

Title: Is singularity resolution trivial?

Speakers: Ding Jia

Collection: Young Researchers Conference

Date: June 22, 2022 - 9:25 AM

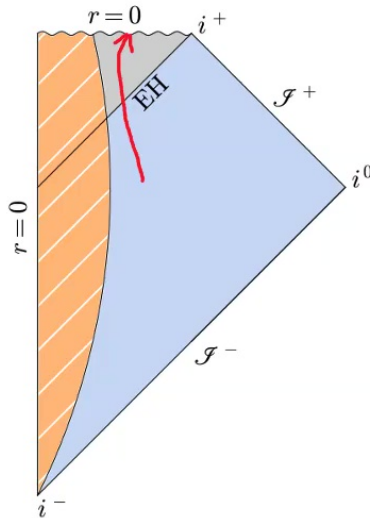
URL: <https://pirsa.org/22060055>

Abstract: Many non-trivial ideas have been proposed to resolve singularities in quantum gravity. In this talk I argue that singularity resolution can be trivial in gravitational path integrals, because geodesically incomplete singular spacetimes are usually not included in the sum. For theories where this holds, there is no need to develop non-trivial ideas on singularity resolution. Instead, efforts should better be directed to understand tunneling processes and complex-valued spacetimes.

# IS SINGULARITY RESOLUTION TRIVIAL?

June 21, 2022, Waterloo  
Young Researchers Conference  
Ding Jia

# PENROSE-HAWKING SINGULARITY THEOREMS FOR CLASSICAL SPACETIME



Establish singularities in the geodesically incomplete sense

-Timelike paths terminate at finite proper time

- $g$  stops existing at singularities

# THERE ARE MANY NON-TRIVIAL IDEAS ABOUT SINGULARITY RESOLUTION IN QUANTUM SPACETIME



The Gordian Knot

Choose variable

(e.g., Loop Quantum Cosmology vs. Wheeler-deWitt)

Choose action

(e.g., regular BH: add terms to Einstein-Hilbert)

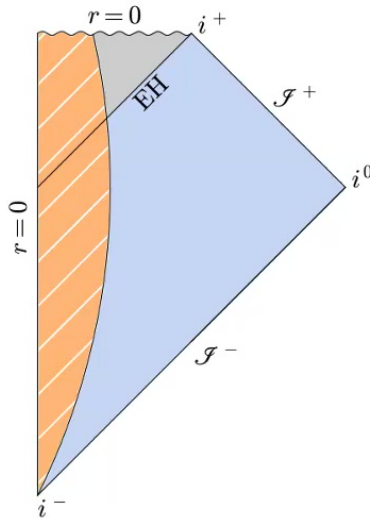
Choose boundary condition

(e.g., final state proposals)

Understand black hole/cosmological solutions in detail  
(e.g., BKL considerations)

...

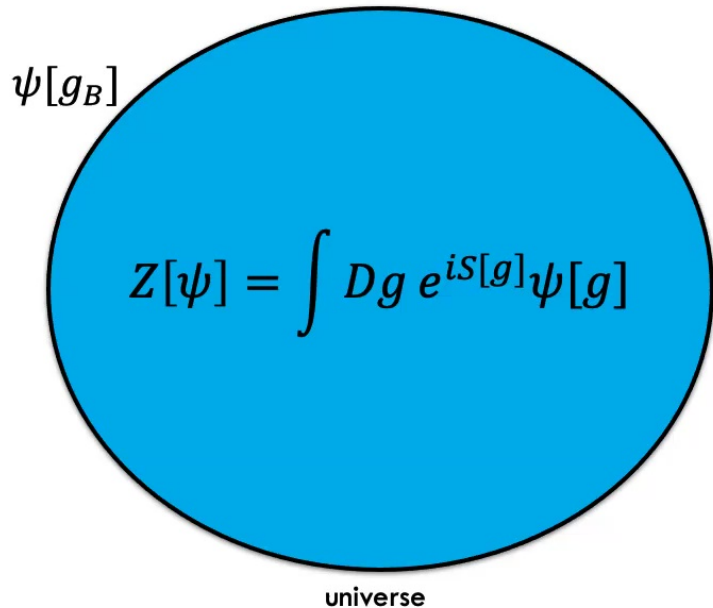
# I WANT TO PRESENT A TRIVIAL IDEA ABOUT SINGULARITY RESOLUTION



Key:

Singular spacetimes don't belong to gravitational path integrals

# ASSUMPTION: GRAVITATIONAL PATH INTEGRAL



$\psi[g_B]$

$$Z[\psi] = \int Dg e^{iS[g]} \psi[g]$$

universe

Pick a theory you like

Can include matter

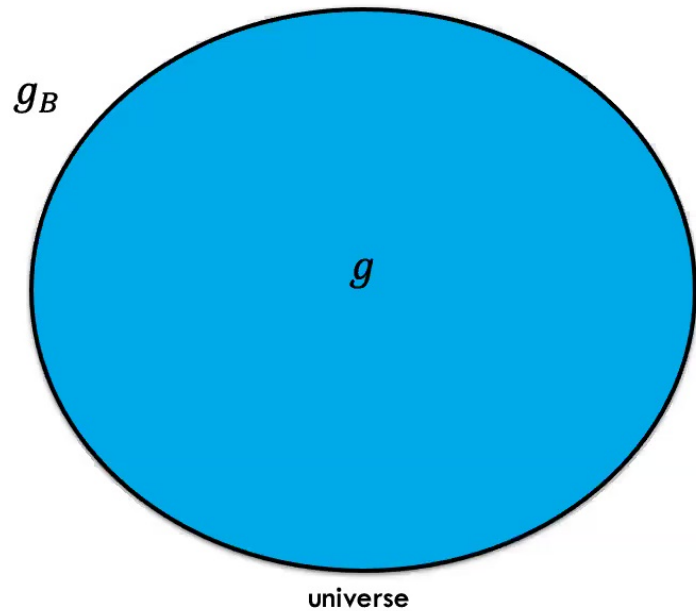
$$Z[\psi] = \int Dg D\phi e^{iS[g,\phi]} \psi[g_B, \phi_B]$$

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$$Z[\psi, w] = \int Dg D\phi e^{iS[g,\phi]} \psi[g_B, \phi_B] w[g, \phi]$$

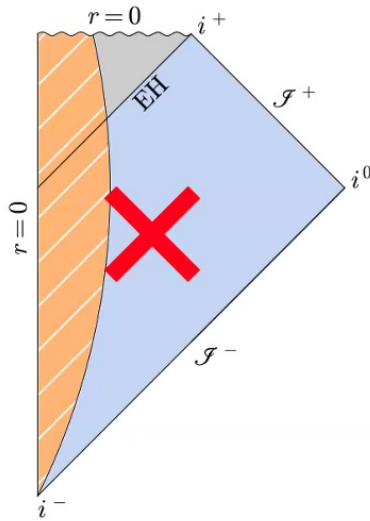
Can include mixedness (double path)

# $g_B$ HAS TO EXIST!



For all  $g$  in  $\int Dg$ ,  
 $g_B$  has to exist as some codimension-1 space(time)  
configuration

# DO SINGULAR SPACETIMES BELONG?

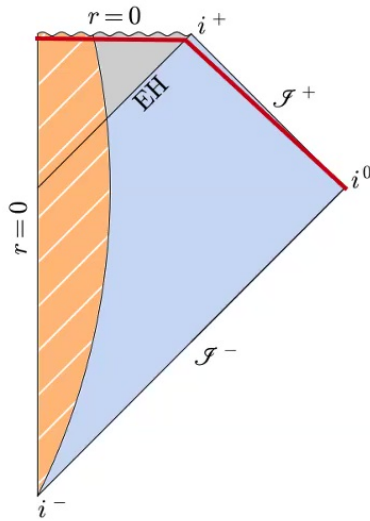


$g$  does not exist at singularities

Hence singular spacetimes do not belong  
to the path integrals

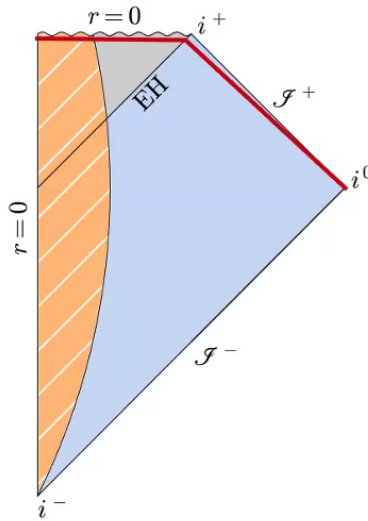


## WHAT IF PICK BOUNDARY CLOSE TO THE SINGULARITY?



Then the restricted portion belongs,  
but there is no singularity!

## WHAT IF PICK BOUNDARY CLOSE TO THE SINGULARITY?



Then the restricted portion belongs,  
but there is no singularity!

The universe ends before any singular  
behavior takes place

## SUMMARY: A TRIVIAL SOLUTION



DJ. "Is singularity resolution trivial?" arXiv:2204.12304

# SUMMARY: A TRIVIAL SOLUTION



In GR, we care about a differential equation

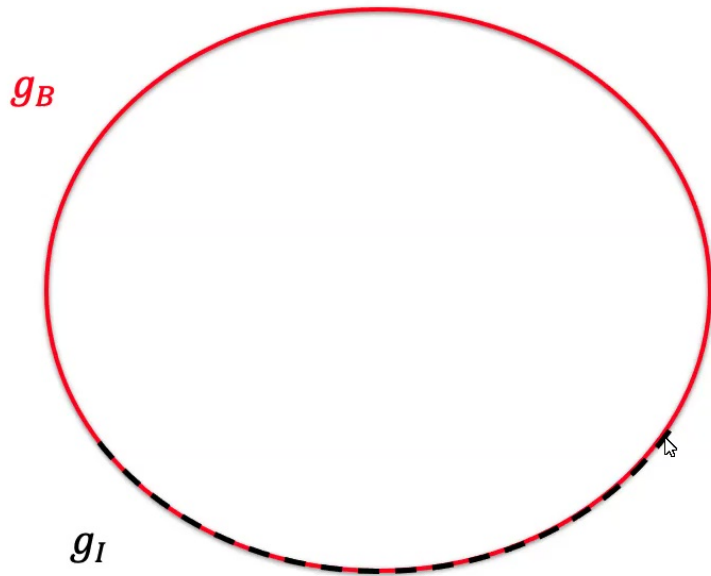
Singular spacetimes are solutions to certain initial value problems, so they constitute an issue

In QG, we care about an integral

Singular spacetimes do not belong to the integral, so they do not constitute an issue

DJ. "Is singularity resolution trivial?" arXiv:2204.12304

# QUESTION 1: DO SINGULAR SPACETIME COME BACK IN CLASSICAL APPROXIMATIONS?



No!

Solve for stationary points:  $\delta S[g] = 0$

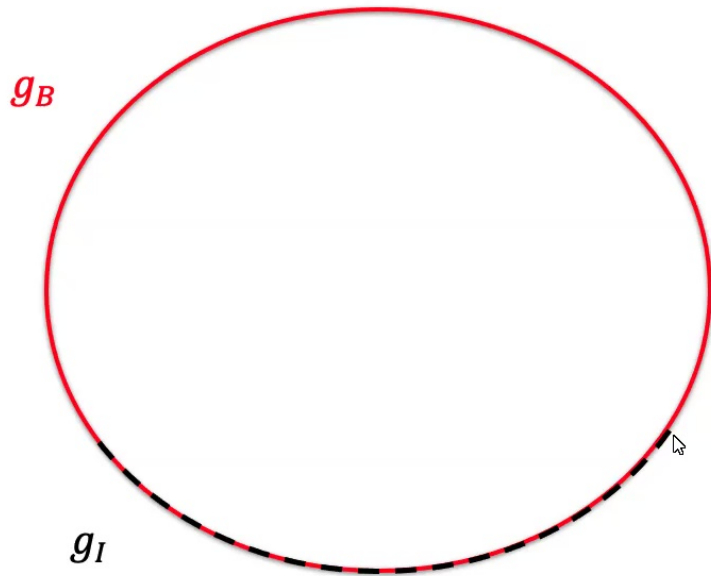
Classical theory:

$$\delta S[g]|_{g_I}$$

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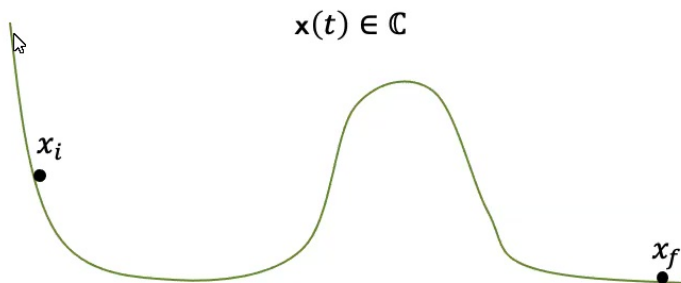
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Classical approximation of quantum theory:

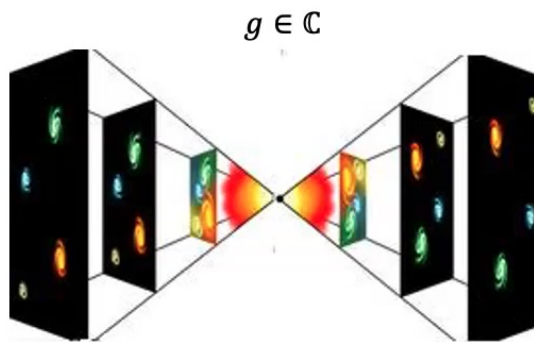
$$\delta S[g]|_{g_B}$$

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## QUESTION 2: WHAT REPLACE SINGULAR SPACETIMES IN CLASSICAL APPROXIMATIONS?

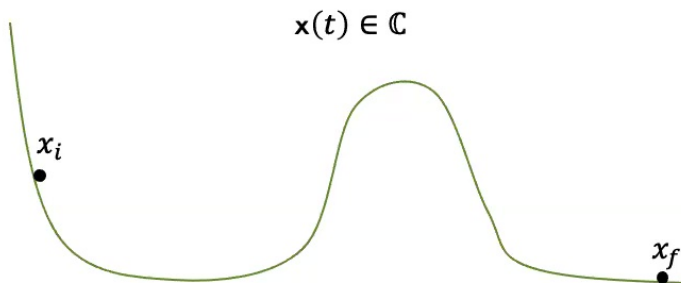


Tunneling solutions!



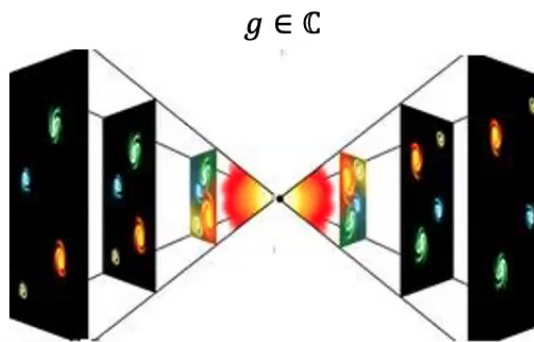
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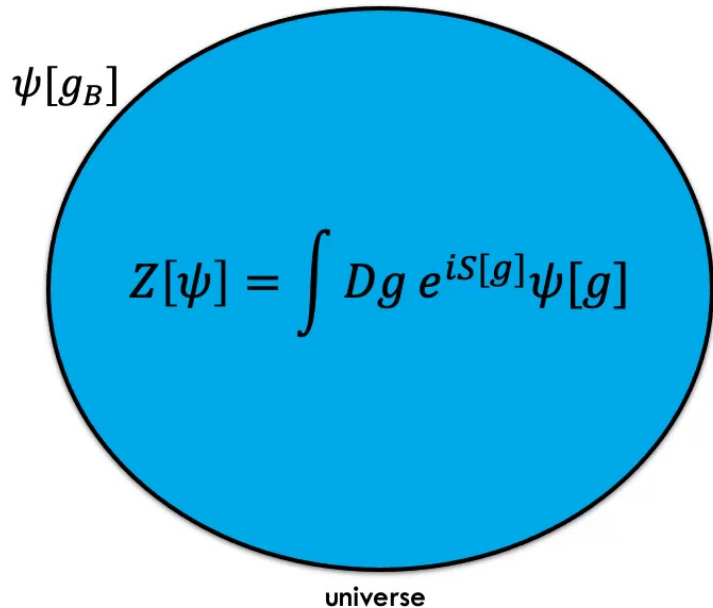
$\delta S = 0$  for black hole initial conditions:  
complex stationary configuration



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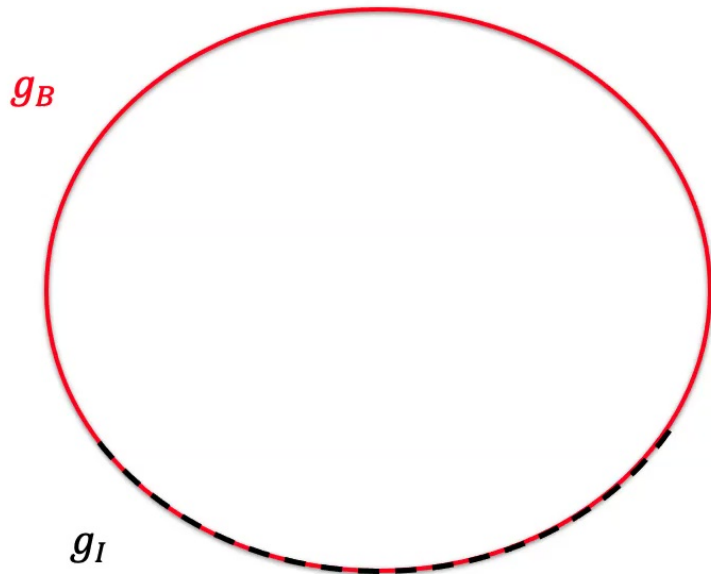
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