Title: Swampland Constraints on Neutrino Masses and Dark Energy

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# Swampland Constraints on Neutrino Masses and Dark Energy



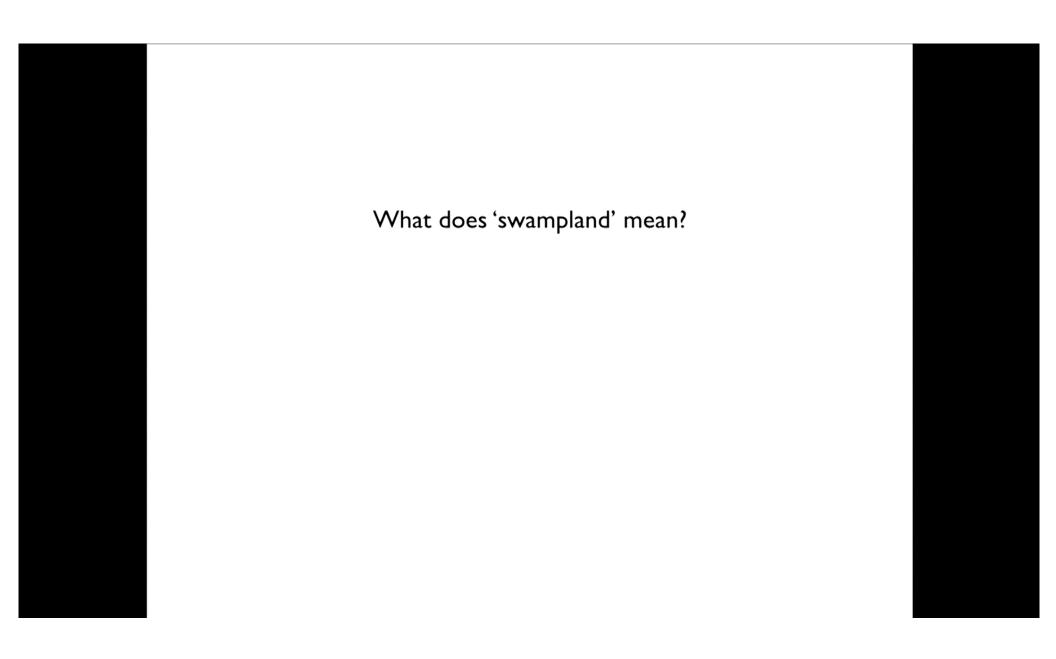
#### Irene Valenzuela

Cornell University

In collaboration with Luis Ibanez and Victor Martin-Lozano (1706.05392, 1707.05811...)

Simplicity III, Perimeter Institute, Sept 2019

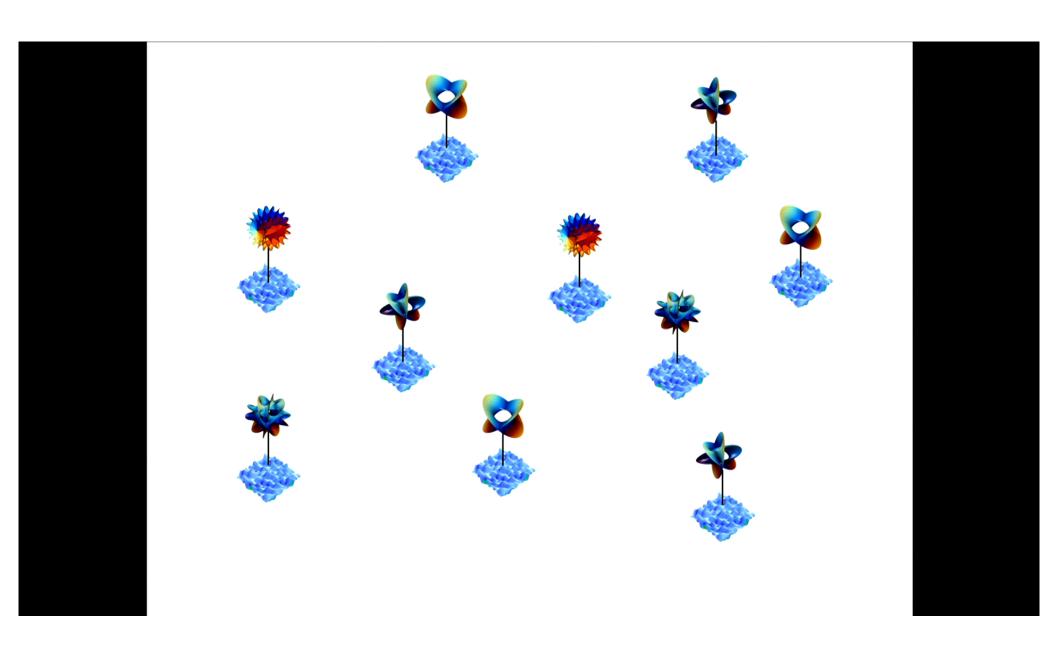
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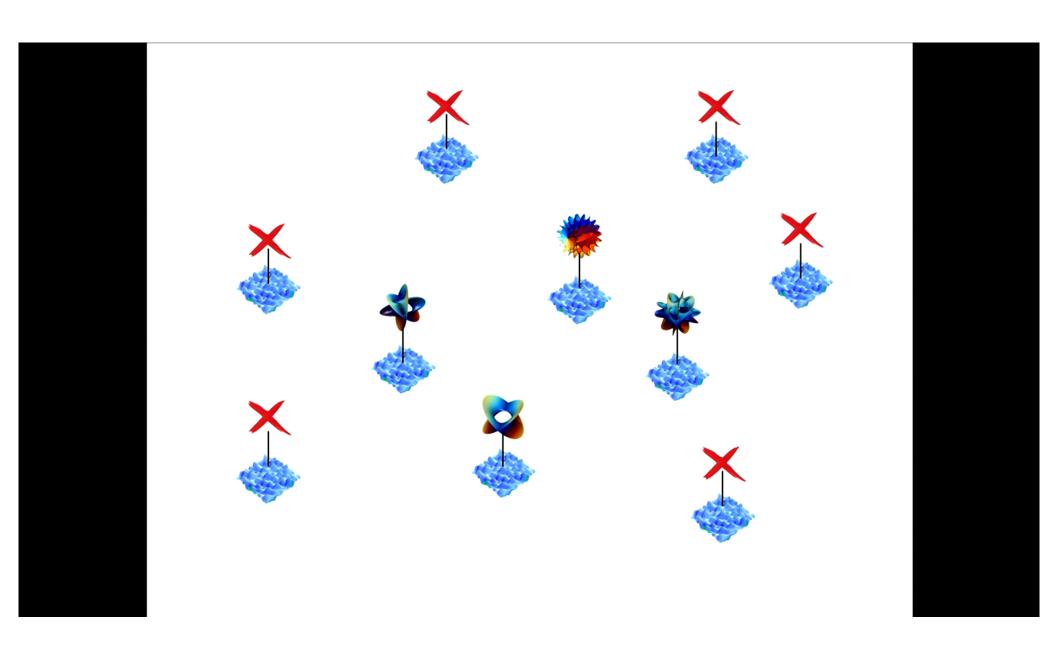


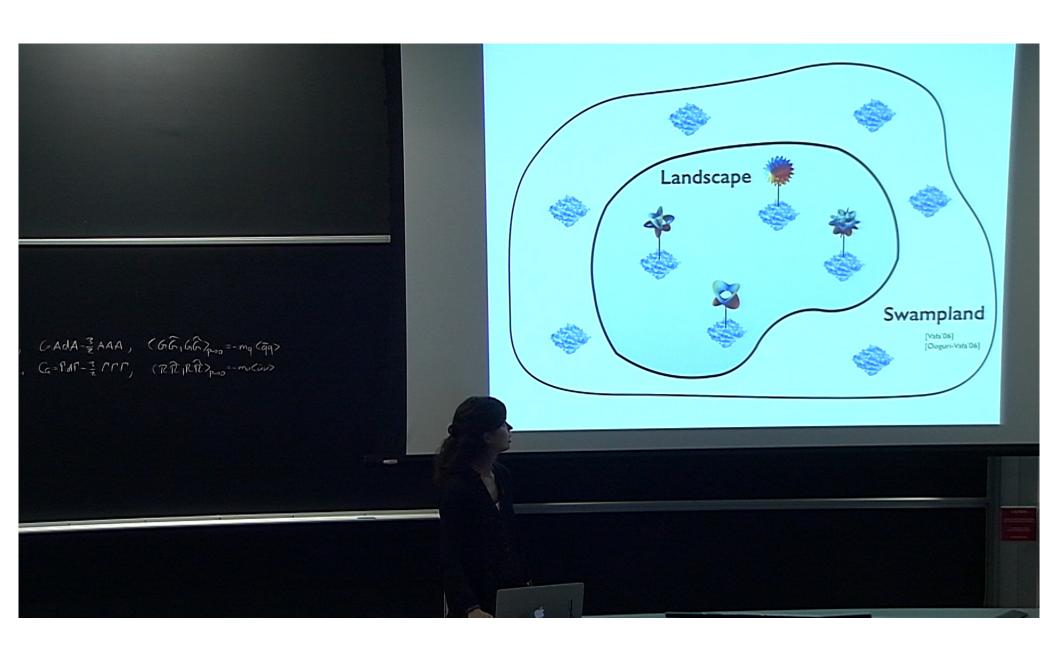
## What does 'swampland' mean?

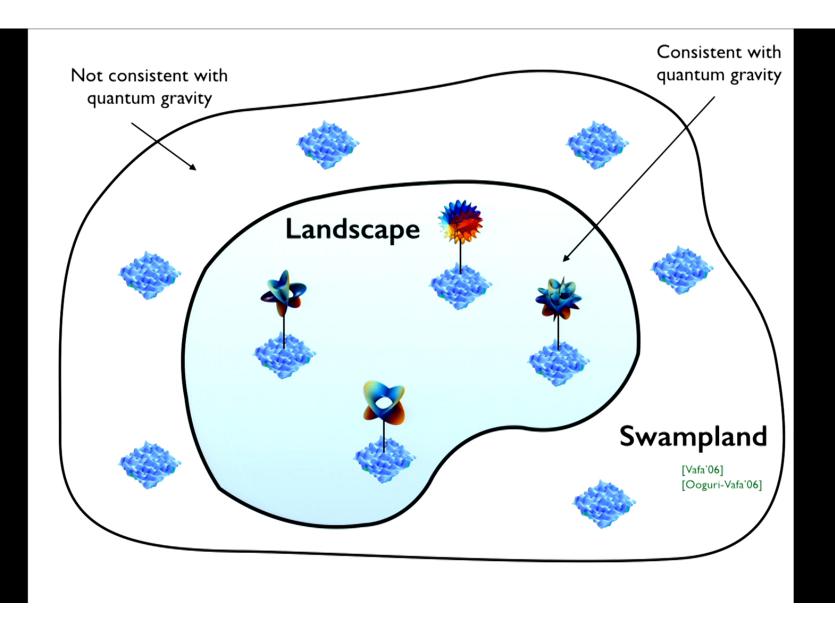












## Not everything is possible in string theory/quantum gravity!!!

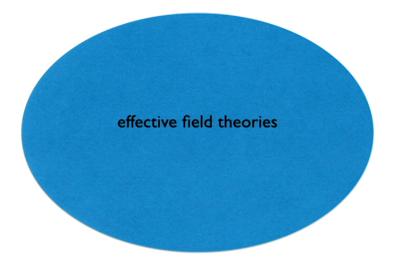
#### (String) Swampland:

Apparently consistent (anomaly-free) quantum effective field theories that cannot be UV embedded in quantum gravity (they cannot arise from string theory)

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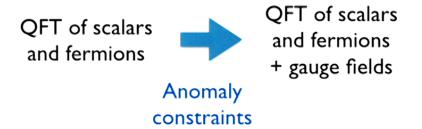
## First guess: Anomalies

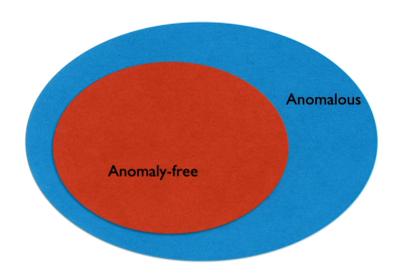
QFT of scalars and fermions



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## First guess: Anomalies

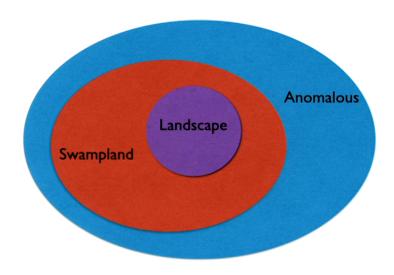




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## First guess: Anomalies





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There are additional (swampland/QG) constraints that any effective QFT must satisfy to be consistent with quantum gravity



UV imprint of quantum gravity at low energies

Outstanding phenomenological implications!

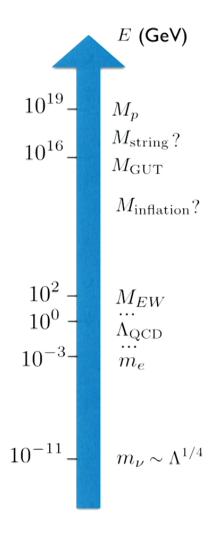
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#### Effective field theories

Modern physics based on a Wilsonian effective field theory approach

Expectation of 'separation of scales':

IR effective theory not very sensitive to UV physics



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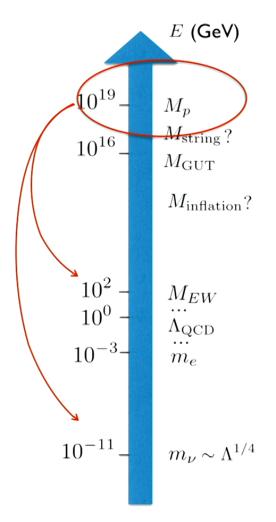
Expectation of 'separation of scales':

IR effective theory not very sensitive to UV physics

This picture can fail!

Quantum gravity induce UV/IR constraints mixing

non-trivial implications at low energies!



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#### Effective field theories

Modern physics based on a Wilsonian effective field theory approach

Expectation of 'separation of scales':

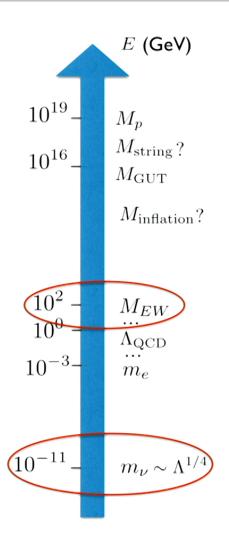
It is already failing...

#### Naturalness problems:

- F Cosmological constant
- EW hierarchy problem

New approach?

Quantum gravity = Missing piece for hierarchy problems?

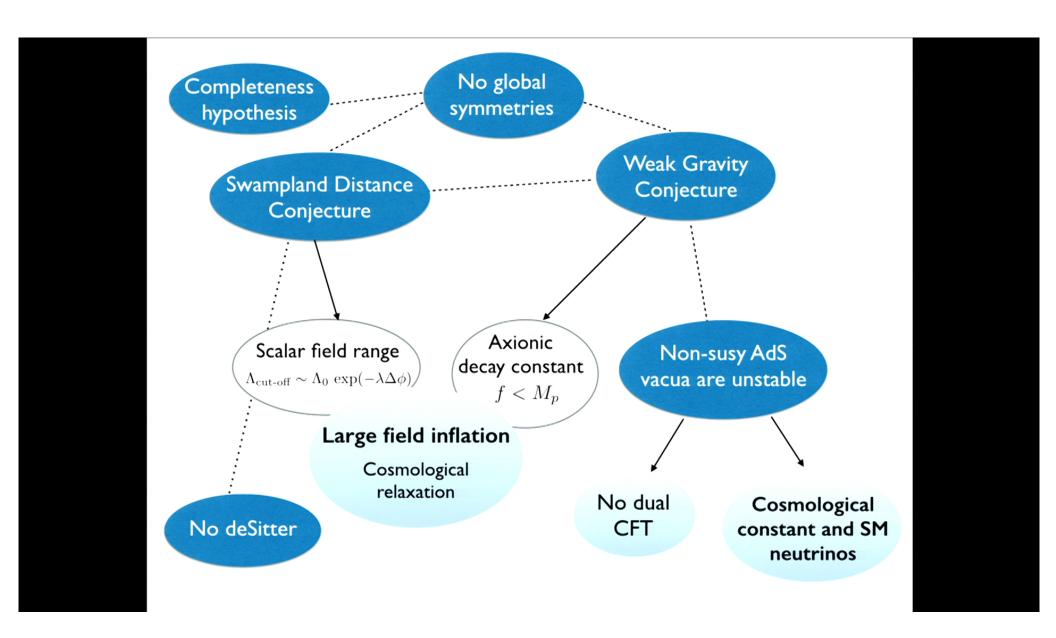


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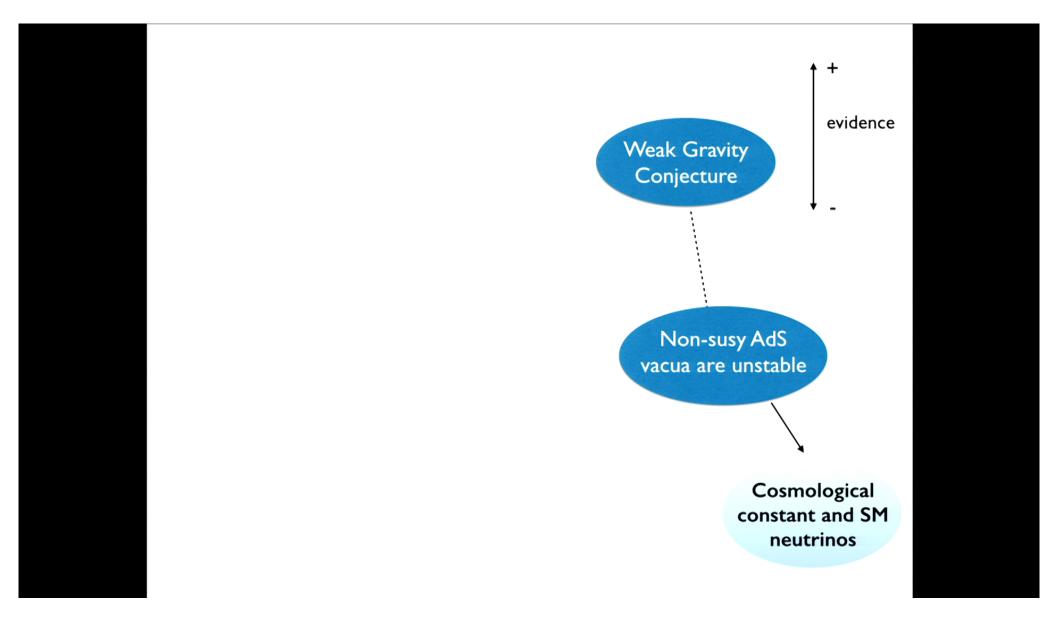
What are the constraints that an effective theory must satisfy to be consistent with quantum gravity?

What distinguishes the landscape from the swampland?

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## Weak Gravity Conjecture

There exist at least a particle in which gravity acts weaker than the gauge force



 $\frac{q}{m} \ge 1$ 

(a superextremal particle)

q : charge m : mass

[Arkani-Hamed, Motl, Nicolis, Vafa'06]

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Evidence:

- Plethora of examples in string theory (not known counter-example)

- Derivation from modular invariance of the 2d CFT [Heidenreich et al'16] [Montero et al'16]

- Relation to entropy bounds, unitarity and causality [Cottrell et al'16] [Andriolo et al'18]

- Derivation from higher derivative corrections to BH's [Fisher et al'17] [Hamada et al'18] [Cheung et al'18] [Charles'19] [Jones et al'19]

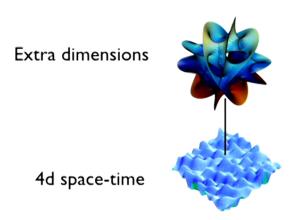
- Relation to cosmic censorship [Crisford et al'17]

- Relation to thermodynamic arguments [Hod'17,] [Urbano'18]

- Relation to entanglement entropy [Montero' 19]

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## Weak Gravity Conjecture for fluxes

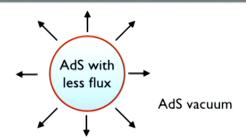


There can be gauge fields propagating in the extra dimensions

$$f_0 \sim \int_{\Sigma_p} F_p$$
 (fluxes in 4d)

WGC applied to the fluxes (in a non-susy vacuum) implies: [Ooguri-Vafa'17]

 $\exists$  Brane (domain wall) with T < Q Bubble instability of the vacuum!



[Maldacena et al.'99]

Non-susy vacua are at best metastable

[Ooguri-Vafa'16] [Freivogel-Kleban'16]

Non-susy stable AdS vacua are in the Swampland (inconsistent with Quantum Gravity)!

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#### Implications for our universe:

Our universe must be metastable



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Non-susy stable AdS vacua are in the Swampland (inconsistent with Quantum Gravity)!

#### Implications for our universe:

Our universe must be metastable

Still, non-susy stable AdS vacua can arise when compactifying SM to lower dimensions !!

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Background independence of QG:

If our 4d SM is consistent with QG



Compactifications of SM should also be consistent

We should not get stable non-susy AdS vacua from compactifying the SM!

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Solution: [Ibanez,Martin-Lozano,IV'17]

We impose the absence of non-susy stable 3d AdS vacua



Constraints on light spectra of SM

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Constraints on light spectra of SM

Assumption: 4d instabilities are not transferred to 3d

$$R_{
m buble} > l_{AdS_3}$$
 (large bubbles)



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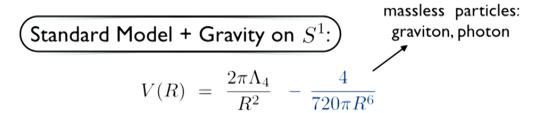


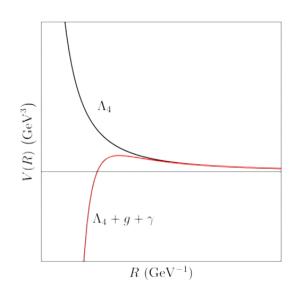
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Standard Model + Gravity on  $S^1$ : [Arkani-Hamed et al.'07] (also [Arnold-Fornal-Wise'10])

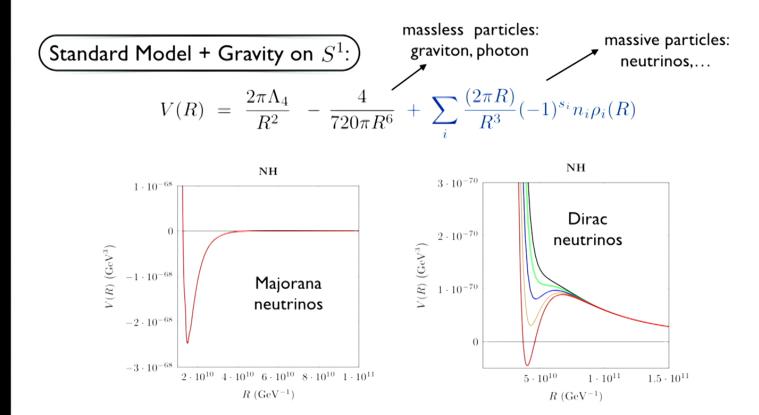
$$V(R) = \frac{2\pi\Lambda_4}{R^2}$$
 + Casimir energy

 $R\equiv$  scalar parametrising the radius of the circle  $S^1$ 





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The more massive the neutrinos, the deeper the AdS vacuum

[Ibanez, Martin-Lozano, IV'17] (see also [Hamada-Shiu'17])

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Standard Model + Gravity on  $S^1$ : $\Big)$ 

Absence of AdS vacua implies:

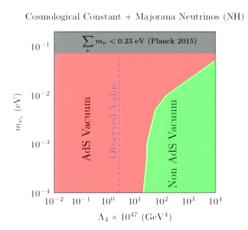
Majorana neutrinos ruled out!

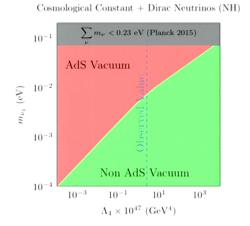
Upper bound for Dirac mass!

$$m_{\nu_1} < 7.7~{\rm meV~(NH)}$$

$$m_{\nu_1} < 2.1 \text{ meV (IH)}$$

#### Lower bound on the cosmological constant





The bound for  $\Lambda_4$  scales as  $m_{
u}^4$ 

(as observed experimentally)

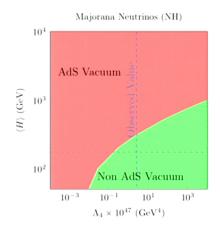
$$\Lambda_4 \ge \frac{a(n_f)30(\Sigma m_i^2)^2 - b(n_f, m_i)\Sigma m_i^4}{384\pi^2}$$

with 
$$a(n_f) = 0.184(0.009) \ b(n_f, m_i) = 5.72(0.29)$$
 for Majorana (Dirac)

First argument (not based on cosmology) to have  $\Lambda_4 \neq 0$ 

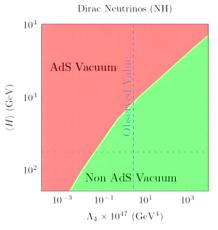
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#### Upper bound on the EW scale



$$M = 10^{10} \text{ GeV}, Y = 10^{-3}$$

 $\underline{\text{Dirac case:}} \quad \langle H \rangle \lesssim 1.6 \frac{\Lambda^{1/4}}{Y_{\nu_1}}$ 



$$Y = 10^{-14}$$

Parameters leading to a higher EW scale do not yield theories consistent with quantum gravity

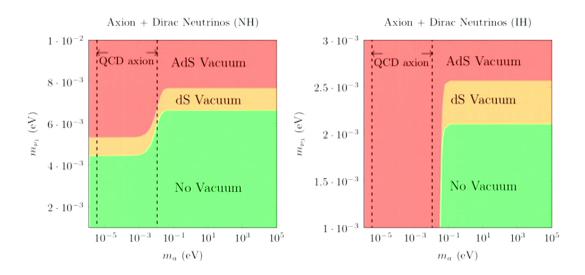
No EW hierarchy problem

## Adding BSM physics

#### Axions

1 axion: negative contribution —— bounds get stronger

Fig. 14 Dirac neutrinos are ruled out in the presence of QCD axion



Multiple axions: can destabilise AdS vacuum bounds disappear

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## Other SM compactifications

ightharpoonup Circle compactification on  $S^1$   $\checkmark$  [Ibanez,Martin-Lozano,IV'17]

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#### Other SM compactifications

There are not new minima (no new bounds) arising at smaller radius when adding the rest of SM particles

... but there can be runaways due to Wilson lines
[Hamada-Shiu'17]

- - Same bounds on neutrino masses
  - New minima that cannot be avoided unless SM is completed with more bosons, like MSSM

(but MSSM at a multi-TeV region at least)

[Gonzalo, Herraez, Ibanez' 18]

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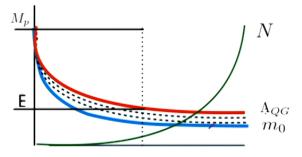
[Gonzalo, Herraez, Ibanez' 18]

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#### Naturalness and Swampland

- Space of parameters is smaller than expected, not everything goes!
  - Provide IR rules to identify non-trivial correlations among parameters of the EFT
- Emergence from integrating out large number of states [Grimm, Palti, IV'18] [Heidenreich,Reece,Rudelius'18]

Swampland conjectures imply the presence of new light states (and a drop-off of the quantum gravity cut-off) whenever we try to recover a global symmetry



N = number of species becoming light (below  $\Lambda_{OG}$  )

$$\Lambda_{QG} = \frac{M_p}{\sqrt{N}}$$

global symmetry limit = (infinitely many massless species) e.g.  $g \rightarrow 0$ 

#### Summary

String Landscape vs Swampland

Not every EFT can be UV embedded in String Theory



Quantum Gravity Constraints

For Consistency with quantum gravity implies constraints on low energy physics:

Conjecture: Non-susy stable AdS vacua are inconsistent with QG



Low energy constraints on SM and BSM light spectra



Upper bound on the EW scale in terms of the cosmological const.

New approach to hierarchy problems? UV/IR mixing?

#### **Assumptions:**

- Conjecture holds (motivated from WGC if there are fluxes in extra dimensions)
- · No additional 3d non-perturbative instabilities transferred from higher dimensions

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