

Title: PSI 2018/2019 - Beyond Standard Model - Lecture 10

Speakers: Latham Boyle

Collection: PSI 2018/2019 - Beyond Standard Model (Boyle)

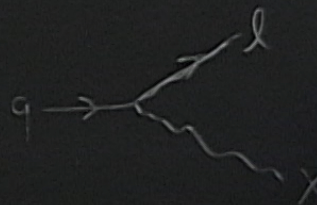
Date: April 08, 2019 - 11:30 AM

URL: <http://pirsa.org/19040026>

L_{SM} : defined by G, R $\leftarrow$ i) g_1, g_2, g_3 all different?
 ii) $\left(\begin{array}{c} \cdot \\ \cdot \end{array} \right) \left(\begin{array}{c} \cdot \\ \cdot \end{array} \right)$
 GUT: embed in G', R'

174 Georgi, Quinn, Weinberg.

EFT/RG:



EFT/RG:

$$\int dx dy e^{-f(x,y)} = \int dx e^{-\tilde{f}(x)}$$

$$Z = \int_{p^2 \leq \Lambda^2} \mathcal{D}[\phi] e^{-S_\Lambda[\phi]} = \int_{p^2 \leq \Lambda'^2} \mathcal{D}[\phi] \underbrace{\int_{\Lambda'^2 \leq p \leq \Lambda^2} \mathcal{D}[\phi] e^{-S_\Lambda[\phi]}}_{e^{-S_\Lambda[\phi]}} = \int_{p^2 \leq \Lambda'^2} \mathcal{D}[\phi] e^{-S_{\Lambda'}[\phi]}$$



EFT/RG:

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$$Z = \int_{p^2 \leq \Lambda^2} \mathcal{D}[\varphi] e^{-S_\Lambda[\varphi]} = \left(\int_{p^2 \leq \Lambda^2} \mathcal{D}[\varphi] \right) \underbrace{\int_{\Lambda^2 \leq p \leq \Lambda^2} \mathcal{D}[\varphi] e^{-S_\Lambda[\varphi]}}_{e^{-S_\Lambda[\varphi]}} = \int_{p^2 \leq \Lambda^2} \mathcal{D}[\varphi] e^{-S_\Lambda[\varphi]}$$

$g(\Lambda)$

$$\beta(g) = \Lambda \frac{dg}{d\Lambda} = \frac{dg}{d(\ln \Lambda)}$$

$$\beta(g_i) = b_i \frac{g_i^3}{16\pi^2}$$

$$\alpha_i = \frac{g_i^2}{4\pi}$$

$$\beta(\alpha_i) = \frac{d\alpha_i}{d\ln\Lambda} = \frac{b_i}{2\pi} \alpha_i^2 \Rightarrow \frac{d\alpha_i}{\alpha_i^2} = \frac{b_i}{2\pi} d\ln\Lambda$$

$$g_i \frac{dg_i}{d\ln\Lambda} = \frac{1}{2} \frac{d(g_i^2)}{d\ln\Lambda} = 2\pi \frac{d\alpha_i}{d\ln\Lambda} = 2\pi \beta(\alpha_i)$$

$$\beta(\alpha_i) = \frac{d\alpha_i}{d\ln\Lambda} = \frac{b_i}{2\pi} \alpha_i^2 \Rightarrow \frac{d\alpha_i}{\alpha_i^2} = \frac{b_i}{2\pi} d\ln\Lambda \rightarrow \alpha_i^{-1}(\Lambda) = \alpha_i^{-1}(M_Z) - \frac{b_i}{2\pi} \ln\left(\frac{\Lambda}{M_Z}\right)$$

$$\alpha_1^{-1}(M_Z) = 59.2$$

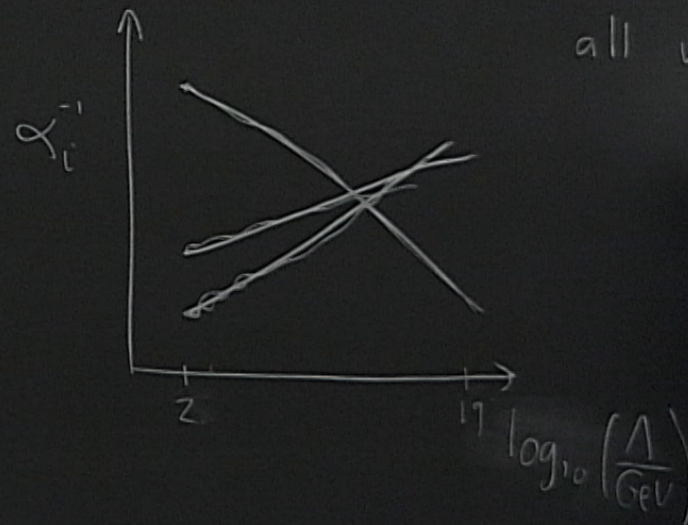
$$\alpha_2^{-1}(M_Z) = 29.6$$

$$\alpha_3^{-1}(M_Z) = 8.5$$

$$b_1 = \frac{41}{10}$$

$$b_2 = -\frac{19}{6}$$

$$b_3 = -7$$



all unified at

$$\frac{d}{d\ln\Lambda} = \frac{b_i}{2\pi} \alpha_i^{-2} \Rightarrow \frac{d\alpha_i}{\alpha_i^2} = \frac{b_i}{2\pi} d\ln\Lambda \rightarrow \alpha_i^{-1}(\Lambda) = \alpha_i^{-1}(M_Z) - \frac{b_i}{2\pi} \ln\left(\frac{\Lambda}{M_Z}\right)$$

59.2

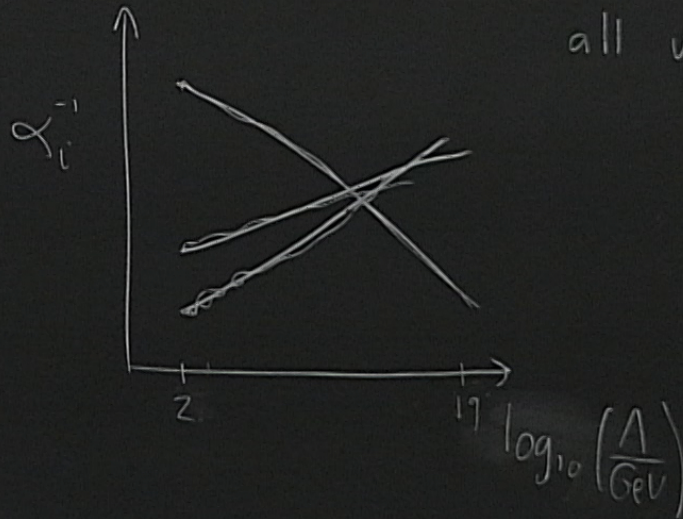
$$b_1 = \frac{41}{10}$$

29.6

$$b_2 = -\frac{19}{6}$$

8.5

$$b_3 = -7$$



all unified at $\Lambda \sim 10^{15} \text{ GeV}$

$$\Gamma_{pt} = \left| \frac{1}{m_i^2} \right|^2 m_p^S$$

Bad news: i) no p^+ decay. rule out $SU(5)$
 ii)

Bad news:

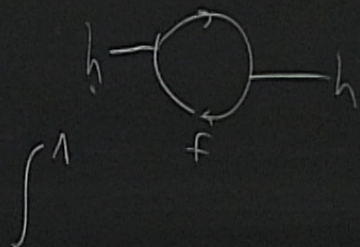
i) no p^+ decay. rule out $SU(5)$

ii) coupling don't unify

iii) $M_{EW} \sim 10^2 \text{ GeV}$, $M_{Pl} \sim 10^{19} \text{ GeV}$

$M_{GUT} \sim 10^{16} \text{ GeV}$

$m^2(h^+h)$



$\leftarrow m^2 \sim \Lambda^2$

Bad news:

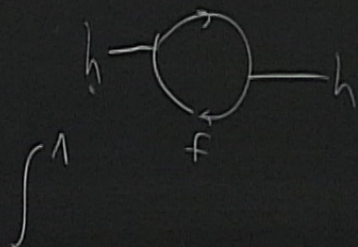
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1981: Dimopoulos/Raby/Wilczek

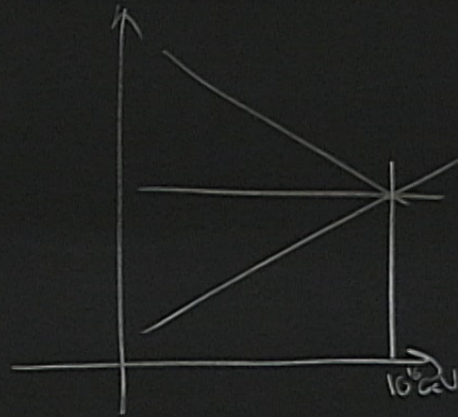
MSSM:

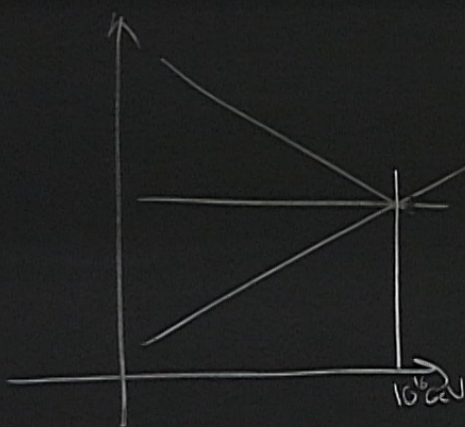
SUSY GUT

$$b_1 = \frac{33}{5}$$

$$b_2 = 1$$

$$b_3 = -3$$





$\mathbb{Z}_2$ sym. R-parity

-1 LSP dark matter

$M \sim \text{GeV}$ "WIMP
miracle"

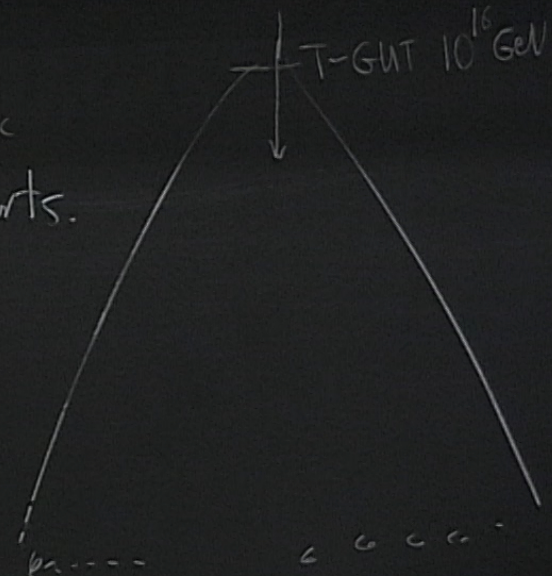
Inflation:

- homogeneous + isotropic
- primordial density perts.
- monopole problem.

$$V(\phi) \sim m^2 \phi^2, \lambda \phi^4$$

$$\Delta_s^2 = \frac{1}{2\epsilon} \left(\frac{V}{M_{Pl}^4} \right) \sim 10^{-9}$$

$$V \sim (10^{16} \text{ GeV})^4$$



bad news

i) no SUSY partners at EW scale.

ii) no SUSY DM

iii) $\Delta_t^2 \sim 8 \left(\frac{V}{M_{\tilde{g}}^4} \right)$

i)