

Title: PSI 2019/2020 - Gravitational Physics - Lecture 10

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LECTURE 10 Black hole thermodynamics

Return to the Kerr metric: $a = J/M$

$$ds^2 = \frac{\Delta}{\Sigma} [dt^2 - a \sin^2 \theta d\phi]^2 - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2$$

$$- \frac{\sin^2 \theta}{\Sigma} [(r^2 + a^2) d\phi - a dt]^2$$

$$\Sigma = r^2 + a^2 \cos^2 \theta$$

$$\Delta = r^2 + a^2 - 2GM/r$$

ics

Area $A = 4\pi(r_+^2 + a^2)$

$$r_+ = GM + \sqrt{G^2M^2 - a^2}$$

$\cos^2\theta$

$-2GM/r$

$$\text{Area } A = 4\pi(r_+^2 + a^2)$$

$$r_+ = GM + \sqrt{G^2M^2 - a^2}$$

What happens if we throw a particle into black hole?

Adding δm , δa - Area

thm tells us $\delta A \geq 0$

$$\text{Area } A = 4\pi(r_+^2 + a^2)$$

$$r_+ = GM + \sqrt{G^2M^2 - a^2}$$

What happens if we throw a particle into black hole?

Adding δm , δa - Area

thm tells us $\delta A \geq 0$

$$\delta A = 8\pi(r_+ \delta r_+ + a \delta a)$$

$$= 8\pi r_+ \left[\delta(GM) + \frac{GM\delta(GM) - a\delta a}{r_+ - GM} \right] + 8\pi a \delta a$$

$$= \frac{8\pi r_+^2 G \delta M}{r_+ - GM} - \frac{8\pi G M a \delta a}{r_+ - GM}$$

Now, $a = J/M$

$$\delta a = \frac{\delta J}{M} - \frac{J \delta M}{M^2}$$

& define $\Omega = \frac{a}{r_+^2 + a^2}$ angular
velocity of
horizon.

δA

$$\delta U = \frac{8\pi}{r_+ - 9M} \left[r_+^2 q \delta M - \frac{qMa\delta J}{M} + qMa \frac{J\delta M}{M^2} \right]$$

$$= \frac{8\pi}{r_+ - 9M} \left[(r_+^2 + a^2) q \delta M - (r_+^2 + a^2) q \Omega \delta J \right]$$

$$\Rightarrow \delta M = \frac{(r_+ - 9M)}{8\pi q (r_+^2 + a^2)} \delta A + \Omega \delta J.$$

$$\delta M - \frac{GMa\delta J}{M} + GMa \frac{\overset{aM}{J}\delta M}{M^2}$$

$$^2) \quad G\delta M - (r_+^2 + a^2)G\Omega\delta J$$

$$\delta A + \Omega\delta J.$$

cf. $dU = TdS - pdV + \mu_i dQ_i$

Suggests thermodynamic interpretation with

$$S = cA/G$$

$$T = \frac{(r_+ - GM)}{8\pi c(r_+^2 + a^2)}$$

Alternative view: "path integral" approach

Partition fn $Z \sim \text{tr} e^{-\beta H}$ $\beta = 1/T$

$$H = \int d^3x \mathcal{H} = \frac{1}{\beta} \int d^3x dt \mathcal{H}$$

Alternative view: "path integral" approach.

$$\text{Partition fn } Z \sim \text{tr} e^{-\beta H} \quad \beta = 1/T$$

$$H = \int d^3x \mathcal{H} = \frac{1}{\beta} \int d^3x dt \mathcal{H}$$

Replace βH by I_E - Euclidean action.

Euclidean geometry?

$$t \rightarrow i\tau \quad g_E = -g_L(i\tau)$$

$$\Rightarrow I_E = \int_M \frac{R\sqrt{g}}{16\pi G} d^4x + \int_{\partial M} \frac{K\sqrt{h}}{8\pi G} d^3x$$

Expect saddle points to dominate path integral, i.e. classical solns.

Euclidean geometry?

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Expect saddle points to dominate path integral, i.e. classical solns with periodic τ .

$$= -g_L(i\tau)$$

$$= \int \frac{K\sqrt{h}}{8\pi G} d^3x$$

to dominate path
solns with periodic τ .

Take SCH

$$ds^2 = \left(1 - \frac{2GM}{r}\right) dt^2 + \frac{dr^2}{\left(1 - \frac{2GM}{r}\right)} + r^2 d\Omega^2$$

1st deal with β . Since $r = 2GM$ is nonsingular in Lorentzian geometry, it should be nonsingular here.

Focus on $r \rightarrow 2GM$.

$$\frac{dr^2}{1 - \frac{2GM}{r}} \sim 8GM d(\sqrt{r-2GM})^2$$

So look for a new coord

$$\rho^2 = \lambda(r - 2GM)$$

Focus on $r \rightarrow 2GM$.

$$\frac{dr^2}{1 - \frac{2GM}{r}} \sim 8GM d(\sqrt{r-2GM})^2$$

So look for a new coord

$$\rho^2 = \lambda(r - 2GM)$$

$$2\rho d\rho = \lambda dr$$

$$4\lambda(r - 2GM)d\rho^2 = \lambda^2 dr^2$$

ie $\frac{2GM}{(r-2GM)}$

$$\text{ie } \frac{2GM dr^2}{(r-2GM)} = \frac{8GM}{\lambda} dp^2$$

So set $\lambda = 8GM$, then
 p is proper distance from $2GM$.

$$\left(1 - \frac{2GM}{r}\right) \sim \frac{p^2}{8GM \cdot 2GM}$$

$$\Rightarrow ds_{\tau,r}^2 = dp^2 + p^2 d\left(\frac{T}{4GM}\right)^2$$

$$\frac{8GM}{\lambda} dp^2$$

M , then
distance from $2GM$

$$M \cdot 2GM$$

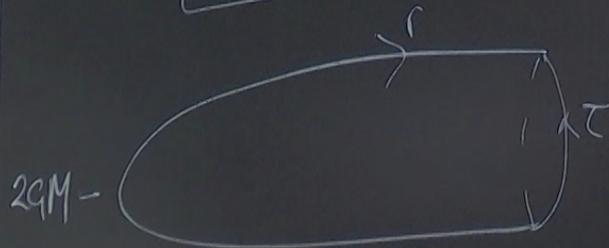
$$\rho^2 d\left(\frac{\tau}{4GM}\right)^2$$

$$\rho^2 d\theta^2$$

Compare to origin of
polar coords: i.e.

τ is periodic &

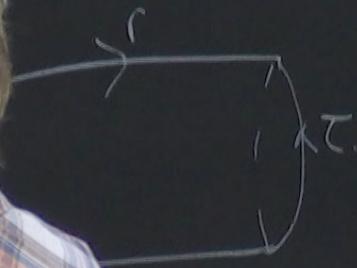
$$\Delta\tau = 8\pi GM$$



BLACK HOLE CIGAR.

polar coords: is
 τ is periodic &

$$A = 8\pi GM$$



CIGAR.

$$8\pi GM$$

Hawking temperature

Gives $c = 1/4$ so

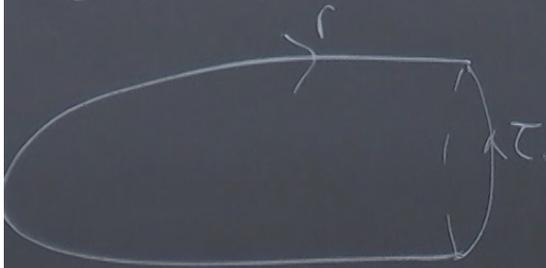
$$S = A/4G$$

$$T_{\text{Kerr}} = \frac{r_+ - r_-}{2\pi(r_+^2 + a^2)}$$

Note $T \rightarrow 0$
as $a \rightarrow GM$.

Compare to origin of
polar coords: i.e.
 τ is periodic &

$$\Delta\tau = 8\pi GM$$



BLACK HOLE CIGAR.

$$T_H = \frac{hc^3}{8\pi GM k_B} \sim 6 \times 10^{-8} K \frac{M_\odot}{M}$$

Hawking temperature

gives $c = 1/4$ so

$$S = A/4G$$

$$T_{\text{Kerr}} = \frac{r_+ - GM}{2\pi(r_+^2 + a^2)}$$

Note $T \rightarrow 0$
as $a \rightarrow GM$.

S from Z ?

$$S = \beta^2 \frac{\partial}{\partial \beta} \left[-\beta^{-1} \underbrace{\ln Z}_{I_E} \right]$$

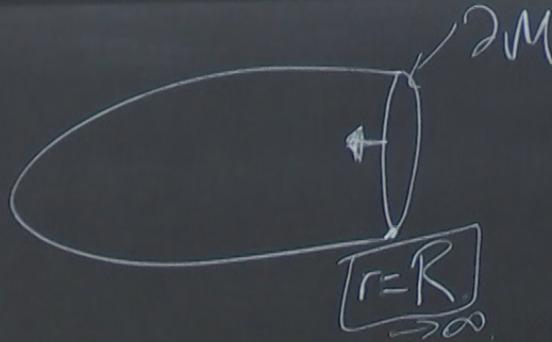
Need to compute I_{sch} . Clearly
bulk term vanishes as $R=0$,
so everything is encoded in
boundary term.

$\beta^{-1} \ln Z$
 I_E

sch. Clearly

$R=0,$

ed in



$$ds_{2n}^2 = \left(1 - \frac{2GM}{r}\right) dt^2 + R^2 d\Omega_{2n}^2$$

$$n = -\sqrt{1 - \frac{2GM}{r}} \frac{\partial}{\partial r}$$

$$K = D_a n^a = \frac{1}{r^2} \partial_r (r^2 n^r)$$

$$= \frac{2}{r} + \frac{GM n^r}{r(r-2GM)}$$

$$\sqrt{h} = R^2 \sin\theta \sqrt{1 - \frac{2GM}{r}}$$

$2M$
 $\frac{M}{R} \rightarrow \infty$

$$\frac{M}{R} dt^2 + R^2 d\Omega^2$$

$$\frac{2GM}{R} \frac{\partial}{\partial r}$$

$$\partial_r (r^2 n^r)$$

$$\frac{1}{r} \frac{\partial}{\partial r} (r^2 n^r)$$

$$\sqrt{h} = R^2 \sin\theta \sqrt{1 - \frac{2GM}{R}}$$

$$\int K \sqrt{h} = \int -R^2 \sin\theta \sqrt{1 - \frac{2GM}{R}} \left(\frac{2}{R} \sqrt{1 - \frac{2GM}{R}} + \frac{GM}{R^2 \sqrt{1 - \frac{2GM}{R}}} \right)$$

$$= -4\pi\beta [2R - 3GM]$$

divergent - even as $M \rightarrow 0$!

Renormalise by subtracting flat space value.

$$n_0 \sqrt{1 - \frac{2GM}{R}}$$

$$\sin \theta \sqrt{1 - \frac{2GM}{r}} \left(\frac{2}{R} \sqrt{1 - \frac{2GM}{R}} + \frac{GM}{R^2 \sqrt{1 - \frac{2GM}{R}}} \right)$$

$$[2R - 3GM]$$

even as $M \rightarrow 0$.

by subtracting
reference value.

$$ds_{\text{shell}}^2 = dt^2 + R^2 d\Omega^2$$

But our boundaries
must match.

$$dT^2 \leftrightarrow \left(1 - \frac{2GM}{R}\right) dt^2$$

$$\Delta T = \sqrt{1 - \frac{2GM}{R}} \Delta \tau$$

boundary term.

1. r $r(r-2GM)$.

$$\eta = -\frac{2}{\partial r} \Rightarrow K_0 = -\frac{2}{R}$$

$$\sqrt{h_0} = R^2 \sin \theta$$

$$\int K \sqrt{h_0} = -\frac{2}{R} \cdot 4\pi R^2 \cdot \beta_0$$

$$= -8\pi R \cdot \sqrt{1 - \frac{2GM}{R}} \beta_0$$

$$= -8\pi \beta R \left[1 - \frac{GM}{R} \right] + \dots$$

$$= \frac{2M}{r} + \frac{GM}{r(r-2GM)}$$

flat space value.

Hence

$$I_{\text{sch}} - I_0 = \frac{1}{8\pi G} \left[-4\pi\beta(2R-3GM) + 8\pi\beta(R-GM) \right]$$

$$= \frac{\beta M}{2} \quad \leftarrow \quad \begin{aligned} \beta &= 8\pi GM \\ &= 4\pi r_f \end{aligned}$$

$$\text{Check } S = \beta^2 \frac{\partial}{\partial \beta} \left[-\frac{1}{\beta} \cdot \frac{\beta M}{2} \right] = \frac{\beta^2}{16\pi G} + \frac{A}{4G}$$

Finally, back to Kerr:

$$GM = \frac{r_+^2 + a^2}{2r_+}$$

$$S = \frac{\pi(r_+^2 + a^2)}{g}$$

$$J = aM$$

$$J^2 = a^2 M^2 = \frac{a^2 (r_+^2 + a^2)^2}{4r_+^2 G^2} = \frac{a^2 S^2}{4\pi^2 r_+^2}$$

$$\begin{aligned} \Rightarrow \frac{4\pi^2 J^2}{S^2} + 1 &= 1 + a^2/r_+^2 \\ &= (r_+^2 + a^2)/r_+^2 \\ &= \frac{4G^2 M^2}{(r_+^2 + a^2)} = \frac{4\pi G M^2}{S} \end{aligned}$$

$$M^2 =$$

$$\frac{a^2 S^2}{4\pi^2 r_+^2}$$

$$M^2 = \frac{S}{4\pi G} \left[1 + \frac{4\pi^2 J^2}{S^2} \right]$$

Christodoulou - Ruffini

$$\begin{aligned} \Omega &= \left. \frac{\partial M}{\partial J} \right|_S = \frac{1}{2GM} \cdot \frac{2\pi J}{S} \\ &= \frac{\pi a M}{\pi(r_+^2 + a^2)M} \end{aligned}$$

$$\frac{S}{\pi G} \left[1 + \frac{4\pi^2 J^2}{S^2} \right]$$

modanlou - Ruffini

$$\begin{aligned} \frac{M}{J|S} &= \frac{1}{2GM} \cdot \frac{2\pi J}{S} \\ &= \frac{\pi a M}{\pi(r_+^2 + a^2)M} \quad \Downarrow \end{aligned}$$

$$\begin{aligned} T &= \frac{\partial M}{\partial S} \Big|_J = \frac{1}{8\pi G M} \left[1 - \frac{4\pi^2 J^2}{S^2} \right] \\ &= \frac{1}{8\pi G M} \left[1 - \frac{a^2}{r_+^2} \right] \\ &= \frac{1}{8\pi G M r_+^2} \left[r_+^2 - (r_+^2 - 2GM r_+) \right] \\ &= \frac{(r_+ - GM)}{2\pi(r_+^2 + a^2)} \quad \Downarrow \end{aligned}$$

"chemical" expressions for b.h. TDs