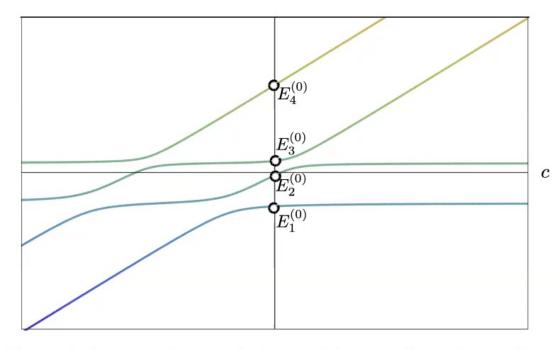
$$B = \sum_{i} b_{i} |v_{i}\rangle \langle v_{i}|$$

Add projectors $\left.b_{i}\left|v_{i}\right\rangle\left\langle v_{i}\right|\right.$ one by one.

As long as the overlap:

$$\langle a_i(b)|v_i\rangle \neq 0$$

-> no level crossing!



Plot of eigenvalues of A+cB as a function of c.

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Summary so far

- Newton's cradle spectra shows how eigenvalues and eigenvectors change when we add a 1-dimensional projector to a Hamiltonian.
 - Eigenvalues move like Newton's cradle.
 - Eigenvector of the highest eigenvalue rotates into the projector subspace as the coupling constant $\mu \to \infty$.
 - We can decompose any Hamiltonian into a sum of 1-dim. projectors ->
 -> new strategy for understanding addition of Hamiltonians.
 - New understanding of Cauchy interlacing, level repulsion...

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Stimulated Unruh Effect and Acceleration-Induced Transparency

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Acceleration effects in quantum physics

- Acceleration doesn't fit naturally into quantum physics.
- An example where it does: Unruh effect (extremely small).
- Is there more?

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Acceleration effects in quantum physics

- Acceleration doesn't fit naturally into quantum physics.
- An example where it does: Unruh effect (extremely small).
- Is there more?

Why?

Acceleration related to gravity, through the equivalence principle.

Here:

- New acceleration-induced phenomena,
- Increased measurability of the Unruh effect,
- Insights into acceleration (gravity).

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Light-matter interactions

Resonant	Non-resonant
The usual	The unusual
 Absorption, emission 	Unruh effect, Hawking effect
Strong	• Weak
Can be stimulated (e.g. laser)	Activated by acceleration or gravity

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Light-matter interactions

	Resonant		Non-resonant
•	The usual	•	The unusual
•	Absorption, emission	•	Unruh effect, Hawking effect
•	Strong	•	Weak
•	Can be stimulated (e.g. laser)	•	Activated by acceleration or gravity
New:	Can be modulated with acceleration/gravity!	New:	Can be stimulated!
	"Acceleration-induced transparency"		→ Non-resonant effects become more measurable!

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The Unruh effect can be stimulated

- First we stimulate the Unruh effect by the presence of photons in the quantum field
- Problem: this also stimulates the resonant terms

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The Unruh effect can be stimulated

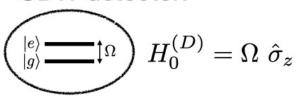
- First we stimulate the Unruh effect by the presence of photons in the quantum field
- Problem: this also stimulates the resonant terms
- Solution: acceleration-induced transparency
 - suppresses the resonant terms to zero
- Outcome: we are left with strong non-resonant terms, in particular:
 Strong Unruh effect

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The standard Unruh effect

· UDW detector:



Scalar quantum field:

$$H_0^{(D)} = \Omega \; \hat{\sigma}_z \; H_0^{(F)} = \int d^3k \, \omega_{f k} \, \hat{a}_{f k}^\dagger \hat{a}_{f k} \; .$$

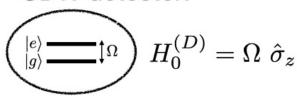
• Interaction Hamiltonian: $H_{int} = G \, \hat{\sigma}_x(\tau) \otimes \hat{\phi}(t(\tau), \mathbf{x}(\tau))$

$$H_{int} \propto \int d\mathbf{k} \left(\underline{\sigma^- a^\dagger_\mathbf{k} + \sigma^+ a_\mathbf{k}} + \underline{\sigma^- a_\mathbf{k} + \sigma^+ a^\dagger_\mathbf{k}} \right)$$
ron-resonant

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The standard Unruh effect

UDW detector:



Scalar quantum field:

$$\mathcal{J}_{\Omega} = 0 \hat{\sigma}_z \qquad H_0^{(F)} = \int d^3k \, \omega_{\mathbf{k}} \, \hat{a}_{\mathbf{k}}^{\dagger} \hat{a}_{\mathbf{k}} \, d_{\mathbf{k}} \, \hat{a}_{\mathbf{k}}^{\dagger} \hat{a}_{\mathbf{k}}$$

• Interaction Hamiltonian: $H_{int} = G \, \hat{\sigma}_x(\tau) \otimes \phi(t(\tau), \mathbf{x}(\tau))$

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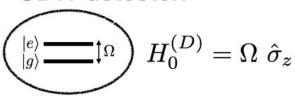
• Amplitude for $|g\rangle\otimes|0\rangle \rightarrow |e\rangle\otimes|1\rangle$: $\mathcal{A}(\mathbf{k})=\langle e|\,\langle 1_{\mathbf{k}}|\,i\,\int d\tau H_{int}(\tau)\,|g\rangle\,|0\rangle$

$$\Rightarrow$$
 $A(\mathbf{k}) = \frac{i G}{(2\pi)^{3/2} \sqrt{\omega_{\mathbf{k}}}} I_{+}(\mathbf{k}),$

$$\Rightarrow \mathcal{A}(\mathbf{k}) = \frac{i G}{(2\pi)^{3/2} \sqrt{\omega_{\mathbf{k}}}} I_{+}(\mathbf{k}), \quad \text{where} \quad I_{\mp}(\Omega, k) = \int d\tau \, e^{i\Omega\tau \mp ik^{\mu}x_{\mu}(\tau)}.$$

The stimulated Unruh effect (Fock state field)

UDW detector:



Scalar quantum field:

- Interaction Hamiltonian: $H_{int} = G \, \hat{\sigma}_x(\tau) \otimes \hat{\phi}(t(\tau), \mathbf{x}(\tau))$
- The transition: $|g\rangle\otimes|n_{\mathbf{k}}\rangle\rightarrow|e\rangle\otimes|\phi'\rangle$
- The transition: $|g\rangle\otimes|n_{\mathbf{k}}\rangle$. The transition: $|g,n_{\mathbf{k}}\rangle\to G\Big[\sqrt{n+1}I_+(k)\,|e,(n+1)_{\mathbf{k}}\rangle\Big]+\sqrt{n}I_-(k)\,|e,(n-1)_{\mathbf{k}}\rangle\Big]$

Standard vs. stimulated Unruh effect

	Standard	Stimulated (Fock state)
Transition:	$ g angle\otimes 0 angle ightarrow e angle\otimes 1 angle$	$ g\rangle\otimes n_{\mathbf{k}}\rangle\rightarrow e\rangle\otimes (n+1)_{\mathbf{k}}\rangle$
Probability:	$p(\mathbf{k}) = \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} I_+(\mathbf{k}) ^2$	$p(\mathbf{k}) = \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} \left(n I_{-}(\mathbf{k}) ^2 + (n+1) I_{+}(\mathbf{k}) ^2 \right)$

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The resonant terms can be suppressed

Standard: $p(\mathbf{k}) \propto \left|I_{+}(\mathbf{k})\right|^{2}$

Resonant term contribution

Stimulated (Fock state):

$$p(\mathbf{k}) \propto (n+1)|I_{+}(\mathbf{k})|^{2} + n|I_{-}(\mathbf{k})|^{2}$$

· Trajectory-dependent time integrals:

$$I_{\mp}(\Omega, k) = \int d\tau \, e^{i\Omega\tau \mp ik^{\mu}x_{\mu}(\tau)}$$

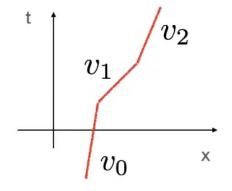
• We know how to eliminate $I_+(\mathbf{k})$: in inertial motion $I_+(\Omega,k)=0$,

$$I_{-}(\Omega, k) \neq 0.$$

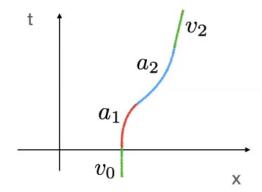
• Can we get rid of $I_{-}(\mathbf{k})$ somehow?

Acceleration-induced transparency

- Q: Can we tune the trajectory so that for some Ω the time integral $I_-(\Omega,k)=\int \mathrm{d}\tau\,e^{i\Omega\tau-ik^\mu x_\mu(\tau)}=0$?
- Yes!

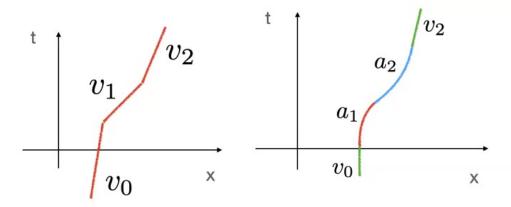


Proved also true for a smoother trajectory:



Acceleration-induced transparency

$$I_{-}(\Omega, k) = \int d\tau \, e^{i\Omega\tau - ik^{\mu}x_{\mu}(\tau)} = 0$$

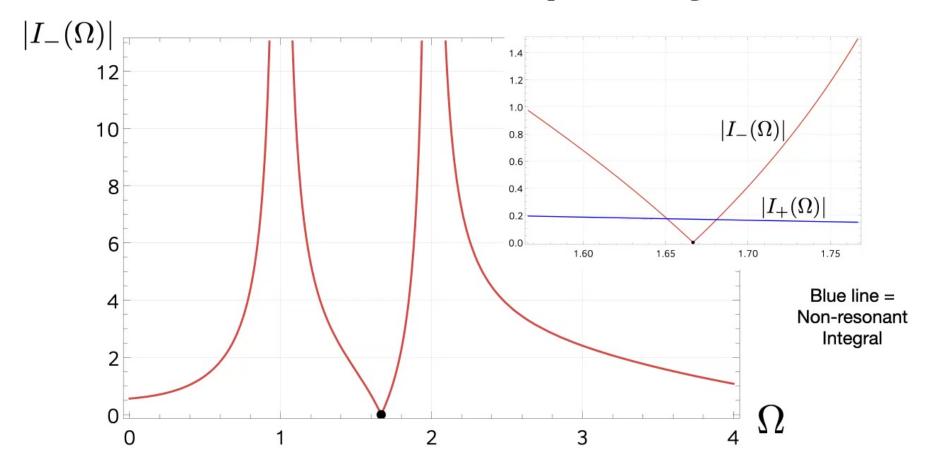


- The detector with gap Ω will not get excited due to resonant terms.
- At the same time: $|I_+(\Omega, \mathbf{k})| > 0$.
- If the detector gets excited, it is due to the non-resonant terms.

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Acceleration-induced transparency



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Summary

[1] Also the non-resonant transitions can be stimulated!

Stimulated transitions are n times stronger :

$$(\sigma^+ a)$$

$$\sigma^+ a^\dagger$$

nulated transitions are
$$n$$
 times stronger:
$$\sigma^+ a$$

$$p(\mathbf{k}) = \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} |I_+(\mathbf{k})|^2 \quad \rightarrow \quad p(\mathbf{k}) = \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} \left(n |I_-(\mathbf{k})|^2 + (n+1) |I_+(\mathbf{k})|^2 \right)$$

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Summary

[1] Also the non-resonant transitions can be stimulated!

Stimulated transitions are n times stronger :



$$p(\mathbf{k}) = \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} |I_+(\mathbf{k})|^2 \longrightarrow p(\mathbf{k}) = (n+1) \frac{G^2}{(2\pi)^3 \omega_{\mathbf{k}}} |I_+(\mathbf{k})|^2$$

[2] Acceleration-induced transparency can suppress the resonant terms!

Strong non-resonant terms dominate.



Strong forward (or time-reversed) Unruh effect.

A new method to measure the Unruh effect in lab.

- New effects:
 - [1] Stimulated forward and time-reversed Unruh effects,
 - [2] Acceleration-induced transparency,
- Combining them leads to improved measurability of the Unruh effect.

Stimulated Unruh effect = $n \times S$ tandard Unruh effect

• E.g. for a 100 mW laser pointer: $n \approx 10^{16}$

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The new effects suggest for gravity:

- Resonant terms are potentially strongly affected by gravity, including the possibility of gravity-induced translucency or even transparency.
- Non-resonant terms, e.g. of Hawking and other horizon radiation, can be stimulated (e.g. by accretion disk luminosity?).

Paper: Acceleration-Induced Effects in Stimulated Light-Matter Interactions, B. Šoda, V. Sudhir, and A. Kempf, Phys. Rev. Lett. 128, 163603

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- Acceleration-induced transparency -> parameters controlled by choice of trajectory.
- Idea: use parametric control of interaction Hamiltonian -> turn off transitions.

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- Idea: use parametric control of interaction Hamiltonian -> turn off transitions.

Setup: Total Hamiltonian: $H = H_0 + H_{int}(t)$

 $H_{\mathrm{int}}=\mu\left(t
ight)\left|v\left(t
ight)
ight
angle\left\langle v\left(t
ight)
ight|.$ (simple, gives results)

Interaction picture: $H_{\text{int}}(t) = e^{iH_0t}H_{\text{int}}e^{-iH_0t}$

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$$H_0 = \sum_{j=1}^{N} s_j |s_j\rangle \langle s_j| \qquad \longrightarrow \qquad H_{\text{int}}(t) = \mu(t) \sum_{j,k} e^{i(s_j - s_k)t} \langle s_j | v(t) \rangle \langle v(t) | s_k \rangle |s_j\rangle \langle s_k|.$$

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Time evolution:
$$U_{I}\left(t\right)=e^{i\int_{t_{0}}^{t}d au H_{\mathrm{int}}\left(au\right)}pprox1+i\int_{t_{0}}^{t}d au H_{\mathrm{int}}\left(au
ight)+...,$$
 (Weak coupling approx.)

Probability of excitation: $p_{\text{exc.}} = \left| \langle e | U_I(t) | g \rangle \right|^2$

$$p_{\mathrm{exc.}} = \left| \int_{t_0}^t d au \mu\left(au
ight) e^{i\Omega au} v_1\left(au
ight) v_0^*\left(au
ight)
ight|^2$$
 (gap = Ω)

Problem: find
$$f(t)$$
 such that $p_{\mathrm{exc.}} = \left| \int_{t_0}^t d\tau e^{i\Omega \tau} f\left(au
ight) \right|^2 = 0.$

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We already know a solution f(t) - from acceleration-induced transparency.

$$f(\tau) = e^{-ik^{\mu}x_{\mu}(\tau)}$$

There,
$$p_{
m exc.} \propto \left| \int d au e^{i\Omega au - ik^\mu x_\mu(au)}
ight|^2 = 0$$
 , for a suitable choice of $x_\mu(au)$.

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Example. Varying the projector in time. $v_i(t) = e^{if_i(au)}v_i^{(0)}$

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Example. Varying the coupling in time: $\mu(t)$

Projector(s) constant in time: $|v_1\rangle \langle v_1|, |v_2\rangle \langle v_2|$

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Summary

Showed two new tools for interactions:

- 1) Newton's cradle spectra
- Parametrically induced decoupling (generalization of acceleration-induced transparency in light-matter int.)

Their purpose:

- 1) New insights into non-perturbative interactions. (Newton's cradle spectra)
- A way to turn off (some) transitions using parametric control over the interaction Hamiltonian.

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