Title: Decoherence vs space-time diffusion: testing the quantum nature of gravity

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Abstract: Consistent dynamics which couples classical and quantum systems exists, provided it is stochastic. This provides a way to study the back-reaction of quantum systems on classical ones and has recently been explored in the context of quantum fields back-reacting on space-time. Since the dynamics is completely positive and circumvents various no-go theorems this can either be thought of as a fundamental theory, or as an effective theory describing the limit of quantum gravity where the gravitational degrees of freedom are taken to be classical. In this talk we explore some of the consequences of complete positivity on the dynamics of classical-quantum systems. We show that complete positivity necessarily results in the decoherence of the quantum system, and a breakdown of predictability in the classical-phase space. We prove there is a trade-off between the rate of this decoherence and the degree of diffusion in the metric: long coherence times require strong diffusion relative to the strength of the coupling, which potentially provides a long-distance experimental test of the quantum nature of gravity We discuss the consequences of complete positivity on preparing superpositions of gravitationally different states. Each state produces different distributions of the gravitational field determined by the constraints of the theory. The overlap of these distributions imposes an upper bound on the degree of coherence of the superposition.

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Decoherence vs space-time diffusion: testing the quantum nature of gravity

Zach Weller-Davies

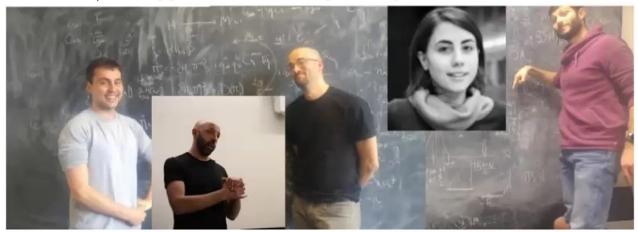
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Quantum foundations seminar, PI, 2020

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W/ J. Oppenheim, J.Camps, C.Sparaciari, B. Šoda



"Decoherence vs space-time diffusion: testing the quantum nature of gravity" $_{\triangleright}$

"Decoherence of quantum fields on classical space-time"

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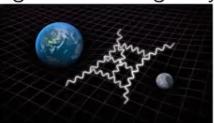
Motivation - why study classical quantum dynamics?



 Revisiting the question of whether or not gravity can be fundamentally classical



• Understanding the regime in which gravity behaves classically



(Feynman 1957 Chapman Hill, Eppley-Hannah 1977, Blanchard Jadczyk 1993, Diosi 1995)

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 We can write down a classical-quantum master equation and use symmetry + complete positivity + physics input input to determine experimental and theoretical consequences

$$rac{\partial arrho}{\partial t} = \int dz' W^{\mu
u} \left(z \mid z'
ight) \mathcal{L}_{\mu} arrho \left(z'
ight) \mathcal{L}_{
u}^{\dagger} - rac{1}{2} \left\{ \int dz' W^{\mu
u} \left(z' \mid z
ight) \mathcal{L}_{
u}^{\dagger} \mathcal{L}_{\mu}, arrho(z)
ight\}$$

- Not the same as semi-classical gravity
- In this talk I will discuss some of the consequences of complete positivity on the dynamics

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Main Messages

Using positivity of the state and complete positivity of the dynamics

$$\int d\Delta d\Delta' \operatorname{Tr}_{A,B} \left[(f_A(\Delta)g_B(\Delta') - g_A(\Delta)f_B(\Delta'))^{\dagger} (f_A(\Delta)g_B(\Delta') - g_A(\Delta)f_B(\Delta'))T_A(\Delta)T_B(\Delta') \right]$$
(22)

A trade off between diffusion and decoherence

We will show there is a trade-off between the decoherence and diffusion: long coherence times require strong diffusion

Trade-off is experimentally bounding

The trade-off between decoherence and diffusion places strong experimental bounds on post-quantum theories of gravity, providing a potential indirect test for the quantum nature of gravity

Positivity of the state and gravitational decoherence

By demanding positivity of the CQ state, we can gain insight into some of the puzzling natures of the Diosi-Penrose decoherence rate

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Classical quantum states $\varrho(q,p)$

The state space consists of a Hilbert space at each point in phase space. Hybrid states are positive, $\rho(q,p) \geq 0$ and normalized $\int dq dp Tr[\varrho(q,p)] = 1$

Often take classical degrees of freedom to live in a phase space and will denote them by \boldsymbol{z}

Example

A hybrid qubit with classical position and momentum q, p

$$\varrho(q,p,t) = \begin{pmatrix} u_0(q,p,t) & \alpha(q,p,t) \\ \alpha^*(q,p,t) & u_1(q,p,t) \end{pmatrix}$$
(1)

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Conditions from positivity



The state must be **positive**, $\varrho(q,p) \geq 0$. For the hybrid qubit,

$$\varrho(q,p,t) = \begin{pmatrix} u_0(q,p,t) & \alpha(q,p,t) \\ \alpha^*(q,p,t) & u_1(q,p,t) \end{pmatrix}$$
(2)

this implies

$$|\alpha(q, p, q)|^2 \le u_0(q, p, t)u_1(q, p, t)$$
 (3)

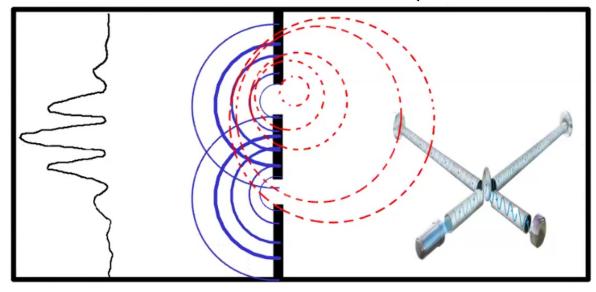
So for **coherence**, the classical distributions $u_0(q, p, t), u_1(q, p, t)$ describing the populations must have overlap.



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This has a somewhat intuitive explanation



$$|\psi\rangle \rightarrow |L\rangle|E_L\rangle + |R\rangle|E_R\rangle$$

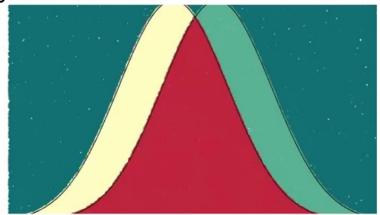
$$\rho = Tr_{E}(|\psi\rangle\langle\psi|) = \begin{pmatrix} \frac{1}{2} & \alpha \\ \alpha^{*} & \frac{1}{2} \end{pmatrix}$$

with $\alpha = \langle E_L | E_R \rangle$, we see interference patterns because the electromagnetic fields are not orthogonal

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Classical states $|E_L\rangle, |E_R\rangle$ are perfectly distinguishable, **unless** they are **probability distributions**



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What kind of dynamics is allowed?



It must preserve the state space

$$\varrho(q,p) \geq 0$$
, $\int dq dp Tr[\varrho(q,p)] = 1$

We ask that it be completely positive on the quantum system

CQ dynamics (Oppenheim 2018)- CQ version of Kraus (87)

$$\varrho(z,t+\delta t) = \sum_{\mu\nu} \int dz' \Lambda^{\mu\nu} \left(z \mid z',\delta t\right) L_{\mu} \varrho\left(z',t\right) L_{\nu}^{\dagger} \qquad (4)$$

where positivity demands $\Lambda^{\mu\nu}(z|z')$ is a positive matrix for each z, z' and

$$\int dz \sum_{\mu\nu} \Lambda^{\mu\nu} \left(z \mid z', \delta t \right) L_{\nu}^{\dagger} L_{\mu} = \mathbb{I}$$
 (5)

due to normalization.

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What kind of dynamics is allowed? continued



Quantum

Dynamics: $\sigma(t) = \sum_{\mu} \lambda^{\mu\nu} L_{\mu} \sigma(0) L_{\nu}^{\dagger}$

Normalization: $\sum_{\mu} \lambda^{\mu\nu} L_{\nu}^{\dagger} L_{\mu} = \mathbb{I}$

Positivity: $\lambda^{\mu\nu}$ positive matrix

Classical

Dynamics: $p(z,t) = \int dz' P(z \mid z',t) p(z',0)$

Normalization: $\int dz P(z \mid z') = 1$

Positivity: P(z|z') positive for each z, z'

CQ

Dynamics: $\varrho(z,t) = \int dz' \Lambda^{\mu\nu} (z \mid z',t) L_{\mu} \varrho(z',0) L_{\nu}^{\dagger}$

Normalization: $\int dz \sum_{\mu\nu} \Lambda^{\mu\nu} (z \mid z', t) L^{\dagger}_{\nu} L_{\mu} = \mathbb{I}$

Positivity: $\Lambda^{\mu\nu}(z|z',t)$ positive matrix for each z,z'

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Quantum

Master equation: $\frac{\partial \sigma}{\partial t} = -i[H, \sigma] + h^{\alpha\beta}L_{\alpha}\sigma L_{\beta}^{\dagger} - \frac{1}{2}\left\{h^{\alpha\beta}L_{\beta}^{\dagger}L_{\alpha}, \sigma\right\}$

Positivity: $h^{\alpha\beta}$ positive matrix

Classical

Master equation:

$$\frac{\partial p}{\partial t} = \int dz'W(z \mid z') p(z') - \int dz'W(z' \mid z) p(z)$$

Positivity: $\delta(z,z') + \delta t W(z|z')$ positive

CQ

Master Equation: (In a basis of Lindblad operators $L_{\mu} = (I, L_{\alpha})$)

$$\frac{\partial \varrho}{\partial t} = \int dz' W^{\mu\nu} \left(z \mid z' \right) L_{\mu} \varrho \left(z' \right) L_{\nu}^{\dagger} - \frac{1}{2} \left\{ \int dz' W^{\mu\nu} \left(z' \mid z \right) L_{\nu}^{\dagger} L_{\mu}, \varrho(z) \right\}$$

Positivity:

$$\Lambda^{\mu\nu}\left(z\mid z',\delta t\right) = \left[\begin{array}{cc} \delta\left(z,z'\right) + \delta t W^{00}\left(z\mid z'\right) & \delta t W^{0\beta}\left(z\mid z'\right) \\ \delta t W^{\alpha0}\left(z\mid z'\right) & \delta t W^{\alpha\beta}\left(z\mid z'\right) \end{array}\right]$$

a positive matrix

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Kramers-Moyal expansion

 In classical dynamics we can perform a Kramers-Moyal expansion of the master equation

$$\frac{\partial p}{\partial t} = \int dz' W (z \mid z') p(z') - \int dz' W (z' \mid z) p(z)$$

$$\longrightarrow \sum_{n=1}^{\infty} (-1)^n \frac{\partial^n}{\partial z^n} [D_n(z) p(z, t)]$$

where

$$D_n(z) = \frac{1}{n!} \int dz' \left(z' - z\right)^n W\left(z' \mid z\right)$$

are the moments of the transition amplitude

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Kramers-Moyal expansion continued

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Example

Take
$$D_{1,q}=rac{\partial H}{\partial p}$$
, $D_{1,p}=-rac{\partial H}{\partial q}$, $D_{n\geq 2}=0$
$$rac{\partial p}{\partial t}=\{H,p\}$$

Example

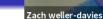
Take
$$D_1=\mu$$
, $D_2=D$, $D_{n\geq 3}=0$
$$\frac{\partial p}{\partial t}=-\frac{\partial}{\partial z}\left[\mu(z)p(z,t)\right]+\frac{\partial^2}{\partial z^2}\left[D(z)p(z,t)\right]$$

the Fokker-Plank equation

- D_1 characterizes the amount of Hamiltonian evolution in the system (more precisely the drift)
- D₂ characterizes the amount of diffusion in the system

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CQ Kramers-Moyal expansion

CQ Kramers-Moyal expansion

$$\begin{split} \frac{\partial \varrho(z,t)}{\partial t} &= \sum_{n=1}^{\infty} (-1)^n \left(\frac{\partial^n}{\partial z^n} \right) \left(D_n^{00}(z) \varrho(z,t) \right) \\ &- i [H(z), \varrho(z)] + D_0^{\alpha\beta}(z) L_{\alpha} \varrho(z) L_{\beta}^{\dagger} - \frac{1}{2} D_0^{\alpha\beta} \left\{ L_{\beta}^{\dagger} L_{\alpha}, \varrho(z) \right\}_{+} \\ &+ \sum_{\mu\nu \neq 00} \sum_{n=1}^{\infty} (-1)^n \left(\frac{\partial^n}{\partial z^n} \right) \left(D_n^{\mu\nu}(z) L_{\mu} \varrho(z,t) L_{\nu}^{\dagger} \right) \end{split}$$

Important moments for the talk

- $D_0^{\alpha\beta}$ characterizes the decoherence
- ullet $D_1^{\mu
 u}$ characterizes the Hamiltonian part of the back-reaction on phase-space (drift)
- $D_2^{\mu\nu}$ characterizes the diffusion (spreading) in the phase-space

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Main Results



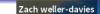
We can use the conditions on positivity to prove the following

- A CQ Pawula theorem: unique (a.s) continuous master equation, or else one must have infinite moments (jumping)
- We must have a Decoherence term: For CQ back-reaction, we must have a decoherence term $D_0^{\alpha\beta}$. Interaction with a classical system necessarily causes decoherence
- We **must** have a Diffusion term $D_2^{\mu\nu}$: Interaction with a classical system necessarily results in a loss of predictability on the classical phase space

A trade off between diffusion and decoherence

There is a trade-off between the amount of diffusion and decoherence: long coherence times require strong diffusion

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A CQ Pawula Theorem

CQ Pawula Theorem

For non-trivial CQ evolution, we must have infinitely many moments in the master equation (specifically, none of the even moments can vanish), or else the master equation takes the form

$$\frac{\partial \varrho(z,t)}{\partial t} = -i[H(z), \varrho(z,t)] + \sum_{n=1}^{n=2} (-1)^n \left(\frac{\partial^n}{\partial z^n}\right) \left(D_n^{00} \varrho(z,t)\right)
+ \frac{\partial}{\partial z} \left(D_1^{0\alpha} \varrho(z,t) L_\alpha^{\dagger}\right) + \frac{\partial}{\partial z} \left(D_1^{\alpha 0} L_\alpha \varrho(z,t)\right)
+ D_0^{\alpha \beta}(z) L_\alpha \varrho(z) L_\beta^{\dagger} - \frac{1}{2} D_0^{\alpha \beta} \left\{L_\beta^{\dagger} L_\alpha, \varrho(z)\right\}_+$$

Furthermore, $2D_{2,ii}^{00} \geq (D_0^{-1})_{\alpha\beta} D_{1,i}^{0a} D_{1,i}^{0\beta*}$

Important fact for rest of the talk

We **must** have a decoherence term $D_0^{\alpha\beta}$ and a diffusion term $D_2^{\mu\nu}$

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A trade-off between decoherence and diffusion



A trade-off between decoherence and diffusion

For all CQ master equations, we derive a trade-off between decoherence and diffusion depending on the drift in the system

- Two sources of drift with back-reaction, $D_{1,i}^{0\alpha}$, $D_{1,i}^{\alpha\beta}$
- The purely classical diffusion is bounded below by $2D_{2,ii}^{00} \geq \left(D_{0,i}^{-1}\right)_{\alpha\beta} D_{1,i}^{0a} D_{1,i}^{0\beta*}$ (Diosi, 1995)
- The CQ diffusion term $D_2^{\alpha\beta}$ must satisfy the bound $\sum_{\alpha} 2D_{2,ii}^{\alpha\alpha} \sum_{\beta} D_0^{\beta\beta}(z) \ge \left| \sum_{\alpha} D_{1,i}^{\alpha\alpha}(z) \right|^2$

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A trade-off between decoherence and diffusion

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Physical input

If we want the dynamics to approximately reproduce Hamiltonian dynamics, we know what the first moment should be! (Oppenheim 2018)

$$Tr[\frac{\partial}{\partial z_i}(D_{1,i}^{\mu\nu}L_{\mu}\varrho(z)\mathcal{L}_{\nu}^{\dagger})] = Tr[\{H_m,\varrho\}]$$

• c.f two classical systems (z_1, z_2) interacting with a H_I

$$\frac{\partial p(z_1, z_2, t)}{\partial t} = \{H_1, \rho\} + \{H_2, \rho\} + \{H_1, \rho\}$$

 Integrating out the second system and defining $\bar{\rho}(z_1) = \int dz_2 \rho(z_1, z_2)$ we get an effective e.o.m

$$\frac{\partial \bar{\rho}(z_1)}{\partial t} = \left\{H_1, \bar{\rho}(z_1)\right\} + \int dz_2 \left\{H_I, \rho(z_1, z_2)\right\}$$

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Trade-off provides an experimental bound

 \bullet There are experimental ${\bf upper\ bounds}$ on the decoherence rate λ

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A trade-off between decoherence and diffusion



A trade-off between decoherence and diffusion

For all CQ master equations, we derive a trade-off between decoherence and diffusion depending on the drift in the system

- Two sources of drift with back-reaction, $D_{1,i}^{0\alpha}$, $D_{1,i}^{\alpha\beta}$
- The purely classical diffusion is bounded below by $2D_{2,ii}^{00} \geq \left(D_0^{-1}\right)_{\alpha\beta} D_{1,i}^{0a} D_{1,i}^{0\beta*}$ (Diosi, 1995)
- ullet The CQ diffusion term $D_2^{lphaeta}$ must satisfy the bound $\sum_{\alpha} 2D_{2,ii}^{\alpha\alpha} \sum_{\beta} D_0^{\beta\beta}(z) \ge \left| \sum_{\alpha} D_{1,i}^{\alpha\alpha}(z) \right|^2$
- Trade-off has important consequences!

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Trade-off provides an experimental bound

- There are experimental upper bounds on the decoherence rate λ
- Using the physical input that we want the dynamics to approximately reproduce Hamiltonian dynamics
- Give us an experimental lower bounds on the amount of diffusion in the classical system – at a relevant energy scale

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Brief aside on post-quantum gravity



- In order to study the trade-off in a concrete setting we shall study a toy model of a non-relativistic quantum field interacting with a classical Newtonian potential
- I will therefore give a very brief tour of post-quantum gravity



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Summary of post-quantum gravity

- Take classical degrees of freedom to be Riemmanian 3-metric g_{ab} and their conjugate momenta π^{ab}
- Couple to QM and consider dynamics of the state $\varrho(g_{ab},\pi^{ab})$ in an ADM formalism
- Can study independence of the dynamics of the lapse and shift to find Momentum and Hamiltonian constraints (Oppenheim, ZWD 2011.15112)

$$\mathcal{L}_{constraint} = \int d^{3}x M_{\alpha} D_{b} N \left[\left\{ g^{ab}(y), \mathcal{H}(y) \varrho \right\} + \left(-2i C_{H}^{ab} h^{\alpha \beta} + \frac{i}{2} C_{N}^{ab} W_{0}^{\alpha \beta} - \frac{i}{2} C_{J}^{ab} W^{\alpha \beta} \right) [L_{\beta}^{\dagger} L_{\alpha}, \varrho] \right]$$

$$+ 2 \left(C_{J}^{ab} h^{\alpha \beta} + C_{H}^{ab} W^{\alpha \beta} \right) L_{\alpha} \varrho L_{\beta}^{\dagger} - \left(C_{N}^{ab} h^{\alpha \beta} + C_{H}^{ab} W_{0}^{\alpha \beta} \right) \left\{ L_{\beta}^{\dagger} L_{\alpha}, \varrho \right\} + \right]$$

$$\int d^{3}x N M_{a} \mathcal{H}_{b} \left(\left\{ g^{ab}, \mathcal{H}_{b} \bar{\mathcal{L}}(\varrho) \right\} - \bar{\mathcal{L}} \left(\left\{ g^{ab}, \mathcal{H}_{b} \varrho \right\} \right)$$

$$+ \left(\frac{i}{2} \mathcal{R}_{NN}^{\alpha \beta} - \frac{i}{2} \mathcal{R}_{JJ}^{\alpha \beta} - 2i \mathcal{R}_{HH}^{\alpha \beta} \right) [L_{\beta}^{\dagger} L_{\alpha}, \varrho] + \left(2 \mathcal{R}_{HJ}^{\alpha \beta} + 2 \mathcal{R}_{JH}^{\alpha \beta} \right) L_{\alpha} \varrho L_{\beta}^{\dagger} - \left(\mathcal{R}_{NH}^{\alpha \beta} + \mathcal{R}_{HN}^{\alpha \beta} \right) \left\{ L_{\beta}^{\dagger} L_{\alpha} \right\} +$$

$$+ \int d^{3}x N M_{a} \left[\left(C_{JN}^{ab} (W_{0}^{\phi \phi} - 2i h^{\phi \phi}) - C_{J}^{ab} W_{0}^{\phi \phi} + C_{N}^{ab} W^{\phi \phi} \right) D_{b} \phi \varrho \phi +$$

$$+ \left(C_{JN}^{ab} (W_{0}^{\phi \phi} + 2i h^{\phi \phi}) + C_{J}^{ab} W_{0}^{\phi \phi} - C_{N}^{ab} W^{\phi \phi} \right) \phi \varrho D_{b} \phi$$

$$+ \left(C_{JN}^{ab} (W_{0}^{\phi \phi} + 2i h^{\phi \phi}) + C_{J}^{ab} W^{\phi \phi} - V_{N}^{ab} W^{\pi \pi} \right) D_{b} \pi \varrho \pi + \left(C_{JN}^{ab} (-W_{0}^{\pi \pi} - 2i h^{\pi \pi}) + W_{0}^{\pi \pi} C_{J}^{ab} - C_{N}^{ab} W^{\pi \pi} \right) \pi \varrho D_{b} \pi$$

$$+ \left(C_{JN}^{ab} (-W_{0}^{\phi \phi} - 2i h^{\phi \phi}) + C_{N}^{ab} W^{ef} - W_{0}^{ef} h^{ef} C_{J}^{ab} \right) \left(D_{e} D_{f} \phi \right) \varrho D_{b} \phi$$

$$+ \left(C_{JN}^{ab} (-W_{0}^{\phi \phi} - 2i h^{\phi \phi}) - C_{N}^{ab} W^{ef} - W_{0}^{ef} h^{ef} C_{J}^{ab} \right) \left(D_{e} D_{f} \phi \right) \varrho D_{b} \phi$$

$$+ \left(C_{JN}^{ab} (-W_{0}^{\phi \phi} - 2i h^{\phi \phi}) - C_{N}^{ab} W^{ef} + W_{0}^{ef} C_{J}^{ab} \right) D_{b} \phi \varrho \left(D_{e} D_{f} \phi \right) \right]$$

$$\int d^{3}x - N M_{b} C_{JN}^{ab} W^{\phi \phi} \left\{ D_{b} \phi \phi, \varrho \right\} + N W^{\pi \pi} \left\{ \pi D_{b} \left(M_{a} \pi C_{JN}^{ab}, \varrho \right) \right\} + M_{a} C_{JN}^{ab} \left\{ D_{b} \phi D_{d} (N W^{cd} D_{c} \phi, \varrho) \right\} +$$

$$- 2 \int d^{3}x \left[N W_{0}^{\pi \pi} \pi D_{b} \left(C_{JN}^{ab} M_{a} \varrho \right) \pi + M_{b} D_{d} N W_{0}^{cd} C_{JN}^{ab} D_{c} \phi \varrho D_{b} \phi \right] \approx 0$$
(97)

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Newtonian limit - state space

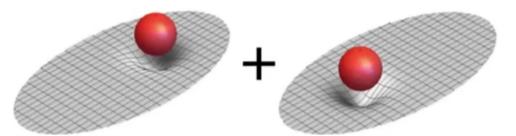
• We consider the case of a non-relativistic quantum field in a partially decohered super-position of approximately orthogonal $|L\rangle$, $|R\rangle$

$$|L/R\rangle = \int d^3x f_{L/R}(x)\psi^{\dagger}(x)|0\rangle$$

$$f_L(x)f_R(x) pprox 0$$
 , $\psi(x) = \int rac{d^3p}{(2\pi)^3} a_{\vec{p}} e^{i\vec{p}\cdot\vec{x}}$

- Take classical d.o.f to be Newtonian potential Φ and its canonical conjugate π_g
- The state space is then

$$\varrho\left(\Phi,\pi_{g},t\right) = \begin{pmatrix} u_{L}\left(\Phi,\pi_{g},t\right) & \alpha\left(\Phi,\pi_{g},t\right) \\ \alpha^{*}\left(\Phi,\pi_{g},t\right) & u_{R}\left(\Phi,\pi_{g},t\right) \end{pmatrix}$$



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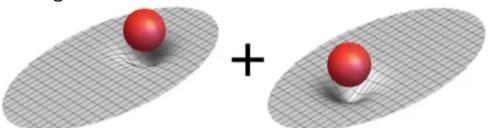


Positivity of the state and gravitational decoherence

By demanding positivity of the CQ state, we can gain insight into some of the puzzling natures of the Diosi-Penrose decoherence rate $\lambda_D = \frac{\Delta E_D}{\hbar}$

$$\Delta E_D = \int d^3x d^3x' \frac{[m_L(x) - m_R(x)][m_L(x') - m_R(x')]}{|x - x'|}$$

• Non-local, mass in the left branch interacts with the mass in the right branch



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Positivity bounds coherence



Classical-quantum decoherence rates

We see that we can arrive at the Diosi-Penrose decoherence using local dynamics. We find it is **not** a dynamical effect, but instead is a constraint imposed by demanding that the density matrix be **positive**.

$$\varrho\left(\Phi, \pi_{g}, t\right) = \begin{pmatrix} u_{L}\left(\Phi, \pi_{g}, t\right) & \alpha\left(\Phi, \pi_{g}, t\right) \\ \alpha^{*}\left(\Phi, \pi_{g}, t\right) & u_{R}\left(\Phi, \pi_{g}, t\right) \end{pmatrix}$$
(6)

At t = 0 we take the marginal distributions for the populations

$$u_{L/R}(\Phi) = \frac{\mathcal{N}}{2} \exp\left[-\int d^3x \frac{\left(\Phi(x) - \Phi_{L/R}(x)\right)^2}{2\sigma^2}\right] \tag{7}$$

i.e, Gaussian's peaked around the value of the Newtonian potential Φ_L , Φ_R which satisfies Poisson's equation

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Positivity bounds coherence continued

• Positivity of the state tells us that at t = 0,

$$|\alpha(\Phi, \pi_g, 0)|^2 \le u_L(\Phi, \pi_g, 0) u_R(\Phi, \pi_g, 0)$$

• For the Gaussian marginal distributions

$$\left|\alpha\left(\Phi, \pi_g, 0\right)\right|^2 \leq \frac{\mathcal{N}^2}{4} \exp\left[-\int d^3x \frac{\left(\Phi_L(x) - \Phi_R(x)\right)^2}{4\sigma^2}\right] \\ \times \exp\left[-\int d^3x \frac{\left(\Phi(x) - \frac{1}{2}\left(\Phi_L(x) + \Phi_R(x)\right)\right)^2}{\sigma^2}\right]$$

ullet If $\Phi_{L/R}$ satisfy Poisson's equation

$$\Phi_{L/R} = -4\pi G \int d^3x' \frac{\mu_{L/R}(x')}{|x-x'|}$$

$$\int d^3x \left(\Phi_L(x) - \Phi_R(x)\right)^2 \sim \int d^3x d^3x' \frac{[\mu_L(x) - \mu_R(x)][\mu_L(x') - \mu_R(x')]}{|x - x'|} \tag{8}$$

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Positivity bounds coherence continued



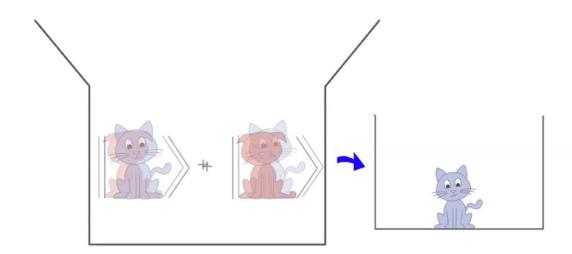
Quantum decoherence on classical space-time

- We see that here DP is **not** a dynamical decoherence rate.
 Instead it is a **non-dynamical** effect which arises due to **positivity** of the density matrix.
- Gives a bound on the allowed coherence one can prepare: two
 masses with very different gravitational fields cannot be
 prepared in a coherent superposition.
- Since it is a condition on the allowed states one is allowed to prepare the non-locality is less of a problem than if this is a dynamical effect.

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 When we include dynamics, we also have a dynamical contribution to the decoherence (model dependent)

$$\int D\Phi D\pi_{g} |\alpha(\Phi, \pi_{g})| = \bar{\alpha}(t) = \frac{1}{4} \exp\left[-\int d^{3}x \frac{(\Phi_{L}(x) - \Phi_{R}(x))^{2}}{8\sigma^{2}}\right]$$
$$\times \exp\left[-\int d^{3}x \left[\frac{\lambda(\Phi)t (\Phi_{L}(x) + \Phi_{R}(x))}{2}\right]\right]$$



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Positivity bounds coherence continued

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• **Positivity** of the state tells us that at t = 0,

$$|\alpha\left(\Phi,\pi_{g},0\right)|^{2} \leq u_{L}\left(\Phi,\pi_{g},0\right)u_{R}\left(\Phi,\pi_{g},0\right)$$

For the Gaussian marginal distributions

$$\left|\alpha\left(\Phi, \pi_{g}, 0\right)\right|^{2} \leq \frac{\mathcal{N}^{2}}{4} \exp\left[-\int d^{3}x \frac{\left(\Phi_{L}(x) - \Phi_{R}(x)\right)^{2}}{4\wp^{2}}\right]$$

$$\times \exp\left[-\int d^{3}x \frac{\left(\Phi(x) - \frac{1}{2}\left(\Phi_{L}(x) + \Phi_{R}(x)\right)\right)^{2}}{\sigma^{2}}\right]$$

• If $\Phi_{L/R}$ satisfy Poisson's equation

$$\Phi_{L/R} = -4\pi G \int d^3x' \frac{\mu_{L/R}(x')}{|x-x'|}$$

$$\int d^3x \left(\Phi_L(x) - \Phi_R(x)\right)^2 \sim \int d^3x d^3x' \frac{[\mu_L(x) - \mu_R(x)][\mu_L(x') - \mu_R(x')]}{|x - x'|} \tag{8}$$

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• We take the pure gravity Hamiltonian to be

$$H_c(\Phi) = \int_x \left(-\frac{\pi G}{3} \pi_g^2 + \frac{(\nabla \Phi)^2}{4\pi G} \right)$$

• The Quantum Hamiltonian is

$$H_m(\Phi,\phi)=m\int d^3x(1+2\Phi(x))\psi^{\dagger}\psi:=H_m^0+H_I(\Phi)$$

 Using the theory of Oppenheim 2018 we consider a master equation

$$\frac{\partial \varrho}{\partial t} \approx \{H_c(\Phi), \varrho\} - i \left[H_m^0, g\right] + m \int d^3 x D_0 \left[\psi \varrho \psi^{\dagger} - \frac{1}{2} \left\{\psi^{\dagger} \psi, \varrho\right\}\right] + 2m \int d^3 x \psi \frac{\delta \varrho}{\delta \pi_g} \psi^{\dagger} + m \int \psi \frac{\delta^2}{\delta \pi_g^2} (D_2 \varrho) \psi^{\dagger} + \cdots$$

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This approximately reproduces the Newtonian interaction

$$\operatorname{Tr}\left[\left\{H_{I},\varrho\right\}\right]=-2m\int d^{3}x\operatorname{Tr}\left[\psi^{\dagger}\psi\frac{\delta\varrho}{\delta\pi_{g}}\right],\ m\psi^{\dagger}(x)\psi(x)=m(x)$$

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Trade-off between decoherence and diffusion

- The drift results in a lower bound for the diffusion in the momenta conjugate to Φ , π_g
- This results in gravitational kinetic energy production

$$\Delta E = \int d^3x \frac{Gc^2\pi}{3} \left\langle \pi_g^2 \right\rangle \ge \int d^3x \frac{2tc^2G\pi |\langle m(x) \rangle|^2}{3\lambda}$$

(where λ is the decoherence rate and $\langle m(x) \rangle$ is the expectation value of the mass density.)

- We can get an order of magnitude estimate for the kinetic energy production from Gerlich et al. 2007
- The decoherence rate is $\lambda < 10^{-5} s^{-1}$ for clusters of nucleons of mass 10^{-24} and typical radius $r \sim 10^{-9}$. Taking the experiment to be conducted on the order of seconds we find gravitational kinetic energy production

$$\Delta E \sim 10^{-11} J, \ mc^2 \sim 10^{-8} J$$

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- Complete positivity gives us a trade-off between decoherence, drift and diffusion
- In toy models, this seems to be a very relevant prediction of treating the gravitational field classically

General lesson

Treating gravity classically leads to diffusion in the gravitational field which should be experimentally testable - but need to understand this more generally

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Summary

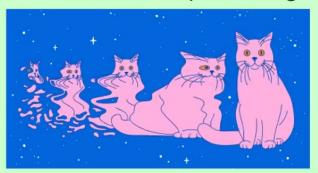
Quantum decoherence in classical space-time

- We have shown that positivity of the CQ state gives rise to upper bounds on the coherence.
- \bullet A dynamical term, due to the decoherence term D_0
- A non-dynamical term, due to positivity of the state, which gives rise to a Diosi-Penrose like bound

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A trade-off between decoherence and diffusion

 We have shown there is a trade-off between decoherence and diffusion: long coherence times require strong diffusion



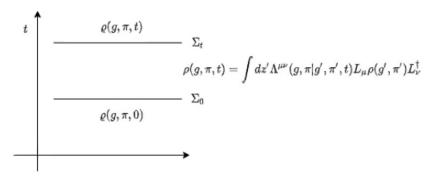
 In simple toy model of quantum fields interaction with classical gravity we have seen this is a non-negligible effect and potentially puts fundamental CQ theories in danger of running afoul of experimental observations.

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Discussion and future outlook



- Better understand the validity of the toy model we use, so that we can strengthen statements and begin thinking about possible experiments
- Understand what general lessons can be learned: a lot of the results rely on complete positivity, are largely independent of the specifics of the dynamics, and we might expect them to generalize to the non-Markovian case.

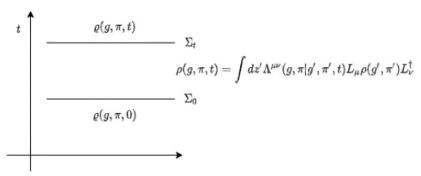


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Discussion and future outlook



- Better understand the validity of the toy model we use, so that we can strengthen statements and begin thinking about possible experiments
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 Better understand when can expect the theory to hold as an effective theory, for example by understanding how we arrive at the CQ limit.

Pirsa: 20120030 Page 40/41 Motivation CQ state CQ dynamics and Kramers-Moyal Expansion CQ Pawula theorem Main results





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