Title: Axion-like Relics: New Constraints from Old Comagnetometer Data

Speakers: Itay Bloch

Series: Particle Physics

Date: October 27, 2020 - 1:00 PM

URL: http://pirsa.org/20100070

Abstract: Abstract: TBD

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## Comagnetometers as a Probe to New Physics

Oct 2020



Itay M. Bloch Tel Aviv University

(Work done in the Rafael Quantum Optics Lab)



#### Collaborators:

T. Volansky, Y. Hochberg, E. Kuflik

Based on work presented in arxiv:1907.03767

(New information based on active works with Gil Ronen and Or Katz)

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#### **Outline**

- Axion Like Particles (ALPs)
  - Axions and Axion Like Particles
  - Coherent Interactions
- The Comagnetometer
  - What is a Comagnetometer?
  - Comagnetometer Dynamics
  - Comagnetometer Experiments
  - Future Prospects (and recent developments)

#### Conclusions

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#### **QCD** Axion

• A solution to the strong CP problem,  $\theta_{QCD} \rightarrow a/f_a$ .

[Peccei, Quinn 1977; Weinberg 1978; Wilczek 1978]

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#### **QCD** Axion

• A solution to the strong CP problem,  $\theta_{QCD} \rightarrow a/f_a$ .

[Peccei, Quinn 1977; Weinberg 1978; Wilczek 1978]

- Pseudo-scalars. Has derivative couplings, e.g.:  $\frac{\partial_{\mu}a}{f_a}\bar{\psi}\gamma_5\gamma^{\mu}\psi$
- $m_a \propto \frac{\Lambda_{\rm QCD}^2}{f_a}$
- Can be a CDM component
- Can be very light and remain CDM candidate:

$$m_a$$
(relevant to talk) <  $10^{-12}$  eV

#### **QCD** Axion

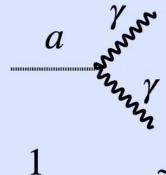
#### **Axion Like Particles (ALPs)**

• A solution to the strang CP problem,  $\theta_{QCD} \to a/f_a$ .

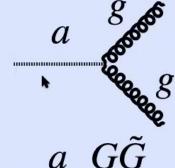
- Pseudo-scalars. Has derivative couplings, e.g.:  $\frac{\partial_{\mu}a}{f_a}\bar{\psi}\gamma_5\gamma^{\mu}\psi$
- $m_a \propto \frac{\Lambda_{\rm QCD}^2}{J_a}$
- Can be a CDM component (we assume all)
- Can be very light and remain CDM candidate:

$$m_a$$
(relevant to talk) <  $10^{-12}$  eV

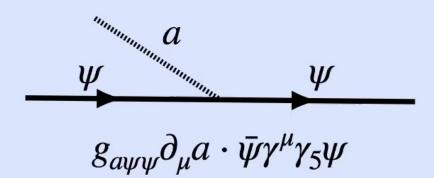
#### **ALP-SM Interactions**



$$-\frac{1}{4}g_{a\gamma\gamma}aF\tilde{F}$$



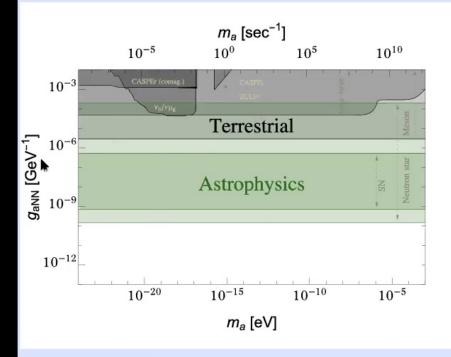
$$-\frac{a}{f_a}\frac{GG}{32\pi^2}$$

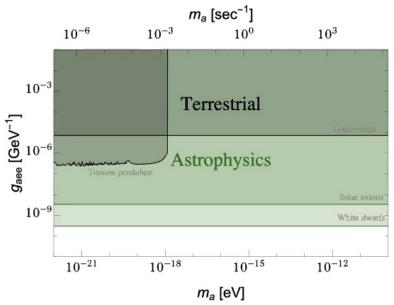


#### **ALP Interactions with Neutrons and Electrons**

$$g_{ann}\partial_{\mu}a\cdot\bar{n}\gamma^{\mu}\gamma_{5}n$$

$$g_{aee}\partial_{\mu}a\cdot \bar{e}\gamma^{\mu}\gamma_5 e$$





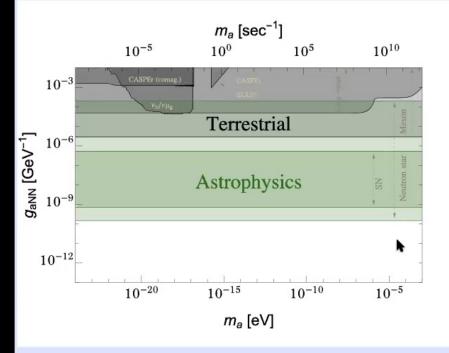
6

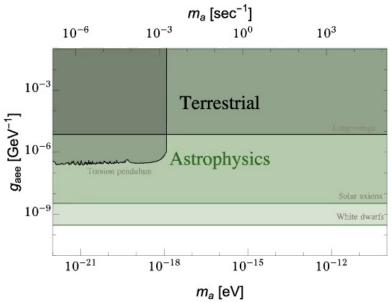
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#### **ALP Interactions with Neutrons and Electrons**

$$g_{ann}\partial_{\mu}a\cdot\bar{n}\gamma^{\mu}\gamma_{5}n$$

$$g_{aee}\partial_{\mu}a\cdot \bar{e}\gamma^{\mu}\gamma_5 e$$





Is there a way to improve on the astrophysical bounds?

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### Coherent Interactions (1)

For light ALPs 
$$(m_a \lesssim 30 \text{ eV})$$
,  $n_a = \frac{0.3 \text{ GeV}}{m_a \cdot \text{cm}^3} > 1/\lambda_{\text{de-Broglie}}^3$ 

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#### Coherent Interactions (1)

For light ALPs (
$$m_a \lesssim 30 \text{ eV}$$
),  $n_a = \frac{0.3 \text{ GeV}}{m_a \cdot \text{cm}^3} > 1/\lambda_{\text{de-Broglie}}^3$ 

$$a = a_0 e^{ix^{\mu}p_{\mu}}$$

In the non-relativistic limit, for  $\mathcal{L}_{a\psi\psi} \sim \partial_{\mu} a \cdot \bar{\psi} \gamma^{\mu} \gamma_5 \psi$ ,

$$H_{a\psi\psi} \sim \partial_i a \cdot \bar{\psi} \gamma^i \gamma_5 \psi \sim a_0 \cdot \overrightarrow{p}_a \cdot \overrightarrow{S}_{\psi}$$

### Coherent Interactions (2)

$$H_{a\psi\psi} = -\,g_{a\psi\psi}\overrightarrow{b}_a\cdot\overrightarrow{S}_\psi = -\,\overrightarrow{b}_{a-\psi}\cdot\overrightarrow{S}_\psi$$

$$\overrightarrow{b}_{a-\psi} = g_{a\psi\psi}\sqrt{2\rho_a}\cos(m_a t) \cdot \overrightarrow{v}_{a-\psi \text{ [astro-ph/9501042]}}$$

This is an effect linear in  $g_{a\psi\psi}$ !

But how to measure it?

## Larmor Precession and Zeeman Splitting

Interaction of a classical magnetic field  $\overrightarrow{B}$  with a spin  $\overrightarrow{S}$ :

$$\frac{\overrightarrow{B}}{\overrightarrow{S}} \Rightarrow \overrightarrow{S} \times \overrightarrow{B} \qquad \qquad H = -\chi \overrightarrow{B} \cdot \overrightarrow{S}$$

Reminder: 
$$H_{a\psi\psi} = -\overrightarrow{b}_{a-\psi} \cdot \overrightarrow{S}_{\psi}$$

Is there a known way to measure magnetic fields?

## Using Alkali Magnetometers to Hunt ALPs

Magnetometers can measure the magnetic-like effect of the ALPs!

But they also measure the actual magnetic fields of the SM...

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#### Differentiating: EM vs. ALPs

Magnetic Fields:

$$H = -\gamma_e \overrightarrow{S_e} \cdot \overrightarrow{B} - \gamma_n \overrightarrow{S_n} \cdot \overrightarrow{B}$$

**Anomalous Fields:** 

$$H = -g_{aee} \overrightarrow{S}_{e} \cdot \overrightarrow{b}_{a} - g_{ann} \overrightarrow{S}_{n} \cdot \overrightarrow{b}_{a}$$

$$\equiv -\overrightarrow{S}_{e} \cdot \overrightarrow{b}_{e} - \overrightarrow{S}_{n} \cdot \overrightarrow{b}_{n}$$

What if we measure with two particles?

### **But That is Not Enough**

(subtracting two numbers is harder than it sounds)

•

Can we **build a device** that subtracts the two splittings from each other?

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    - Future Prospects
  - Conclusions

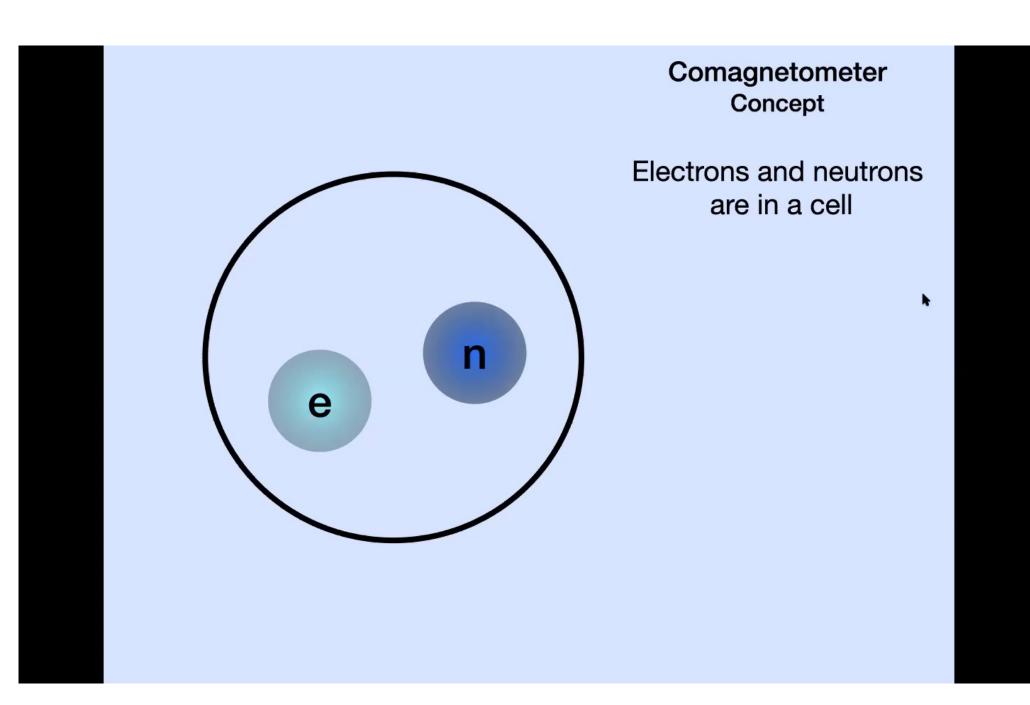
ALP DM generate a magnetic-like field, but it couples differently to different particles compared to the SM.

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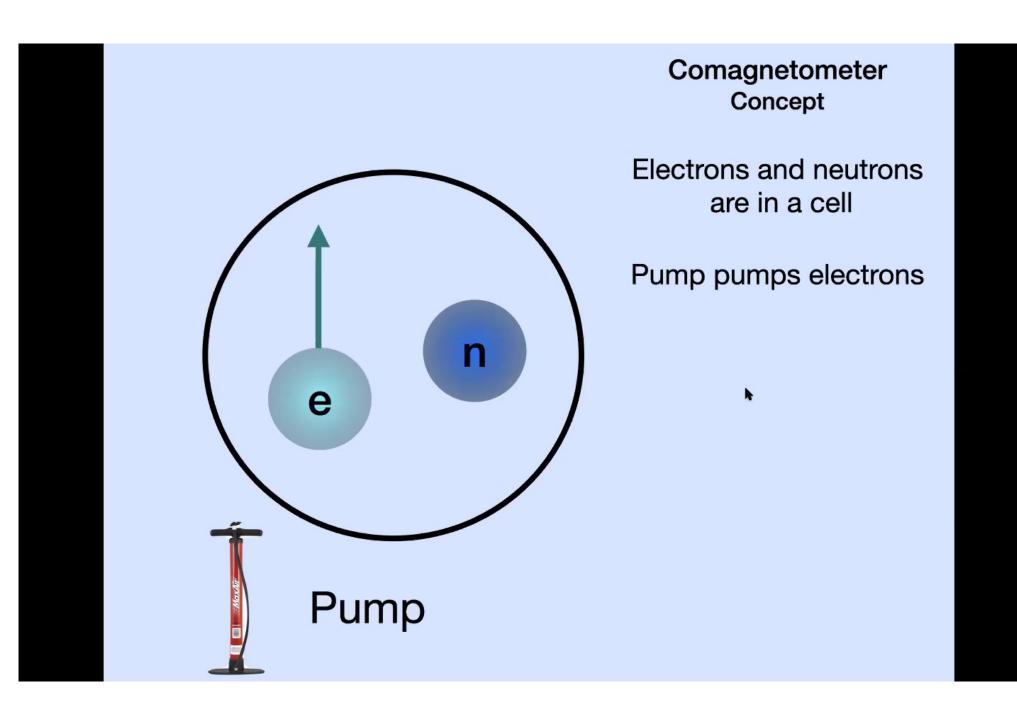
## Comagnetometer Concept

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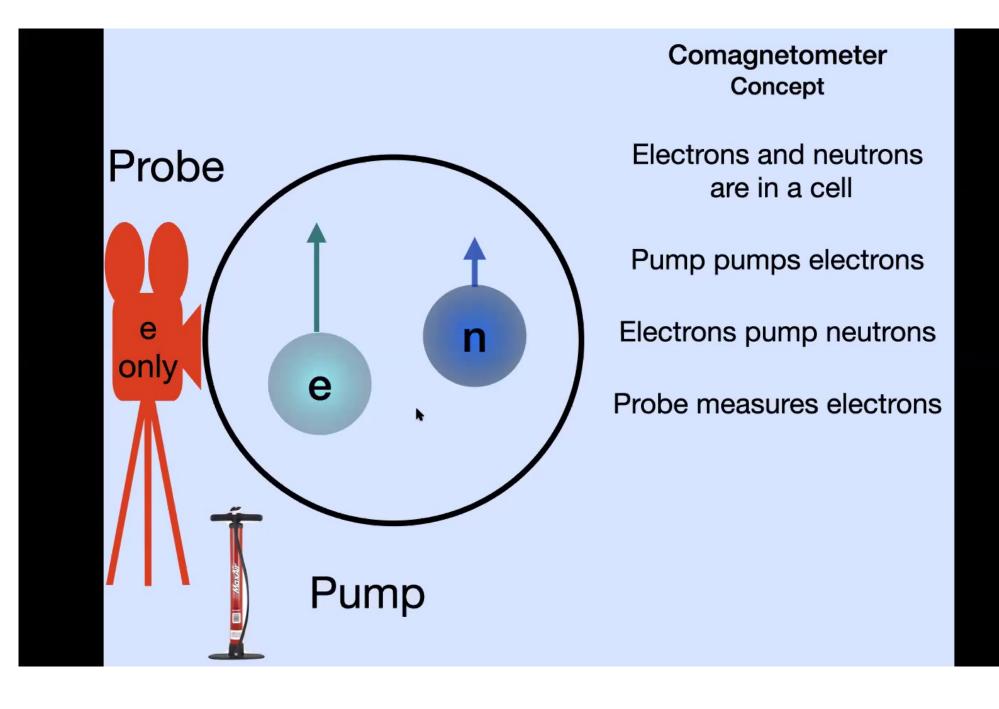
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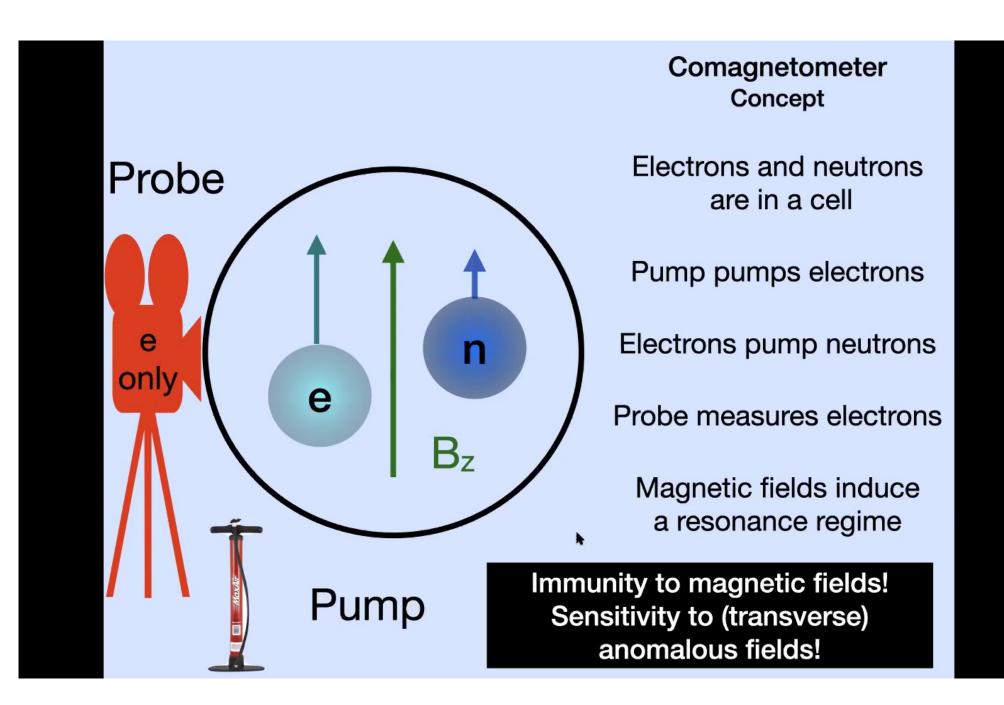
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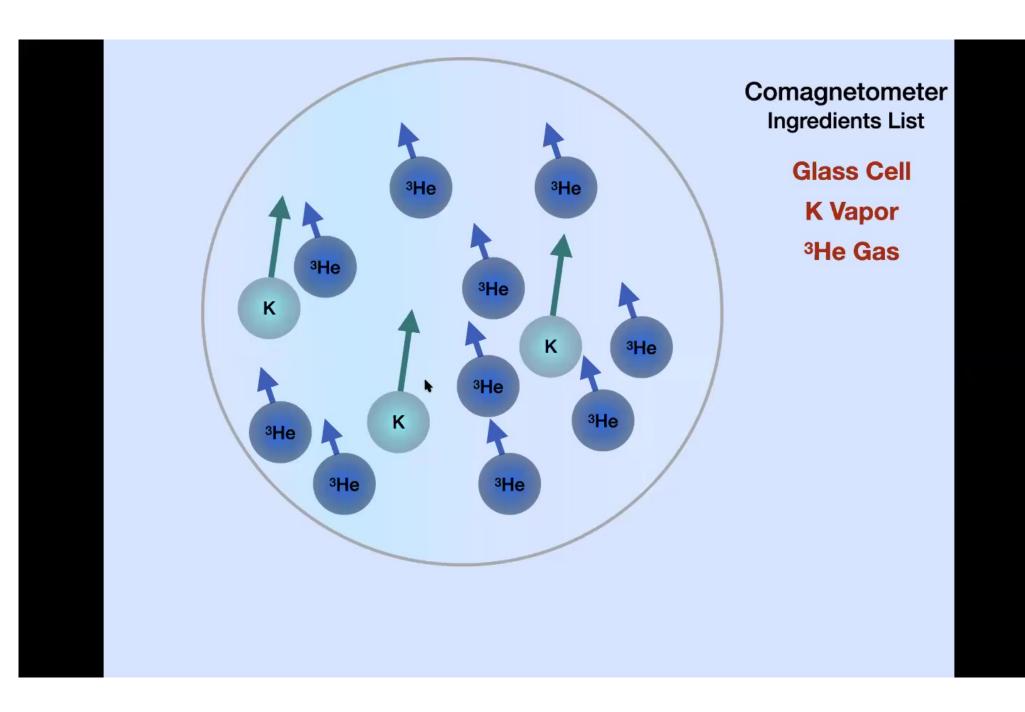
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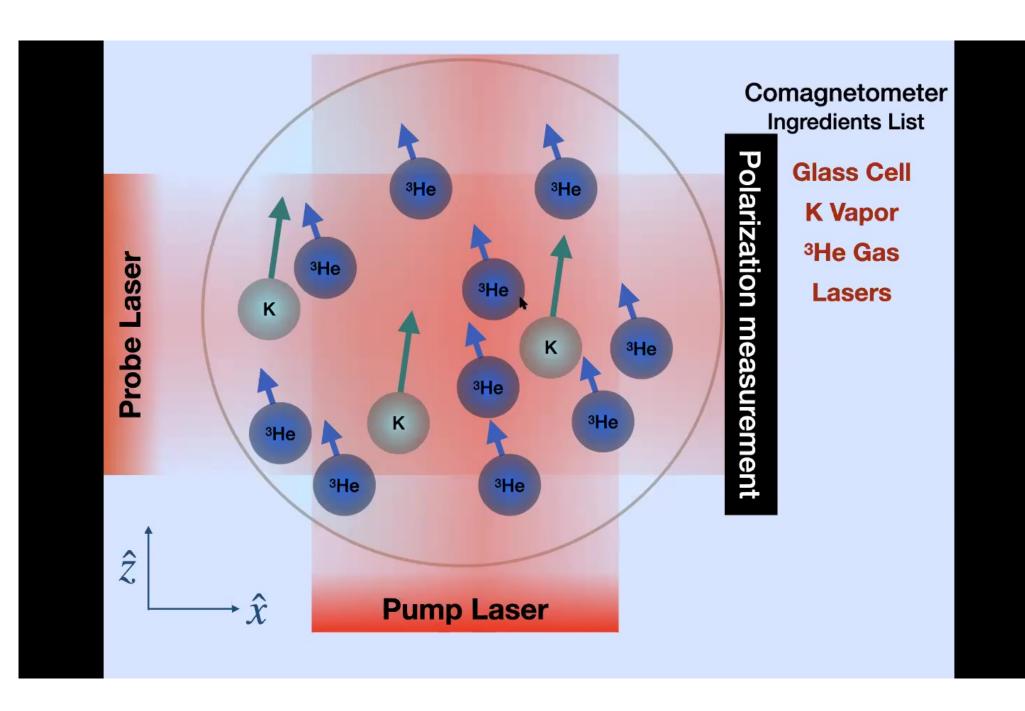
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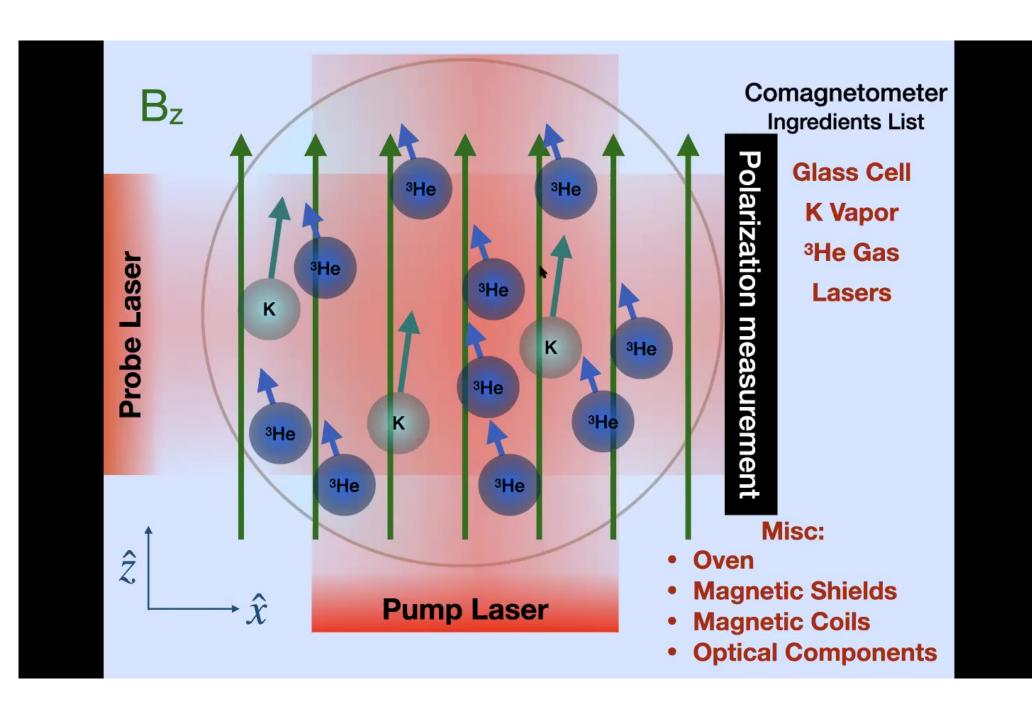
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# First Magnetometer: Alkali

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#### Alkali Magnetometer Dynamics

$$t = 0$$

 $t \approx 0.02 \text{ [sec]}$ 

Laser Pumps, Relaxation "steadies"





$$\overrightarrow{S}_K = \frac{1}{2} R_{\text{pu}} \overrightarrow{S}_{\text{pu}} - R_K \overrightarrow{S}_K$$

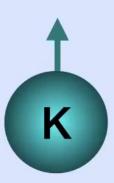
### Alkali Magnetometer Dynamics

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$$\overrightarrow{S}_{K} = \frac{1}{2} R_{\text{pu}} \overrightarrow{S}_{\text{pu}} - R_{K} \overrightarrow{S}_{K}$$

$$\downarrow \qquad \qquad \downarrow$$

$$\overrightarrow{S}_{K} \sim S_{z} \hat{z} + e^{-R_{k}t} \cdot \overrightarrow{\delta S}$$

### Alkali Magnetometer Dynamics

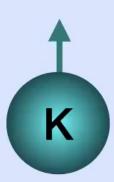
$$t = 0$$

$$t \approx 0.02 \text{ [sec]}$$

Laser Pumps, Relaxation "steadies"  $t \gtrsim 0.02 [sec]$ 

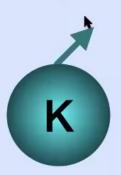
Anomalous and Regular Magnetic fields tilt the steady state solution





$$\overrightarrow{S}_{K} = \frac{1}{2} R_{pu} \overrightarrow{S}_{pu} - R_{K} \overrightarrow{S}_{K}$$

$$\overrightarrow{S}_{K} \sim S_{z} \hat{z} + e^{-R_{k}t} \cdot \overrightarrow{\delta S}$$



(Keeps wiggling with the fields)

$$+\gamma_K \left(\overrightarrow{B} + \frac{\overrightarrow{b}_K}{\gamma_K}\right) \times \overrightarrow{S}_K$$

#### Alkali Steady State

\* To leading order in important stuff.

$$\dot{\vec{S}}_K = \gamma_K \left( \overrightarrow{B} + \frac{\overrightarrow{b}_K}{\gamma_K} \right) \times \overrightarrow{S}_K - R_K \overrightarrow{S}_K + \frac{1}{2} R_{\text{pu}} \vec{s}_{\text{pu}} = 0$$

$$S^z = \underbrace{\frac{R_{pu}}{2R_K}}$$

z spin determined by the pump. No sensitivity to b.

$$S^{x} = S^{z} \frac{(\gamma_{K} B + b_{K})_{y}}{R_{K}}$$

$$S^{y} = -S^{z} \frac{(\gamma_{K}B + b_{K})_{x}}{R_{K}}$$

### Reducing Magnetic Noise

One problem was that  $\gamma_K$  is very large

can we find another atom with smaller  $\gamma$ ?

Noble Gases\*!

\* Mostly <sup>3</sup>He,Xe,<sup>21</sup>Ne.

(BONUS):

Almost no chemical interactions



Can have higher density (quieter measurements)

#### (Naive) Noble Magnetometer Dynamics

$$t = 0$$

$$t \approx \mathcal{O}(1)$$
 [hr]

Alkali Pumps, Relaxation "steadies"



Anomalous and Regular Magnetic fields tilt the steady state solution





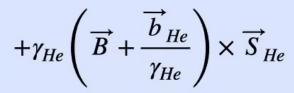


(Keeps wiggling with the fields if they vary slowly)

$$\dot{\overrightarrow{S}}_{He} = -R_{He}\overrightarrow{S}_{He} + R_{\text{pu,eff}}\overrightarrow{S}_{K}$$



$$\overrightarrow{S_{\text{He}}} \sim S_z \hat{z} + e^{-R_{\text{He}}t} \cdot \overrightarrow{\delta S}$$



#### **Noble Steady State**

\*To leading order in important stuff.

$$S^{z} = \frac{R_{pu,eff}}{R_{He}} S_{K}^{z}, S^{x} = S^{z} \frac{(\gamma_{He}B + b_{He})_{x}}{\gamma_{He}B_{z}}, S^{y} = S^{z} \frac{(\gamma_{He}B + b_{He})_{y}}{\gamma_{He}B_{z}}$$

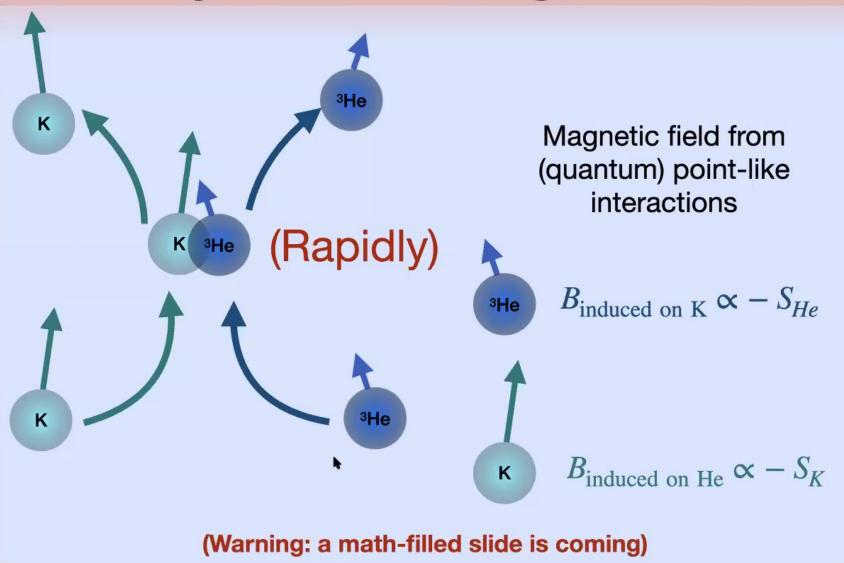
Magnetic noise is still too big!

And since  $b_{\rm He} \sim e^{im_a t}$ , we can only look at masses which are  $m_a \lesssim 1/[{\rm hr}]$ 

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#### Mutually Induced Magnetic Fields



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#### The Mixed Alkali-Noble Equations (1)

- \* To leading order in important, relevant stuff.
- \* From now on we will only discuss the transverse state, with  $v^{\perp} = v^x + i \cdot v^y$ .

$$\dot{\overrightarrow{S}}_{K} = \gamma_{K} \left( \overrightarrow{B} + \frac{\overrightarrow{b}_{K}}{\gamma_{K}} \right) \times \overrightarrow{S}_{K} - R_{K} \overrightarrow{S}_{K}$$

$$\dot{\overrightarrow{S}}_{He} = \gamma_{He} \left( \overrightarrow{B} + \frac{\overrightarrow{b}_{He}}{\gamma_{He}} \right) \times \overrightarrow{S}_{He}$$

$$\begin{pmatrix} \dot{S}_{K}^{\perp} \\ \dot{S}_{He}^{\perp} \end{pmatrix} = \begin{pmatrix} i\omega_{K} - R_{K} \\ & & i\omega_{He} \end{pmatrix} \begin{pmatrix} S_{K}^{\perp} \\ S_{He}^{\perp} \end{pmatrix}$$

$$\omega_K = \gamma_K B_z$$

$$i\omega_{He}$$
  $\left(\begin{array}{c} S_K^{\perp} \\ S_{He}^{\perp} \end{array}\right)$ 

$$\omega_{He} = \gamma_{He} B_z$$

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- \* To leading order in important, relevant stuff.
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$$\dot{\overrightarrow{S}}_{K} = \gamma_{K} \left( \overrightarrow{B} + \frac{\overrightarrow{b}_{K}}{\gamma_{K}} \right) \times \overrightarrow{S}_{K} - R_{K} \overrightarrow{S}_{K} + 2\gamma_{K} \lambda M_{\text{He}} \overrightarrow{S}_{He} \times \overrightarrow{S}_{K}$$

$$\dot{\overrightarrow{S}}_{He} = \gamma_{He} \left( \overrightarrow{B} + \frac{\overrightarrow{b}_{He}}{\gamma_{He}} \right) \times \overrightarrow{S}_{He} + 2\gamma_{He} \lambda M_{K} \overrightarrow{S}_{K} \times \overrightarrow{S}_{He}$$

$$\begin{pmatrix} \dot{S}_{K}^{\perp} \\ \dot{S}_{He}^{\perp} \end{pmatrix} = \begin{pmatrix} i\omega_{K} - R_{K} & -i\omega_{He-K} \\ -i\omega_{K-He} & i\omega_{He} \end{pmatrix} \begin{pmatrix} S_{K}^{\perp} \\ S_{He}^{\perp} \end{pmatrix}$$

$$\omega_K = \gamma_K B_z + \omega_{K0} \qquad \qquad \omega_{He} = \gamma_{He} B_z + \omega_{He0}$$

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#### The Mixed Alkali-Noble Equations (2)

\* To leading order in important stuff.

$$\begin{pmatrix} \dot{S}_{K}^{\perp} \\ \dot{S}_{He}^{\perp} \end{pmatrix} = \begin{pmatrix} i\omega_{K} - R_{K} & -i\omega_{He-K} \\ -i\omega_{K-He} & i\omega_{He} \end{pmatrix} \begin{pmatrix} S_{K}^{\perp} \\ S_{He}^{\perp} \end{pmatrix}$$

ODE with two modes and two eigenvalues.

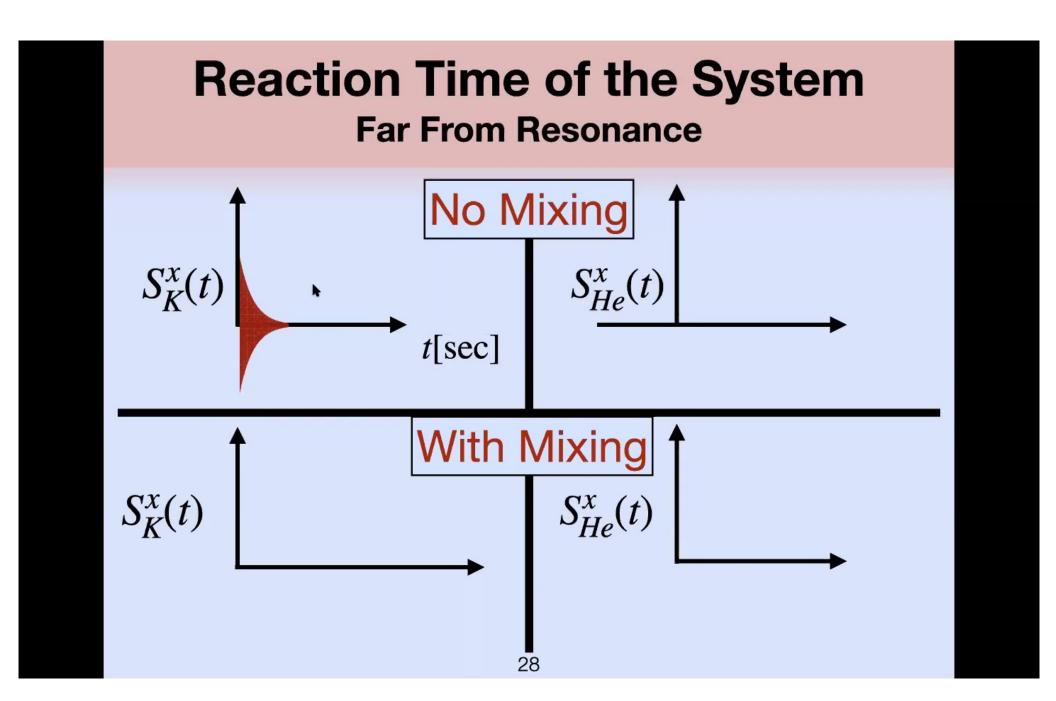
$$\omega_{K} = \gamma_{K} B_{z} + \omega_{K0} \qquad \qquad \omega_{He} = \gamma_{He} B_{z} + \omega_{He0}$$

Can we resonantly couple the two modes?

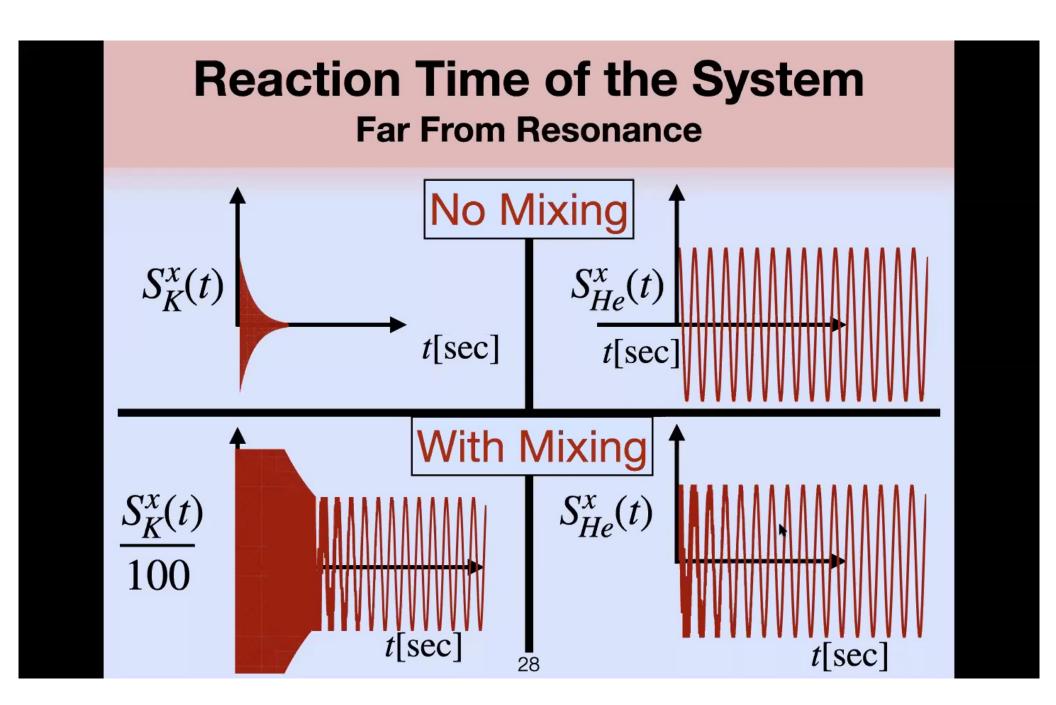
What would that do?

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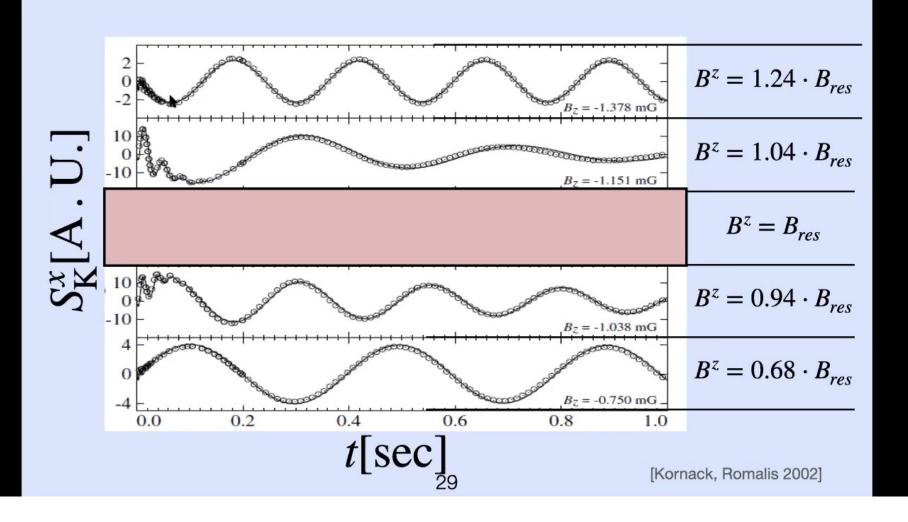


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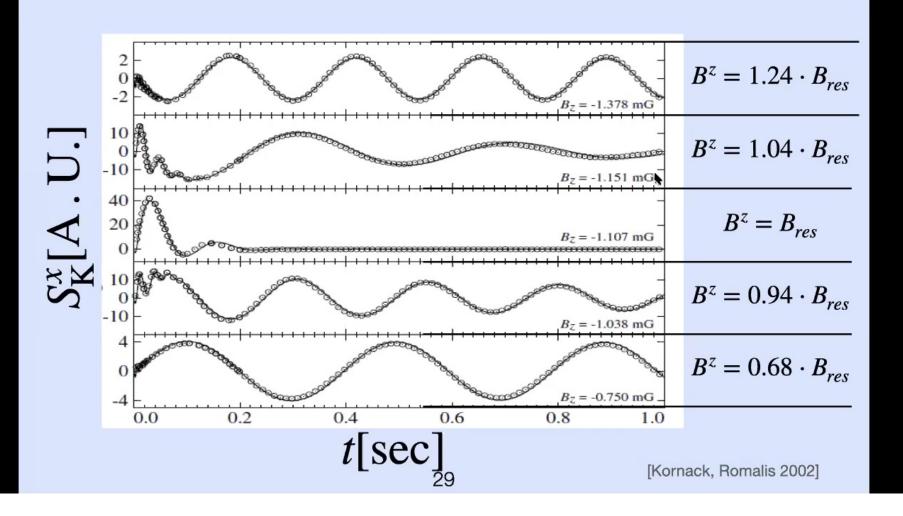
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#### The Resonance Regime (1)



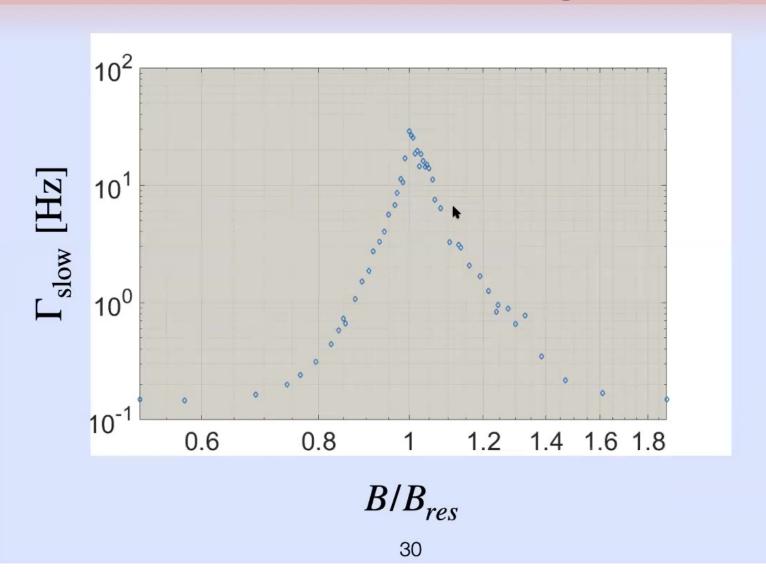
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#### The Resonance Regime (1)



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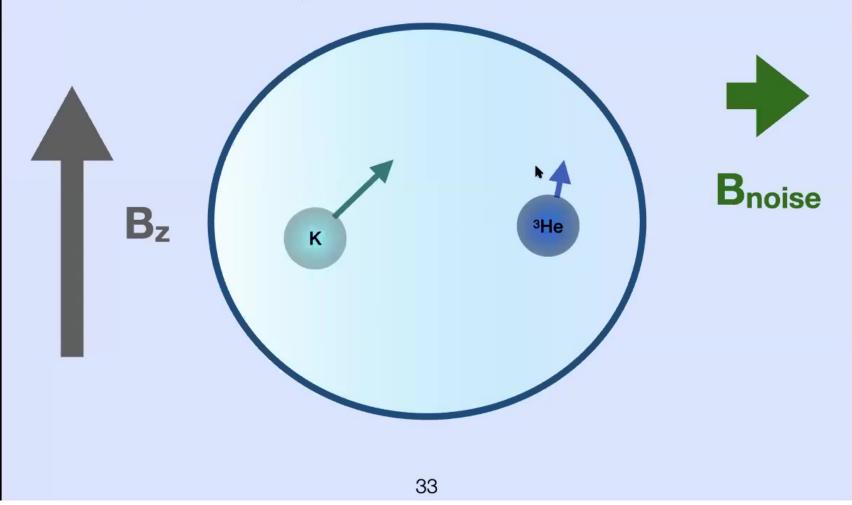
# The Resonance Regime (2)



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# Compensation Point Illustration (B)

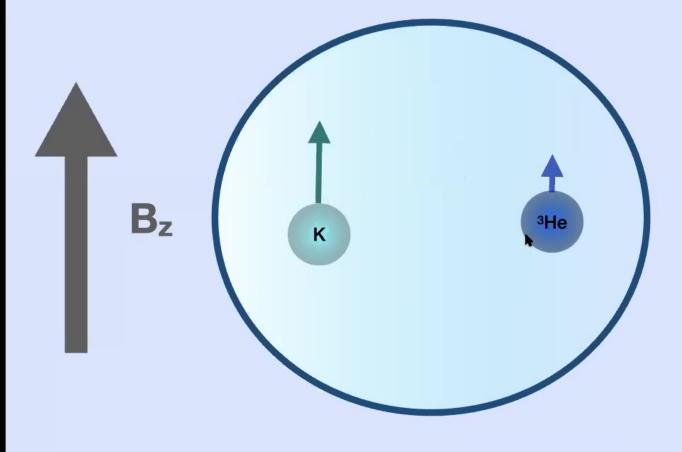
\* A 2D heuristic illustration, so some artistic freedom was taken.



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#### Compensation Point Illustration (be)

\* A 2D heuristic illustration, so some artistic freedom was taken.

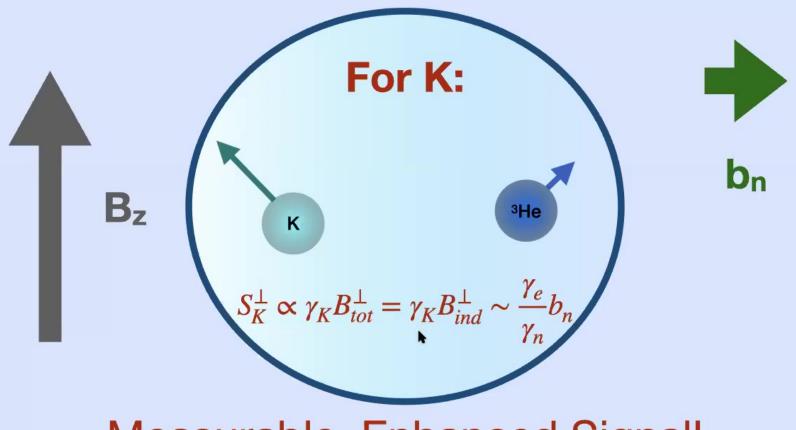


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#### Compensation Point Illustration (b<sub>n</sub>)

\* A 2D heuristic illustration, so some artistic freedom was taken.



Measurable, Enhanced Signal!

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- **Conclusions**

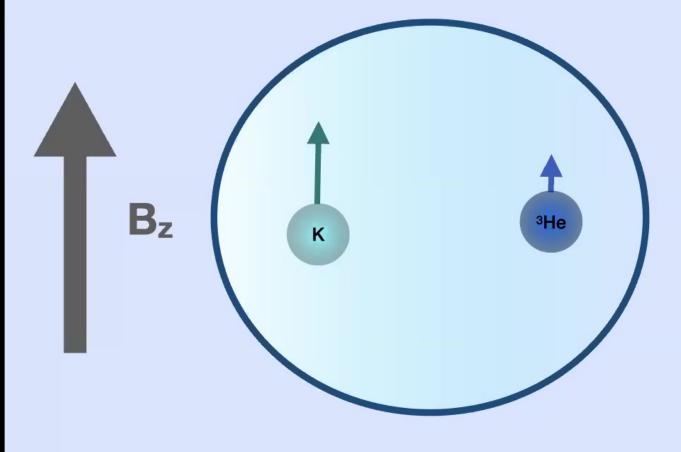
**ALP DM generate a** magnetic-like field, but it couples differently to different particles compared to the SM.

Through a special tuning of the controlled magnetic fields, comagnetometers are sensitive to anomalous fields, while immune to magnetic noise.

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#### Compensation Point Illustration (B)

\* A 2D heuristic illustration, so some artistic freedom was taken.

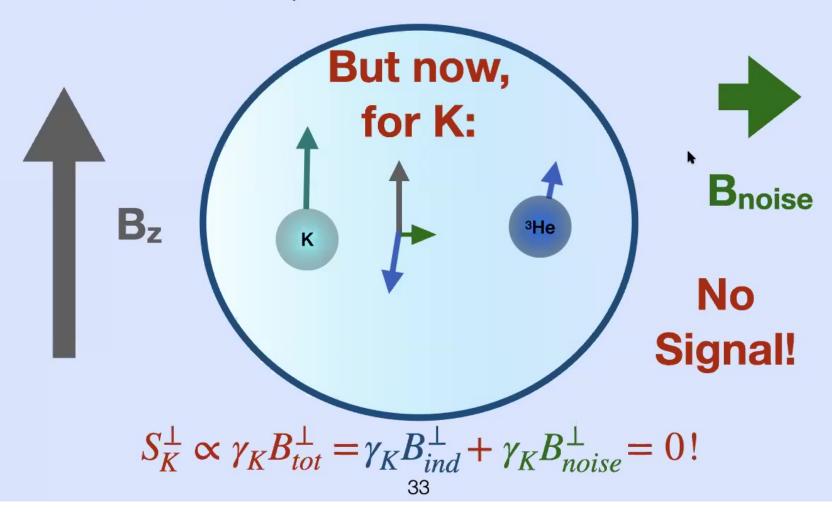


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#### Compensation Point Illustration (B)

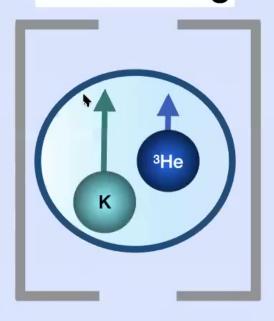
\* A 2D heuristic illustration, so some artistic freedom was taken.



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#### **Direct Detection with the Comag**

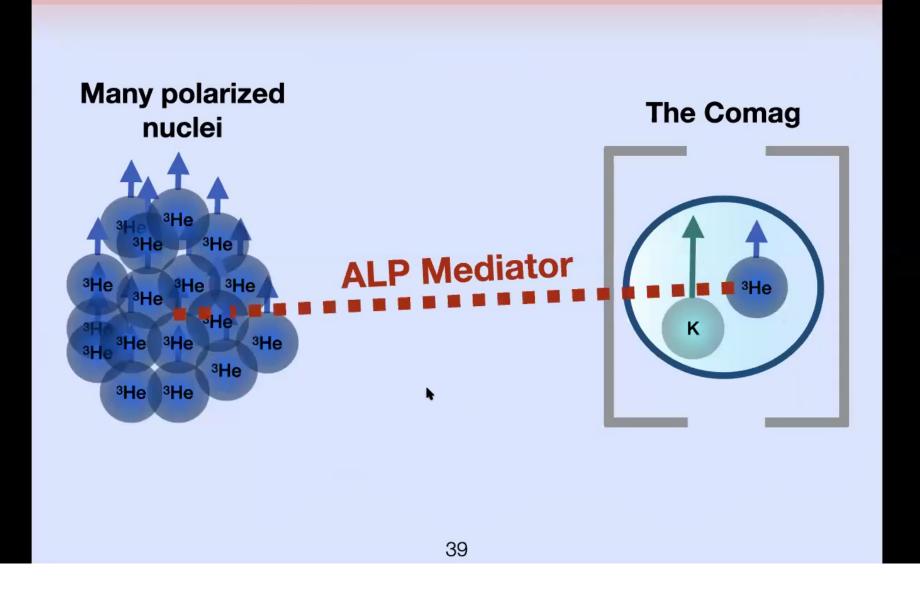
#### **The Comag**



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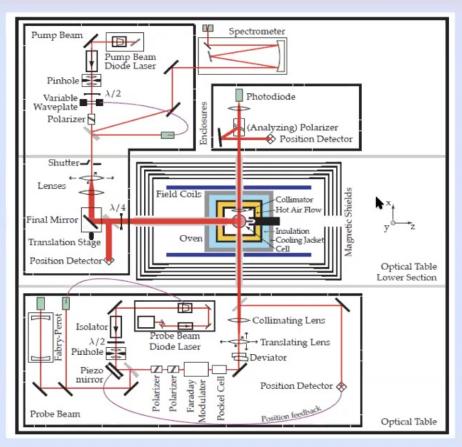
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#### Fifth Force Search



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# **Old Systems**



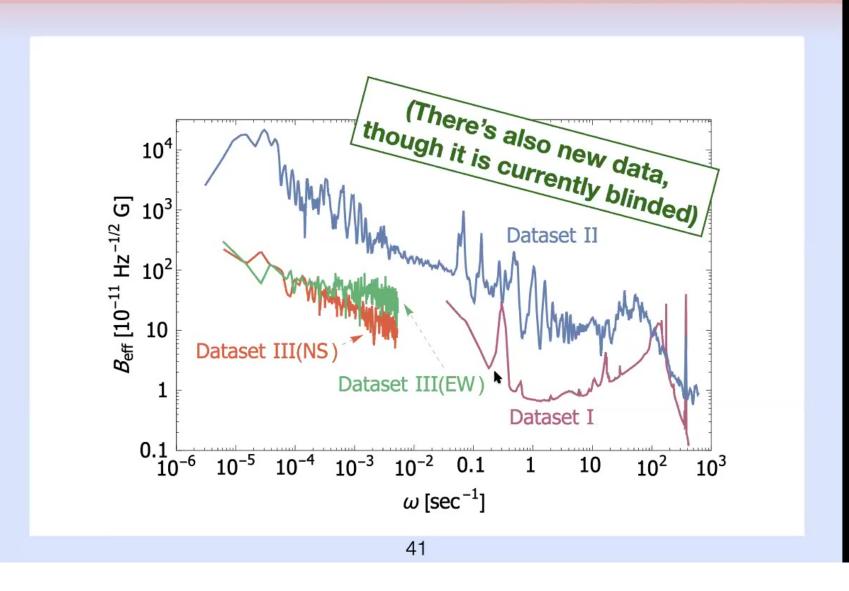




[Gergoios Vasilakis Dissertation 2011], [Justin M. Brown Dissertation 2011], [Thomas W. Kornack Dissertation 2005]

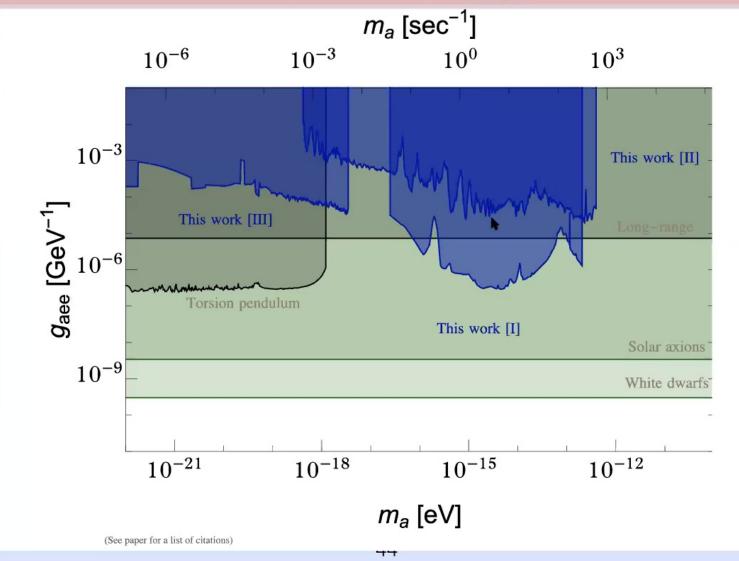
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#### The Existing Data

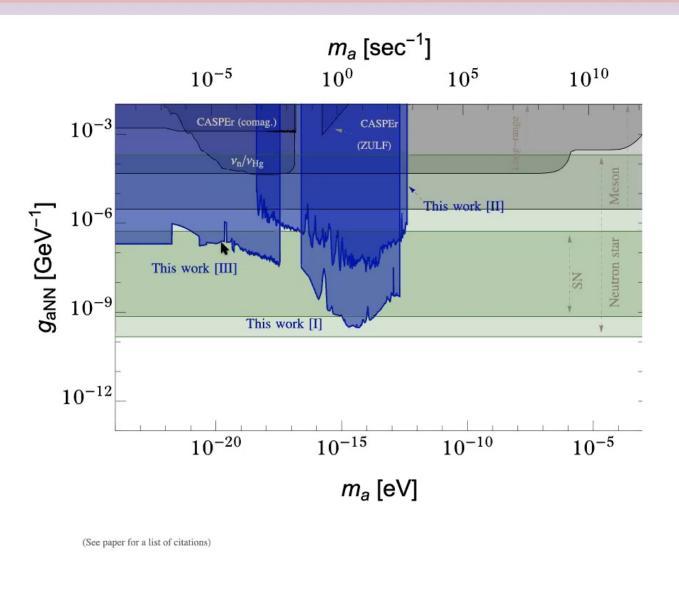


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# Results (e)



# Results and Projections (n)



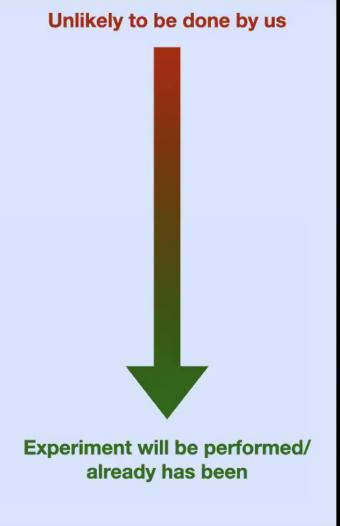
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#### **Future Prospects:**

Cold Xenon
Long Range Search

A Dedicated
Comagnetometer Search

Dedicated "Comagnetometer -Adjacent" searches



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# Four experiments should occur in the near future



1. High frequency "comagnetometer"

(Data is waiting to be analyzed)

2. Rb Xe-Xe "Co-Comagnetometer"

(Proof of concept stage, no clear schedule for full experiment)

3. Rb-Xe mid-frequency "Comangetometry"

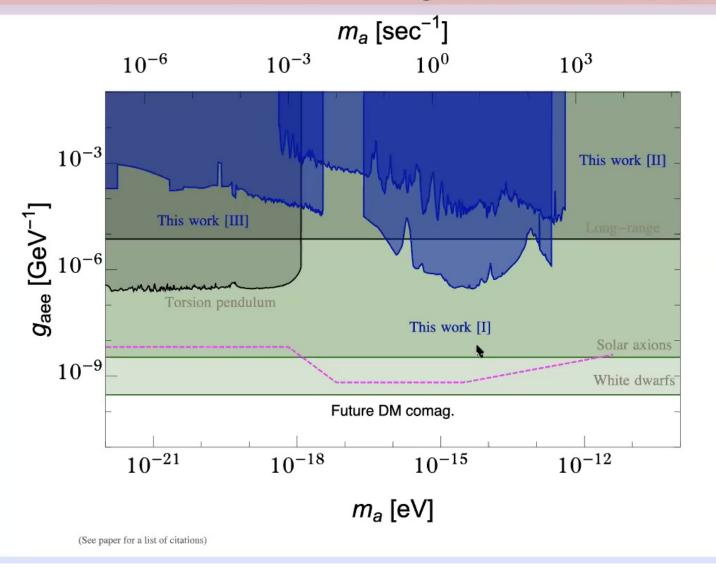
(Initial measurements taken, full data in a few months)

4. Rb-Xe low-frequency "Comangetometry"

(Proof of Concept stage, true experiment in a few months)

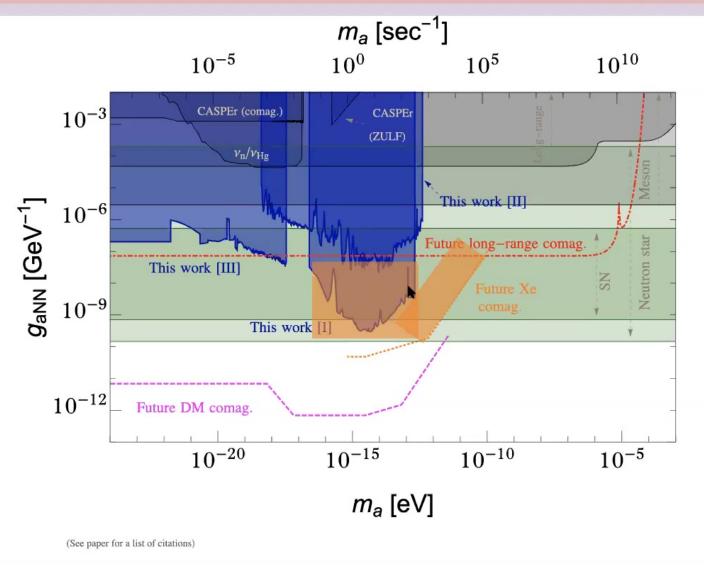
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# Results and Projections (e)



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# Results and Projections (n)



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#### Conclusions

- Comagnetometers offer a strong probe to the ALPs parameter space.
- Future experiments with comagnetometers could supersede the best existing bounds.
- If unexpected improvements occur in the experimental direction, comagnetometers might be sensitive to (asymmetric) cosmic neutrinos.

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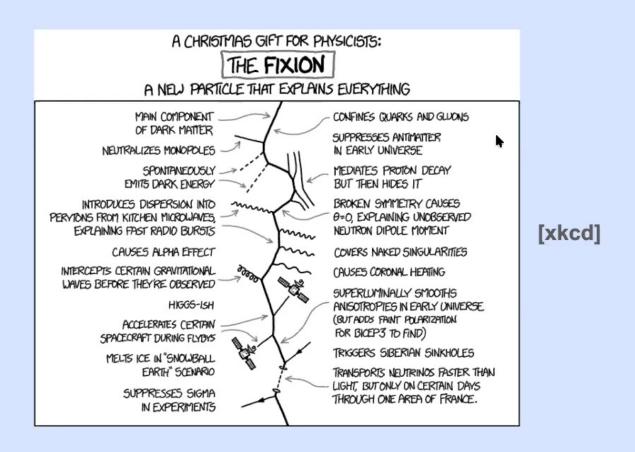
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#### Conclusions

- Comagnetometers offer a strong probe to the ALPs parameter space.
- Future experiments with comagnetometers could supersede the best existing bounds.
- If unexpected improvements occur in the experimental direction, comagnetometers might be sensitive to (asymmetric) cosmic neutrinos.
- ALPs in these masses could not only be so-called fuzzy DM, but might in fact be the QCD axion.

[L. Hui et al.:1610.08297; Marques-Tavares et al.:1803.07575]

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# **Thanks For Listening!**

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