

Title: Dark Matter meets Quantum Gravity

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Manuel Reichert

# Dark Matter meets Quantum Gravity

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Quantum Gravity 2020, 17. July 2020

CP3-Origins, SDU Odense, Denmark

MR, Juri Smirnov: PRD 101 (2020) 6; arXiv:1911.00012

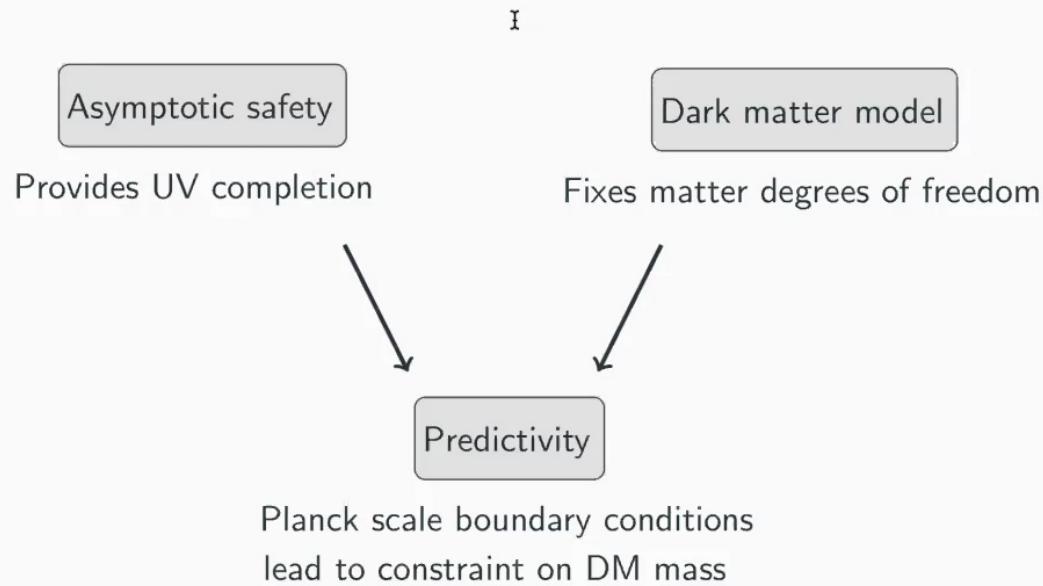
CP<sup>3</sup> Origins  
Cosmology & Particle Physics

CP3

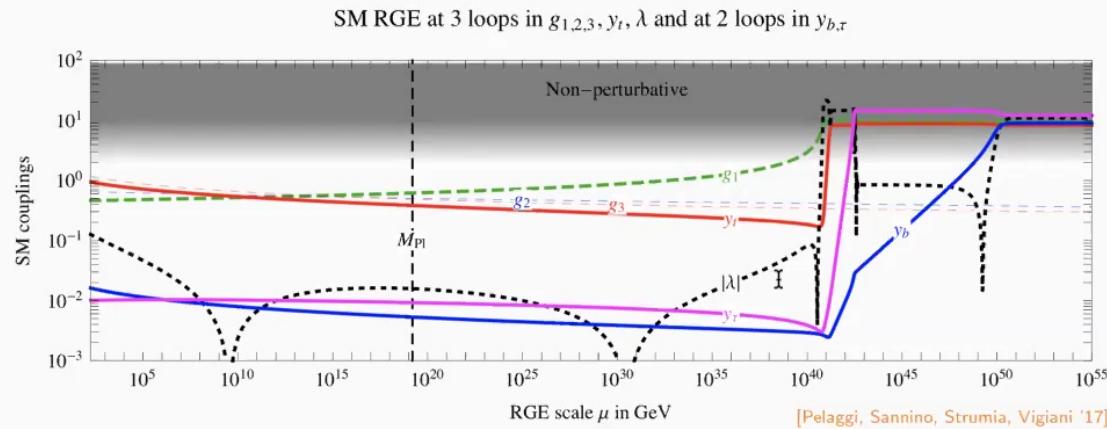
## Constraining physics beyond the SM with quantum gravity



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# The Standard Model

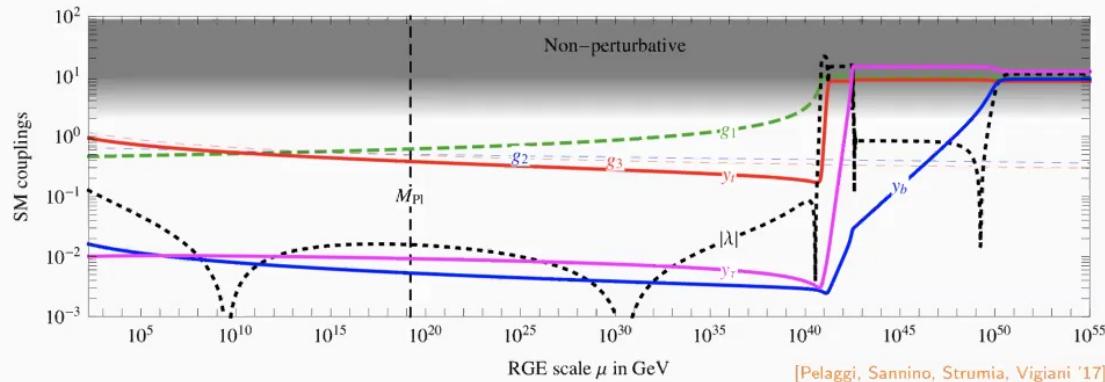


- Perturbative up to  $M_{Pl}$
- Landau pole beyond  $M_{Pl}$

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# The Standard Model

SM RGE at 3 loops in  $g_{1,2,3}, y_t, \lambda$  and at 2 loops in  $y_{b,\tau}$



- Perturbative up to  $M_{\text{Pl}}$
- Landau pole beyond  $M_{\text{Pl}}$

Asymptotic safety seems to be able  
to incorporate SM matter & cure  
Landau poles

[Daum, Harst, Reuter '09, '10]

[Doná, Eichhorn, Percacci '13; Meibohm, Pawłowski, MR '15]

[Christiansen, Litim, Pawłowski, MR '17; Eichhorn, Held '17, '18]

[Biemans, Platania, Saueressig '17]

[Alkhofer, Eichhorn, Held, Nieto, Percacci '20; ...]

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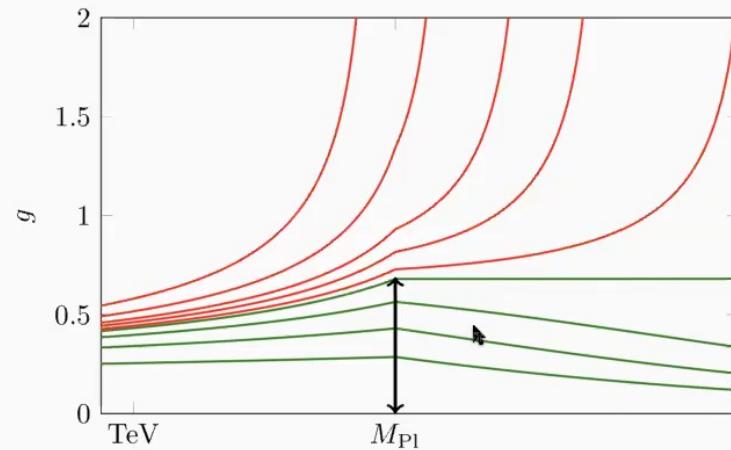


## Avoiding Landau poles



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$$\beta_g = \beta_{g,\text{matter}} - f_g g$$



Avoiding a Landau pole leads to a constrain/prediction

## Planck scale boundary conditions

Quartic scalar coupling with  $0 \leq f_\lambda$

$$\beta_\lambda = \beta_{\lambda, \text{matter}} + f_\lambda \lambda$$

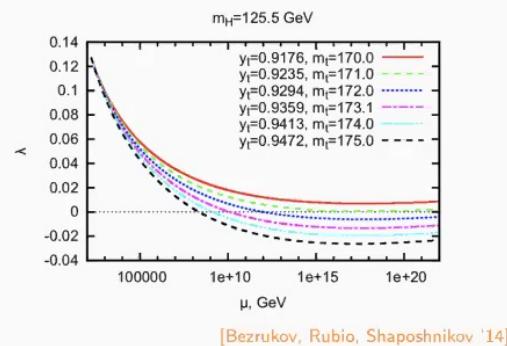
Boundary condition:  $\lambda(M_{\text{Pl}}) \approx 0$

[Eichhorn, et al '17; Pawłowski, MR, et al '18; ...]

Higgs mass

[Shaposhnikov, Wetterich '09]

$$m_h = 126 - 136 \text{ GeV}$$



[Bezrukov, Rubio, Shaposhnikov '14]

$U(1)$  gauge coupling

$$\beta_g = \beta_{g, \text{matter}} - f_g g$$

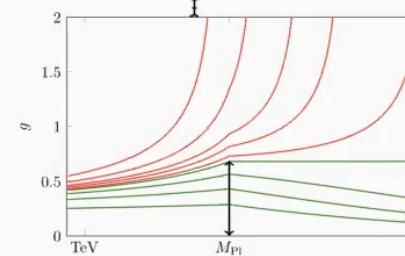
$0 \leq f_g \lesssim 0.04$  [Daum, et al '10; Folkerts, et al '12]

[Christiansen, MR, et al '17; Eichhorn, et al '17; ...]

Boundary condition at one loop

$$g(M_{\text{Pl}}) \leq \sqrt{\frac{f_g}{\beta_{g, \text{1-loop}}}}$$

[Eichhorn, Versteegen '17]



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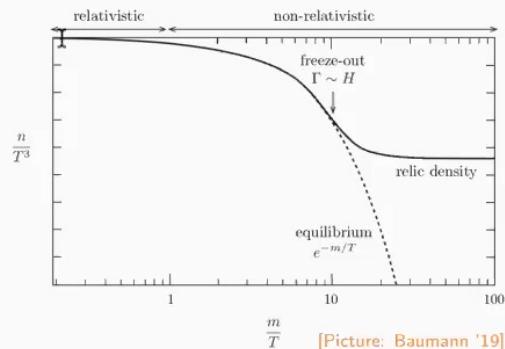


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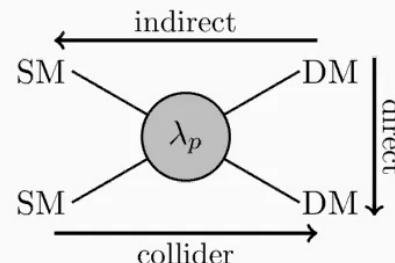
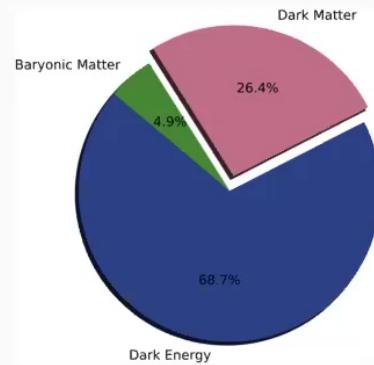
# Dark Matter

## Assumptions

- DM is a particle
- Production via freeze-out



Determines annihil. cross section



Example: Higgs portal  $\lambda_p H^\dagger H S^* S$

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## Lagrangian of the dark sector

$$\mathcal{L}_D \sim \frac{1}{2} D_\mu S D^\mu S^* + \lambda_p H^\dagger H S S^* + \lambda_S (S S^*)^2 + \frac{m_S^2}{2} S S^*$$

$$+ i \bar{\psi} \not{D} \psi + M_\psi \bar{\psi} \psi + y_\psi S \bar{\psi} \psi^c$$

$$+ \frac{1}{4} F_{\mu\nu}^X F_X^{\mu\nu} + \frac{\epsilon}{2} F_{\mu\nu}^Y F_X^{\mu\nu} + \frac{M_{Z'}^2}{2} (Z'_\mu - \partial_\mu \zeta)^2$$

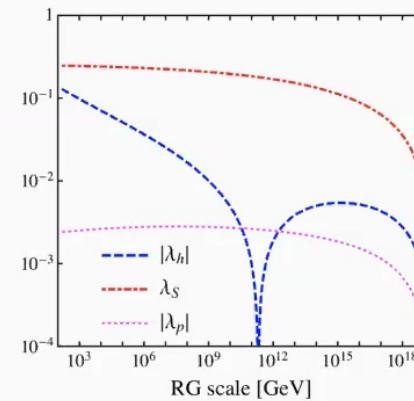
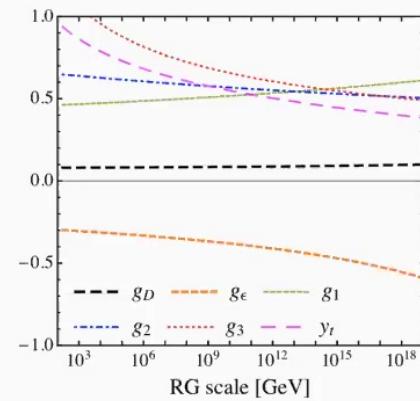
$\sim$  Higgs-portal-scalar + dark-fermion + dark- $U(1)_X$ -with-dark-photon

- Dark matter candidate:  $S$  or  $\psi$  (depending on mass hierarchy)
- Dark  $U(1)_X$  with coupling  $g_D$  generates  $\lambda_p$  below  $M_{\text{Pl}}$
- Vector-like fermion  $\psi$  for vacuum stability of  $S$
- Mass of dark photon  $Z'$  via Stueckelberg mechanism
- Kinetic mixing between  $U(1)_Y$  and  $U(1)_X$  via coupling  $g_\epsilon$

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## Constraining the portal coupling



Predictivity:

- $\lambda_i(M_{\text{Pl}}) \approx 0$
- $g_D(M_{\text{Pl}})$  and  $g_\epsilon(M_{\text{Pl}})$  bounded

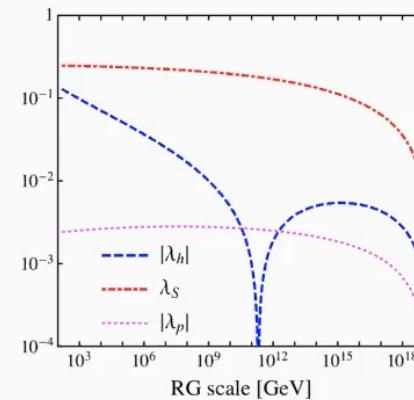
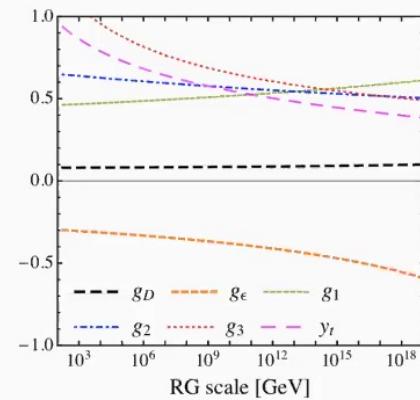


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## Constraining the portal coupling



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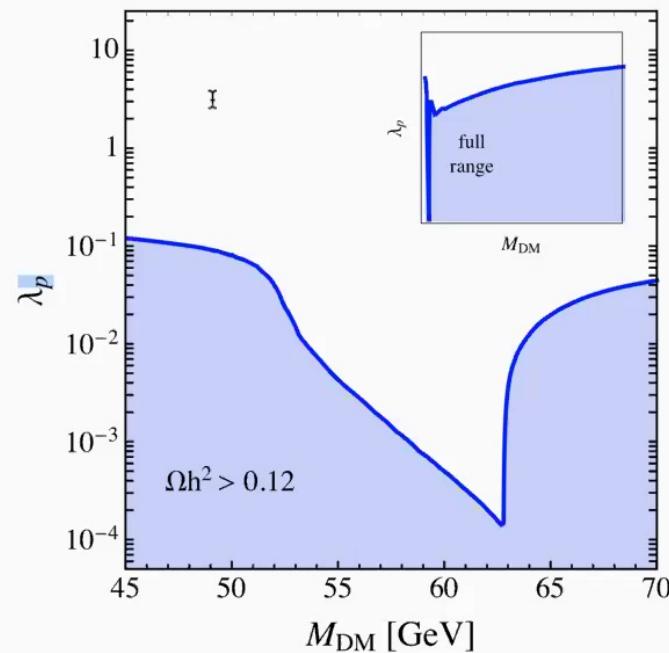
$$m_h = (125 \pm 10) \text{ GeV} \quad \longleftrightarrow \quad |\lambda_p(\text{TeV})| \leq 0.15$$

$$m_h = (125 \pm 1) \text{ GeV} \quad \longleftrightarrow \quad |\lambda_p(\text{TeV})| \leq 0.10$$

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## Scalar dark matter

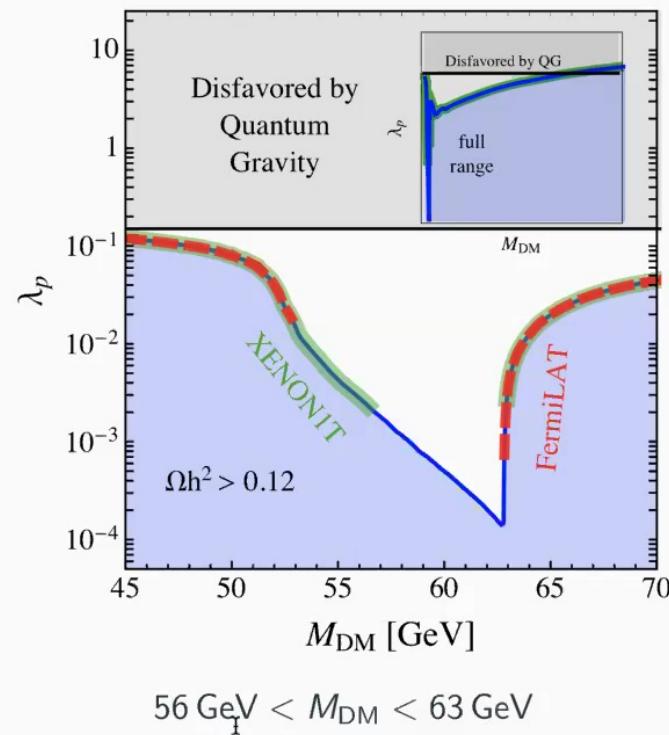


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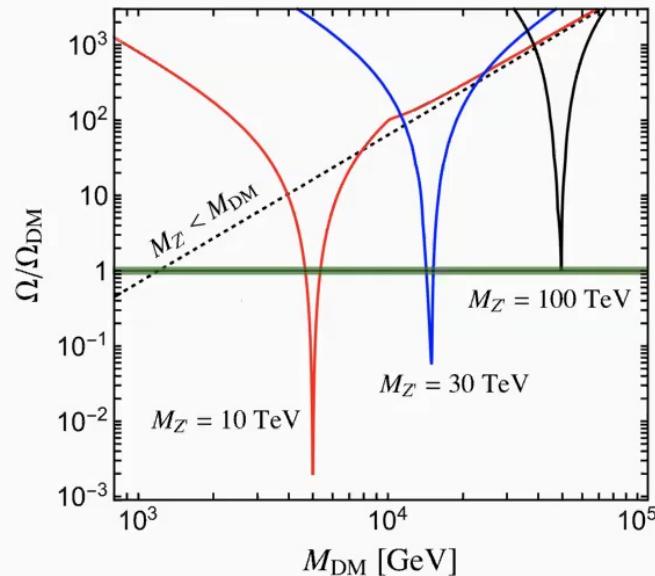


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## Scalar dark matter



## Fermionic dark matter



Non-resonant  $M_{Z'} < M_{\text{DM}}$ :  $M_{\text{DM}} < 2 \text{ TeV}$

Resonant  $M_{Z'} > M_{\text{DM}}$ :  $M_{\text{DM}} < 50 \text{ TeV}$

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## Summary

- Search for simple DM models compatible with asymptotic safety
- Boundary conditions at  $M_{\text{Pl}}$  lead to constraints on the DM mass
- Scalar Higgs portal  $56 \text{ GeV} < M_{\text{DM}} < 63 \text{ GeV}$
- Fermionic dark matter  $M_{\text{DM}} < 50 \text{ TeV}$
- All models are experimentally testable
- What can we learn about quantum gravity from the experimental tests of these models?

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Thank you for your attention