

Title: Quantum Reference Frames for Superpositions of Spacetimes

Speakers: Anne-Catherine de la Hamette

Series: Quantum Foundations

Date: May 31, 2022 - 3:30 PM

URL: <https://pirsa.org/22050065>

Abstract: The current theories of quantum physics and general relativity on their own do not allow us to study situations in which spacetime is in a quantum superposition. In this talk, I propose a general strategy to determine the dynamics of objects on an indefinite spacetime metric, using an extended notion of quantum reference frame transformations. First, we study the situation of the gravitational source mass being in a spatial superposition state and, using a generalized principle of covariance, show how to transform to a frame in which the standard theories of GR and QFT allow to determine the dynamics. In the second part, we consider superpositions of conformally equivalent metrics inhabited by a massive quantized Klein-Gordon field. By requiring invariance of the KG equation under quantum conformal transformations, we find that the superposition is transferred to the quantum field in the form of an effective, spacetime dependent mass term. Overall, the proposed strategy allows to construct the respective explicit quantum frame change operators, and to study physical phenomena such as time dilation and cosmological particle production in different quantum frames.

Zoom Link: <https://pitp.zoom.us/j/96903859307?pwd=aEtLUy9tME5GL25nTjBVNXVmb2N3Zz09>



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Quantum Reference Frames for Superpositions of Spacetimes

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University of Vienna, IQOQI Vienna

in collaboration with V. Kabel, E. Castro-Ruiz, and Č. Brukner
arXiv:2112.11473 + work in preparation

Perimeter Institute, Waterloo, ON, May 31st, 2022





- ▶ Falling through masses in superposition: quantum reference frames for indefinite metrics, arXiv:2112.11473
- ▶ Spacetimes in superposition: Extended symmetries of the Klein-Gordon field, *to appear*



Viktoria Kabel



Esteban Castro-Ruiz

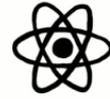


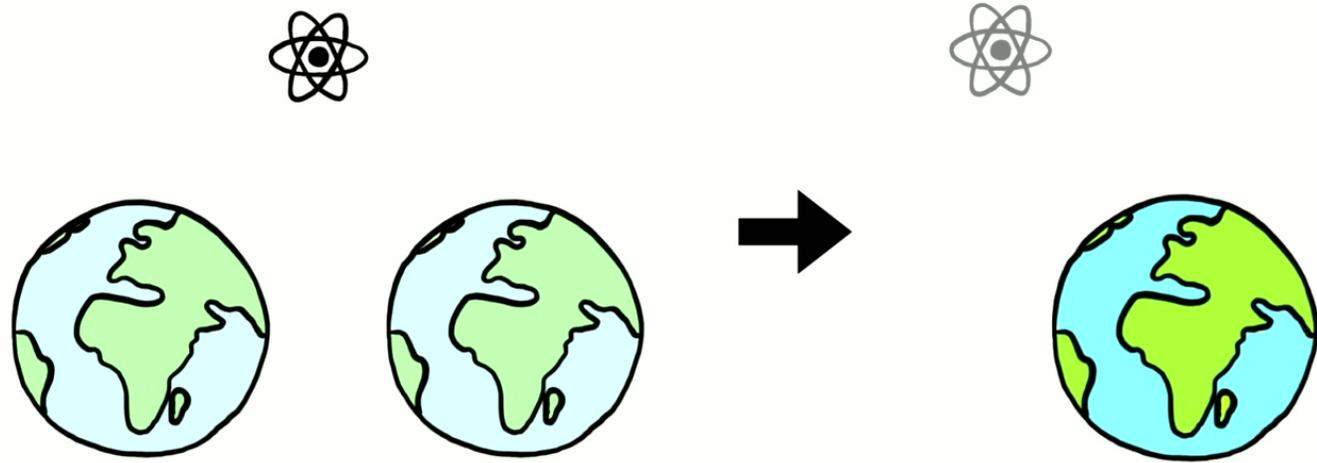
Časlav Brukner

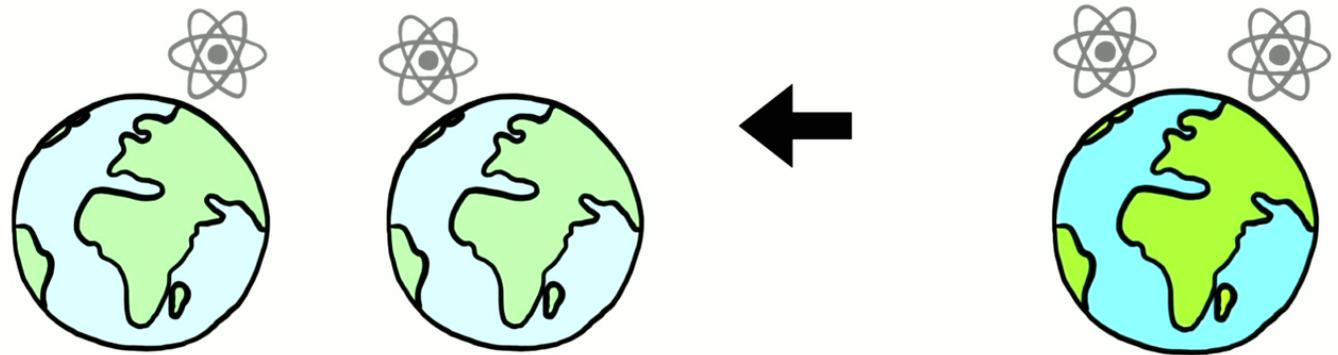












Overview

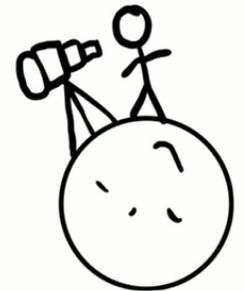
Falling through masses in superposition

1. Quantum reference frames
2. Generalized principle of covariance
3. Motion of a test particle and time dilation
4. Generalizations

Extended symmetries of KG field in superposition of spacetimes

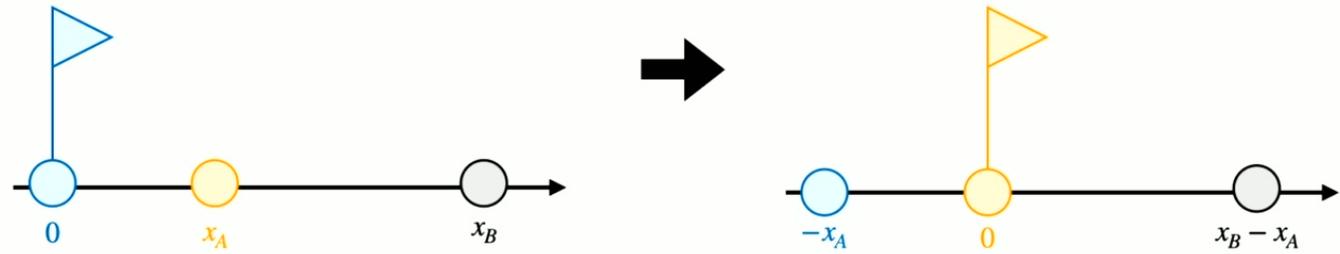
1. Symmetries of the KG field in an expanding universe
2. Quantum conformal transformations
3. Cosmological particle production

Superpositions of spacetimes: extended symmetry principles



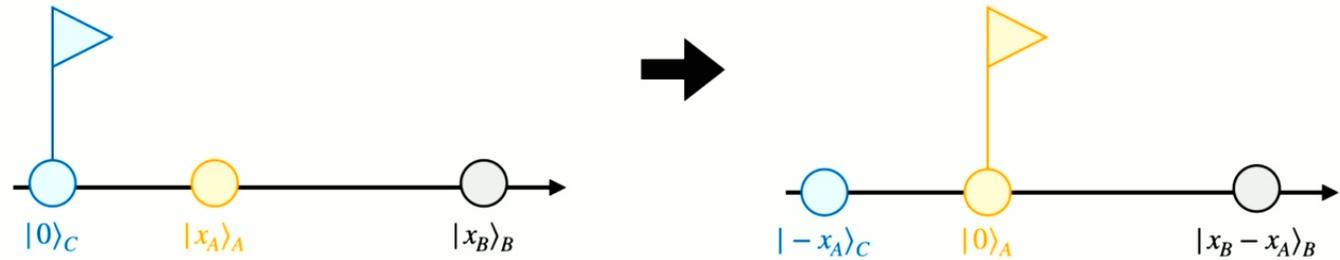
Quantum Reference Frames

Formalism



Quantum Reference Frames

Formalism



$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C |x_A\rangle_A |x_B\rangle_B$$

$$\in \mathcal{H}_{ABC}^{(C)} = \mathcal{H}_A^{(C)} \otimes \mathcal{H}_B^{(C)} \otimes \mathcal{H}_C^{(C)}$$

$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A |-x_A\rangle_C |x_B - x_A\rangle_B$$

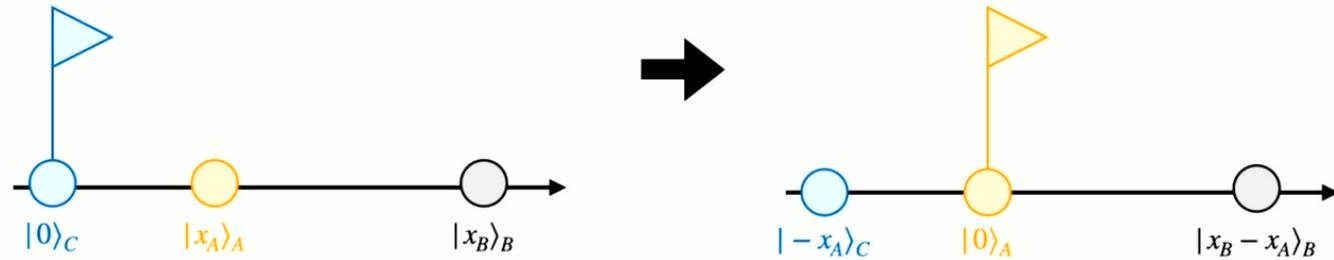
$$\in \mathcal{H}_{ABC}^{(A)} = \mathcal{H}_A^{(A)} \otimes \mathcal{H}_B^{(A)} \otimes \mathcal{H}_C^{(A)}$$

F. Giacomini, E. Castro-Ruiz, and Č. Brukner, *Nature Comm.* 10, 494 (2019).
 A.-C. de la Hamette, T. Galley, *Quantum* 4, 367 (2020).



Quantum Reference Frames

Formalism



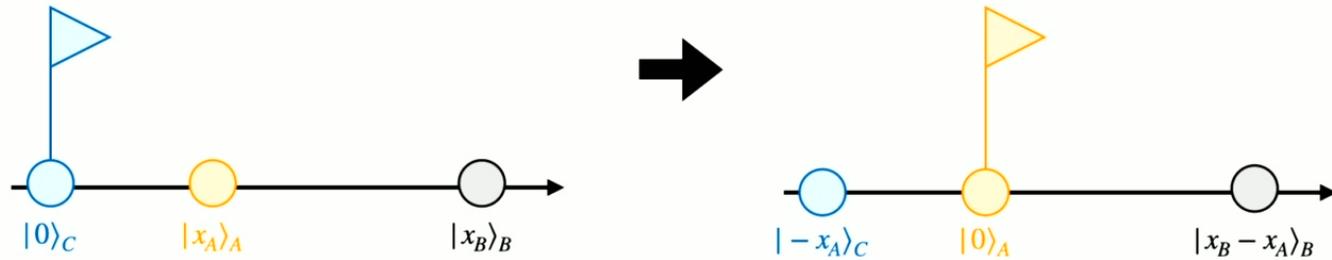
$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C |x_A\rangle_A |x_B\rangle_B$$

$$\begin{aligned} |\psi\rangle_{ABC}^{(A)} &= |0\rangle_A |-x_A\rangle_C |x_B - x_A\rangle_B \\ &= \mathcal{P}_{CA} |0\rangle_C |x_A\rangle_A \boxed{e^{\frac{i}{\hbar} x_A \hat{p}_B} |x_B\rangle_B} \end{aligned}$$



Quantum Reference Frames

Formalism



$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C |x_A\rangle_A |x_B\rangle_B$$

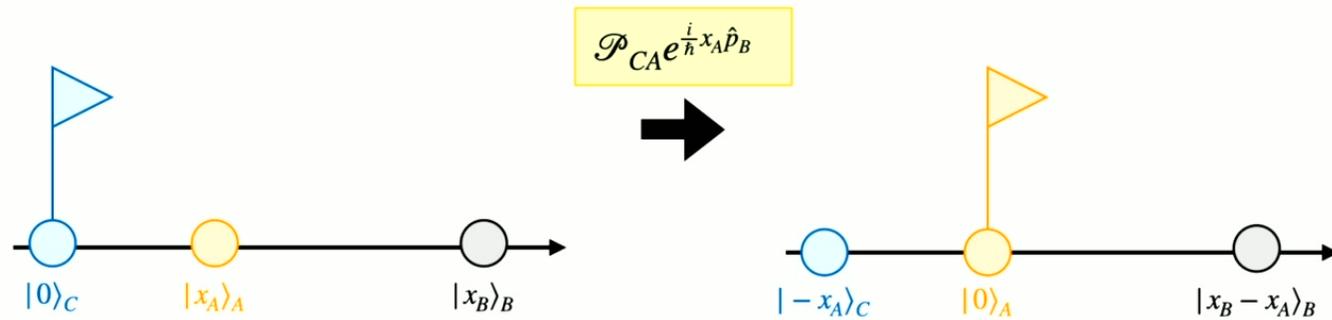
$$\begin{aligned} |\psi\rangle_{ABC}^{(A)} &= |0\rangle_A |-x_A\rangle_C |x_B - x_A\rangle_B \\ &= \mathcal{P}_{CA} |0\rangle_C |x_A\rangle_A e^{\frac{i}{\hbar} x_A \hat{p}_B} |x_B\rangle_B \end{aligned}$$

$$\mathcal{P}_{CA} = \text{SWAP}_{CA} \circ \int |-x_A\rangle\langle x_A|_A dx_A$$



Quantum Reference Frames

Formalism



$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C |x_A\rangle_A |x_B\rangle_B$$

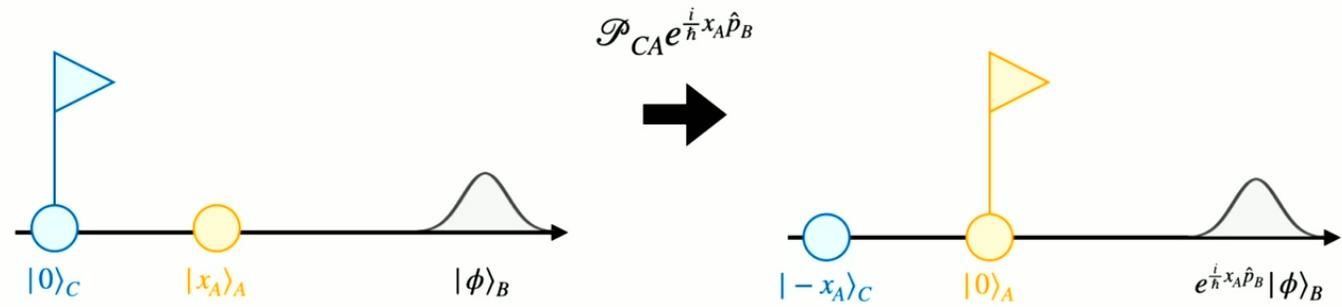
$$\begin{aligned} |\psi\rangle_{ABC}^{(A)} &= |0\rangle_A |-x_A\rangle_C |x_B - x_A\rangle_B \\ &= \mathcal{P}_{CA} |0\rangle_C |x_A\rangle_A e^{\frac{i}{\hbar} x_A \hat{p}_B} |x_B\rangle_B \\ &= \mathcal{P}_{CA} e^{\frac{i}{\hbar} x_A \hat{p}_B} |\psi\rangle_{ABC}^{(C)} \end{aligned}$$

$$\mathcal{P}_{CA} = \text{SWAP}_{CA} \circ \int |-x_A\rangle\langle x_A|_A dx_A$$



Quantum Reference Frames

Formalism



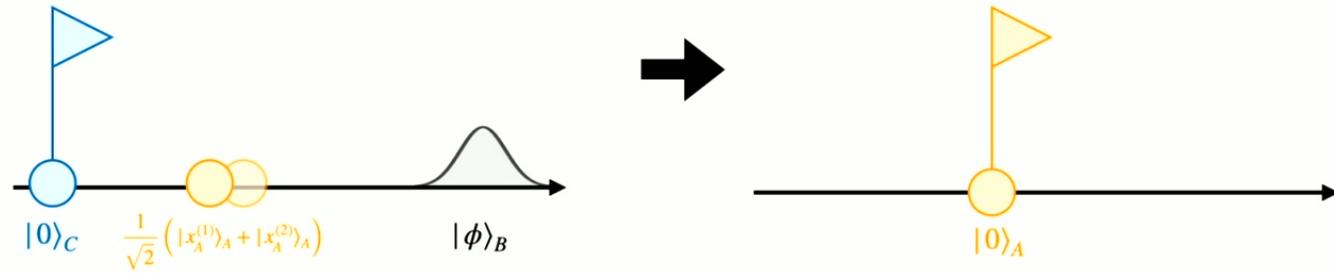
$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C |x_A\rangle_A |\phi\rangle_B$$

$$|\psi\rangle_{ABC}^{(A)} = \mathcal{P}_{CA} e^{\frac{i}{\hbar} x_A \hat{p}_B} |0\rangle_C |x_A\rangle_A |\phi\rangle_B$$



Quantum Reference Frames

Formalism

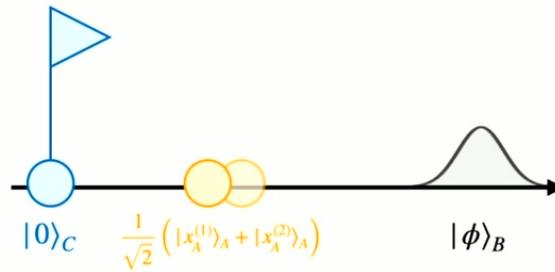


$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C \frac{1}{\sqrt{2}} \left(|x_A^{(1)}\rangle_A + |x_A^{(2)}\rangle_A \right) |\phi\rangle_B$$

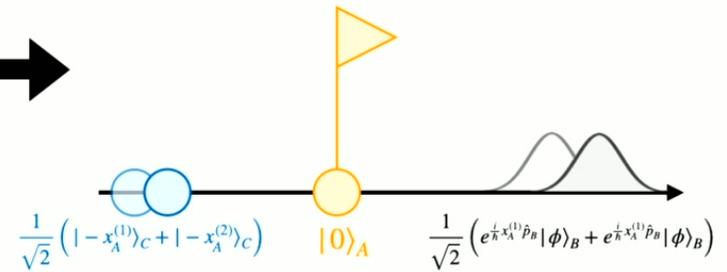


Quantum Reference Frames

Formalism



$$|\psi\rangle_{ABC}^{(C)} = |0\rangle_C \frac{1}{\sqrt{2}} (|x_A^{(1)}\rangle_A + |x_A^{(2)}\rangle_A) |\phi\rangle_B$$

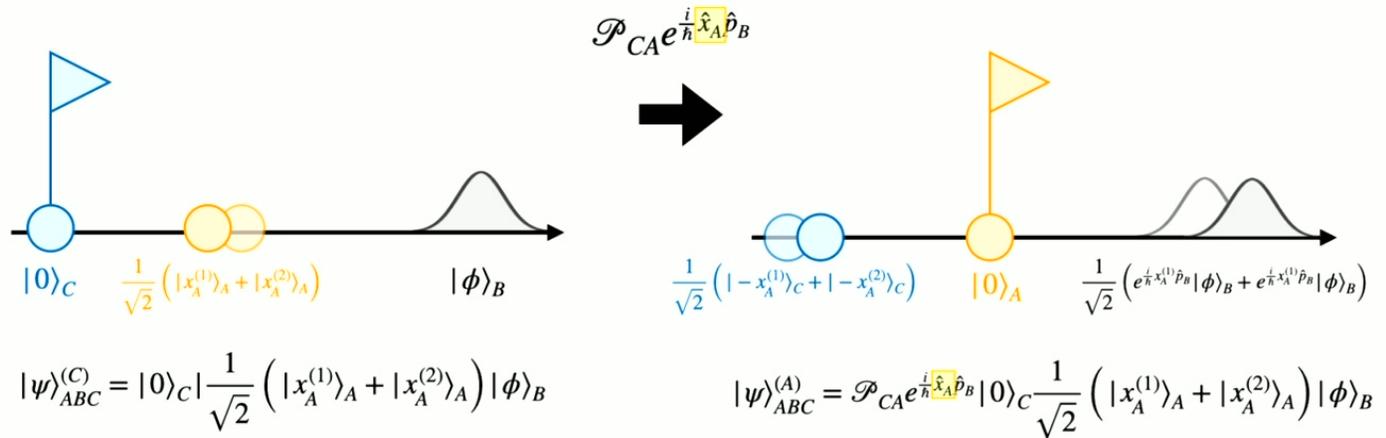


$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A \frac{1}{\sqrt{2}} (|-x_A^{(1)}\rangle_C e^{\frac{i}{\hbar}x_A^{(1)}\hat{p}_B}|\phi\rangle_B + |-x_A^{(2)}\rangle_C e^{\frac{i}{\hbar}x_A^{(2)}\hat{p}_B}|\phi\rangle_B)$$



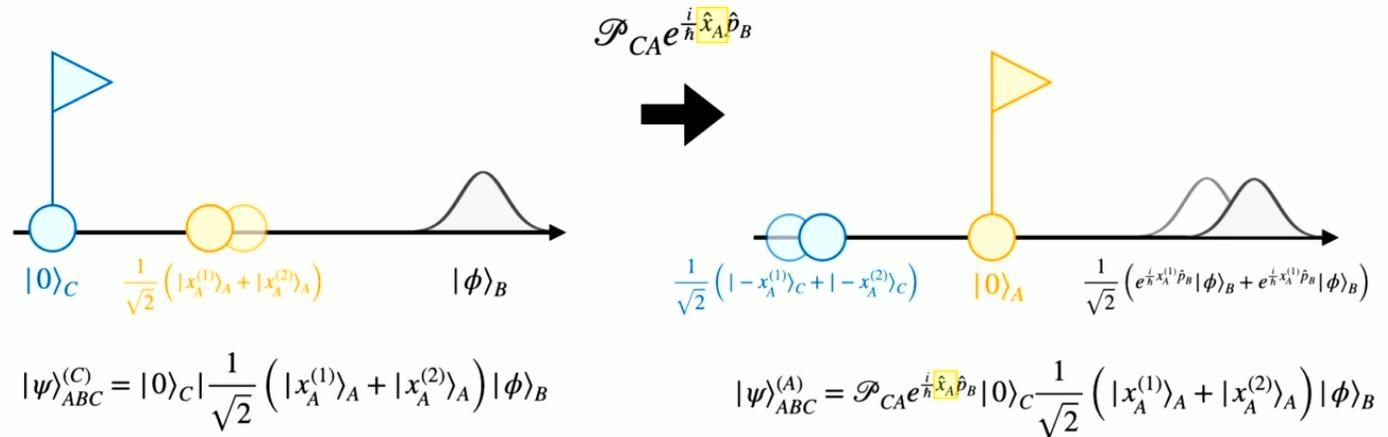
Quantum Reference Frames

Formalism



Quantum Reference Frames

Formalism

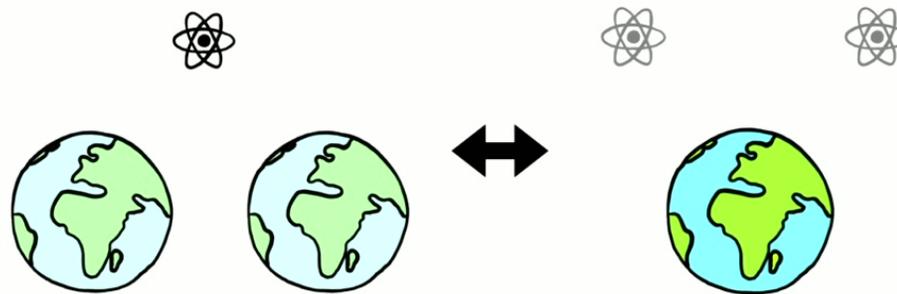


Quantum controlled translation

$$\mathcal{P}_{CA} e^{\frac{i}{\hbar} \hat{x}_A \hat{p}_B}$$


Quantum Reference Frames

Generalized Principle of Covariance

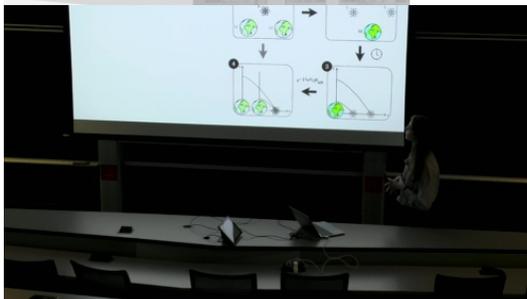
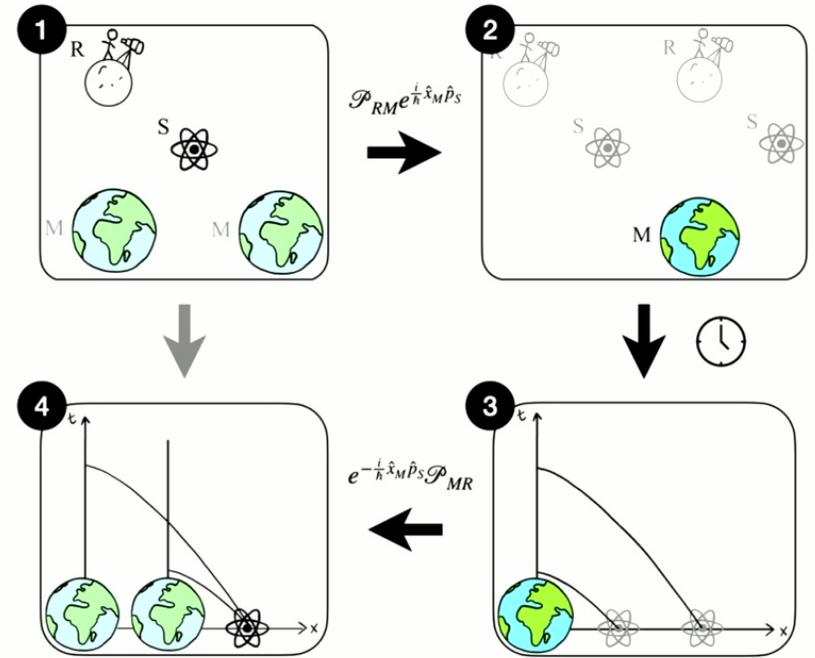


Covariance of dynamical laws under quantum coordinate transformations:
Physical laws retain their form under *quantum* coordinate transformations.



Applications

Motion of a Test Particle

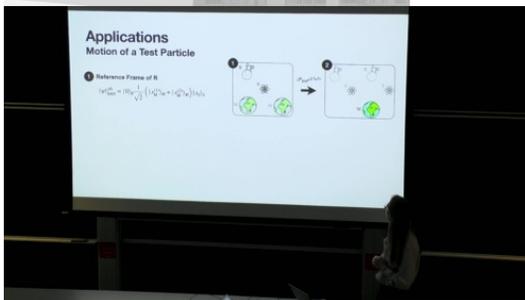
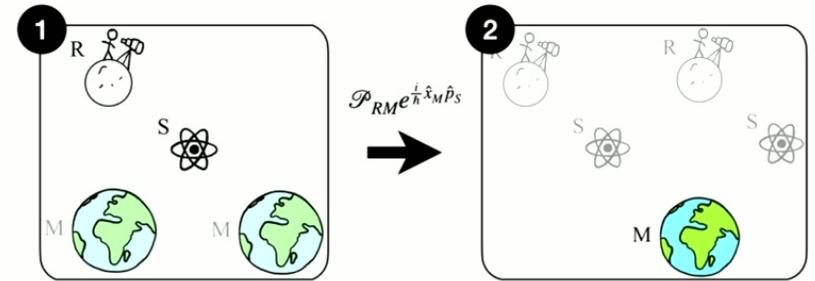


Applications

Motion of a Test Particle

1 Reference Frame of R

$$|\psi\rangle_{RMS}^{(R)} = |0\rangle_R \frac{1}{\sqrt{2}} \left(|x_M^{(1)}\rangle_M + |x_M^{(2)}\rangle_M \right) |x_S\rangle_S$$



Applications

Motion of a Test Particle

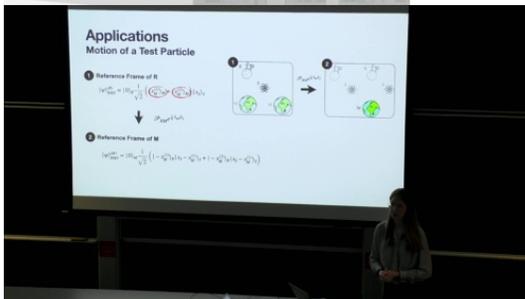
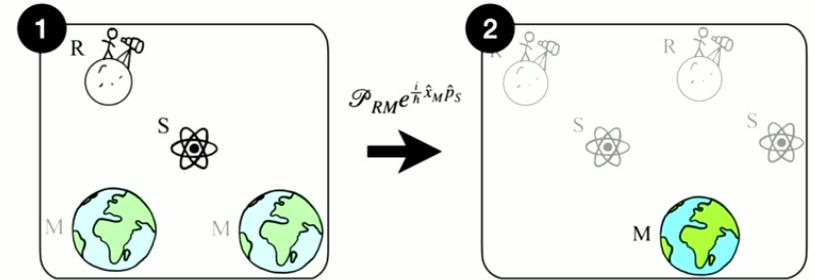
1 Reference Frame of R

$$|\psi\rangle_{RMS}^{(R)} = |0\rangle_R \frac{1}{\sqrt{2}} \left(|x_M^{(1)}\rangle_M + |x_M^{(2)}\rangle_M \right) |x_S\rangle_S$$

$$\downarrow \mathcal{P}_{RM} e^{\frac{i}{\hbar} \hat{x}_M \hat{p}_S}$$

2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(| -x_M^{(1)}\rangle_R |x_S - x_M^{(1)}\rangle_S + | -x_M^{(2)}\rangle_R |x_S - x_M^{(1)}\rangle_S \right)$$



Applications

Motion of a Test Particle

2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(| -x_M^{(1)} \rangle_R |x_S - x_M^{(1)}\rangle_S + | -x_M^{(2)} \rangle_R |x_S - x_M^{(1)}\rangle_S \right)$$



3

$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} | -x_M^{(1)} \rangle_R |\tilde{x}_S^{(1)}\rangle_S + e^{-\frac{i}{\hbar}\Phi^{(2)}} | -x_M^{(2)} \rangle_R |\tilde{x}_S^{(2)}\rangle_S \right)$$

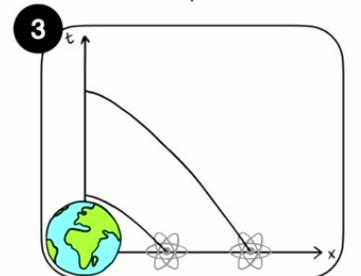
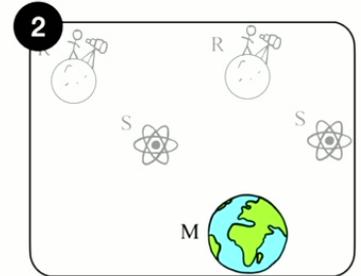
geodesic motion

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\nu\rho}^\mu \frac{dx^\nu}{d\tau} \frac{dx^\rho}{d\tau} = 0$$

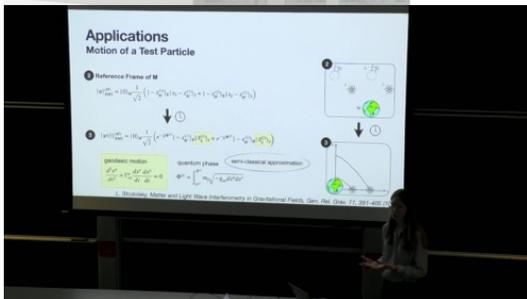
quantum phase

$$\Phi^{(i)} = \int_{A^{(i)}}^{B^{(i)}} m_S \sqrt{-g_{\mu\nu}} dx^\mu dx^\nu$$

semi-classical approximation



L. Stodolsky, *Matter and Light Wave Interferometry in Gravitational Fields*, *Gen. Rel. Grav.* 11, 391-405 (1979).



Applications

Motion of a Test Particle

2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(| -x_M^{(1)} \rangle_R |x_S - x_M^{(1)} \rangle_S + | -x_M^{(2)} \rangle_R |x_S - x_M^{(1)} \rangle_S \right)$$



3

$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} | -x_M^{(1)} \rangle_R |\tilde{x}_S^{(1)} \rangle_S + e^{-\frac{i}{\hbar}\Phi^{(2)}} | -x_M^{(2)} \rangle_R |\tilde{x}_S^{(2)} \rangle_S \right)$$

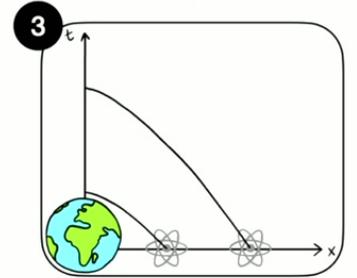
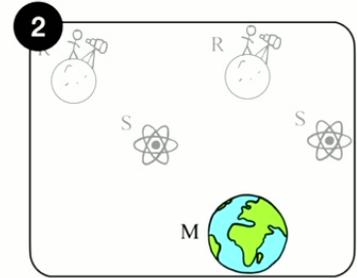
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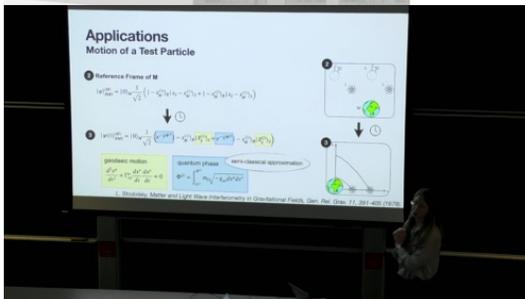
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Applications

Motion of a Test Particle

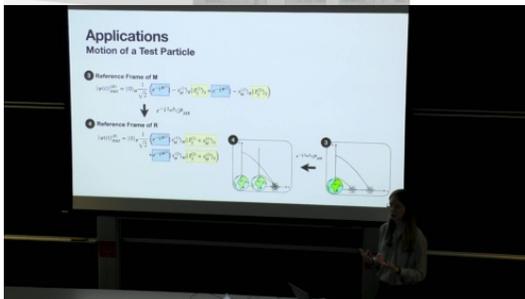
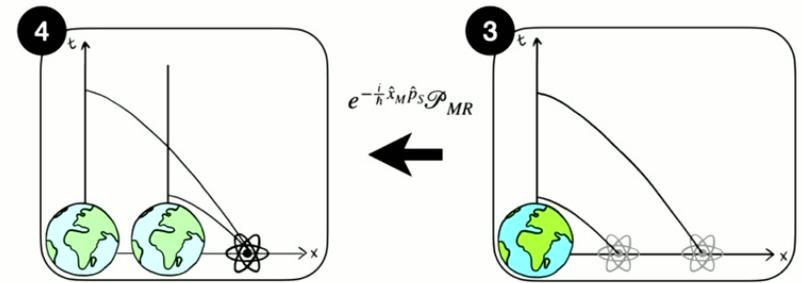
3 Reference Frame of M

$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} | -x_M^{(1)} \rangle_R |\tilde{x}_S^{(1)}\rangle_S + e^{-\frac{i}{\hbar}\Phi^{(2)}} | -x_M^{(2)} \rangle_R |\tilde{x}_S^{(2)}\rangle_S \right)$$

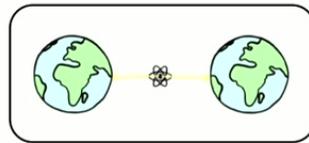
$$\Downarrow e^{-\frac{i}{\hbar}\hat{x}_M \hat{p}_S} \mathcal{P}_{MR}$$

4 Reference Frame of R

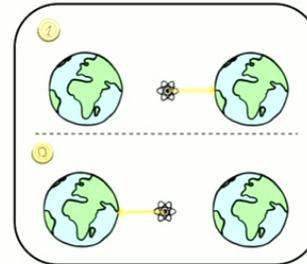
$$|\psi(t)\rangle_{RMS}^{(R)} = |0\rangle_R \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} |x_M^{(1)}\rangle_M |\tilde{x}_S^{(1)} + x_M^{(1)}\rangle_S + e^{-\frac{i}{\hbar}\Phi^{(2)}} |x_M^{(2)}\rangle_M |\tilde{x}_S^{(2)} + x_M^{(2)}\rangle_S \right)$$



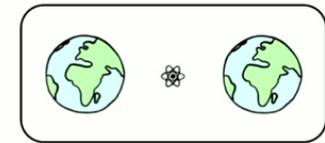
Comparison with other Approaches



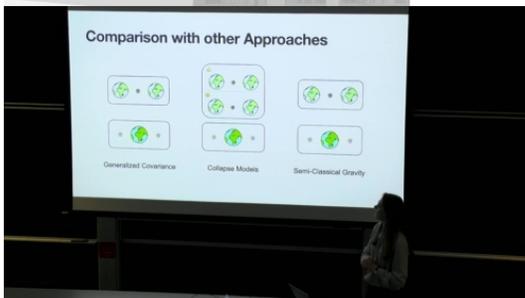
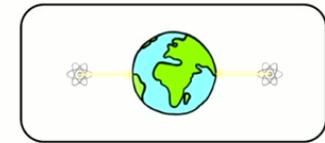
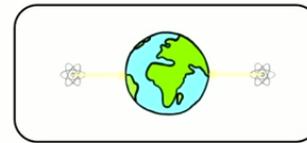
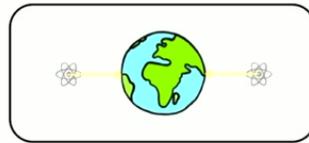
Generalized Covariance



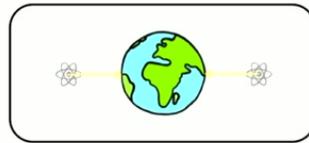
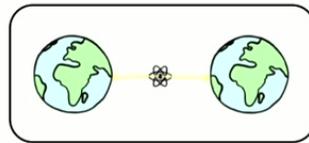
Collapse Models



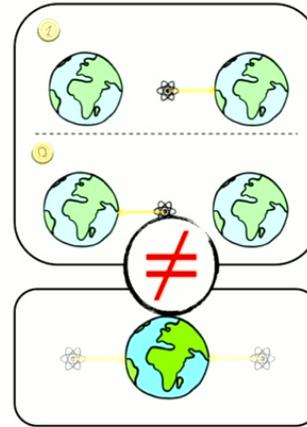
Semi-Classical Gravity



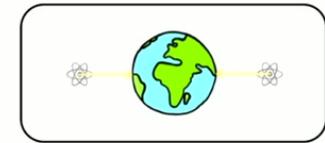
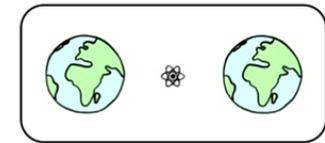
Comparison with other Approaches



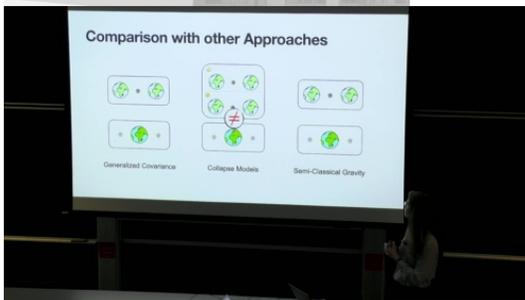
Generalized Covariance



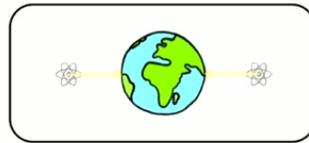
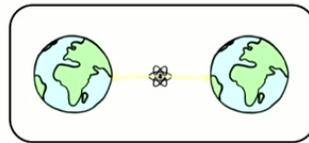
Collapse Models



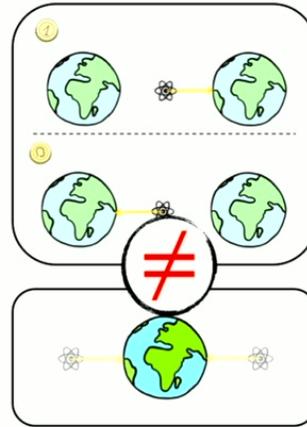
Semi-Classical Gravity



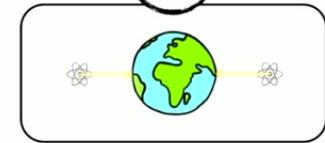
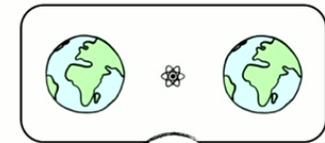
Comparison with other Approaches



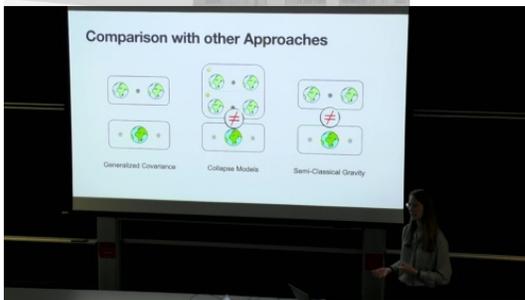
Generalized Covariance



Collapse Models

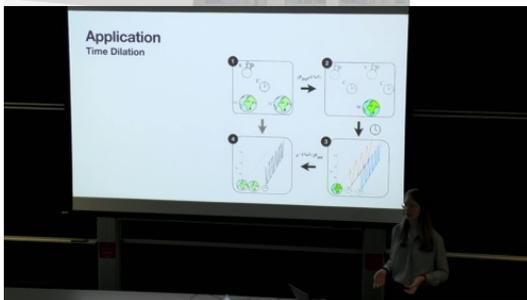
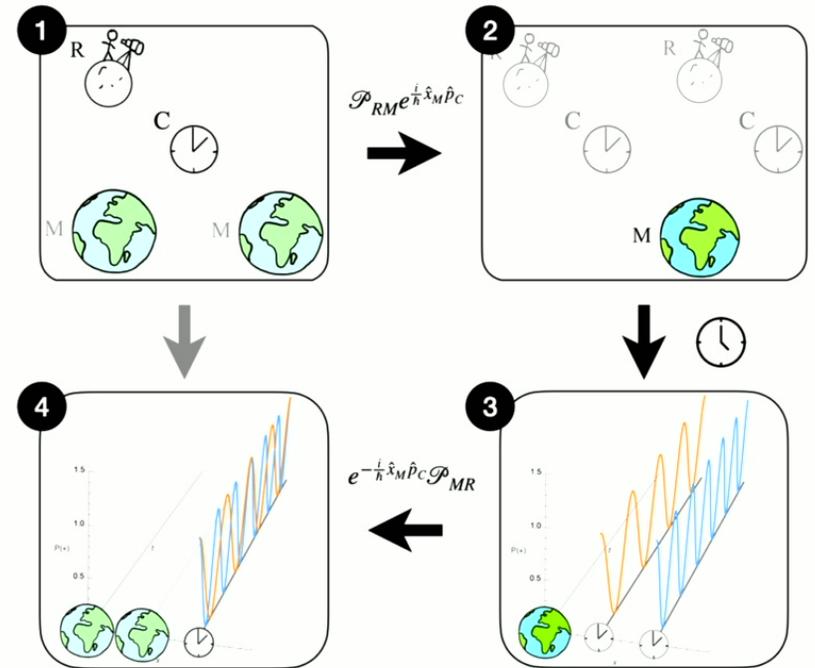


Semi-Classical Gravity



Application

Time Dilation



Application

Time Dilation

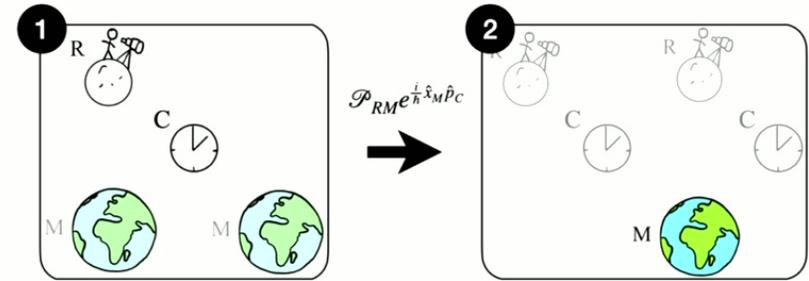
1 Reference Frame of R

$$|\psi\rangle_{RMC}^{(R)} = |0\rangle_R \frac{1}{\sqrt{2}} \left(|x_M^{(1)}\rangle_M + |x_M^{(2)}\rangle_M \right) |x_C\rangle_{C_{ext}} |s(\tau_0)\rangle_{C_{int}}$$

$$\downarrow \mathcal{P}_{RM} e^{\frac{i}{\hbar} \hat{x}_M \hat{p}_C}$$

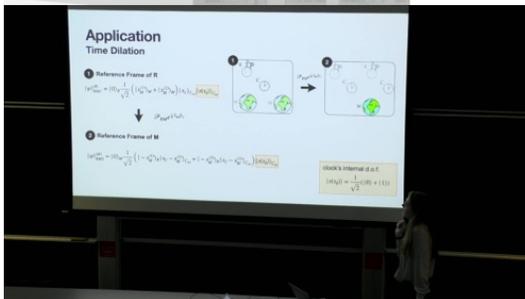
2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(|-x_M^{(1)}\rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} + |-x_M^{(2)}\rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} \right) |s(\tau_0)\rangle_{C_{int}}$$



clock's internal d.o.f.

$$|s(\tau_0)\rangle = \frac{1}{\sqrt{2}} (|0\rangle + |1\rangle)$$



Application

Time Dilation

2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(|-x_M^{(1)}\rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} + |-x_M^{(2)}\rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} \right) |s(\tau_0)\rangle_{C_{int}}$$



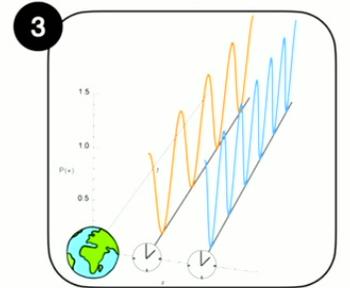
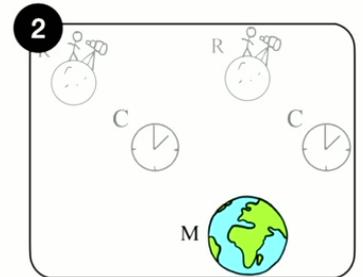
$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} |-x_M^{(1)}\rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(1)})\rangle_{C_{int}} + e^{-\frac{i}{\hbar}\Phi^{(2)}} |-x_M^{(2)}\rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(2)})\rangle_{C_{int}} \right)$$

proper time

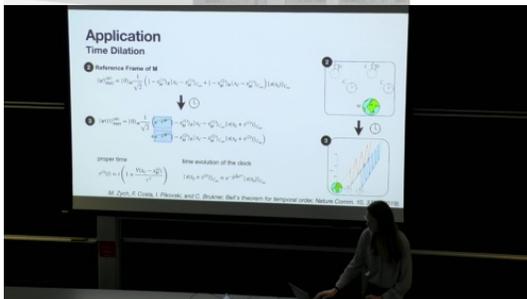
$$\tau^{(i)}(t) = t \left(1 + \frac{V(x_C - x_M^{(i)})^2}{c^2} \right)$$

time evolution of the clock

$$|s(\tau_0 + \tau^{(i)})\rangle_{C_{int}} = e^{-\frac{i}{\hbar}\hat{\Omega}\tau^{(i)}} |s(\tau_0)\rangle_{C_{int}}$$



M. Zych, F. Costa, I. Pikovski, and C. Brukner, Bell's theorem for temporal order, Nature Comm. 10, 3772 (2019).



Application

Time Dilation

2 Reference Frame of M

$$|\psi\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(|-x_M^{(1)}\rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} + |-x_M^{(2)}\rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} \right) |s(\tau_0)\rangle_{C_{int}}$$



$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} |-x_M^{(1)}\rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(1)})\rangle_{C_{int}} + e^{-\frac{i}{\hbar}\Phi^{(2)}} |-x_M^{(2)}\rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(2)})\rangle_{C_{int}} \right)$$

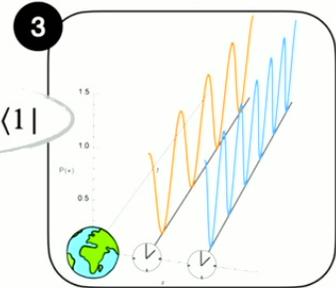
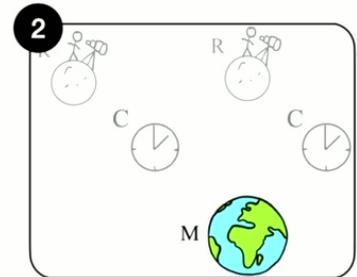
proper time

$$\tau^{(i)}(t) = t \left(1 + \frac{V(x_C - x_M^{(i)})}{c^2} \right)$$

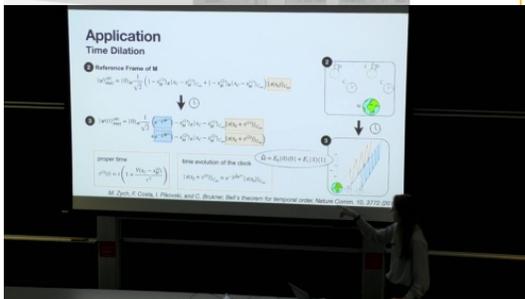
time evolution of the clock

$$|s(\tau_0 + \tau^{(i)})\rangle_{C_{int}} = e^{-\frac{i}{\hbar}\hat{\Omega}\tau^{(i)}} |s(\tau_0)\rangle_{C_{int}}$$

$$\hat{\Omega} = E_0|0\rangle\langle 0| + E_1|1\rangle\langle 1|$$



M. Zych, F. Costa, I. Pikovski, and C. Brukner, Bell's theorem for temporal order, Nature Comm. 10, 3772 (2019).



Application

Time Dilation

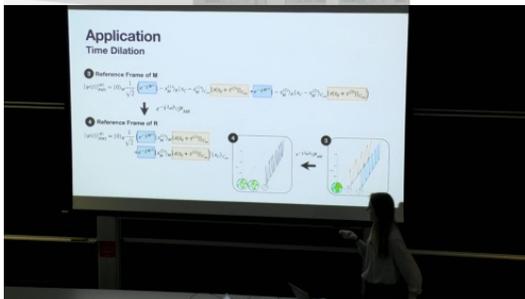
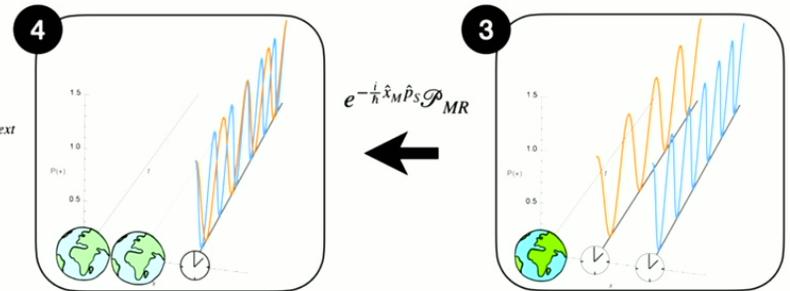
3 Reference Frame of M

$$|\psi(t)\rangle_{RMS}^{(M)} = |0\rangle_M \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} | -x_M^{(1)} \rangle_R |x_C - x_M^{(1)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(1)})\rangle_{C_{int}} + e^{-\frac{i}{\hbar}\Phi^{(2)}} | -x_M^{(2)} \rangle_R |x_C - x_M^{(2)}\rangle_{C_{ext}} |s(\tau_0 + \tau^{(2)})\rangle_{C_{int}} \right)$$

$$\downarrow e^{-\frac{i}{\hbar}\hat{x}_M \hat{p}_S \mathcal{P}_{MR}}$$

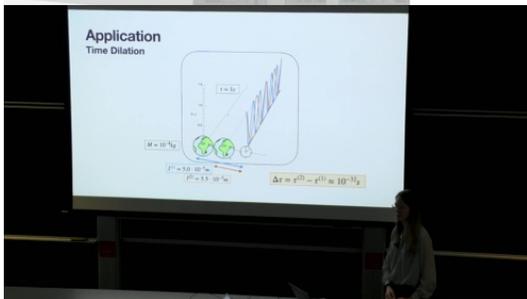
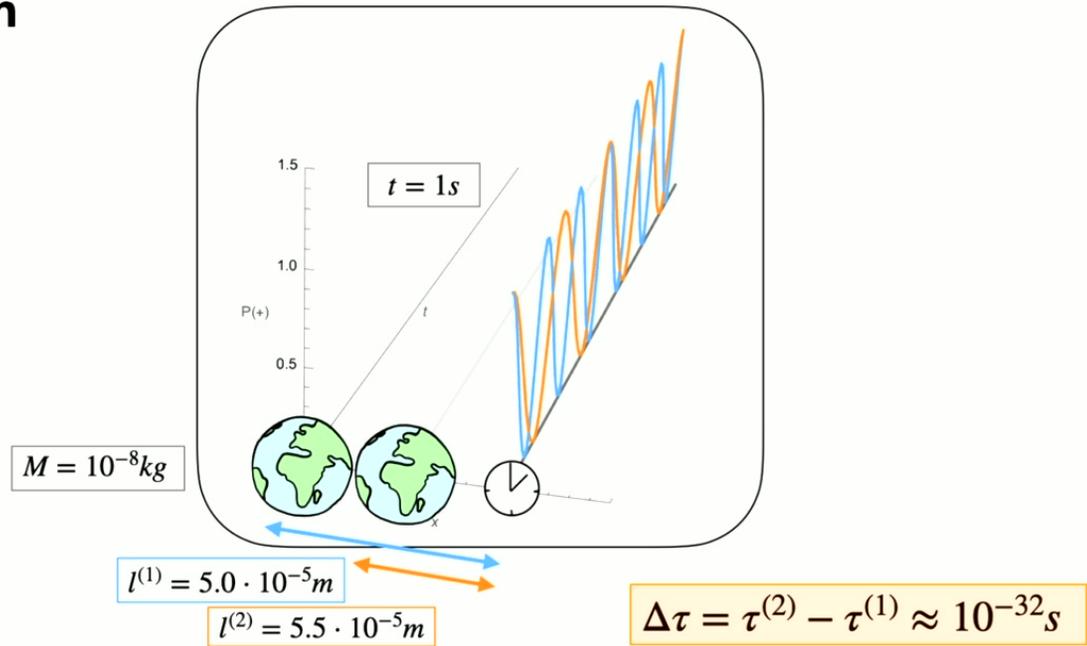
4 Reference Frame of R

$$|\psi(t)\rangle_{RMS}^{(R)} = |0\rangle_R \frac{1}{\sqrt{2}} \left(e^{-\frac{i}{\hbar}\Phi^{(1)}} |x_M^{(1)}\rangle_M |s(\tau_0 + \tau^{(1)})\rangle_{C_{int}} + e^{-\frac{i}{\hbar}\Phi^{(2)}} |x_M^{(2)}\rangle_M |s(\tau_0 + \tau^{(1)})\rangle_{C_{int}} \right) |x_C\rangle_{C_{ext}}$$



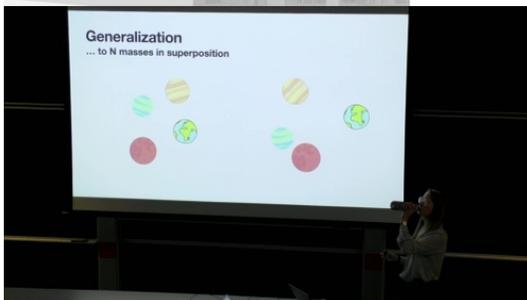
Application

Time Dilation



Generalization

... to N masses in superposition

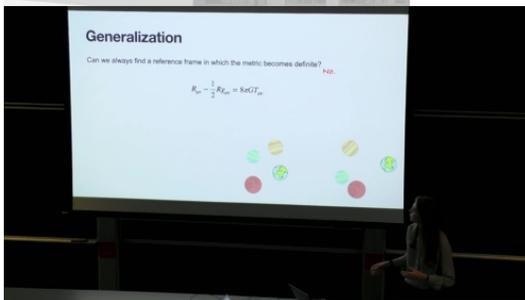


Generalization

Can we always find a reference frame in which the metric becomes definite?

No.

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi GT_{\mu\nu}$$

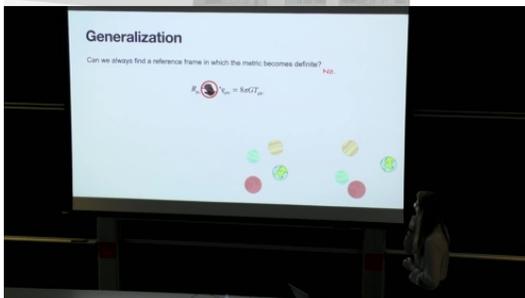


Generalization

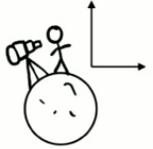
Can we always find a reference frame in which the metric becomes definite?

No.

$$R_{\mu\nu} \text{ (with a hand icon crossed out) } g_{\mu\nu} = 8\pi G T_{\mu\nu}$$



Generalization



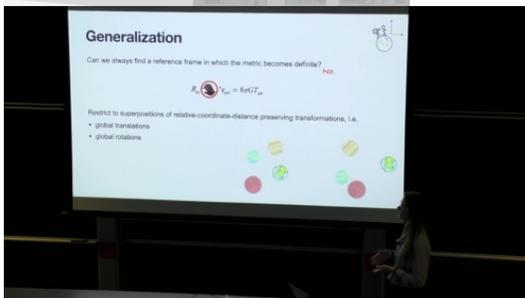
Can we always find a reference frame in which the metric becomes definite?

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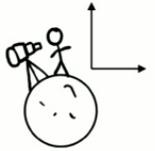
$$R_{\mu\nu} \text{ (with a hand icon crossed out) } g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

Restrict to superpositions of relative-coordinate-distance preserving transformations, i.e.

- global translations
- global rotations



Generalization

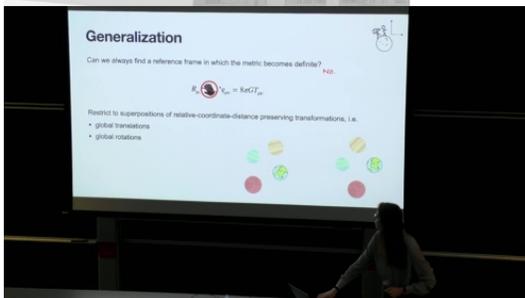


Can we always find a reference frame in which the metric becomes definite? **No.**

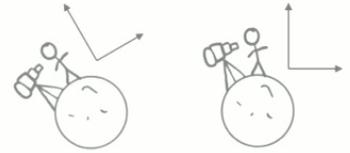
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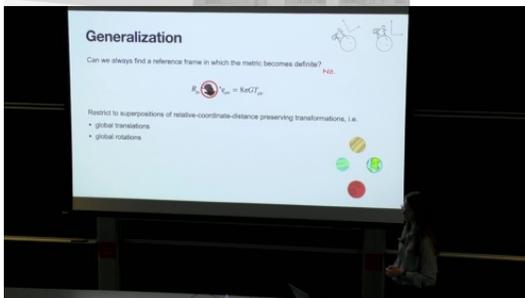
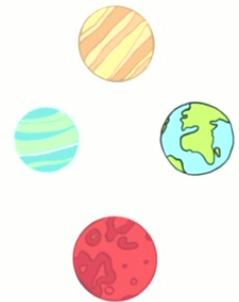
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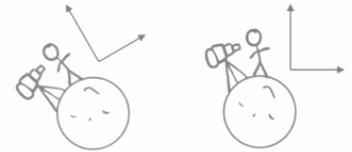
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Generalization



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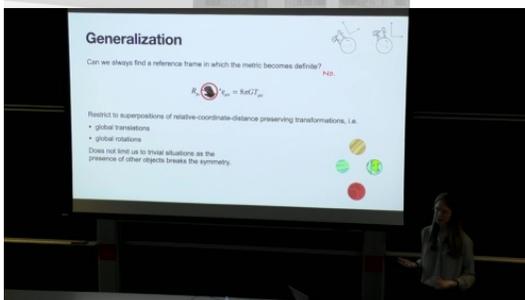
No.

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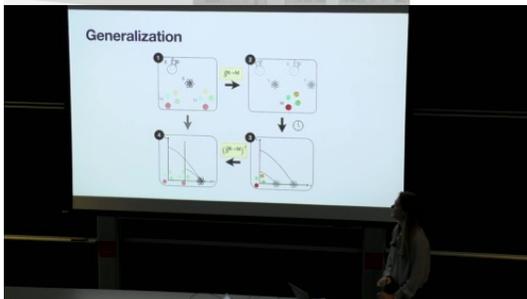
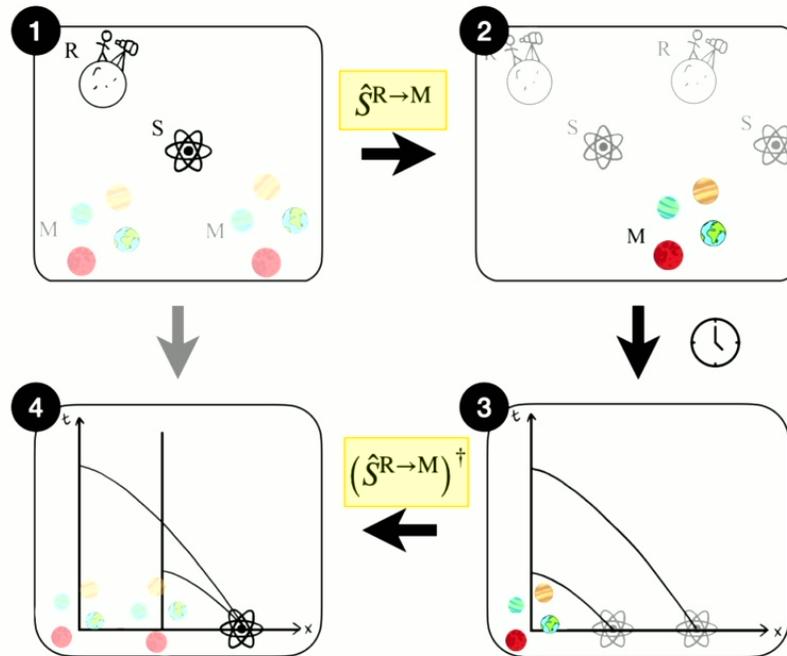
Restrict to superpositions of relative-coordinate-distance preserving transformations, i.e.

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- global rotations

Does not limit us to trivial situations as the presence of other objects breaks the symmetry.

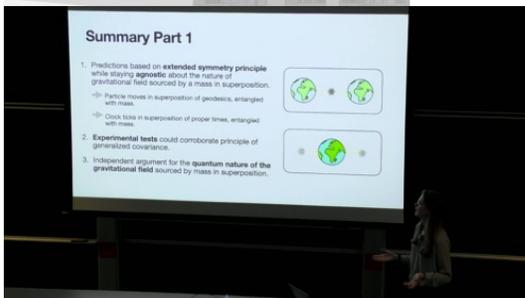
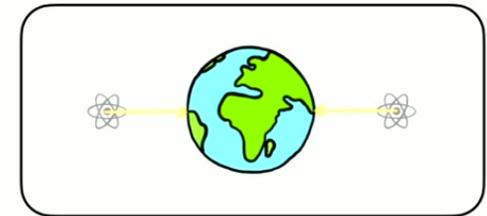
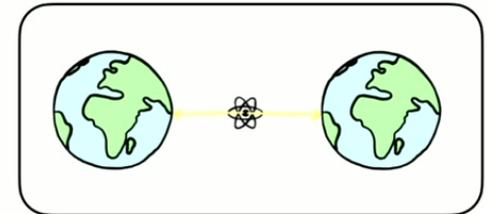


Generalization

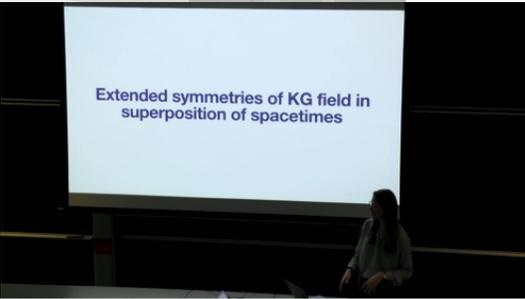


Summary Part 1

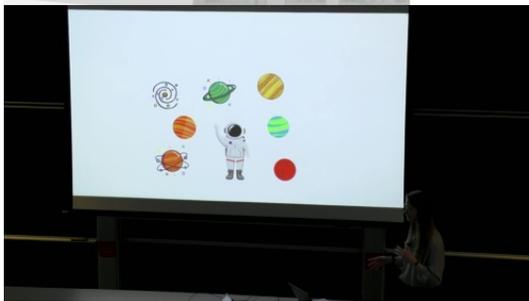
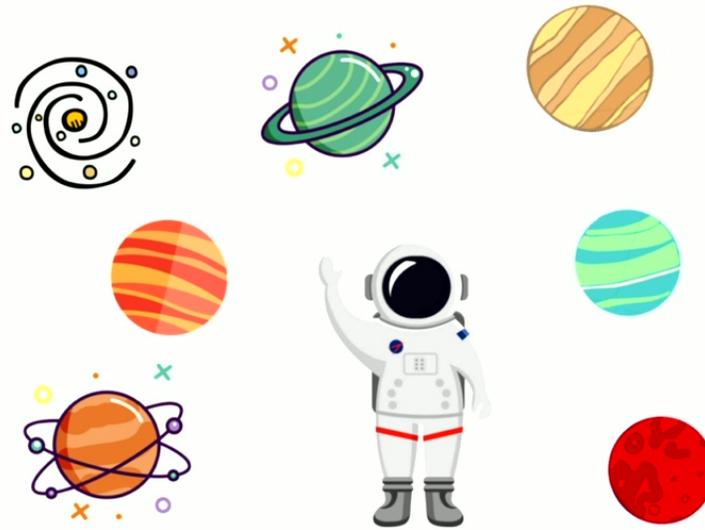
1. Predictions based on **extended symmetry principle** while staying **agnostic** about the nature of gravitational field sourced by a mass in superposition.
 - ➔ Particle moves in superposition of geodesics, entangled with mass.
 - ➔ Clock ticks in superposition of proper times, entangled with mass.
2. **Experimental tests** could corroborate principle of generalized covariance.
3. Independent argument for the **quantum nature of the gravitational field** sourced by mass in superposition.

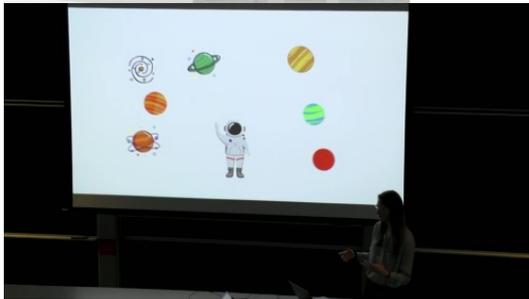
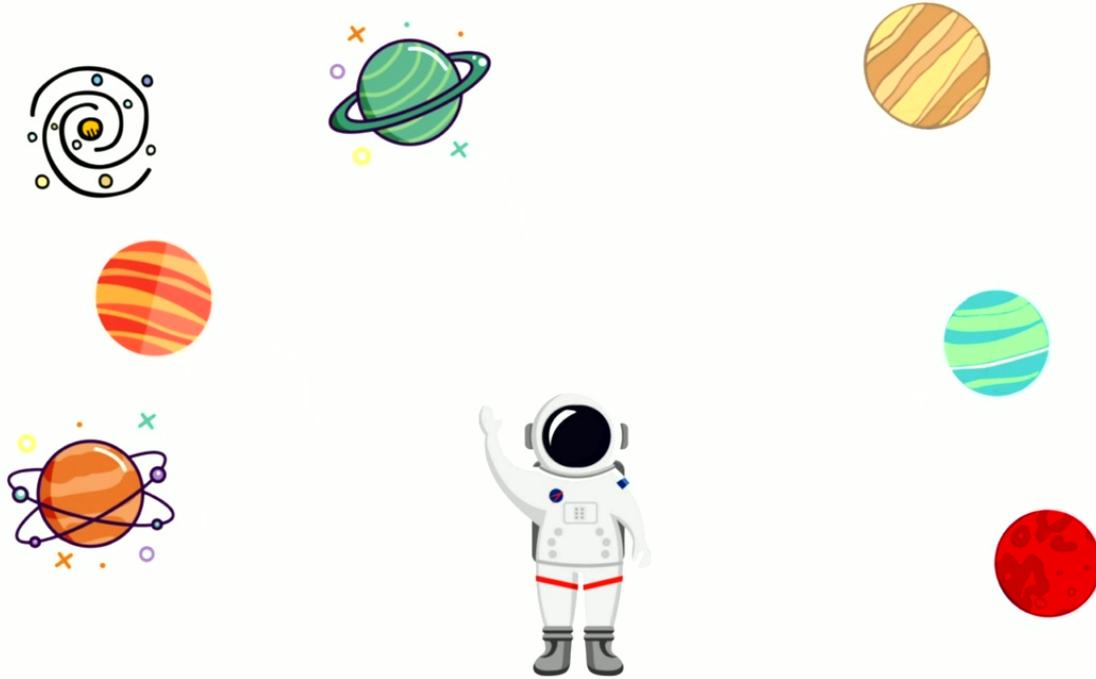
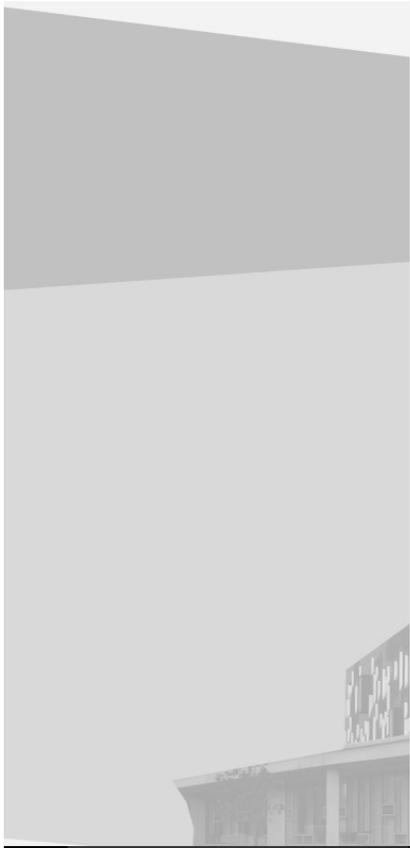


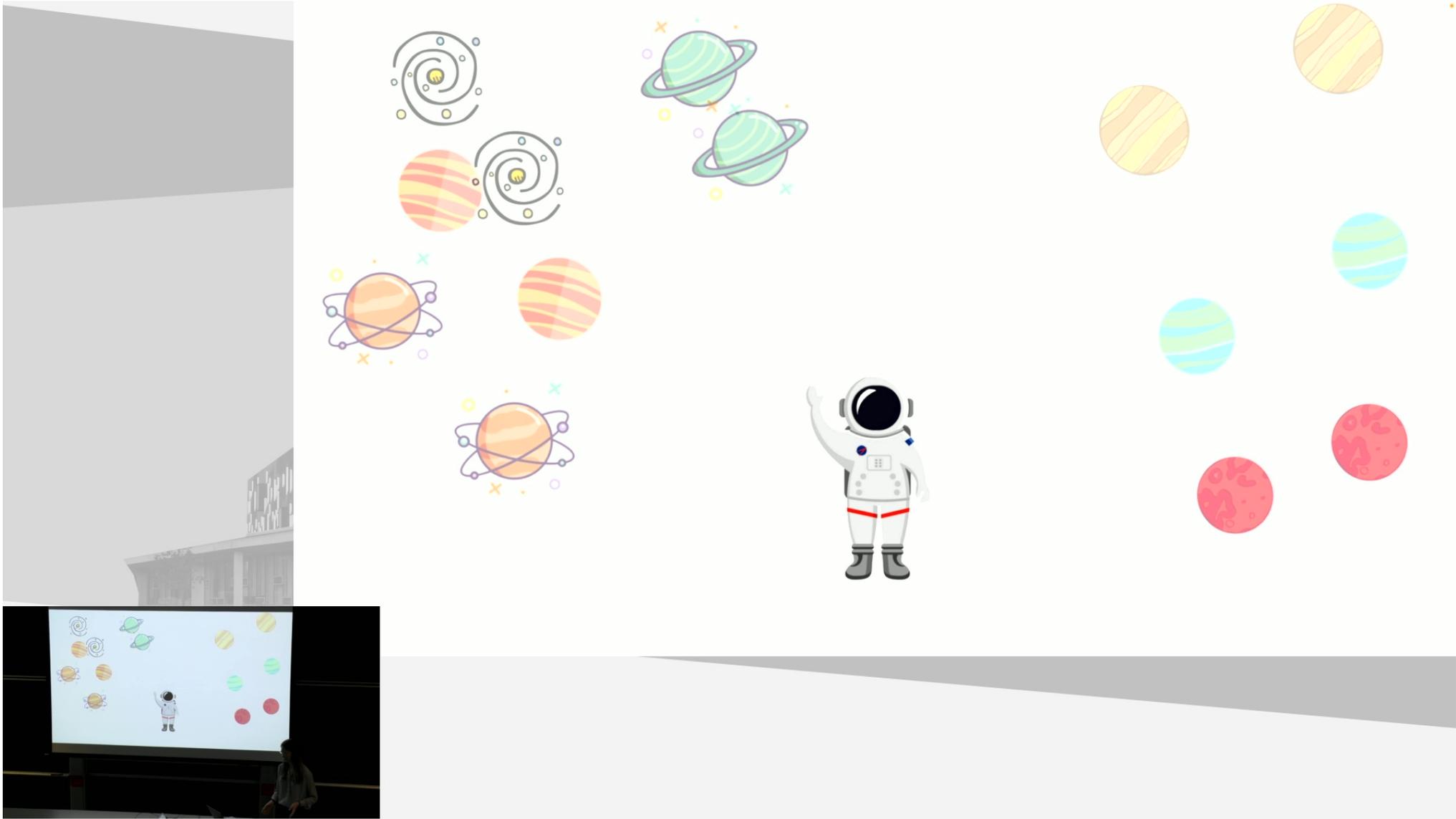
Extended symmetries of KG field in superposition of spacetimes



Extended symmetries of KG field in
superposition of spacetimes







Motivation

Build intuition and techniques to deal with “true” **quantum metric**, including **fuzziness of spacetime**



Motivation

Build intuition and techniques to deal with “true” **quantum metric**, including **fuzziness of spacetime**



So far:

- ▶ Superpositions of **diffeomorphic metrics**
- ▶ Restricted to superpositions of metrics related by **relative coordinate distance preserving transformations**
- ▶ Semiclassical states of **particles** in the vicinity of gravitational source

Now:

- ▶ Going **beyond** coordinate transformations (**diffeomorphisms**) to conformal transformations
- ▶ Staying in the **semi-classical** approximation for superposition state of metric
- ▶ **Quantum fields** inhabiting a spacetime in superposition

Similar route: Find **transformation** that maps situation to one with **definite** metric



Symmetries of the KG field...

... in an expanding universe

Take a spatially flat ($k = 0$) FLRW spacetime:

$$ds^2 = -dt^2 + a^2(t)\delta_{ij}dx^i dx^j$$

and the Klein-Gordon equation in curved spacetime

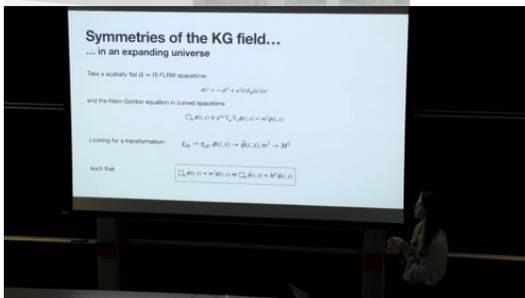
$$\square_g \phi(t, x) \equiv g^{\mu\nu} \nabla_\mu \nabla_\nu \phi(t, x) = m^2 \phi(t, x)$$

Looking for a transformation:

$$g_{ab} \rightarrow \eta_{ab}, \phi(t, x) \rightarrow \tilde{\phi}(t, x), m^2 \rightarrow M^2$$

such that

$$\square_g \phi(t, x) = m^2 \phi(t, x) \Leftrightarrow \square_\eta \tilde{\phi}(t, x) = M^2 \tilde{\phi}(t, x)$$



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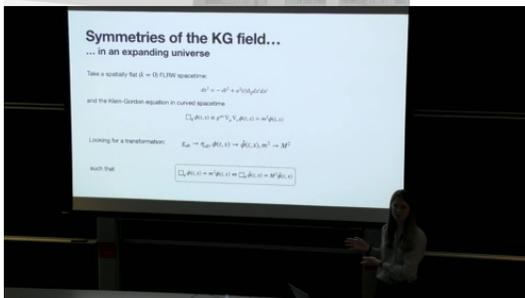
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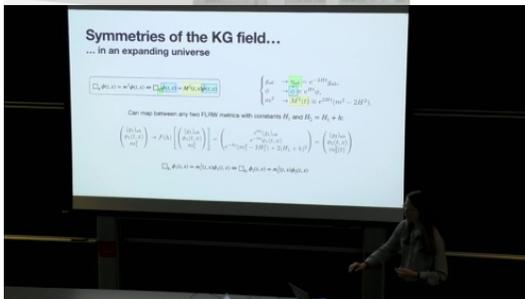
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$$\begin{cases} g_{ab} & \rightarrow \eta_{ab} = e^{-2Ht} g_{ab}, \\ \phi & \rightarrow \tilde{\phi} \equiv e^{Ht} \phi, \\ m^2 & \rightarrow M^2(t) \equiv e^{2Ht} (m^2 - 2H^2). \end{cases}$$

Can map between any two FLRW metrics with constants H_1 and $H_2 = H_1 + h$:

$$\begin{pmatrix} (g_1)_{ab} \\ \phi_1(t, x) \\ m_1^2 \end{pmatrix} \rightarrow \mathcal{F}(h) \left[\begin{pmatrix} (g_1)_{ab} \\ \phi_1(t, x) \\ m_1^2 \end{pmatrix} \right] = \begin{pmatrix} e^{ht} (g_1)_{ab} \\ e^{-ht} \phi_1(t, x) \\ e^{-ht} (m_1^2 - 2H_1^2) + 2(H_1 + h)^2 \end{pmatrix} = \begin{pmatrix} (g_2)_{ab} \\ \phi_2(t, x) \\ m_2^2(t) \end{pmatrix}$$

$$\square_{g_1} \phi_1(t, x) = m_1^2(t, x) \phi_1(t, x) \Leftrightarrow \square_{g_2} \phi_2(t, x) = m_2^2(t, x) \phi_2(t, x)$$



Symmetries of the KG field...

... in an expanding universe

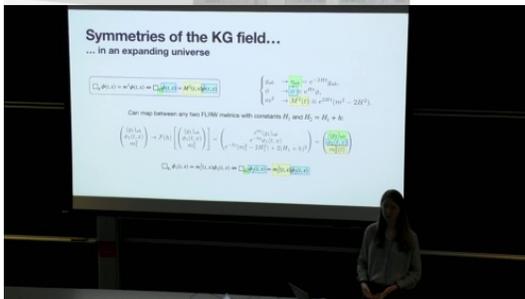
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Symmetries of the KG field...

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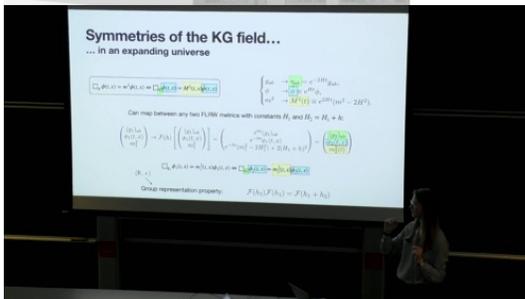
$$\begin{pmatrix} (g_1)_{ab} \\ \phi_1(t, x) \\ m_1^2 \end{pmatrix} \rightarrow \mathcal{F}(h) \left[\begin{pmatrix} (g_1)_{ab} \\ \phi_1(t, x) \\ m_1^2 \end{pmatrix} \right] = \begin{pmatrix} e^{ht} (g_1)_{ab} \\ e^{-ht} \phi_1(t, x) \\ e^{-ht} (m_1^2 - 2H_1^2) + 2(H_1 + h)^2 \end{pmatrix} = \begin{pmatrix} (g_2)_{ab} \\ \phi_2(t, x) \\ m_2^2(t) \end{pmatrix}$$

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$(\mathbb{R}, +)$

Group representation property:

$$\mathcal{F}(h_2) \mathcal{F}(h_1) = \mathcal{F}(h_1 + h_2)$$



Symmetries of the KG field...

... for conformally related spacetimes

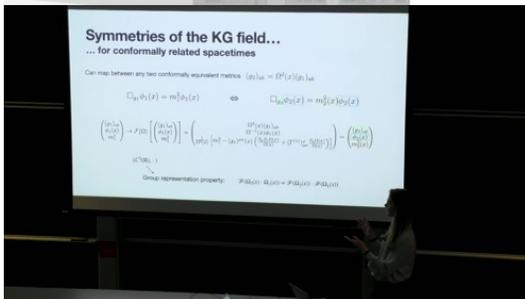
Can map between any two conformally equivalent metrics $(g_2)_{ab} = \Omega^2(x)(g_1)_{ab}$

$$\square_{g_1} \phi_1(x) = m_1^2 \phi_1(x) \quad \Leftrightarrow \quad \square_{g_2} \phi_2(x) = m_2^2(x) \phi_2(x)$$

$$\begin{pmatrix} (g_1)_{ab} \\ \phi_1(x) \\ m_1^2 \end{pmatrix} \rightarrow \mathcal{F}(\Omega) \left[\begin{pmatrix} (g_1)_{ab} \\ \phi_1(x) \\ m_1^2 \end{pmatrix} \right] = \begin{pmatrix} \Omega^2(x)(g_1)_{ab} \\ \Omega^{-1}(x)\phi_1(x) \\ \frac{1}{\Omega^2(x)} \left[m_1^2 - (g_1)^{\mu\nu}(x) \left(\frac{\partial_\mu \partial_\nu \Omega(x)}{\Omega(x)} + (\Gamma^{(1)})^\rho_{\mu\nu} \frac{\partial_\rho \Omega(x)}{\Omega(x)} \right) \right] \end{pmatrix} = \begin{pmatrix} (g_2)_{ab} \\ \phi_2(x) \\ m_2^2(x) \end{pmatrix}$$

$(C^2(\mathbb{R}), \cdot)$

Group representation property: $\mathcal{F}(\Omega_2(x) \cdot \Omega_1(x)) = \mathcal{F}(\Omega_2(x)) \cdot \mathcal{F}(\Omega_1(x))$



quantized

Symmetries of the KG field...

... in a conformally flat universe

Curved spacetime



inner product: $(\phi_1, \phi_2)_g$

$$\phi(x) = \int d^3k \left(\overbrace{(u_k, \phi)_g}^{a_k} u_k + \overbrace{(u_k^*, \phi)_g}^{a_k^*} u_k^* \right)$$

$$\mathcal{Q} \downarrow \begin{array}{l} a_k \rightarrow \hat{a}_k \\ a_k^* \rightarrow \hat{a}_k^\dagger \end{array}$$

$$\hat{\phi}(x) = \int d^3k \hat{a}_k u_k + \hat{a}_k^\dagger u_k^*$$

vacuum state: $\hat{a}_k |0\rangle_g = 0$

Minkowski spacetime

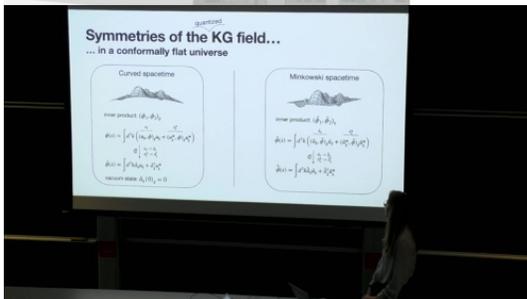


inner product: $(\tilde{\phi}_1, \tilde{\phi}_2)_\eta$

$$\tilde{\phi}(x) = \int d^3k \left(\overbrace{(\tilde{u}_k, \tilde{\phi})_\eta}^{\tilde{a}_k} \tilde{u}_k + \overbrace{(\tilde{u}_k^*, \tilde{\phi})_\eta}^{\tilde{a}_k^*} \tilde{u}_k^* \right)$$

$$\mathcal{Q} \downarrow \begin{array}{l} \tilde{a}_k \rightarrow \hat{\tilde{a}}_k \\ \tilde{a}_k^* \rightarrow \hat{\tilde{a}}_k^\dagger \end{array}$$

$$\hat{\tilde{\phi}}(x) = \int d^3k \hat{\tilde{a}}_k \tilde{u}_k + \hat{\tilde{a}}_k^\dagger \tilde{u}_k^*$$

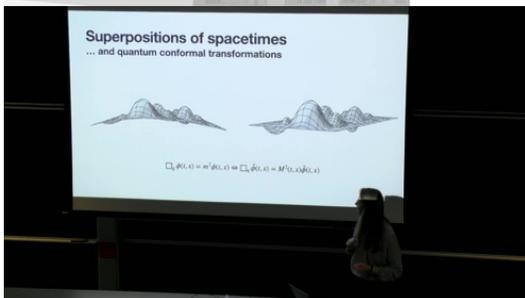


Superpositions of spacetimes

... and quantum conformal transformations

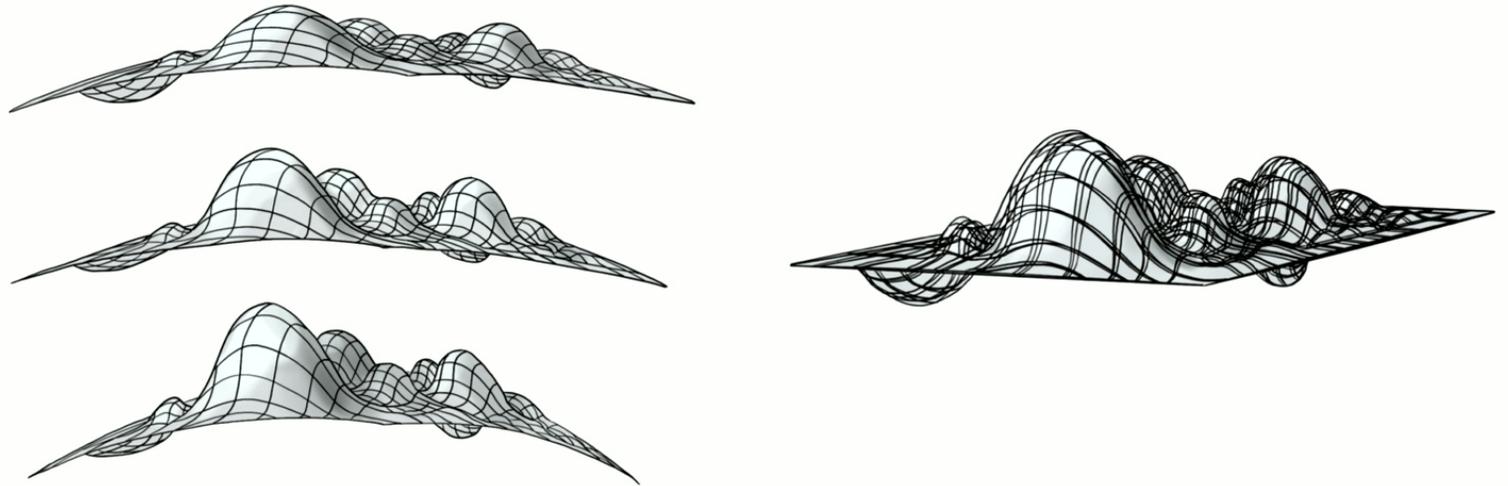


$$\square_g \phi(t, x) = m^2 \phi(t, x) \Leftrightarrow \square_\eta \tilde{\phi}(t, x) = M^2(t, x) \tilde{\phi}(t, x)$$



Superpositions of spacetimes

... and quantum conformal transformations



$$|\Psi\rangle^{(m)} = \left(\sum_i \alpha_i |g_i\rangle \otimes |\phi^{(g, m^2)}\rangle \right) \otimes |m^2\rangle \Leftrightarrow |\Psi\rangle^{(g)} = |g\rangle \otimes \left(\sum_i \alpha_i |\phi^{(g, m_i^2)}\rangle \otimes |m_i^2\rangle \right)$$

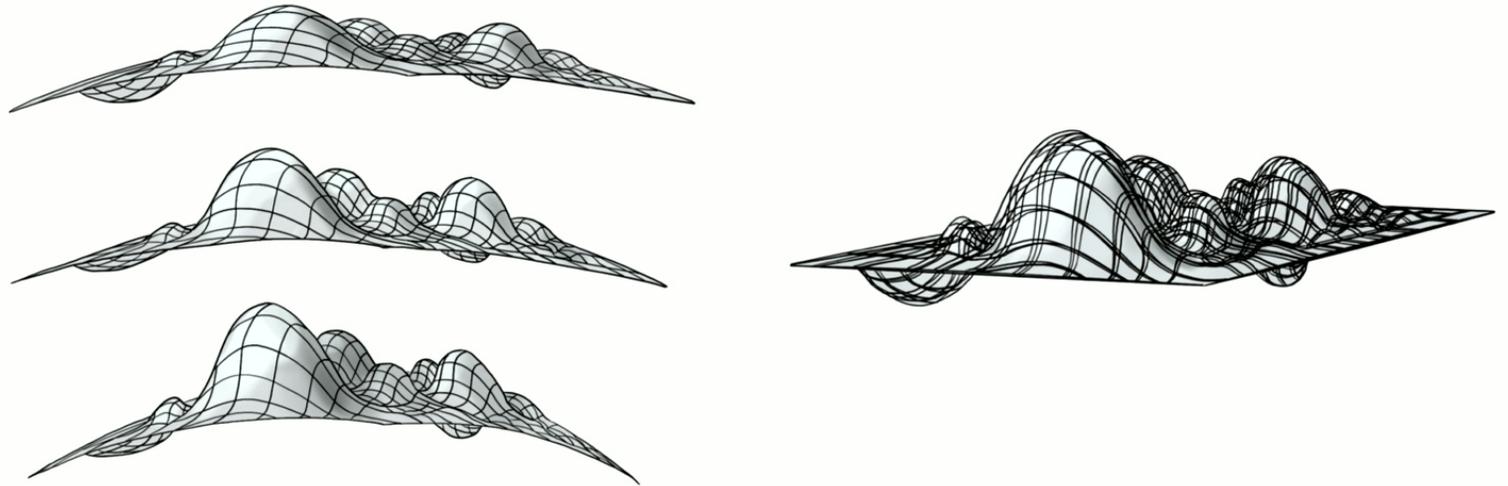
Superpositions of spacetimes
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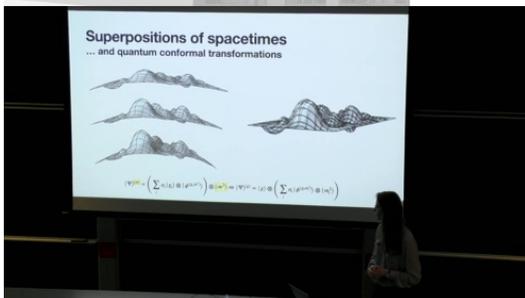
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Superpositions of spacetimes

... and quantum conformal transformations

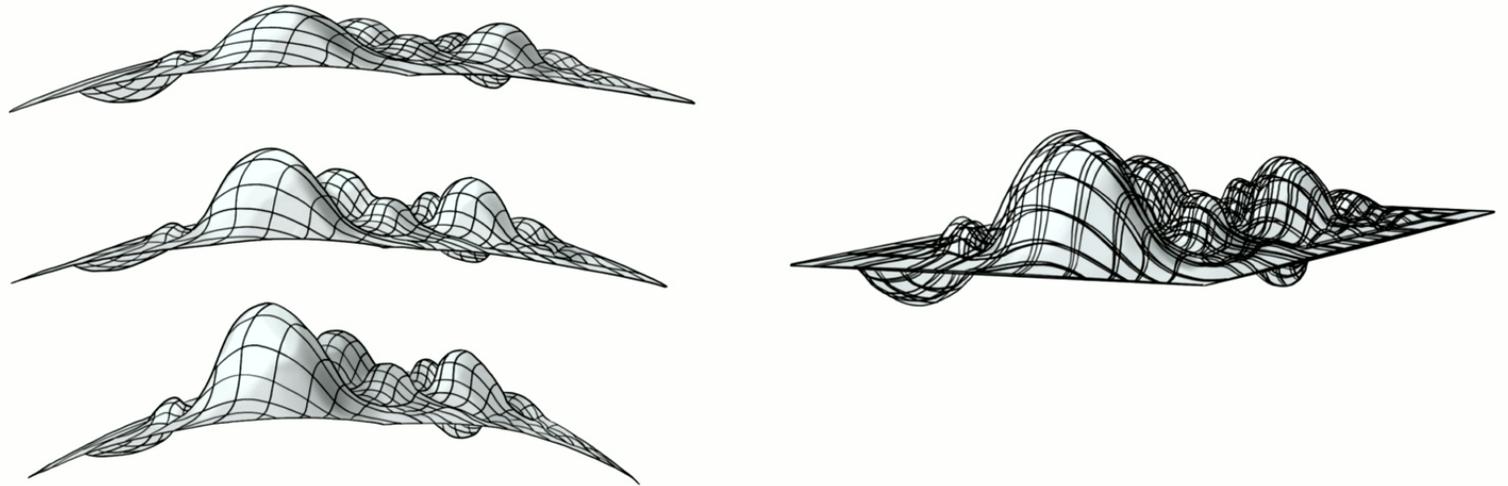


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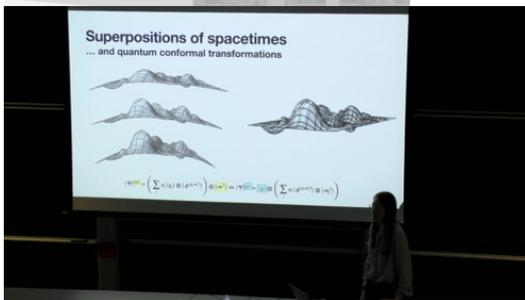


Superpositions of spacetimes

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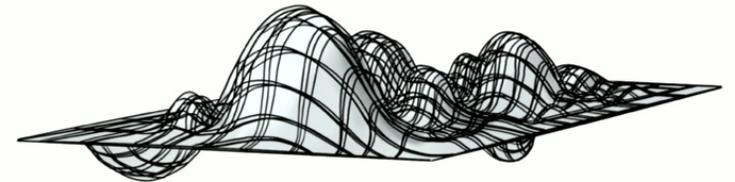
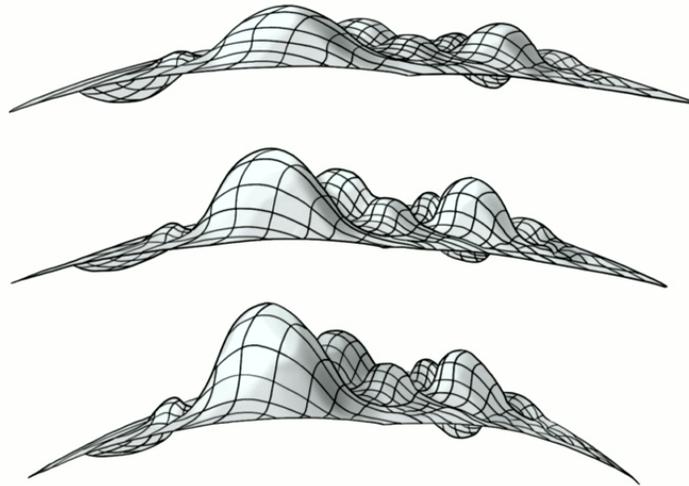


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Superpositions of spacetimes

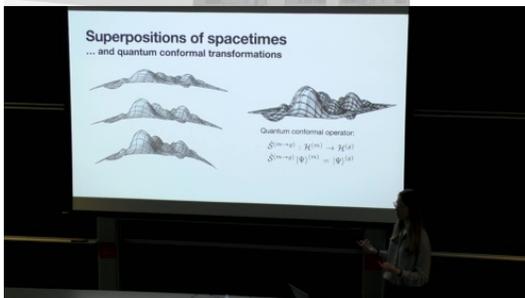
... and quantum conformal transformations



Quantum conformal operator:

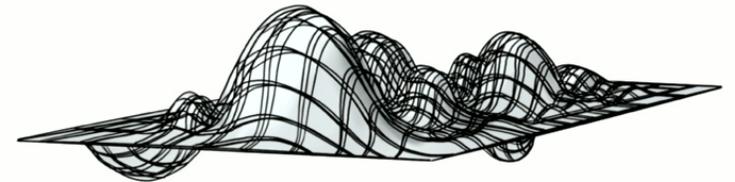
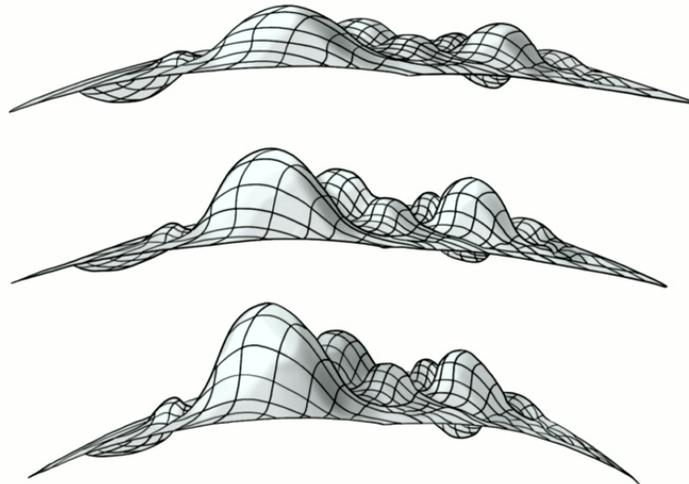
$$\hat{S}^{(m \rightarrow g)} : \mathcal{H}^{(m)} \rightarrow \mathcal{H}^{(g)}$$

$$\hat{S}^{(m \rightarrow g)} |\Psi\rangle^{(m)} = |\Psi\rangle^{(g)}$$



Superpositions of spacetimes

... and quantum conformal transformations

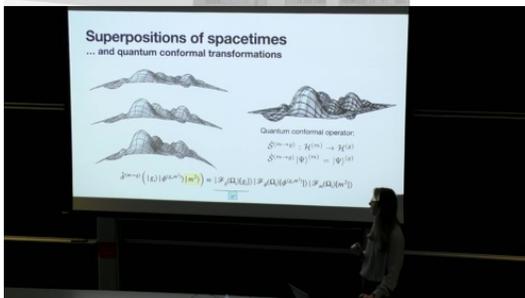


Quantum conformal operator:

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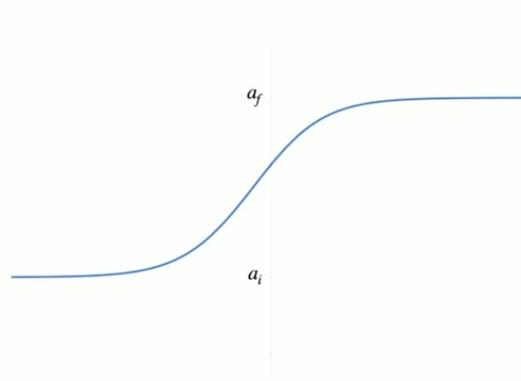
$$\hat{\mathcal{S}}^{(m \rightarrow g)} |\Psi\rangle^{(m)} = |\Psi\rangle^{(g)}$$

$$\hat{\mathcal{S}}^{(m \rightarrow g)} \left(|g_i\rangle |\phi^{(g_i, m^2)}\rangle |m^2\rangle \right) = \underbrace{|\mathcal{F}_g(\Omega_i)[g_i]\rangle |\mathcal{F}_\phi(\Omega_i)[\phi^{(g_i, m^2)}]\rangle}_{|g\rangle} |\mathcal{F}_m(\Omega_i)[m^2]\rangle$$



Cosmological particle production

... in modified Minkowski?



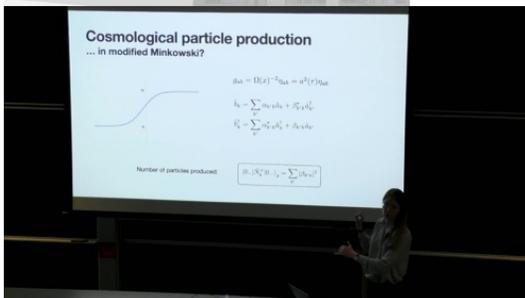
$$g_{ab} = \Omega(x)^{-2} \eta_{ab} = a^2(\tau) \eta_{ab}$$

$$\hat{b}_k = \sum_{k'} \alpha_{k'k} \hat{a}_k + \beta_{k'k}^* \hat{a}_{k'}^\dagger$$

$$\hat{b}_k^\dagger = \sum_{k'} \alpha_{k'k}^* \hat{a}_k^\dagger + \beta_{k'k} \hat{a}_{k'}$$

Number of particles produced:

$$\langle 0_- | \hat{N}_k^+ | 0_- \rangle_g = \sum_{k'} |\beta_{k'k}|^2$$

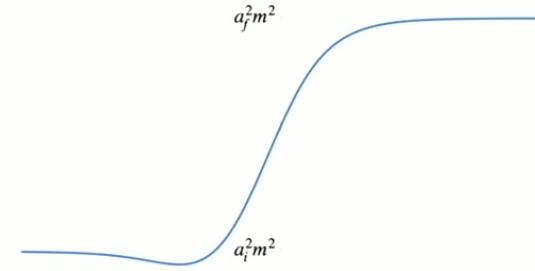
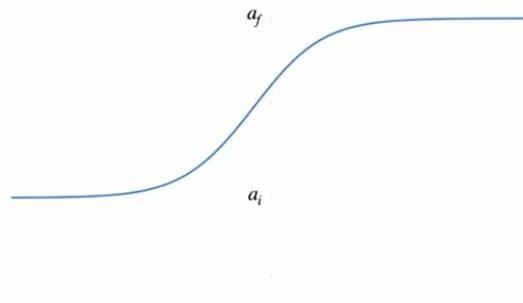


Cosmological particle production

... in modified Minkowski?

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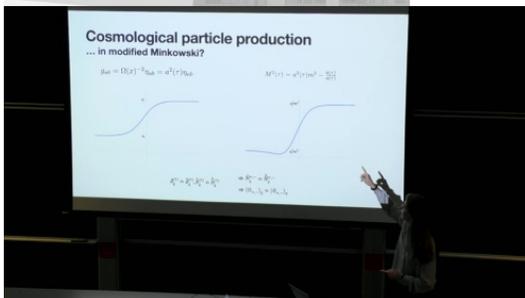
$$M^2(\tau) = a^2(\tau) m^2 - \frac{\ddot{a}(\tau)}{a(\tau)}$$



$$\hat{a}_k^{(+)} = \hat{\tilde{a}}_k^{(+)}, \hat{b}_k^{(+)} = \hat{\tilde{b}}_k^{(+)}$$

$$\Rightarrow \hat{N}_k^{+,-} = \hat{\tilde{N}}_k^{+,-}$$

$$\Rightarrow |0_{+,-}\rangle_g = |0_{+,-}\rangle_\eta$$



Cosmological particle production

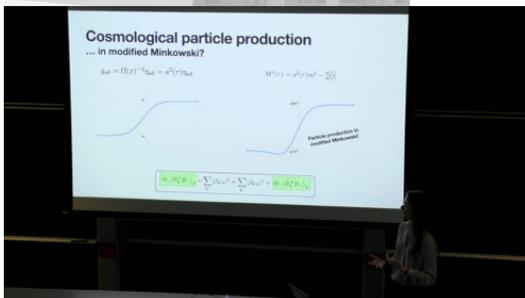
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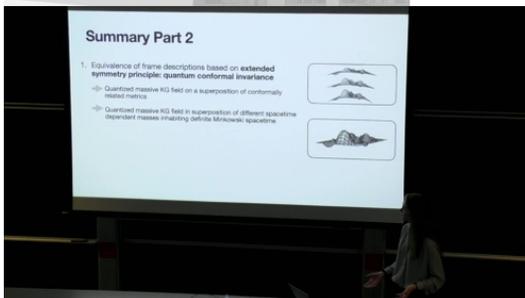
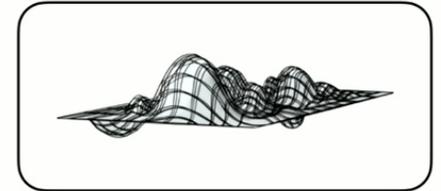
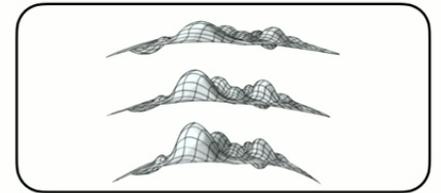
$$\langle 0_- | \hat{N}_k^+ | 0_- \rangle_g = \sum_{k'} |\beta_{k'k}|^2 = \sum_{k'} |\tilde{\beta}_{k'k}|^2 = \langle 0_- | \hat{N}_k^+ | 0_- \rangle_\eta$$



Summary Part 2

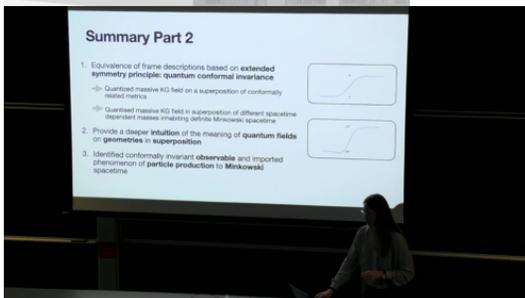
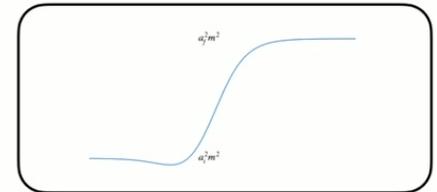
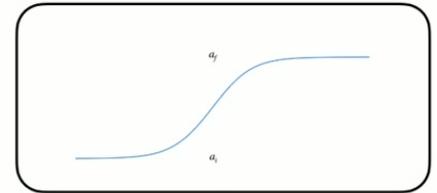
1. Equivalence of frame descriptions based on **extended symmetry principle: quantum conformal invariance**

- ➔ Quantized massive KG field on a superposition of conformally related metrics
- ➔ Quantized massive KG field in superposition of different spacetime dependent masses inhabiting definite Minkowski spacetime



Summary Part 2

1. Equivalence of frame descriptions based on **extended symmetry principle: quantum conformal invariance**
 - ➔ Quantized massive KG field on a superposition of conformally related metrics
 - ➔ Quantised massive KG field in superposition of different spacetime dependent masses inhabiting definite Minkowski spacetime
2. Provide a deeper **intuition** of the meaning of **quantum fields on geometries in superposition**
3. Identified conformally invariant **observable** and imported phenomenon of **particle production to Minkowski spacetime**



Extended symmetry principles

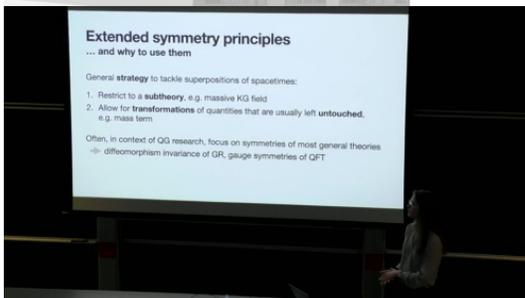
... and why to use them

General **strategy** to tackle superpositions of spacetimes:

1. Restrict to a **subtheory**, e.g. massive KG field
2. Allow for **transformations** of quantities that are usually left **untouched**, e.g. mass term

Often, in context of QG research, focus on symmetries of most general theories

➔ diffeomorphism invariance of GR, gauge symmetries of QFT



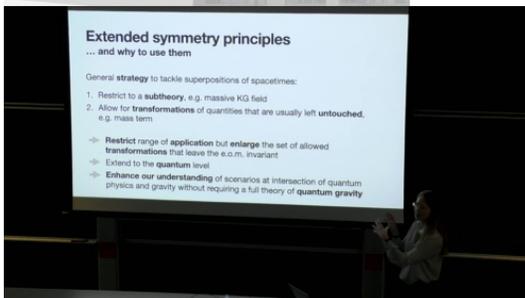
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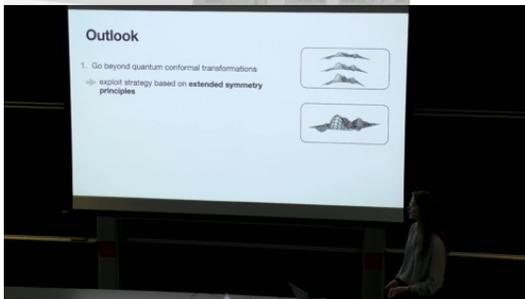
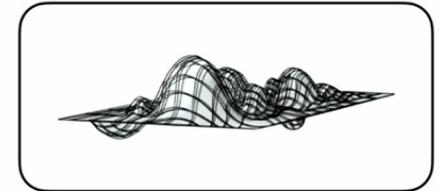
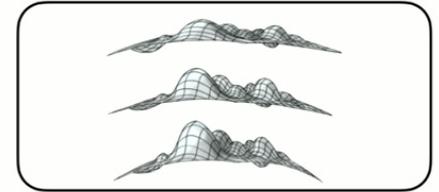
1. Restrict to a **subtheory**, e.g. massive KG field
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- ➔ **Restrict** range of **application** but **enlarge** the set of allowed **transformations** that leave the e.o.m. invariant
- ➔ Extend to the **quantum** level
- ➔ **Enhance our understanding** of scenarios at intersection of quantum physics and gravity without requiring a full theory of **quantum gravity**



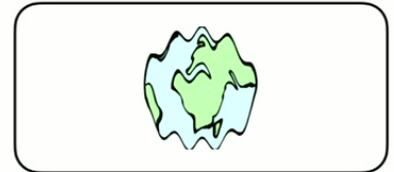
Outlook

1. Go beyond quantum conformal transformations
 - exploit strategy based on **extended symmetry principles**



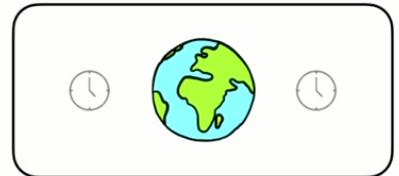
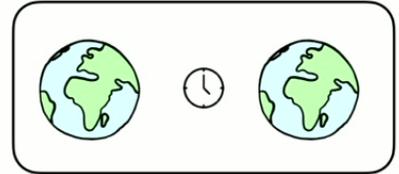
Outlook

1. Go beyond quantum conformal transformations
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2. Take into account **quantum fuzziness** of spacetime
 - ➔ Go **beyond semi-classical approximation** (for massive objects and states of metric)



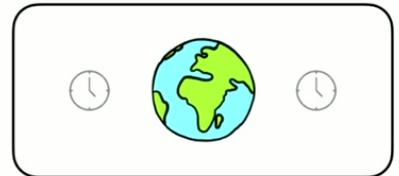
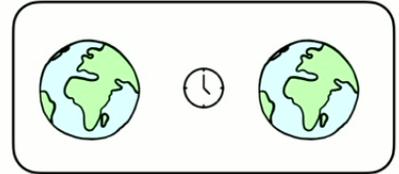
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Thank you!

For more details, see [arXiv:2112.11473](https://arxiv.org/abs/2112.11473) + work to appear soon

