

Title: Twistor action for GR - Atul Sharma, University of Oxford

Speakers:

Series: Quantum Gravity

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Abstract: It has been a long-standing dream of twistor theorists to understand gravity without ever talking about gravitons in space-time. To this end, I will describe the recent discovery of a twistor action formulation of perturbative general relativity. This takes the form of a theory governing complex structure deformations on twistor space. It reduces to Plebanski's formulation of GR on performing the Penrose transform to space-time. Some promising applications include finding on-shell recursion relations like MHV rules for graviton scattering amplitudes, studying the quantum integrability of self-dual GR, etc.

Zoom Link: <https://pitp.zoom.us/j/99235001602?pwd=QVN6b2ZPbTM2SkFWNkxYTEhzd0tsdz09>



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Twistor action for GR

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Based on [Sharma '21], [Adamo, Mason, Sharma '21].
Ongoing work with David Skinner, Roland Bittleston.

MHV amplitudes

- Gluons

$$\frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \cdots \langle n1 \rangle}$$

[Parke, Taylor '86]

- Gravitons

$$\frac{\langle 12 \rangle^8}{\langle 12 \rangle^2 \langle 2r \rangle^2 \langle r1 \rangle^2} \det(\mathbb{H}_{12r}^{12r})$$

[Hodges '12]

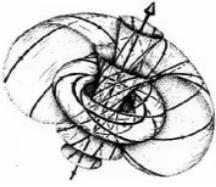
$$\mathbb{H}_{ij} = \begin{cases} \frac{[ij]}{\langle ij \rangle}, & i \neq j \\ -\sum_{k \neq i} \frac{[ik] \langle 1k \rangle \langle 2k \rangle}{\langle ik \rangle \langle 1i \rangle \langle 2i \rangle}, & i = j \end{cases}$$

- With great simplicity comes great duality.



Twistor strings and actions

Twistor actions provide target space descriptions of twistor strings.

	Gauge theory + Conformal gravity	General relativity
Twistor strings	Witten, Berkovits, Roiban, Spradlin, Volovich,...	Skinner, Cachazo, Mason, Geyer,...
Twistor actions	Mason, Skinner, Boels, Bullimore, Adamo, ..., Costello, Bittleston, ...	Today!

Constructive origins of Parke-Taylor formula



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Summary

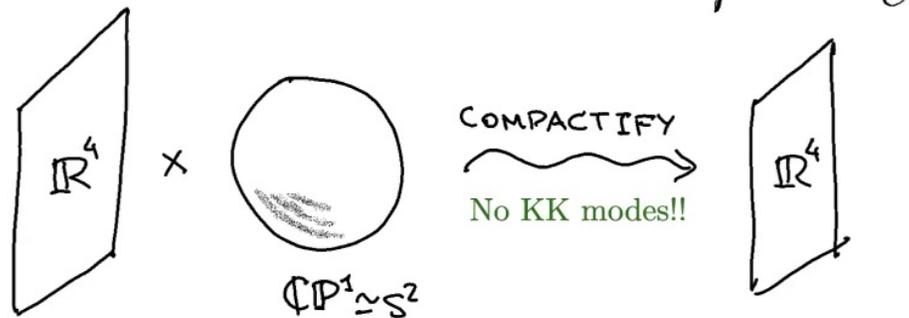
- Motivations and connections
- Twistors for curved space-time
- Twistor action for gravity
- Graviton scattering



Motivations

Weak-weak duality and MHV formalism

$$\mathbb{P}^1 \simeq \mathbb{R}^4 \times \mathbb{P}^1 \simeq \mathcal{O}_{\mathbb{P}^1}(1) \oplus 2$$



$$\int_{\mathbb{P}^1} b \bar{\partial} a + \sum_{n=2}^{\infty} \int_{\mathbb{R}^4} \int_{(\mathbb{P}^1)^n} b^2 a^{n-2} \longrightarrow \int F \wedge *F - \frac{1}{4} F \wedge F \quad [\text{Mason '05}]$$

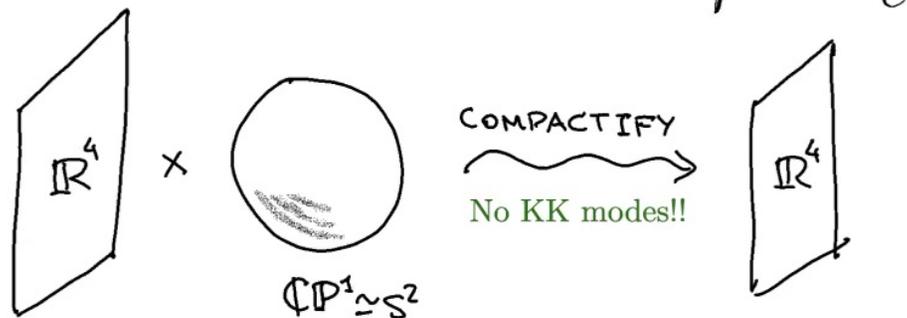


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**CSW recursion/
MHV rules**

- $b^2 a^{n-2}$ interaction \longrightarrow n -point MHV vertex
- $\bar{\partial}^{-1}$ \longrightarrow MHV propagator [Cachazo, Svrcek, Witten '04]
[Boels, Mason, Skinner '07]



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Broader connections

Integrable systems and celestial holography

- Holomorphic Chern-Simons twistor actions for self-dual Yang-Mills.

[Bittleston, Skinner '20] [Costello]

$$\int_{\mathbb{PT}} \frac{\Omega}{\langle \lambda 1 \rangle^2 \langle \lambda 2 \rangle^2} \wedge \text{tr} \left(a \wedge \bar{\partial} a + \frac{2}{3} a \wedge a \wedge a \right)$$



Exactly reduces to SD YM action on compactification.
Related to heterotic $N = 2$ strings.

- Costello applied *twisted holography* to relate twistor actions to *celestial holography*.

[Costello, Strings '21]



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Twistor space for \mathbb{R}^4

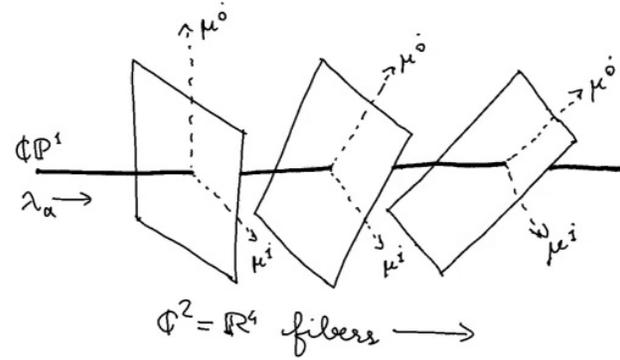
[Woodhouse '85]

Twistor space

$$\mathbb{P}\mathbb{T} = \mathbb{P}^3 - \mathbb{P}^1 \xrightarrow{\text{diffeo}} \mathbb{R}^4 \times \mathbb{P}^1$$

\swarrow (blue arrow) \downarrow (green arrow) \searrow (orange arrow)

$$Z^A = (\mu^{\dot{\alpha}}, \lambda_{\alpha}) \quad \{\lambda_{\alpha} = 0\} \quad (x^{\alpha\dot{\alpha}}, \lambda_{\alpha})$$

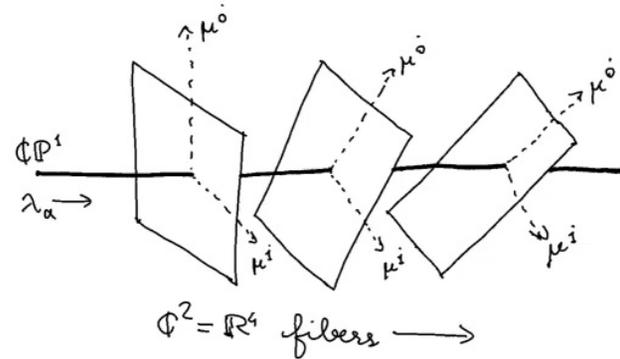
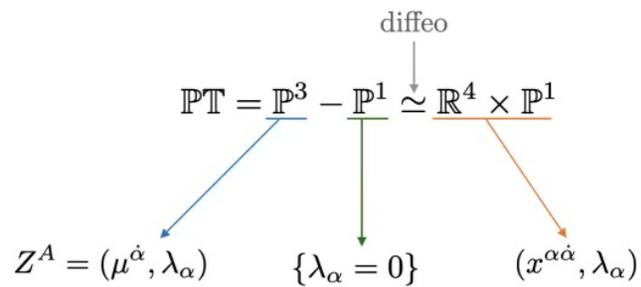


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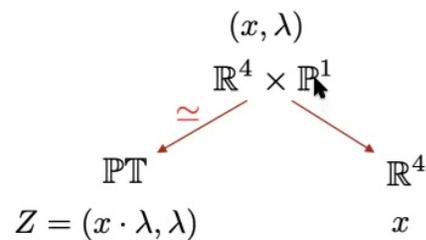
Twistor space for \mathbb{R}^4

[Woodhouse '85]

Twistor space



Double fibration



Incidence relations/ twistor lines

$$x \longleftrightarrow X \simeq \mathbb{P}^1 \subset \text{PT} : \quad \mu^{\dot{\alpha}} = x^{\alpha\dot{\alpha}} \lambda_{\alpha}$$



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Correspondences

Twistors	Space-time
$\mathbb{P}\mathcal{T}$ “Curved” twistor space	M Curved space-time
$X \simeq \mathbb{P}^1 \hookrightarrow \mathbb{P}\mathcal{T}$ Deformed incidence relations	$x^{\alpha\dot{\alpha}}$ Points of space-time
h Complex structure deformation: dynamical field in twistor action	$g = \epsilon_{\alpha\beta} \epsilon_{\dot{\alpha}\dot{\beta}} e^{\alpha\dot{\alpha}} e^{\beta\dot{\beta}}$ Metric and tetrad: dynamical field in GR action
B Auxiliary field in twistor action	$\Gamma_{\alpha\beta}$ ASD graviton: auxiliary field in GR <u>action</u>

We use a 1st order action for GR due to *Palatini* and *Plebanski*



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Almost complex structures



Easiest to work in homogeneous coordinates: $Z^A = (\mu^\alpha, \lambda_\alpha) \in \mathbb{T} = \mathbb{C}^4 - \mathbb{C}^2$

**Basis of complex
1-forms**

$$dZ^A \in T^{*1,0}\mathbb{T}, \quad d\bar{Z}^{\bar{A}} \in T^{*0,1}\mathbb{T}$$

**Deformation of
complex structure**

$$\mathbb{P}\mathbb{T} \longrightarrow \mathbb{P}\mathcal{T}$$

Deformed twistor space

$$dZ^A \mapsto \theta^A = dZ^A - \underbrace{V^A}$$

$$V^A \equiv d\bar{Z}^{\bar{B}} V_{\bar{B}}^A$$

Beltrami differential

**Nijenhuis
tensor**

$$N^A = \bar{\partial}V^A + V^B \wedge \partial_B V^A$$

*Newlander, Nirenberg,
Kodaira, Spencer,...*

$$\bar{\partial} \equiv d\bar{Z}^{\bar{A}} \frac{\partial}{\partial \bar{Z}^{\bar{A}}}$$

$$\partial_B \equiv \mathcal{L}_{\partial/\partial Z^B}$$

$N^A = 0$ yields
integrable complex
structure deformations

Hamiltonian deformations

Hamiltonian complex
structure deformation

$$V^A = I^{AB} \partial_B h, \quad h \equiv h_{\bar{A}} d\bar{Z}^{\bar{A}}$$

Infinity twistor

$$I^{AB} = \begin{pmatrix} \epsilon^{\dot{\alpha}\dot{\beta}} & 0 \\ 0 & \Lambda \epsilon_{\alpha\beta} \end{pmatrix}$$

$\Lambda \longrightarrow$ Cosmological constant

Today: stay in the
 $\Lambda = 0$ case.

$$V^A \equiv (V^{\dot{\alpha}}, V_{\alpha}) = \left(\frac{\partial h}{\partial \mu_{\dot{\alpha}}}, 0 \right)$$

$$\theta^A \equiv dZ^A - V^A \lrcorner = (d\mu^{\dot{\alpha}} - V^{\dot{\alpha}}, d\lambda_{\alpha})$$

$$N^A \equiv (N^{\dot{\alpha}}, N_{\alpha}) = \left(\frac{\partial}{\partial \mu_{\dot{\alpha}}} \left(\bar{\partial} h + \frac{1}{2} \frac{\partial h}{\partial \mu^{\dot{\beta}}} \wedge \frac{\partial h}{\partial \mu_{\dot{\beta}}} \right), 0 \right)$$



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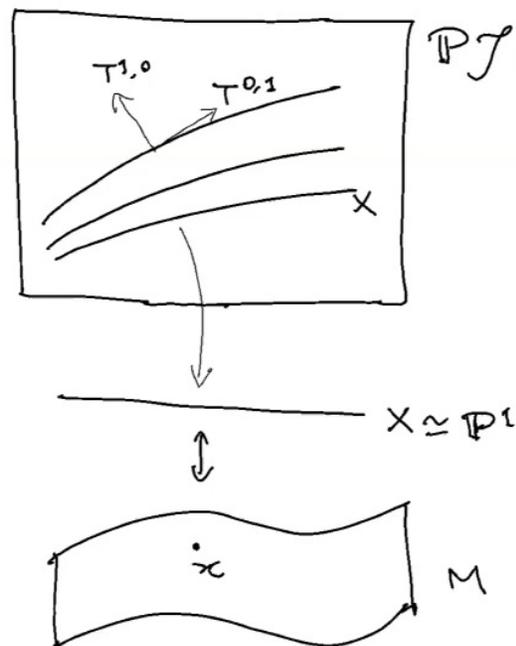
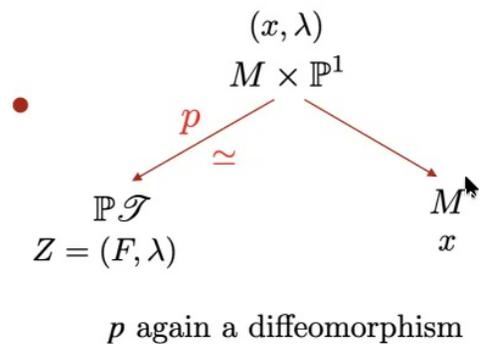
Deformed incidence relations



$$X \simeq \mathbb{P}^1 : \quad \mu^{\dot{\alpha}} = x^{\alpha\dot{\alpha}} \lambda_{\alpha} \quad \longrightarrow \quad \mu^{\dot{\alpha}} = F^{\dot{\alpha}}(x, \lambda)$$

Rational curves in twistor space:

- $F^{\dot{\alpha}}(x, t\lambda) = t F^{\dot{\alpha}}(x, \lambda)$ Degree 1
- $T^{*1,0}\mathbb{P}\mathcal{T}|_X \simeq T^{*1,0}X$ Pseudo-holomorphic
[McDuff, Salamon]



Twistor action for GR

$$S[B, V] = \int_{\mathbb{P}\mathcal{T}} \Omega \wedge B_{\dot{\alpha}} \wedge N^{\dot{\alpha}} \quad \Bigg| \longrightarrow \quad \text{Self-dual (SD) GR}$$

$$+ \frac{\kappa^2}{4} \int_{M \times (\mathbb{P}^1)^2} \langle \lambda_1 \lambda_2 \rangle H_1^{\dot{\alpha}\dot{\gamma}} p_1^*(B_{1\dot{\alpha}} \wedge \Omega_1) \wedge H_2^{\dot{\beta}\dot{\gamma}} p_2^*(B_{2\dot{\beta}} \wedge \Omega_2)$$

→ Anti-self-dual (ASD) interactions

- $\Omega(Z) := \langle \lambda d\lambda \rangle \wedge \theta^{\dot{\alpha}} \wedge \theta_{\dot{\alpha}} \longrightarrow$ “Holomorphic” top-form
- $B_{\dot{\alpha}}(Z) \equiv d\bar{Z}^{\bar{A}} B_{\bar{A}\dot{\alpha}}(Z) \longrightarrow$ Will give rise to ASD graviton perturbations
- $H^{\dot{\alpha}}_{\dot{\beta}}(x, \lambda) \in \text{SL}(2, \mathbb{C}) \longrightarrow$ Spin frame for dotted indices (see paper for def’n).
- $p_i : (x, \lambda_i) \mapsto Z_i^A = (F^{\dot{\alpha}}(x, \lambda_i), \lambda_{i\dot{\alpha}}), \quad B_i \equiv B(Z_i), \dots$



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Twistor action for GR

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Equivalence with GR



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$$S[B, V] = \int_{M \times \mathbb{P}^1} p^* \Omega \wedge p^* B_{\dot{\alpha}} \wedge p^* N^{\dot{\alpha}} - \frac{\kappa^2}{4} \int_{M \times (\mathbb{P}^1)^2} \langle \lambda_1 \lambda_2 \rangle H_1^{\dot{\alpha}\dot{\gamma}} p_1^* B_{1\dot{\alpha}} \wedge H_2^{\dot{\beta}\dot{\gamma}} p_2^* B_{2\dot{\beta}} \wedge \underbrace{p_1^* \Omega_1 \wedge p_2^* \Omega_2}_{\propto \text{dvol}_M \wedge \dots}$$

$p^* B_{\dot{\alpha}} = p^* B_{\dot{\alpha}|_{\mathbb{P}^1}} + p^* B_{\dot{\alpha}|_M}$

Integrating out $p^* B|_M$ imposes a constraint: $\frac{\partial}{\partial \bar{\lambda}_{\dot{\beta}}} \lrcorner p^* N^{\dot{\alpha}} = 0$

- This constraint can be solved exactly!**
- $p^* \theta^{\dot{\alpha}} = \lambda_{\beta} H^{\dot{\alpha}}_{\dot{\beta}}(x, \lambda) e^{\beta\dot{\beta}}(x)$ at fixed λ_{α}
 - $\Gamma_{\alpha\beta}(x) = \int_X \langle \lambda d\lambda \rangle \wedge \lambda_{\alpha} \lambda_{\beta} p^* B_{\dot{\alpha}} \wedge p^* \theta^{\dot{\alpha}}$
- [Sharma '21]

\mathbb{P}^1 -compactification gives Plebanski-Abou-Zeid-Hull action for GR

$$S[e, \Gamma] = \int_M e^{\alpha\dot{\alpha}} \wedge e^{\beta\dot{\beta}} \wedge (d\Gamma_{\alpha\beta} + \kappa^2 \Gamma_{\alpha}{}^{\gamma} \wedge \Gamma_{\gamma\beta})$$

[Abou-Zeid, Hull '05]

MHV generating functional

$$\mathcal{G}(1^-, 2^-) = \kappa^2 \int_M e^{\alpha\dot{\alpha}} \wedge e^{\beta\dot{\beta}} \wedge \Gamma_{1\alpha\dot{\alpha}} \wedge \Gamma_{2\beta\dot{\beta}}$$

[Mason, Skinner '08]

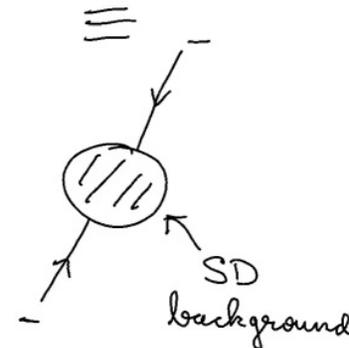
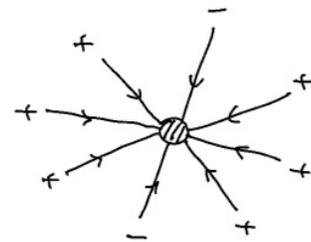
$$= \frac{\kappa^2}{4} \int_{M \times (\mathbb{P}^1)^2} \langle \lambda_1 \lambda_2 \rangle H_1^{\dot{\alpha}\dot{\gamma}} p_1^*(B_{1\dot{\alpha}} \wedge \Omega_1) \wedge H_2^{\dot{\beta}\dot{\gamma}} p_2^*(B_{2\dot{\beta}} \wedge \Omega_2)$$

Interpretation:

- Collection of positive helicity graviton perturbations can be treated as an SD vacuum background.

$$g_{\mu\nu} \rightarrow \eta_{\mu\nu} + \sum_{i=3}^n \epsilon_i h_{i\mu\nu}^+ \text{ asymptotically}$$

- SD vacuum space-times are hyperkahler.
- \mathcal{G} gives amplitude for scattering of two negative helicity graviton perturbations off an SD background.



Hodges' formula

$$\mathcal{G}(1^-, 2^-) = \frac{\kappa^2}{4} \int_{\mathbb{R}^4} d^4x e^{i(k_1+k_2)\cdot x} S_K|_{\text{on-shell}}$$

MHV amplitude extracted from the piece in S_K that is multilinear in each ϵ_i

$$\begin{aligned} A(1^-, 2^-, 3^+, \dots, n^+) &= \frac{\kappa^2}{4} \int_{\mathbb{R}^4} d^4x e^{i(k_1+k_2)\cdot x} \left(\prod_{i=3}^n \frac{\partial}{\partial \epsilon_i} \right) S_K|_{\text{on-shell}} \Big|_{\epsilon_i=0} \\ &= \frac{\kappa^2}{4} \int_{\mathbb{R}^4} d^4x e^{i(k_1+k_2)\cdot x} \langle V_{h_3} V_{h_4} \dots V_{h_n} \rangle_{\text{conn, tree}}^0 \end{aligned}$$

With the “tree-level” connected correlator computed using the free worldsheet theory and vertex operators:

$$S_K^0[m] = \int_X \frac{\langle \lambda d\lambda \rangle}{\langle \lambda 1 \rangle^2 \langle \lambda 2 \rangle^2} \wedge [m \bar{\partial}|_X m] \quad V_{h_i} = \int_X \frac{2 \langle \lambda d\lambda \rangle \wedge h_i|_X}{\langle \lambda 1 \rangle^2 \langle \lambda 2 \rangle^2}$$



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Future: what is this good for?

- MHV formalism for gravity???

[Bjerrum-Bohr, Dunbar, Ita, Perkins, Risager '06] [Adamo, Mason '13]

- Extending gravitational twistor string technology to *non-supersymmetric, loop-level* graviton scattering.
- Covariantization and quantization? [Mason, Wolf '09]
- SD gravity, quantum integrability, celestial/twisted holography...

$$S[h] = \int_{\mathbb{P}\mathcal{I}} \frac{\Omega}{\langle \lambda 1 \rangle^4 \langle \lambda 2 \rangle^4} \wedge \left(h \wedge \bar{\partial} h + \frac{1}{3} h \wedge \frac{\partial h}{\partial \mu^{\dot{\alpha}}} \wedge \frac{\partial h}{\partial \mu_{\dot{\alpha}}} \right)$$

Ongoing work with Skinner and Bittleston

