

Title: Counting the microstates of the cosmic horizon

Speakers: Vasudev Shyam

Series: Quantum Gravity

Date: November 10, 2022 - 2:30 PM

URL: <https://pirsa.org/22110078>

Abstract: I will describe a holographic model for the three dimensional de Sitter static patch where the boundary theory is the so called $T\bar{T} + \Lambda_2$ deformation of the conformal field theory dual to AdS₃ quantum gravity. This identification allows us to obtain the cosmic horizon entropy from a microstate count, and the microstates themselves are a dressed version of those that account for the entropy of certain black holes in AdS space. I will also show how the effect of this dressing at the cosmic horizon is to replace the spacetime dependence of the fields of the undeformed holographic CFT with dependence on the indices of large matrices.

Zoom link: <https://pitp.zoom.us/j/95396921570?pwd=NGFoOGlGY1ZDU2pnNFRwWit3b2w0Zz09>

Counting the micro-states of the cosmic horizon

Vasu Shyam



The
Branco Weiss
Fellowship
Society in Science

Based on:

arXiv > hep-th > arXiv:2106.10227

$T\bar{T} + \Lambda_2$ Deformed CFT on the Stretched dS_3 Horizon

Vasudev Shyam

arXiv > hep-th > arXiv:2110.14670

de Sitter Microstates from $T\bar{T} + \Lambda_2$ and the Hawking–Page Transition

Evan Coleman, Edward A. Mazenc, Vasudev Shyam, Eva Silverstein, Ronak M Soni, Gonzalo Torroba, Sungyeon Yang

arXiv > hep-th > arXiv:2209.11753

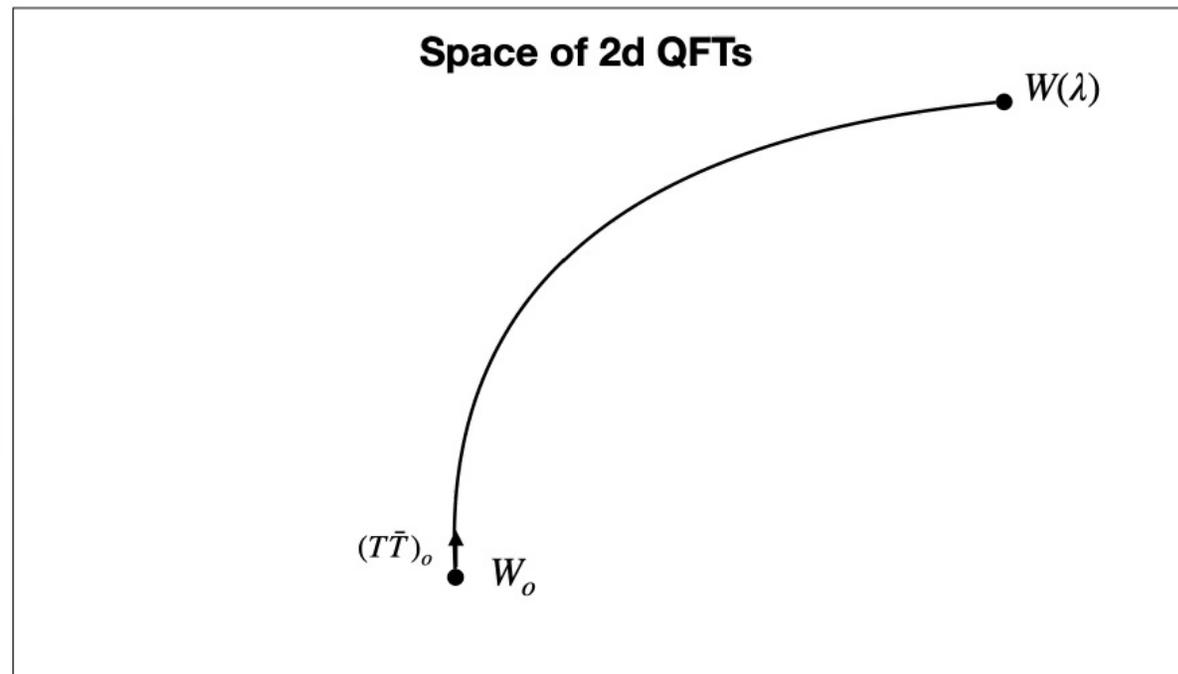
$T\bar{T}$ deformed scattering happens within matrices

Vasudev Shyam, Yigit Yargic

Deforming QFTs in $d = 2$ by $T\bar{T}$

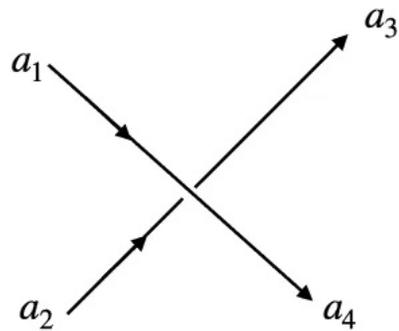
Smirnov & Zamolodchikov, 16'
Cavaglia et. al. 16'

$$\partial_\lambda W(\lambda) = \int_x \langle T\bar{T} \rangle \equiv \int d^2x \det f \left\langle \left(T^{\mu\nu} T_{\mu\nu} - (T_\alpha^\alpha)^2 \right) \right\rangle, \quad W(\lambda = 0) = W_o$$

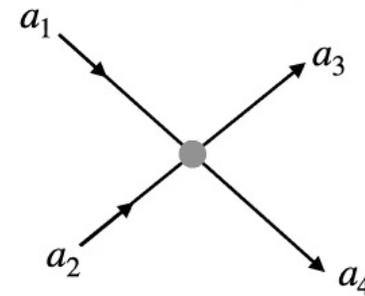


What $T\bar{T}$ does

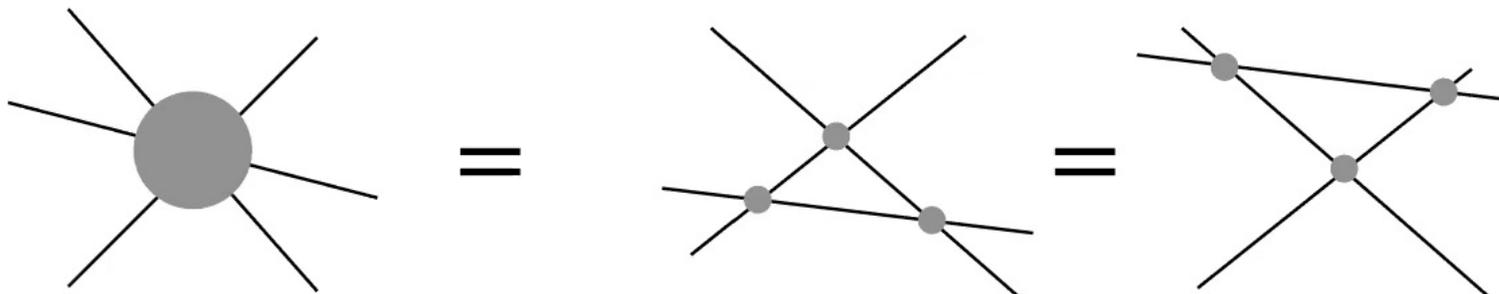
$$S(\{p_{a_1}, p_{a_2}\} \rightarrow \{p_{a_3}, p_{a_4}\}) = 1$$



$$S(\{p_{a_1}, p_{a_2}\} \rightarrow \{p_{a_1}, -p_{a_2}\}) = \exp(\lambda(p_{a_1} \wedge p_{a_2}))$$



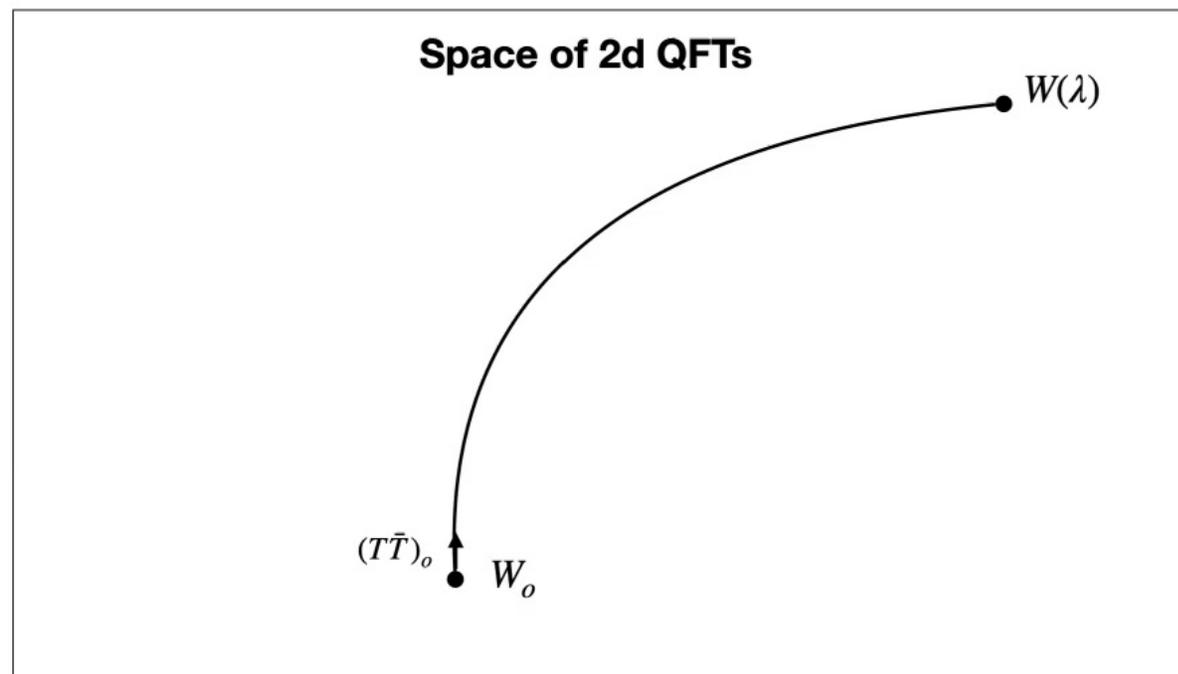
Integrability:



Deforming QFTs in $d = 2$ by $T\bar{T}$

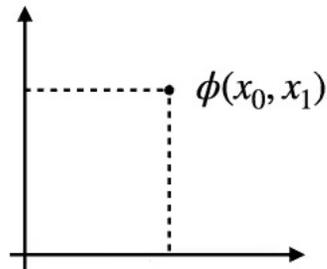
Smirnov & Zamolodchikov, 16'
Cavaglia et. al. 16'

$$\partial_\lambda W(\lambda) = \int_x \langle T\bar{T} \rangle \equiv \int d^2x \det f \left\langle \left(T^{\mu\nu} T_{\mu\nu} - (T_\alpha^\alpha)^2 \right) \right\rangle, \quad W(\lambda = 0) = W_o$$



$T\bar{T}$ turns fields into matrices

Yigit Yagrig, V.S.



$$\begin{pmatrix} \phi_{11} & \cdots & \phi_{1N} \\ \vdots & & \vdots \\ \phi_{NN} \end{pmatrix}$$

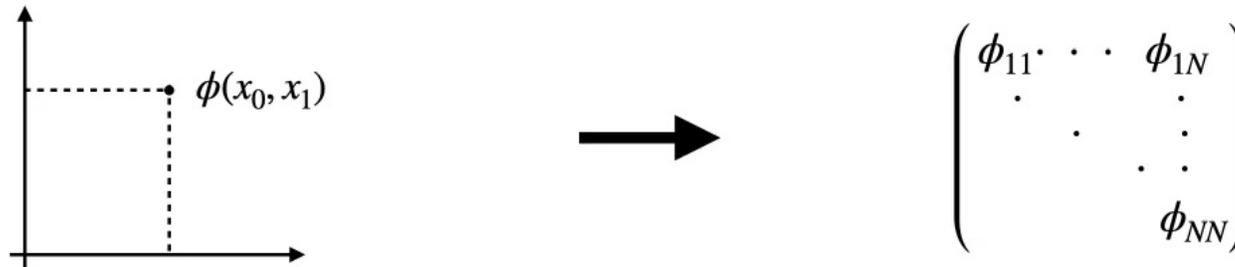
$$\phi_{ij} = \phi_{\lambda^{-1}\left(\frac{i-j}{2}, \frac{i+j}{2}\right)} \equiv \phi\left(p_0 = \frac{1}{\lambda} \left(\frac{i-j}{2}\right), x_1 = \frac{i+j}{2}\right) \quad \phi(p_0, p_1) = \int dx_1 e^{ip_1 x_1} \phi(p_0, x_1)$$

$$\phi \cdot \psi \rightarrow \phi \star \psi$$

$$\phi(p_0^a, p_1^b) \star \phi(p_0^b, p_1^b) = e^{i\lambda p^a \wedge p^b} \phi(p_0^a, p_1^b) \cdot \phi(p_0^b, p_1^b)$$

$T\bar{T}$ turns fields into matrices

Yigit Yagrig, V.S.



$$\phi_{ij} = \phi_{\lambda^{-1}\left(\frac{i-j}{2}, \frac{i+j}{2}\right)} \equiv \phi\left(p_0 = \frac{1}{\lambda} \left(\frac{i-j}{2}\right), x_1 = \frac{i+j}{2}\right) \quad \phi(p_0, p_1) = \int dx_1 e^{ip_1 x_1} \phi(p_0, x_1)$$

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Dressing the number 1

$$S_n(p_{a_1}, \dots, p_{a_n}) = \prod_i (p^2 + m^2)_i \langle \Omega | \mathcal{T} \left(\varphi(p_{a_1}) \cdots \varphi(p_{a_n}) \right) | \Omega \rangle \Big|_{p_i^2 + m_i^2 = 0, \forall i} = 1$$

$T\bar{T}$

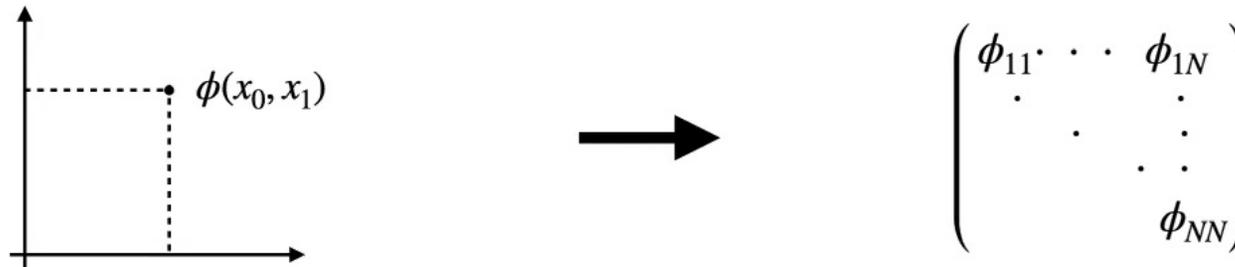
$$\prod_i (p^2 + m^2)_i \langle \Omega | \mathcal{T} \left(\varphi(p_{a_1}) \star \cdots \star \varphi(p_{a_n}) \right) | \Omega \rangle \Big|_{p_i^2 + m_i^2 = 0, \forall i} =$$

$$= \exp \left(-i\lambda \sum_{i < j=1}^l (p_{a_i} \wedge p_{a_j}) \right) \times \exp \left(i\lambda \sum_{k < r=l}^n (p_{a_k} \wedge p_{a_r}) \right) \times 1$$

[Dubovsky, Gorbenko, Mirbabayi 18']

$T\bar{T}$ turns fields into matrices

Yigit Yagrig, V.S.



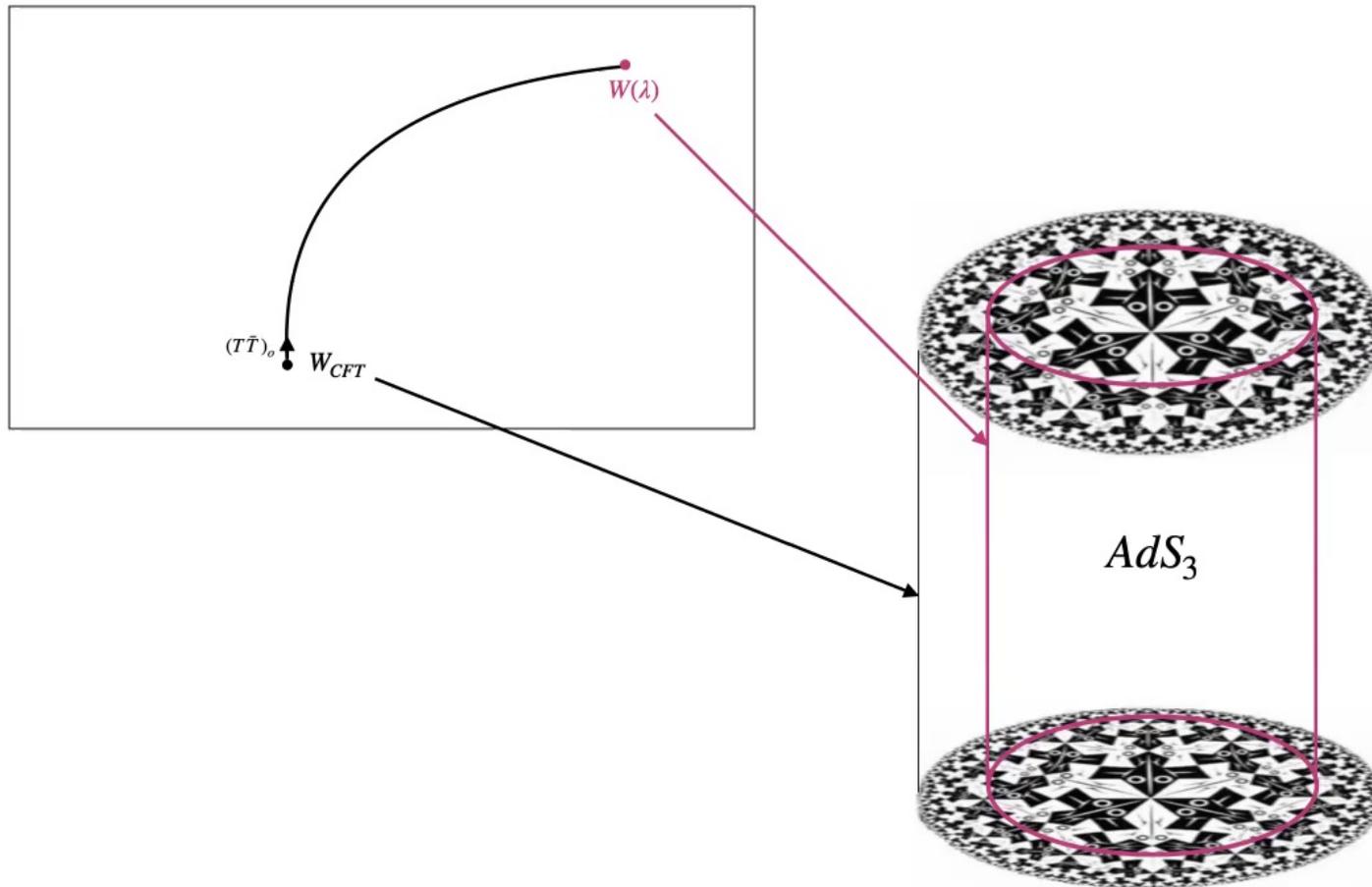
$$\phi_{ij} = \phi_{\lambda^{-1}\left(\frac{i-j}{2}, \frac{i+j}{2}\right)} \equiv \phi\left(p_0 = \frac{1}{\lambda} \left(\frac{i-j}{2}\right), x_1 = \frac{i+j}{2}\right) \quad \phi(p_0, p_1) = \int dx_1 e^{ip_1 x_1} \phi(p_0, x_1)$$

$$\phi \cdot \psi \rightarrow \phi \star \psi$$

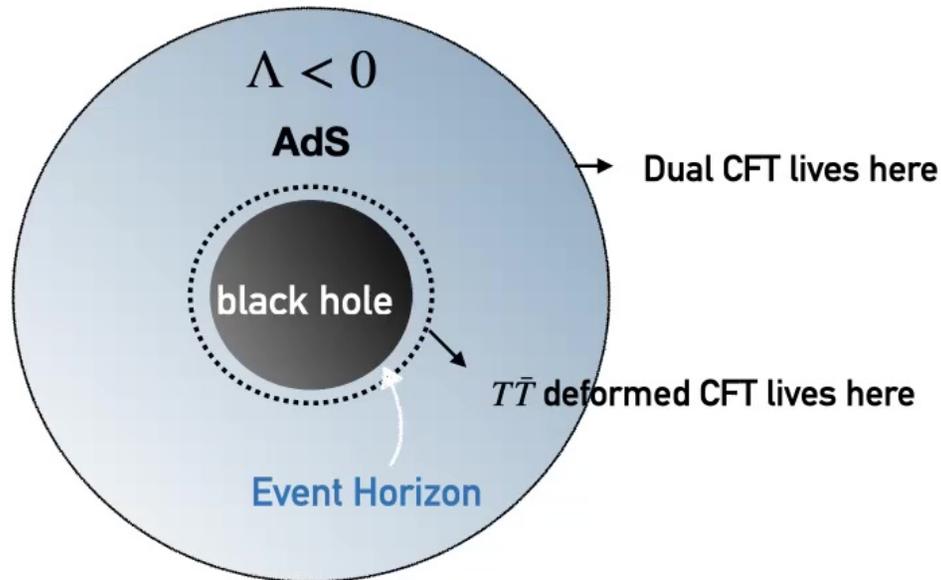
$$\phi(p_0^a, p_1^b) \star \phi(p_0^b, p_1^b) = e^{i\lambda p^a \wedge p^b} \phi(p_0^a, p_1^b) \cdot \phi(p_0^b, p_1^b)$$

$T\bar{T}$ holography: moving “into” the bulk

McGough, Mezei, Verlinde 16'

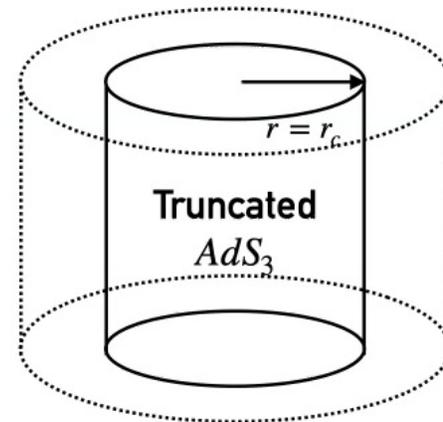


Quasi local energy of cutoff AdS_3



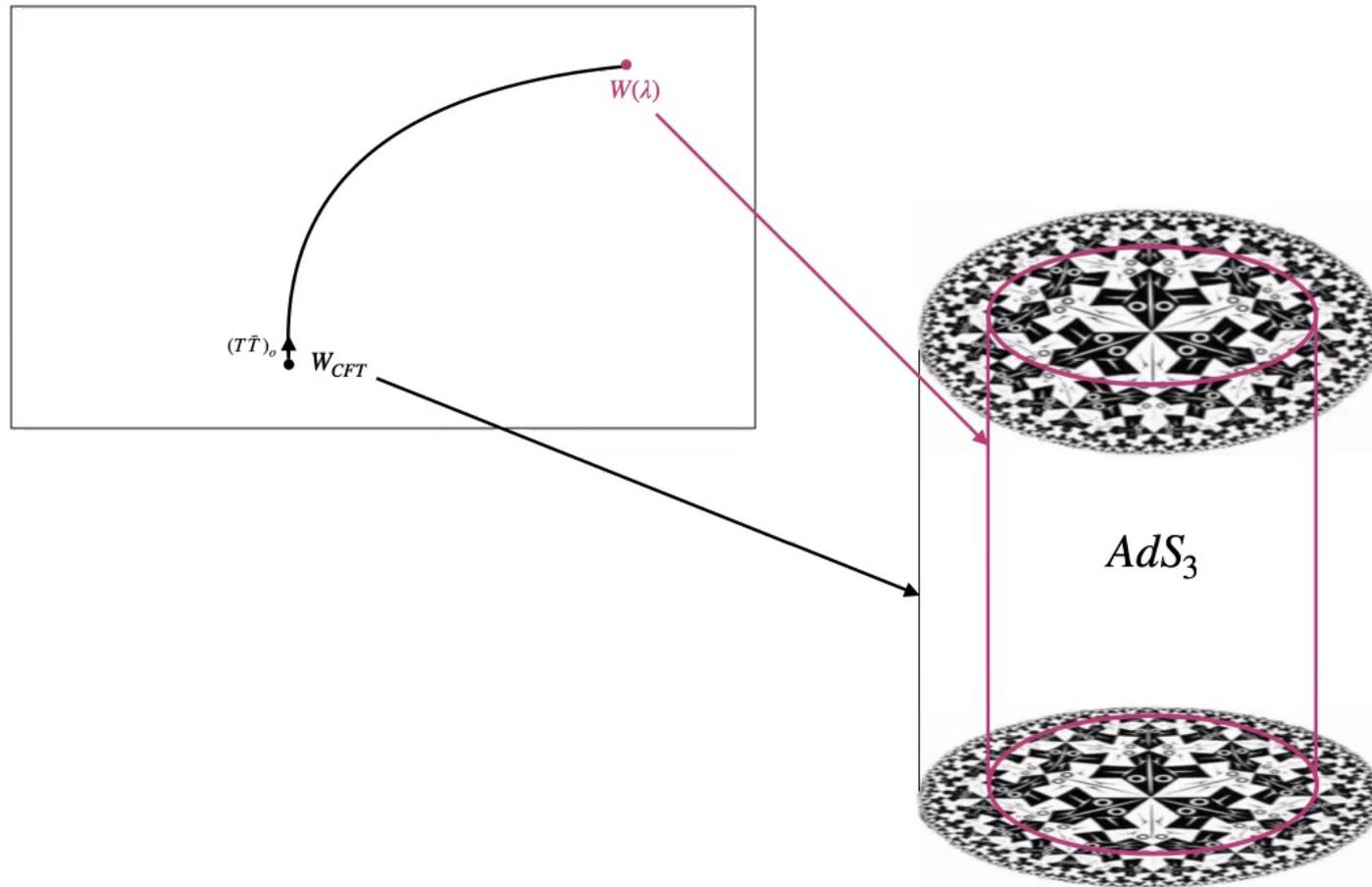
Quasi Local energy:

$$E = \frac{r_c}{4G} \left[1 - \sqrt{1 - \frac{8GM}{r_c^2} + \frac{16G^2 J^2}{r_c^4}} \right]$$

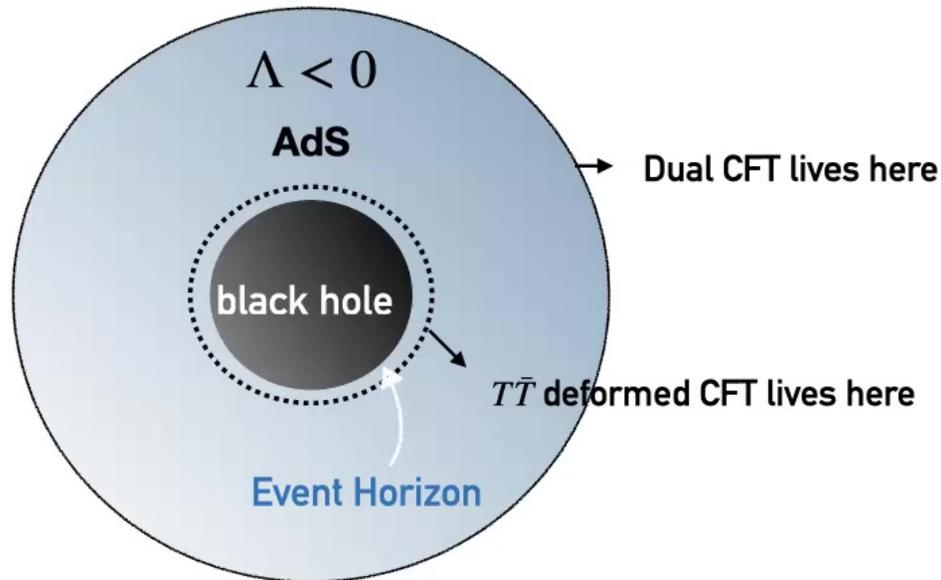


$T\bar{T}$ holography: moving “into” the bulk

McGough, Mezei, Verlinde 16'

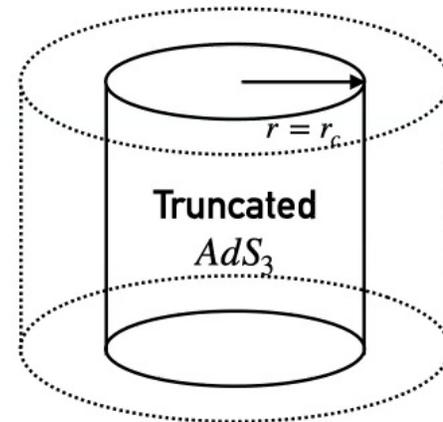


Quasi local energy of cutoff AdS_3

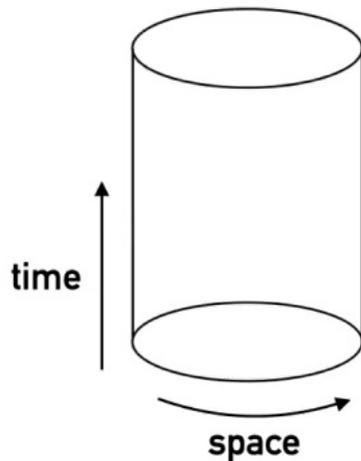


Quasi Local energy:

$$E = \frac{r_c}{4G} \left[1 - \sqrt{1 - \frac{8GM}{r_c^2} + \frac{16G^2 J^2}{r_c^4}} \right]$$



Energy levels flow under $T\bar{T}$



$$\partial_{\mu} E_n(\mu, L) = L \langle n | T\bar{T} | n \rangle = -\frac{1}{4} \left(E_n \partial_L E_n + \frac{P_n^2}{L^2} \right)$$

$$E_n(\mu, L) = \frac{2L}{\mu} \left(1 - \sqrt{1 \pm 2\pi\mu \left(\frac{\Delta_n + \bar{\Delta}_n - \frac{c}{12}}{L^2} \right) + \pi^2 \mu^2 \left(\frac{(\Delta_n - \bar{\Delta}_n)^2}{L^4} \right)} \right)$$

$$E_n^o = \frac{1}{L} \left(\Delta_n + \bar{\Delta}_n - \frac{c}{12} \right), \quad j_n = \frac{1}{L} (\Delta_n - \bar{\Delta}_n)$$

Matching

[McGough, Mezei, Verlinde 16']

$$E_n(\mu, L) = \frac{2L}{\mu} \left(1 - \sqrt{1 \pm 2\pi\mu \left(\frac{\Delta_n + \bar{\Delta}_n - \frac{c}{12}}{L^2} \right) + \pi^2\mu^2 \left(\frac{(\Delta_n - \bar{\Delta}_n)^2}{L^4} \right)} \right)$$

$$\Delta_n + \bar{\Delta}_n - \frac{c}{12} = M, \quad \Delta_n - \bar{\Delta}_n = J$$

$$L = \frac{r_c}{4G}$$

$$E = \frac{r_c}{4G} \left[1 - \sqrt{1 - \frac{8GM}{r_c^2} + \frac{16G^2J^2}{r_c^4}} \right]$$

Honorary mention: Laurent's formula

Freidel 08'

$$\Psi[f] = \int \mathcal{D}e e^{-\frac{2}{\lambda} \int d^2x \det(e-f)} Z_{CFT}[e]$$

$$f_\mu^a \frac{\delta \Psi[f]}{\delta f_\mu^a} = \lambda \epsilon^{ab} \epsilon_{\mu\nu} \frac{\delta^2 \Psi[f]}{\delta f_\mu^a \delta f_\nu^b} + R[f] \det f \Psi[f]$$

$$\equiv \langle T\bar{T} \rangle \Psi[f]$$

Deformation

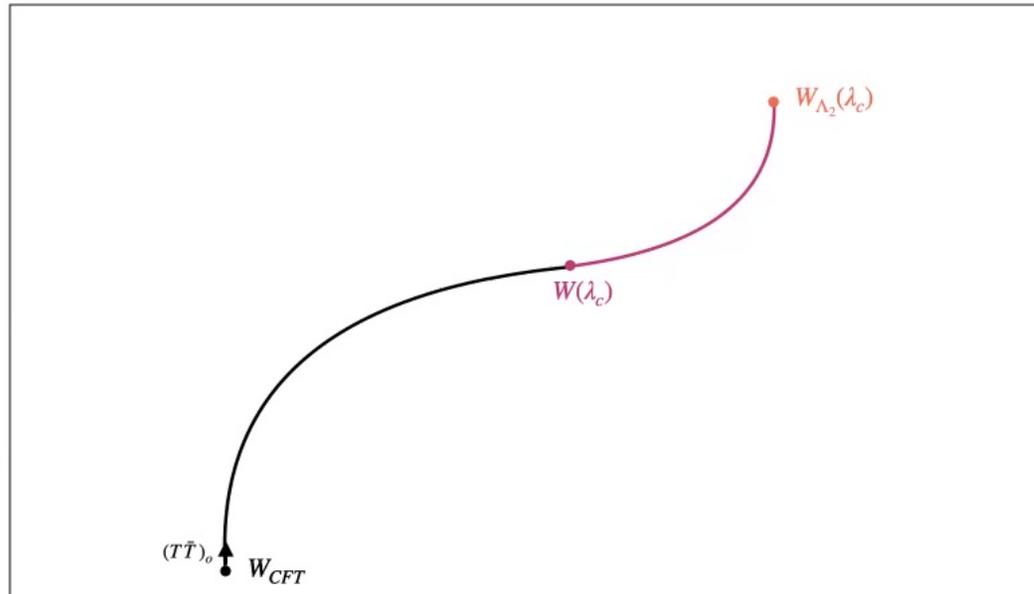


$$e_\mu^a \frac{\delta}{\delta e_\mu^a} Z_{CFT}[e] = \frac{c}{24\pi} \det e R(e) Z_{CFT}[e]$$

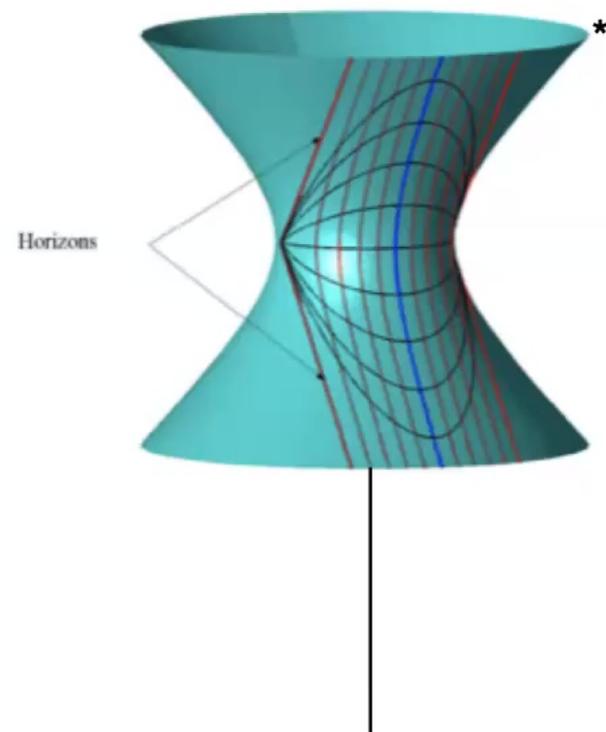
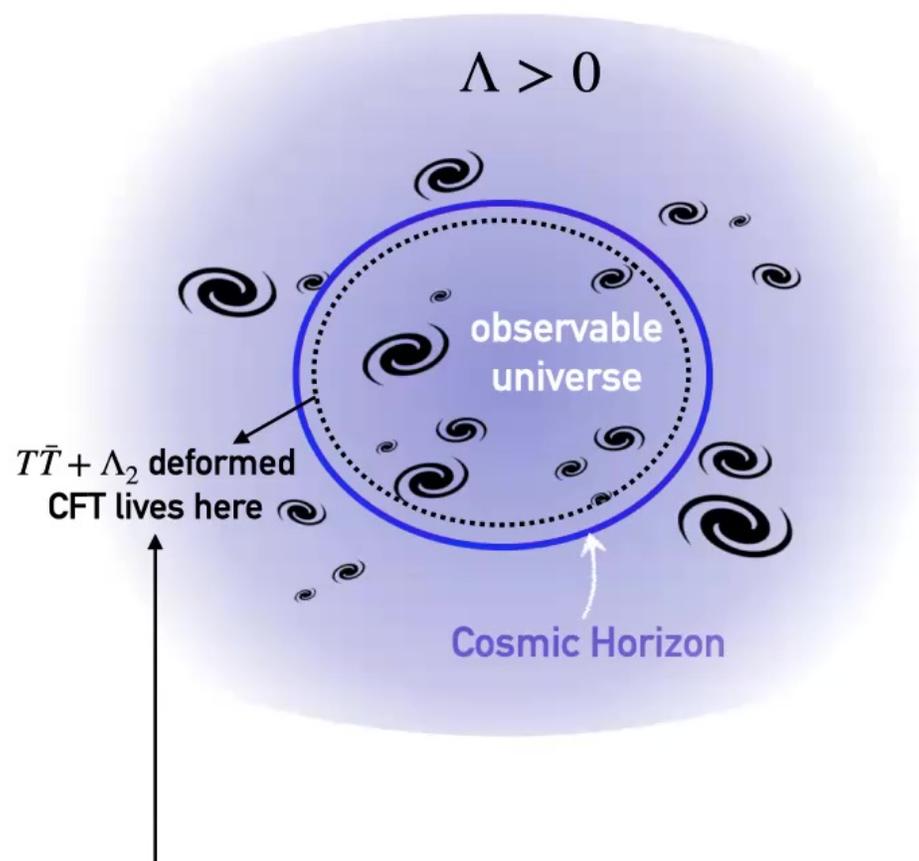
The $T\bar{T} + \Lambda_2$ deformation

Gorbenko, Silverstein, Torroba 18'

$$\partial_\lambda W_{\Lambda_2}(\lambda) = \int d^2x \det f \left\langle \left(T^{\mu\nu} T_{\mu\nu} - (T_\alpha^\alpha)^2 \right) \right\rangle - \frac{1-\eta}{4\pi\lambda^2} \int d^2x \det f, \quad W_{\Lambda_2}(\lambda = \lambda_c) = W(\lambda_c)$$

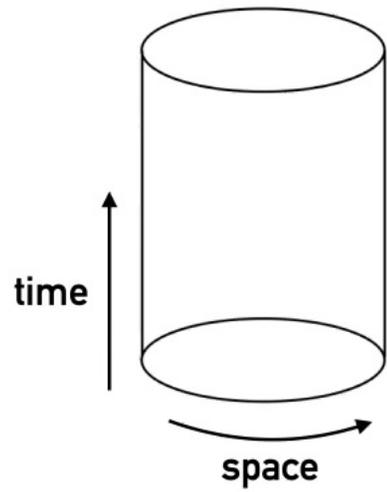


Cutoff Holography in dS_3

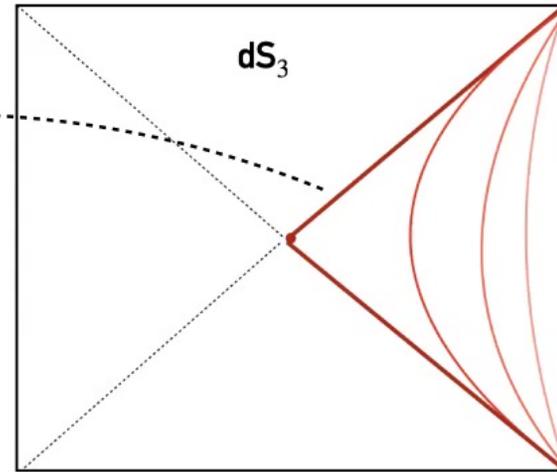


*from U.Moschella,
Prog.Math.Phys.
47 (2006) 120-133,
[10.1007/3-7643-7436-5_4](https://doi.org/10.1007/3-7643-7436-5_4)

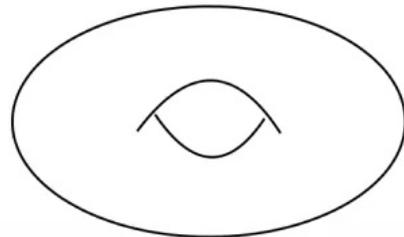
Boundary



Bulk

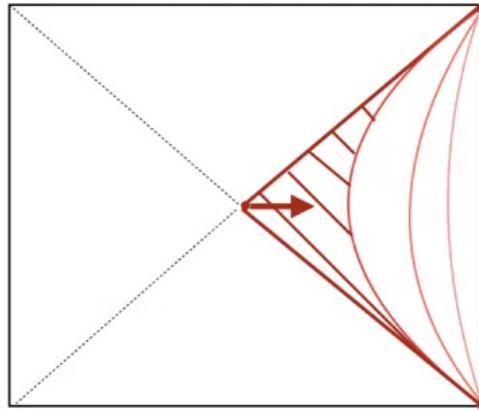


In Euclidean signature

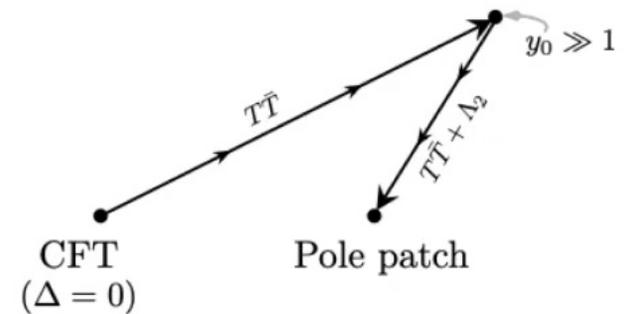
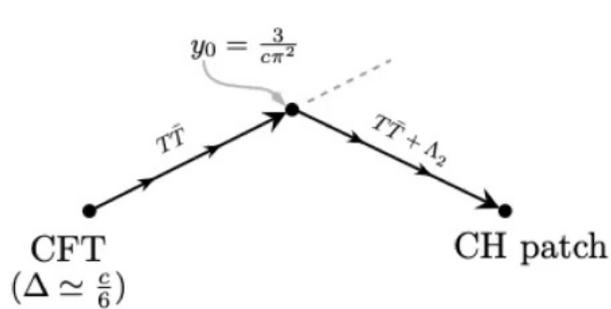
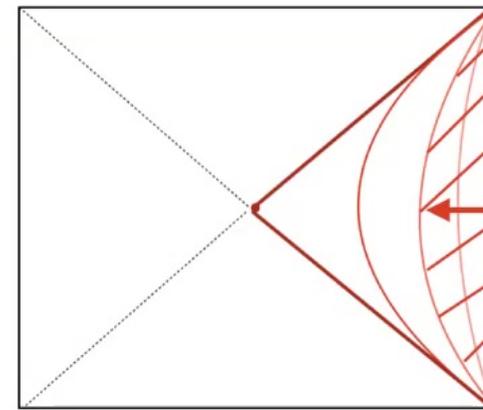


$\eta = -1$: Patches of dS

Cosmic Horizon Patch



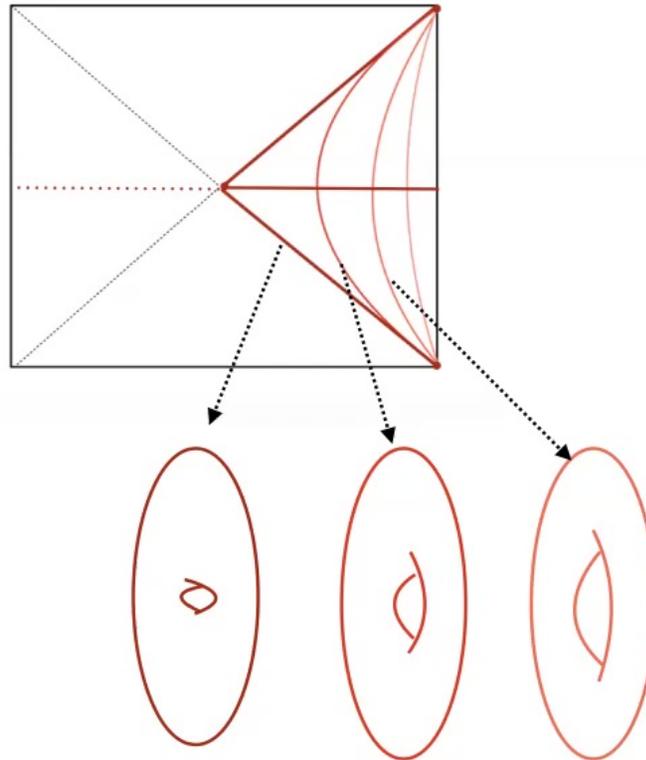
Pole Patch



$$\mathcal{E} = \frac{1}{\pi y} \left(1 + \sqrt{\eta + \dots} \right) \quad \leftarrow \text{related by } \pm\sqrt{} \quad \rightarrow \quad \mathcal{E} = \frac{1}{\pi y} \left(1 - \sqrt{\eta + \dots} \right)$$

Euclidean dS_3 Horizon

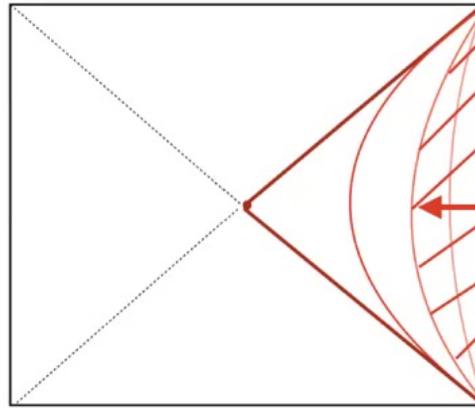
$$ds^2 = \left(1 - \frac{R^2}{\ell^2}\right) dt_E^2 + \frac{1}{\left(1 - \frac{R^2}{\ell^2}\right)} dR^2 + R^2 d\phi^2$$



∴ Approaching the de Sitter horizon
< = >
Taking the $\beta \rightarrow 0$ limit in the dual theory

[Varadarajan 02']

Pole Patch

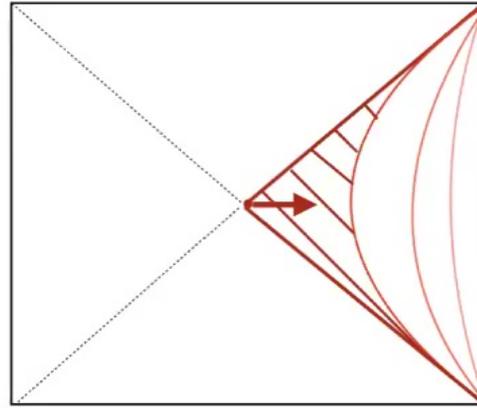


$$\mathcal{E} = \frac{L^2}{\pi\mu} \left(1 - \sqrt{-1 - \frac{4\pi^2\mu}{L^2} \left(\Delta - \frac{c}{12} \right) + 4\pi^4 J^2 \frac{\mu^2}{L^4}} \right)$$

Initial condition fixed by:

$$\mathcal{E}_{\eta=-1}(y \gg 1) = \mathcal{E}_{\eta=1}(y \gg 1)$$

Cosmic Horizon Patch



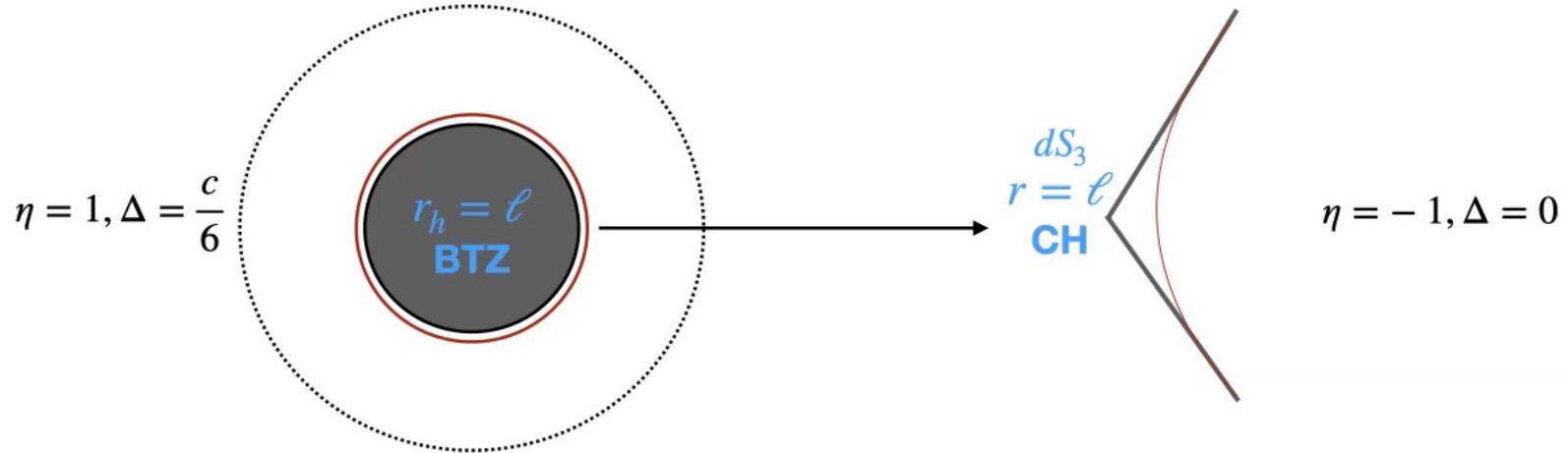
$$\mathcal{E} = \frac{1}{\pi y} \left(1 + \sqrt{-1 + \frac{2y}{y_0} - 4\pi^2 y \left(\Delta - \frac{c}{12} \right) + 4\pi^4 y^2 J^2} \right) \quad y_0 = \frac{1}{4\pi^2 (\Delta_* - c/12)}, \Delta_* = \frac{c}{6} + \delta$$

Initial condition fixed by:

$$\mathcal{E}_{\eta=-1}(y = y_0) = \mathcal{E}_{\eta=1}(y = y_0)$$

$$ds_3^2 = - \left(\frac{r^2 - r_h^2}{l^2} \right) d\tau^2 + \left(\frac{l^2}{r^2 - r_h^2} \right) dr^2 + r^2 d\phi^2,$$

$$ds_3^2 = - \left(\frac{\ell^2 - r^2}{\ell^2} \right) d\tau^2 + \left(\frac{\ell^2}{\ell^2 - r^2} \right) dr^2 + r^2 d\phi^2$$



$$r \equiv r_h + \delta r = \ell + \delta r, \quad \delta r = \frac{w^2}{2\ell} \ll r_h = \ell$$

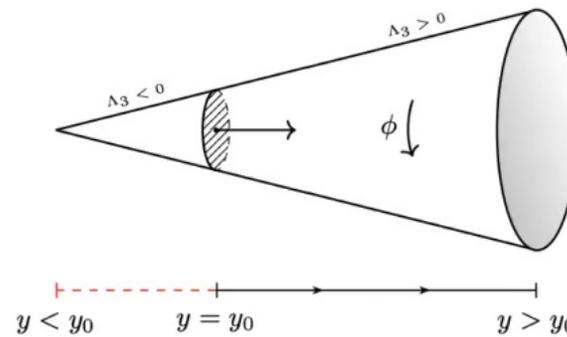
$$r \equiv \ell - \delta r, \quad \delta r = \frac{w^2}{2\ell} \ll \ell$$

Near Horizon: $ds_3^2 = - \frac{w^2}{\ell^2} d\tau^2 + dw^2 + (\ell^2 + \eta w^2) d\phi^2$

$$\mathcal{E} = \frac{L^2}{\pi\mu} \left(1 + \sqrt{-1 - \frac{4\pi^2\mu}{L^2} \left(-\frac{c}{12} \right)} \right) = \frac{1}{\pi y} \left(1 + \sqrt{-1 + \frac{y}{y_0}} \right)$$

Real Excited States

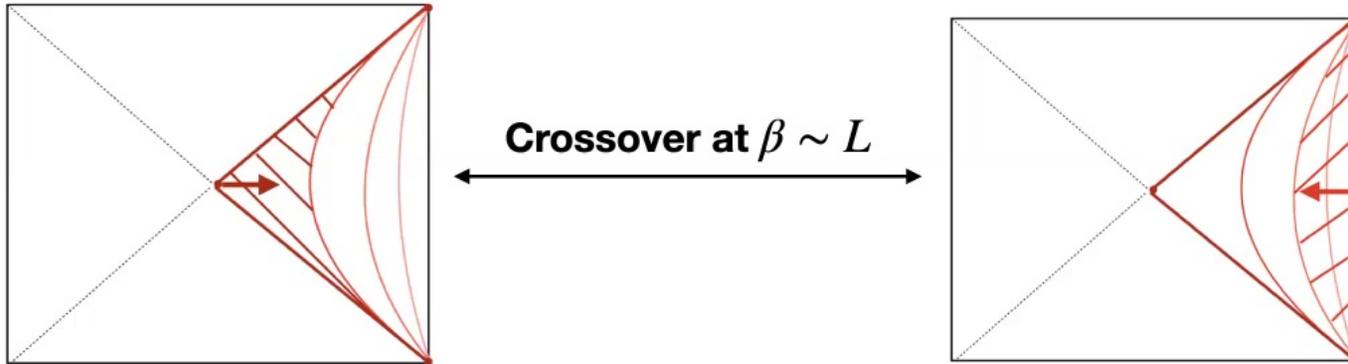
$$\Delta < c/6$$



$$\mathcal{E}_{\eta=-1} = \frac{1}{\pi y} \left(1 + \sqrt{-1 + 4\pi^2 y \left(\Delta_* - \frac{c}{12} + (\Delta_* - \Delta) \right)} \right) \quad \Delta_* = \frac{c}{6} + \delta$$

Hawking-Page like transition

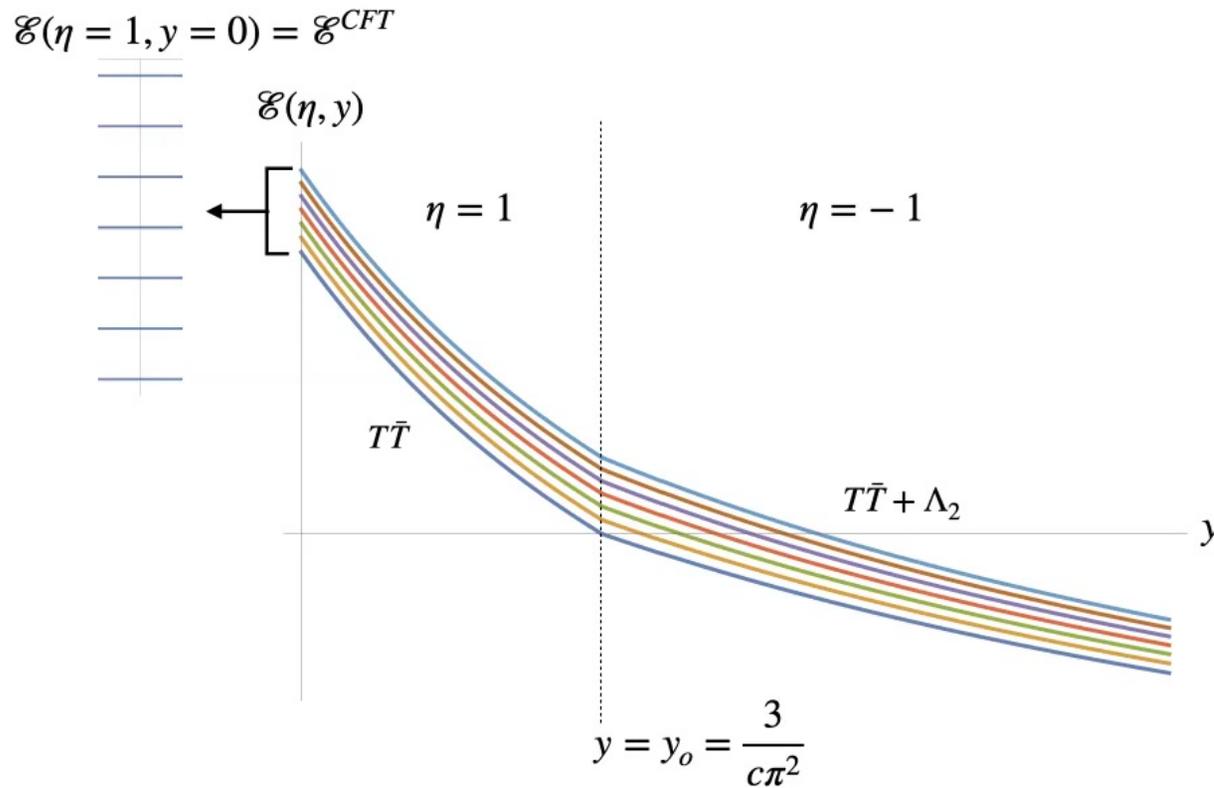
$$\log Z \simeq \max \left\{ -\beta E_{vac}(L), -LE_{vac}(\beta) \right\} = -\frac{\beta L}{\pi\lambda} \min \left\{ 1 - \frac{\beta}{L}, 1 - \frac{L}{\beta} \right\}$$



When $\beta < L$:

$$\log Z \Big|_{\beta < L} \simeq -LE_{vac}(\beta) = S_{Cardy}(\Delta = c/6) - \beta E_{\Delta=c/6}(L)$$

Integrable flow - No level crossing!

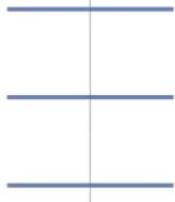


Truncated Hilbert Space

Cutoff @ $\Delta = \frac{c}{6} + \delta$

$\Delta = \frac{c}{6}$  $O(1)$ $\rho(\Delta) \propto \exp\left(\sqrt{\frac{c\pi}{3}}\left(\Delta - \frac{c}{12}\right)\right)$

sparse light spectrum
(sub-leading,
model dependent)

 $\Delta = 0$

$\rho(\Delta) \propto \sum_n \delta(\Delta - \Delta_{n_i})$

Flowing the Density of States

$$\log Z(\beta) = \min_{\mathcal{E}} (\log \rho(\mathcal{E}) - \beta \mathcal{E})$$

Deformation: $\mathcal{E} \rightarrow f_y(\mathcal{E})$

$$\log Z(\beta, y) = \min_{\mathcal{E}} \left(\log \rho(\underbrace{f_y^{-1}(\mathcal{E})}_{\text{undeformed d.o.s.}}) - \beta f_y(\mathcal{E}) \right)$$

Undeformed d.o.s. written in terms of deformed energy levels

Final Result

$$S_{\eta=-1}(\mathcal{E}, y) = \sqrt{\frac{2c\pi}{3} \left(\frac{y - y_o}{\pi y} + \mathcal{E} - \frac{\pi y \mathcal{E}}{2} \right)} - 3 \log \left(\sqrt{\frac{2c\pi}{3} \left(\frac{y - y_o}{\pi y} + \mathcal{E} - \frac{\pi y \mathcal{E}}{2} \right)} \right) + \log c + \dots$$

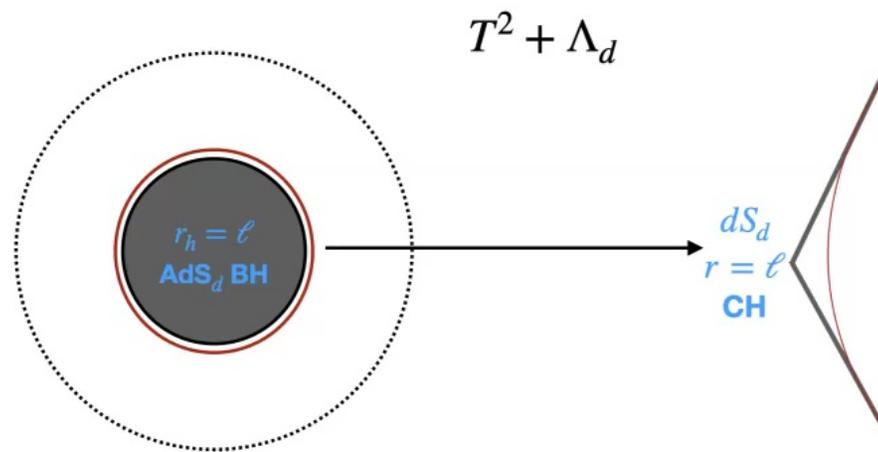
Near horizon limit: $y = y_o = \frac{3}{c\pi^2}$, $\mathcal{E} = \frac{1}{\pi y}$, where:

$$S = \frac{c\pi}{3} - 3 \log \left(\frac{c\pi}{3} \right) + \dots$$

In bulk terms:

$$S_{hor} = \frac{\pi \ell}{2G} - 3 \log \left(\frac{\pi \ell}{2G} \right) + \dots$$

Higher dimensions



Thank you for your attention!