

Title: Central extension in gravity

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Series: Quantum Gravity

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Abstract: The asymptotic symmetry charge algebra of certain class of spacetimes could have a nontrivial central extension, which measures the non-equivariance of the charges of the large gauge transformations. The Cardy formula of the dual CFT has been famously used to derive black hole entropy. However, it remains obscure from the point of view of gravity why such a constant on the gravitational phase space could encode the information about the density of black hole micro-states, and what the degrees of freedom accounting for the black hole entropy truly are. I will discuss the ongoing efforts of understanding these questions in the covariant phase space formalism.



Central Extension in Gravity

with Slava Lysov

- There has been surprising effectiveness of assuming the validity of Cardy formula for computing BH entropy
- ? How to understand the appearance of c and the d.o.f from the phase space of gravity?

Using Cardy formula to compute BH entropy.

• Strominger 97'

Near horizon geometry locally AdS $\Lambda = -\frac{1}{l^2}$

Semiclassical $l \gg G$ large c

$$S_{\text{cardy}} = \ln Z = 2\pi \left(\sqrt{\frac{C_{\Delta}}{6}} + \sqrt{\frac{C_{\bar{\Delta}}}{6}} \right)$$

$\Delta, \bar{\Delta}$ eigenvalue L_0, \bar{L}_0

$$\begin{cases} M = \frac{1}{l} (L_0 + \bar{L}_0) = \frac{r_+^2 + r_-^2}{8Gl} \\ J = L_0 - \bar{L}_0 = \frac{r_+ r_-}{4Gl} \end{cases}$$

$$\rightarrow S = \frac{2\pi r_+}{4G}$$

200

$$S_{\text{cardy}} = \ln f = 2\pi \left(\sqrt{\frac{c}{6}} + \sqrt{\frac{c}{6}} \right) \quad \rightarrow \quad S = \frac{c}{4a}$$

2008 Kerr/CFT

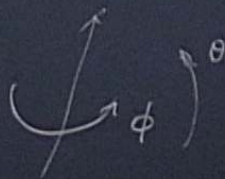
Strominger Guic Hartman Saeg etc ...

Extreme Kerr $J = GM^2$ $r_+ = M$

$$S = 2\pi M^2 = 2\pi J$$

1999 Bardeen Horowitz $\hat{r} \sim M$

$$r = \frac{\hat{r} - M}{\lambda M} \quad \lambda \rightarrow 0$$



$$\phi \sim \phi + 2\pi$$

$$\mathcal{L} = \underline{E(\phi)} \partial \phi - r \underline{E'(\phi)} \partial r$$

$$\{L_m, L_n\} = (m-n) L_{m+n} + \frac{c}{12} m^3 \delta_{m+n,0}$$

$$c = 12J \quad S = \frac{\pi^2}{3} c T = 2\pi M r_+$$

Covariant Phase Space Formalism

$$S = \int_M L(\phi, \partial_\mu \phi, \dots) + \int_{\partial M} l \quad \text{d-dim} \quad \text{Solution space}$$

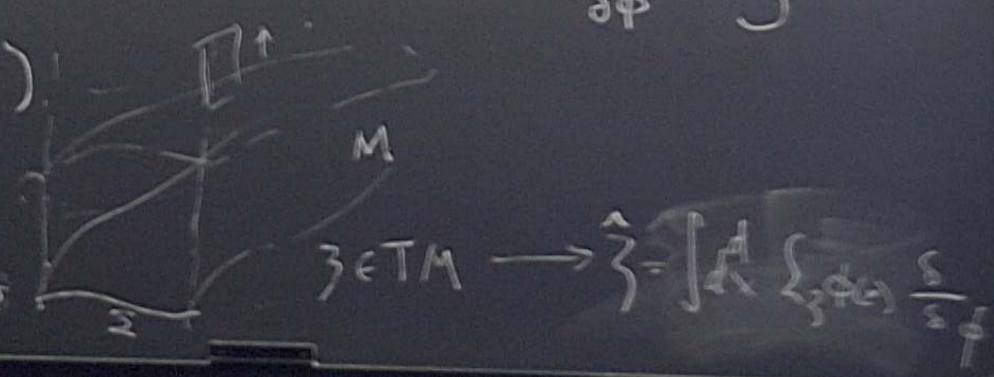
$$\text{E.O.M} + \int_M d\theta(\phi, \delta\phi) \quad \delta\phi|_{\partial M} = 0$$

$\delta\phi$ S

$$W = \delta(\phi, \delta\phi) = \delta_{ij} \theta(\phi, \delta\phi) \frac{\delta\phi^j}{\delta\phi^i}$$

eg. QED: $W = \delta A_\mu + \delta F$

from the phase space of gravity!



semiclassical $l \gg \hbar$ $\xrightarrow{\text{large } c}$

$$S_{\text{ordy}} = \ln f = 2\pi i \left(\sqrt{\frac{c_1 \Delta}{6}} + \sqrt{\frac{c_2 \Delta}{6}} \right) \rightarrow S = \frac{2\pi V + 4\pi \ell}{4a}$$

$$j_3 = \theta(\phi, \Sigma_3 \phi) - \tau_3 L$$

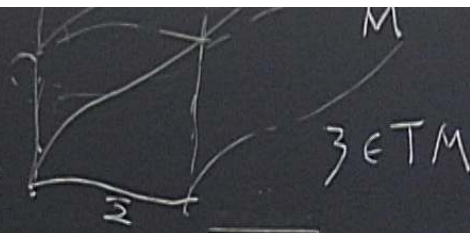
$$\downarrow \delta_3 L = \tau_3 L$$

$$d j_3 = 0$$

$$\downarrow \exists Q_3^N \quad j_3 = dQ_3^N$$

$$\omega = \delta(\phi, \delta\phi) = \delta_{\mathbb{L}} \theta(\phi, \delta\phi)$$

eg: QED: $\omega = \delta A_{\mu} + \delta T$

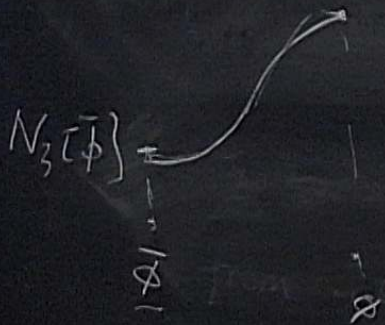


$$Z \in TM \rightarrow \hat{Z} = \int d^d x \left\{ \phi(x) \frac{\delta}{\delta \phi} \right\}$$

$$\delta Q_3 = \int_{\Sigma} \mathcal{L}_3 \omega = \int_{\Sigma} d \left(\delta Q_3^N - \mathcal{L}_3 \theta(\phi, \delta\phi) \right)$$

$$k_3 [\delta\phi, \phi]$$

$$\exists Q_3 \quad \delta Q_3 = \delta Q_3$$



$$Q_3[\phi; \bar{\phi}] = \int_{\bar{\phi}}^{\phi} \int_{\Sigma} k_3 [\delta\phi, \phi] + N_3[\bar{\phi}]$$

Sufficient condition that $K_{3,\eta}$ trivial

• \mathfrak{g} - semi-simple

• $\Omega = -\delta \oplus$ & $\underline{\Sigma}_3 \oplus = 0$

$$\{ \mathbb{Q}_3, \mathbb{Q}_\eta \} = -L_{\hat{3}} L_{\hat{\eta}} \delta \oplus = -L_{\hat{3}} \Sigma_{\hat{\eta}} \oplus + L_{\hat{3}} \delta L_{\hat{\eta}} \oplus$$

$$= -L_{\hat{3}} L_{\hat{\eta}} \oplus + L_{\hat{\eta}} L_{\hat{3}} \oplus = L_{[L_{\hat{3}}, L_{\hat{\eta}}]} \oplus$$

ion that $K_3 \cdot \eta$ trivial

$$\& \underline{\sum_3 \oplus} = 0$$

$$= - \sum_3 \sum_{\hat{\eta}} \delta \oplus = - \sum_3 \sum_{\hat{\eta}} \oplus + \sum_3 \delta \sum_{\hat{\eta}} \oplus$$

$$= - \sum_3 \sum_{\hat{\eta}} \oplus + \sum_{\hat{\eta}} \sum_3 \oplus = \sum_{[\hat{\eta}, 3]} \oplus = Q[\eta, 3]$$

eg: QED: $\omega = \delta A_\mu + \delta T$ $\frac{\partial}{\partial z}$

If $\Phi_g(\Theta(\phi, \delta\phi)) \neq \Theta(\Phi_g\phi, \delta\Phi_g\phi)$

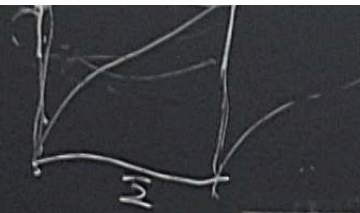
$\exists \parallel \partial \bar{z} \quad \int_{\Sigma} (\Theta + \Theta_{\partial \bar{z}}) = 0$

$\int_{\Sigma} \neq 0$

$\int_{\Sigma} = \int \varepsilon(\phi) \cdot r \partial_r$

$$W = \int (\phi, \delta\phi) = \int_{\Sigma} \Theta(\phi, \delta\phi)$$

eg: QED: $W = \int A \wedge *F$



$$Z \in TM \longrightarrow \hat{Z} = \int d^d x \{ Z \phi \}$$

When $\int_{\Sigma} \Theta \neq 0$

$$\int_{\Sigma} \Omega = 0$$

$\int_{\Sigma} \Theta = g$ we can do a canonical trans. $\Theta \rightarrow \Theta'$ if

\exists field space 1 form $\varphi = \delta\alpha$ $\mathcal{G} = -\delta(\tau_{\Sigma} \delta\alpha)$

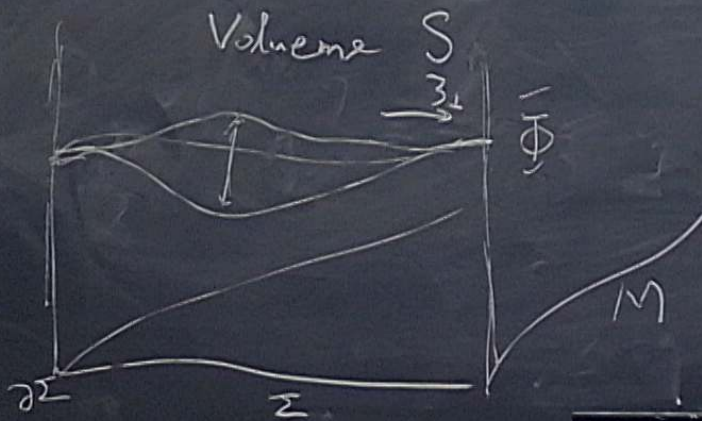
$$\Theta \rightarrow \Theta + \delta\alpha$$

$$\int_{\Sigma} \Theta = 0 \Rightarrow \delta \tau_{\Sigma} \Theta = -\tau_{\Sigma} \delta\Theta = -\tau_{\Sigma} \Omega, \quad Q_{\Sigma} = \tau_{\Sigma} \Theta$$

$$\mathcal{L}_3 \textcircled{-1} = \mathcal{L}_3 + \mathcal{I}_3$$



$$K_{3,\eta} = \int_{\Sigma} \hat{\Gamma}_3 \cdot \hat{\eta} - \int_{\Sigma} \hat{\Gamma}_3 \mathcal{I}_3 + \int_{\eta} \mathcal{I}_\eta$$



Non-equivariance

Lie group G acts on \mathcal{P} $\bar{\Phi}_g$

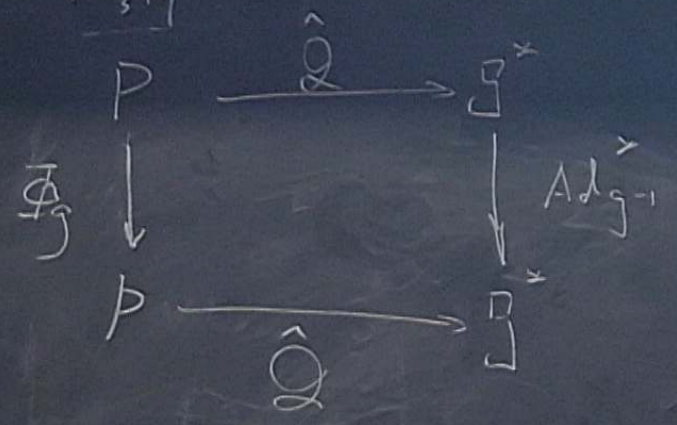
Lie algebra element $\zeta \in \mathfrak{g} \rightarrow \zeta_p$ on \mathcal{P}

Under $\bar{\Phi}_g^{-1} \zeta_p = (\text{Ad}_{g^{-1}} \zeta)_p$

$Q_\zeta[\phi] : \hat{Q} : \mathcal{P} \rightarrow \mathfrak{g}^*$ $(\forall \zeta \in \mathfrak{g}, -\delta \langle \hat{Q}, \zeta \rangle = \tau_\zeta \Omega)$

Casimir fn.

$\exists K_{3,9}$



$$Q_3[\Phi_g \cdot \phi] \neq Q_{\text{Ad}_g^{-1}3}[\phi]$$

$$Q_3[\Phi_g \cdot \phi] - Q_3[\phi] - (Q_{\text{Ad}_g^{-1}3}[\phi] - Q_3[\phi])$$

$\neq 0$

