

Title: Superrotations and Flat Space Holography

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Abstract: We consider implications of superrotations as an asymptotic symmetry of asymptotically flat spacetimes. Beginning with a review of the rich structure of interconnections between soft theorems, asymptotic symmetries, and memory effects, we describe the superrotation iteration. The subleading soft graviton theorem can be cast as a Ward identity for this asymptotic symmetry in 4D, and also as one for the stress tensor of a putative CFT₂. We detail the change of scattering basis motivated by this asymptotic symmetry and discuss recent progress.

Superrotations and Flat Space Holography

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Motivation

- ❖ What can IR physics teach us about gravitational scattering?

More Symmetries \Rightarrow More Constraints

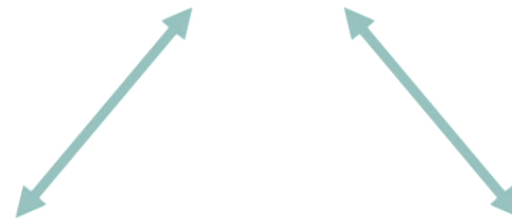
The **asymptotic symmetry group** of *asymptotically flat* spacetimes is much larger than Poincare

A Triangle of Relations

❖ What can IR physics teach us about gravitational scattering?

There exists a generic pattern of connections between asymptotic symmetries, soft theorems, and memory effects...

Soft Theorems



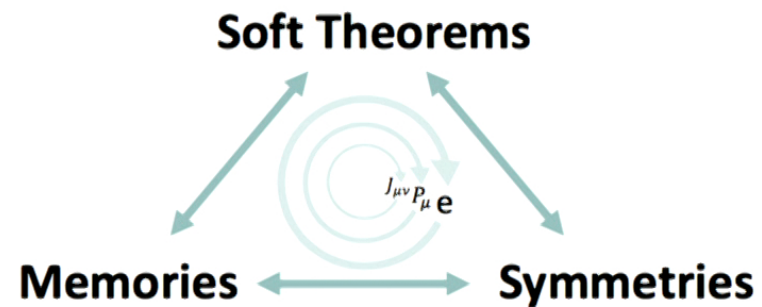
A Triangle of Relations

❖ What can IR physics teach us about gravitational scattering?

There exists a generic pattern of connections between asymptotic symmetries, soft theorems, and memory effects...

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Such that by understanding simpler examples we can identify missing components of new iterations...



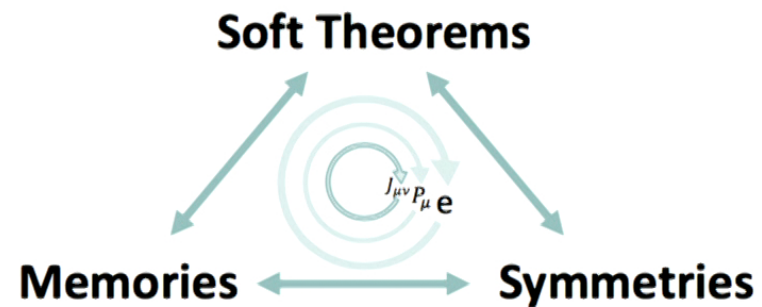
A Triangle of Relations

❖ What can IR physics teach us about gravitational scattering?

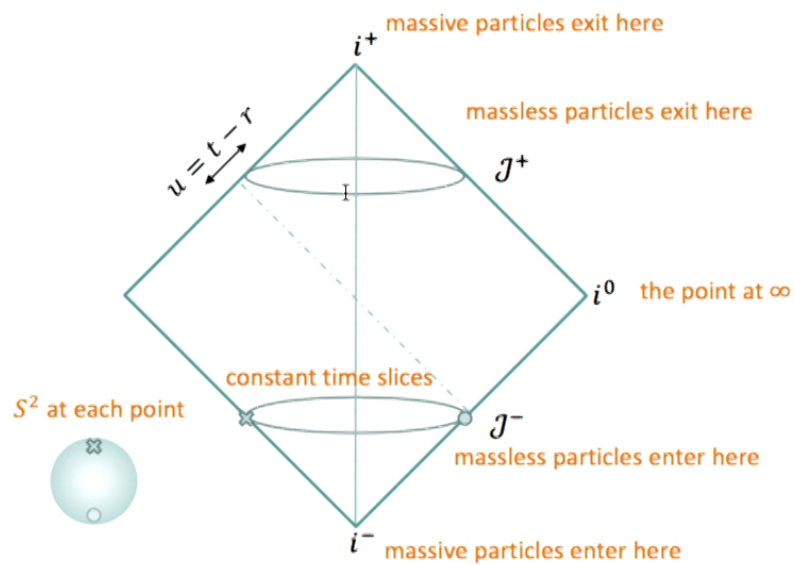
There exists a generic pattern of connections between asymptotic symmetries, soft theorems, and memory effects...

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In this manner a brand new iteration was completed corresponding to **superrotations**. This iteration is related to a generalization of Lorentz transformations and has motivated looking at \mathcal{S} -matrix elements in a new basis with definite $SL(2, \mathbb{C})$ weights

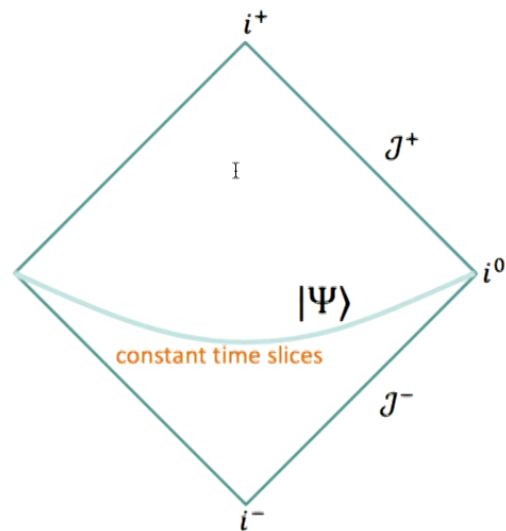


A Simple Example

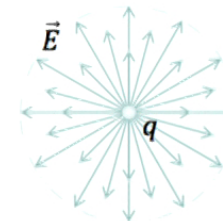


- ❖ Consider the conformal compactification of Minkowski spacetime

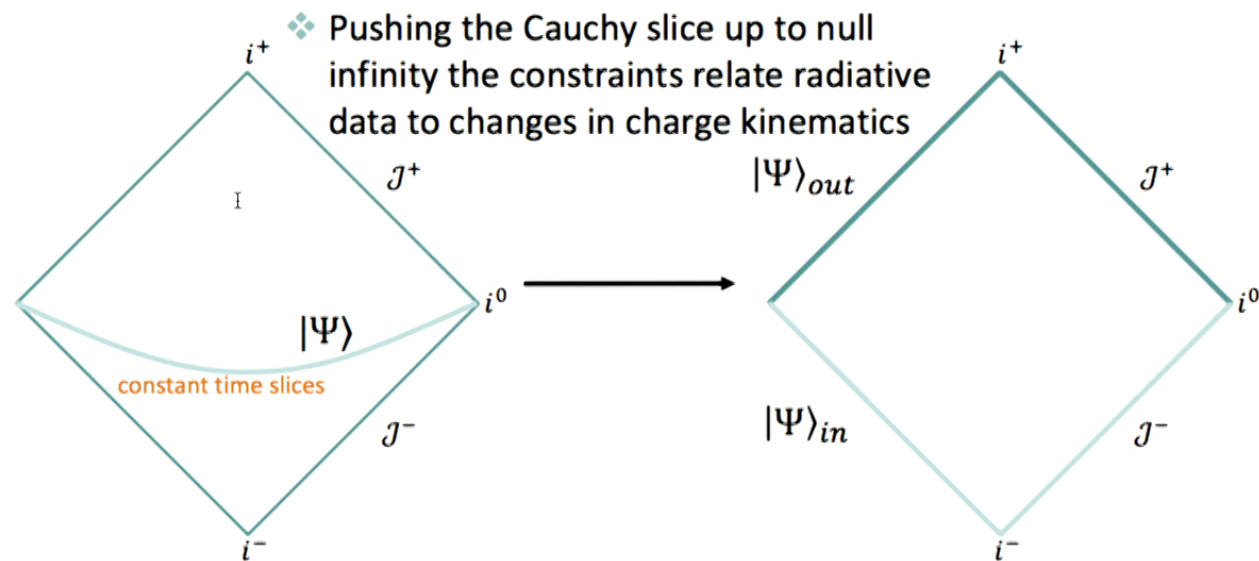
A Simple Example



- ❖ In gauge theories there are constraints that need to be satisfied for the initial data on a Cauchy slice

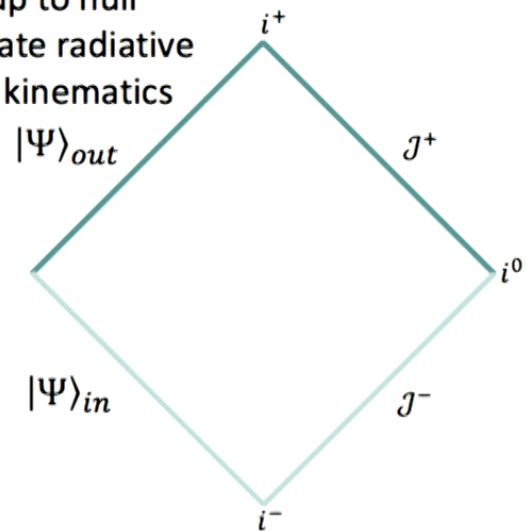
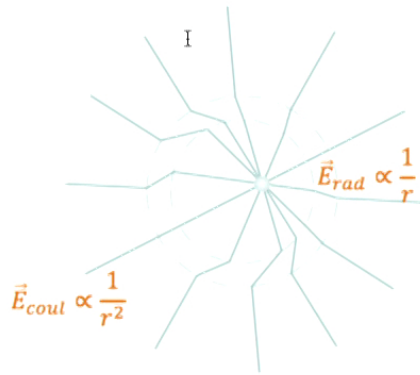


A Simple Example



A Simple Example

- ❖ Pushing the Cauchy slice up to null infinity the constraints relate radiative data to changes in charge kinematics



A Simple Example

Some more details:

- Radial Expansion:

$$\mathcal{A}_z(r, u, z, \bar{z}) = A_z(u, z, \bar{z}) + \sum_{n=1}^{\infty} \frac{A_z^{(n)}(u, z, \bar{z})}{r^n}$$

$$\mathcal{A}_u(r, u, z, \bar{z}) = \frac{1}{r} A_u(u, z, \bar{z}) + \sum_{n=1}^{\infty} \frac{A_u^{(n)}(u, z, \bar{z})}{r^{n+1}}$$

$$F_{ur} = A_u$$

$$F_{z\bar{z}} = \partial_z A_{\bar{z}} - \partial_{\bar{z}} A_z$$

$$F_{uz} = \partial_u A_z$$

- ASG that preserves this expansion:

$$\delta_\epsilon A_z(u, z, \bar{z}) = \partial_z \epsilon(z, \bar{z})$$

- Mode Expansion:

$$\mathcal{A}_\mu(x) = e \sum_{\alpha=\pm} \int \frac{d^3 q}{(2\pi)^3} \frac{1}{2\omega_q} \left[\epsilon_\mu^{\alpha*}(\vec{q}) a_\alpha(\vec{q}) e^{iq \cdot x} + \epsilon_\mu^\alpha(\vec{q}) a_\alpha(\vec{q})^\dagger e^{-iq \cdot x} \right]$$

- Constraint Equation:

[\[arXiv:1407.3789\]](#) $\partial_u A_u = \partial_u (D^z A_z + D^{\bar{z}} A_{\bar{z}}) + e^2 j_u$

Coordinate Conventions:

$$ds^2 = -du^2 - 2dudr + 2r^2 \gamma_{z\bar{z}} dz d\bar{z}$$

$$z = e^{i\phi} \tan \frac{\theta}{2} \quad \gamma_{z\bar{z}} = \frac{2}{(1 + z\bar{z})^2}$$

A Simple Example

Two key points:

- Saddle point at large r picks out a gauge boson momentum pointing in the same direction as where an observer near null infinity would detect it. As a result, one ends up with a mode expansion where the angular integral localizes, and (u, ω) remain as Fourier conjugates.

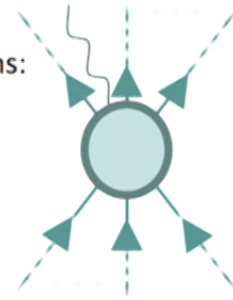
$$e^{iq \cdot x} = e^{-i\omega u - i\omega r(1 - \hat{q} \cdot \hat{x})} \Rightarrow A_z(u, z, \bar{z}) = -\frac{i}{8\pi^2} \frac{\sqrt{2}e}{1 + z\bar{z}} \int_0^\infty d\omega [a_+(\omega \hat{x}) e^{-i\omega u} - a_-(\omega \hat{x})^\dagger e^{i\omega u}]$$

- $\int du$ picks out $\omega \rightarrow 0$. As such we can relate the soft factors to the constraint equations:
Fourier transform of a pole $\frac{1}{\omega}$ is a step function

$$S^{(0)-} = \sum_k eQ_k \frac{p_k \cdot \epsilon^-}{p_k \cdot q}$$

$$\langle z_{n+1}, z_{n+2}, \dots | a_-(q) \mathcal{S} | z_1, z_2, \dots \rangle = S^{(0)-} \langle z_{n+1}, z_{n+2}, \dots | \mathcal{S} | z_1, z_2, \dots \rangle + \mathcal{O}(1)$$

[[arXiv:1407.3789](#), [arXiv:1505.00716](#)]



A Simple Example

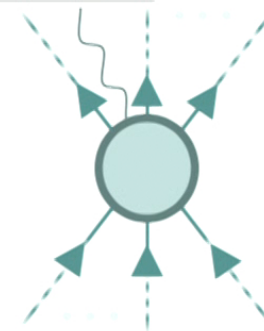
Integrate the constraint equation along u

$$E_r = \frac{Q}{4\pi r^2} \frac{1}{\gamma^2(1 - \vec{\beta} \cdot \hat{n})^2}$$



$$\Delta A_u = 2D^z \Delta A_z + e^2 \int du j_u$$

$$-\frac{e}{4\pi} \lim_{\omega \rightarrow 0} \omega [D^z \hat{\epsilon}_z^{*+} S_p^{(0)+} + D^z \hat{\epsilon}_z^{*-} S_p^{(0)-}] = -e^2 \frac{Q}{4\pi} \frac{1}{\gamma^2(1 - \vec{\beta} \cdot \hat{n})^2}$$



The soft factor indicates that typical scattering processes will produce a nonzero u integrated electric field.

Some Conventions:

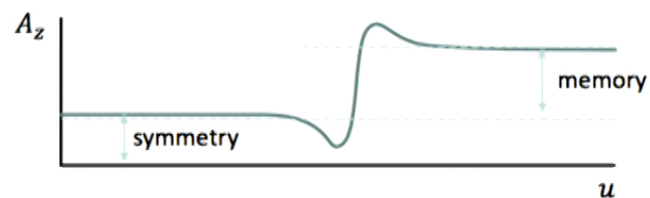
$$p^\mu = m\gamma(1, \vec{\beta}) \quad S_p^{(0)\pm} = eQ \frac{p \cdot \epsilon^\pm}{p \cdot q}$$

$$\Delta A_z = -\frac{e}{4\pi} \hat{\epsilon}_z^{*+} \omega S^{(0)+}$$

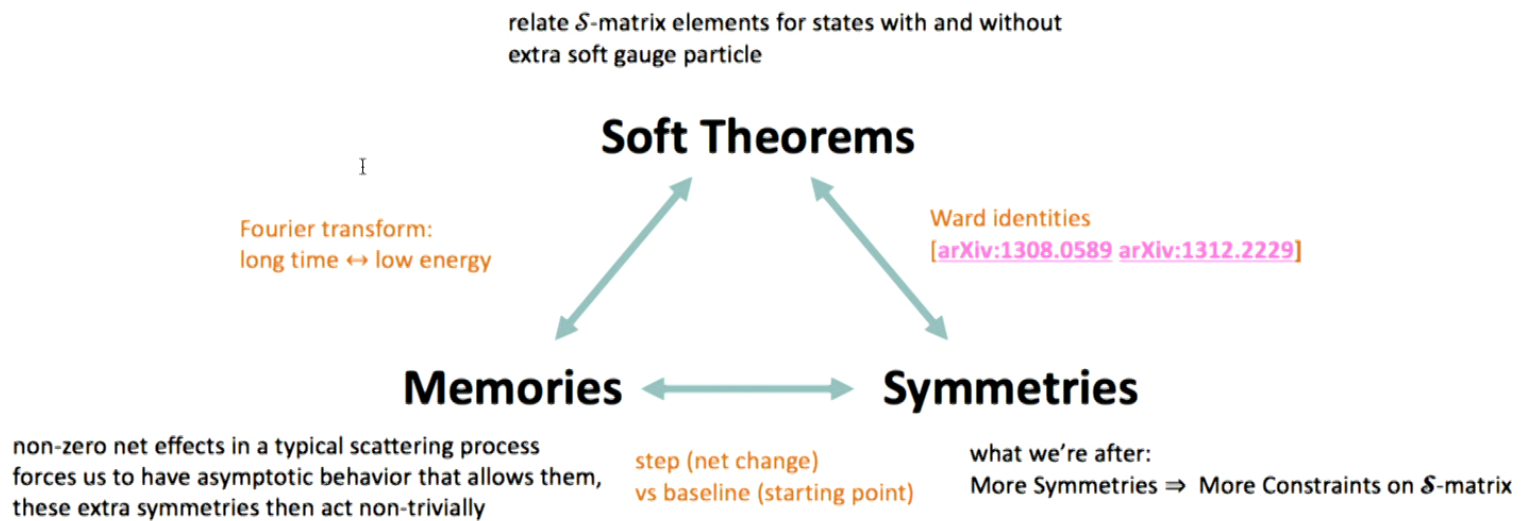
[[arXiv:1505.00716](https://arxiv.org/abs/1505.00716)]

A Simple Example

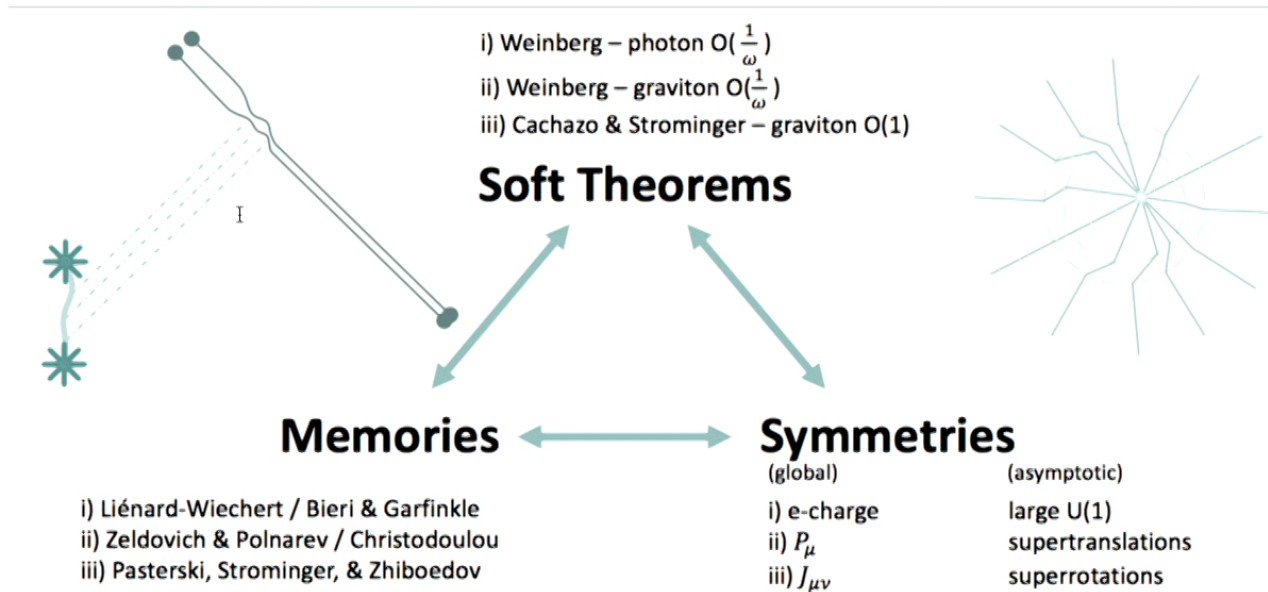
- Upshot: The residue of the Weinberg pole indicates a nonzero value for certain low-energy radiation observables aka **“memory effects”**
- Since setting these modes to zero would trivialize the allowed scattering events, we get with this class of boundary conditions a larger class of gauge transformations that preserve the radial order of the falloffs while shifting the boundary values aka **“large gauge transformations”**



A Triangle of Relations



A Triangle of Relations



Asymptotically Flat Spacetimes



$$ds^2 = -dt^2 + dx^2 + dy^2 + dz^2$$

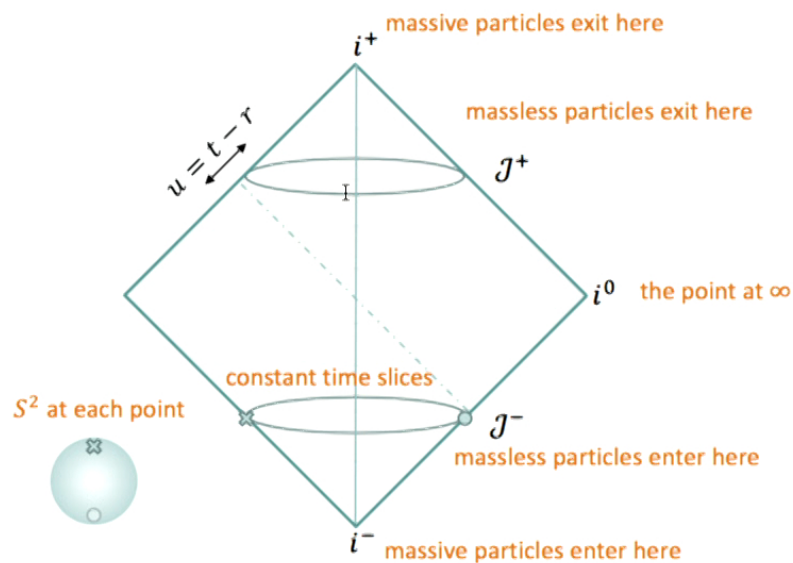
❖ Want to consider non-trivial gravitational backgrounds that are “close” to being flat



BMS 1960's

- Approach flat spacetime far away from sources

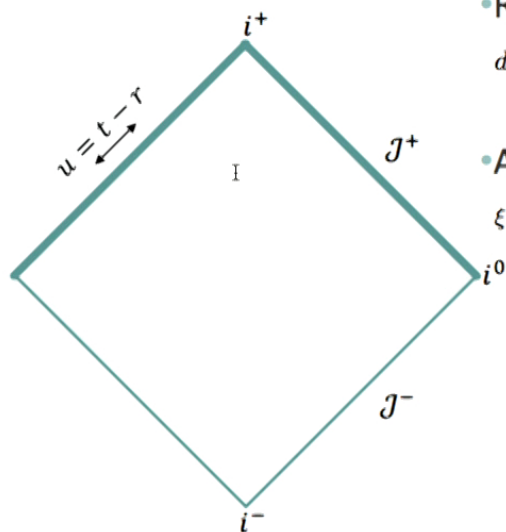
Asymptotically Flat Spacetimes



❖ Interested in set of diffeomorphisms that preserve class of asymptotically flat metrics, characterized by radial fall-off near null infinity

❖ $ASG = \frac{\text{allowed gauge symmetries}}{\text{trivial gauge symmetries}}$

Asymptotically Flat Spacetimes



- Radial Expansion:

$$ds^2 = -du^2 - 2dudr + 2r^2\gamma_{z\bar{z}}dzd\bar{z} + 2\frac{m_B}{r}du^2 + (rC_{zz}dz^2 + D^z C_{zz}dudz + \frac{1}{r}(\frac{4}{3}N_z - \frac{1}{4}\partial_z(C_{zz}C^{zz}))dudz + c.c.) + \dots$$

- ASG that preserves this expansion:

$$\xi^+ = (1 + \frac{u}{2r})Y^{+z}\partial_z - \frac{u}{2r}D^z D_z Y^{+z}\partial_{\bar{z}} - \frac{1}{2}(u+r)D_z Y^{+z}\partial_r + \frac{u}{2}D_z Y^{+z}\partial_u + c.c. + f^+\partial_u - \frac{1}{r}(D^z f^+\partial_z + D^{\bar{z}} f^+\partial_{\bar{z}}) + D^z D_z f^+\partial_r$$

Coordinate Conventions:

$$z = e^{i\phi} \tan \frac{\theta}{2} \quad \gamma_{z\bar{z}} = \frac{2}{(1+z\bar{z})^2}$$

$$f^+ = f^+(z, \bar{z}) \quad \partial_{\bar{z}} Y^{+z} = 0$$

Superrotation Charge

- ❖ We can demonstrate a semiclassical Ward identity for superrotations using the subleading soft graviton theorem [[arXiv:1406.3312](#)].

$$\begin{aligned} \langle out | Q^+[Y] \mathcal{S} - \mathcal{S} Q^-[Y] | in \rangle &= 0 \\ 8\pi G Q^+[Y] &= \int du \int d^2z \sqrt{\gamma} \partial_u [-u Y^A D_A m_B + Y^A N_A + \dots] \\ \partial_u m_B &= \frac{1}{4} \partial_u [D_z^2 C^{zz} + D_{\bar{z}}^2 C^{\bar{z}\bar{z}}] - T_{uu} \\ \partial_u N_z &= \frac{1}{4} \partial_z [D_z^2 C^{zz} - D_{\bar{z}}^2 C^{\bar{z}\bar{z}}] + \partial_z m_B - T_{uz} \end{aligned}$$

Superrotation Charge

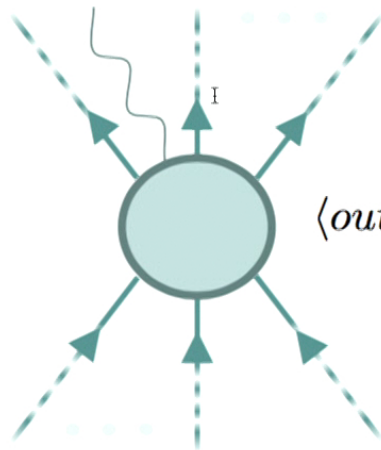
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$$Q^+[Y] = Q_S^+[Y] + Q_H^+[Y]$$

$$Q_S^+[Y] = \frac{1}{2} \int_{\mathcal{I}^+} \underline{du} d^2z D_z^3 Y^z u \partial_u \underline{C}_{\bar{z}} \quad Q_H^+[Y] = \lim_{\Sigma \rightarrow \mathcal{I}^+} \int_{\Sigma} d\Sigma \xi^\mu n_\Sigma^\nu T_{\mu\nu}^M$$

Superrotation Charge

- ❖ We can demonstrate a semiclassical Ward identity for superrotations using the subleading soft graviton theorem [[arXiv:1406.3312](#)].



as Fourier mode
of field operator

$$e^{iq \cdot x} = e^{-i\omega u - i\omega r(1 - \hat{q} \cdot \hat{x})}$$

$$\hat{q}_{pos} \Leftrightarrow \hat{q}_{mom} \quad \& \quad \lim_{\omega \rightarrow 0} \Leftrightarrow \int du$$

$$\langle out | \underline{a_-}(q) \mathcal{S} | in \rangle = \left(S^{(0)-} + S^{(1)-} \right) \langle out | \mathcal{S} | in \rangle + \mathcal{O}(\omega)$$

$$S^{(0)-} = \sum_k \frac{(p_k \cdot \epsilon^-)^2}{p_k \cdot q} \quad S^{(1)-} = -i \sum_k \frac{p_{k\mu} \epsilon^{-\mu\nu} q^\lambda J_{k\lambda\nu}}{p_k \cdot q}$$

Superrotation Charge

❖ Looking again at the superrotation vector field near null infinity, we notice we have two copies of the Witt algebra:

$$\xi^+|_{\mathcal{I}^+} = Y^{+z}\partial_z + \frac{u}{2}D_z Y^{+z}\partial_u + c.c.$$

❖ Moreover, for a particular choice of $Y^z \sim \frac{1}{z-w}$ we find that the soft part of the charge takes the form of a putative 2D stress tensor [[arXiv:1609.00282](https://arxiv.org/abs/1609.00282)].

$$T_{zz} \equiv \frac{i}{8\pi G} \int d^2w \frac{1}{z-w} D_w^2 D^{\bar{w}} \int du u \partial_u C_{\bar{w}\bar{w}}$$

$$\langle T_{zz} \mathcal{O}_1 \cdots \mathcal{O}_n \rangle = \sum_{k=1}^n \left[\frac{h_k}{(z-z_k)^2} + \frac{\Gamma_{z_k z_k}}{z-z_k} h_k + \frac{1}{z-z_k} (\partial_{z_k} - |s_k| \Omega_{z_k}) \right] \langle \mathcal{O}_1 \cdots \mathcal{O}_n \rangle$$

Weight Conventions:

$$h = \frac{1}{2}(s+1+iE_R) \quad \bar{h} = \frac{1}{2}(-s+1+iE_R)$$

$$\Delta = h + \bar{h} \quad s = h - \bar{h}$$

Superrotation Charge

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We need Rindler energy eigenstates!

Weight Conventions:

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$$\Delta = h + \bar{h} \quad s = h - \bar{h}$$

Constructing A Conformal-Primary Basis

❖ Using that the Lorentz group $SO(1,d+1)$ in $\mathbf{R}^{1,d+1}$ acts as the conformal group on \mathbf{R}^d define the *massive scalar conformal primary wavefunction* to:

- satisfy the $(d+2)$ -dimensional massive Klein-Gordon equation of mass m :

$$\left(\frac{\partial}{\partial X^\nu} \frac{\partial}{\partial X_\nu} - m^2 \right) \phi_\Delta(X^\mu; \vec{w}) = 0$$

- transform covariantly as a scalar conformal primary operator in d dimensions under an $SO(1,d+1)$ transformation:

$$\phi_\Delta(\Lambda^\mu_\nu X^\nu; \vec{w}'(\vec{w})) = \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{-\Delta/d} \phi_\Delta(X^\mu; \vec{w}) \quad [\text{arXiv:1705.01027}]$$

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Constructing A Conformal-Primary Basis

$$ds_{H_{d+1}}^2 = \frac{dy^2 + d\vec{z} \cdot d\vec{z}}{y^2}$$

\hat{p} \mathbb{I}




$m \neq 0$
 $p = m\hat{p}$

$$\hat{p}(y, \vec{z}) = \left(\frac{1 + y^2 + |\vec{z}|^2}{2y}, \frac{\vec{z}}{y}, \frac{1 - y^2 - |\vec{z}|^2}{2y} \right)$$

$$\hat{p}^\mu(y', \vec{z}') = \Lambda^\mu{}_\nu \hat{p}^\nu$$

q

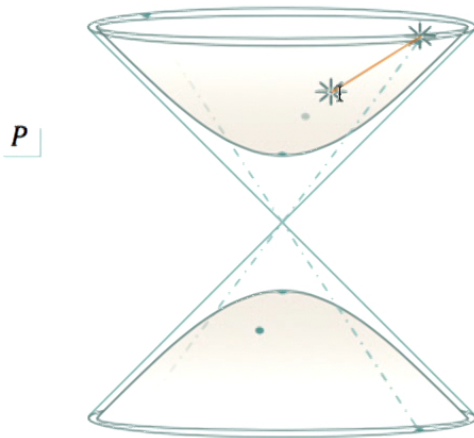
$$q^\mu(\vec{w}) = (1 + |\vec{w}|^2, 2\vec{w}, 1 - |\vec{w}|^2)$$

$$q^\mu(\vec{w}') = \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{1/d} \Lambda^\mu{}_\nu q^\nu(\vec{w})$$


$\vec{w} \in \mathbb{R}^d$

Constructing A Conformal-Primary Basis

- ❖ The desired properties are met by the convolution:



$$\phi_{\Delta}^{\pm}(X^{\mu}; \vec{w}) = \int_{H_{d+1}} [d\hat{p}] G_{\Delta}(\hat{p}; \vec{w}) \exp [\pm i m \hat{p} \cdot X]$$

- ❖ Interpretation as bulk-to-boundary propagation in momentum space
- ❖ Have **plane wave** \Rightarrow **highest-weight**, what about reverse?

Constructing A Conformal-Primary Basis

❖ The orthogonality conditions $\int_{-\infty}^{\infty} d\nu \mu(\nu) \int d^d \vec{w} G_{\frac{d}{2}+i\nu}(\hat{p}_1; \vec{w}) G_{\frac{d}{2}-i\nu}(\hat{p}_2; \vec{w}) = \delta^{(d+1)}(\hat{p}_1, \hat{p}_2)$

$$\int_{H_{d+1}} [d\hat{p}] G_{\frac{d}{2}+i\nu}(\hat{p}; \vec{w}_1) G_{\frac{d}{2}+i\bar{\nu}}(\hat{p}; \vec{w}_2) =$$

$$2\pi^{d+1} \frac{\Gamma(i\nu)\Gamma(-i\nu)}{\Gamma(\frac{d}{2}+i\nu)\Gamma(\frac{d}{2}-i\nu)} \delta(\nu + \bar{\nu}) \delta^{(d)}(\vec{w}_1 - \vec{w}_2) + 2\pi^{\frac{d}{2}+1} \frac{\Gamma(i\nu)}{\Gamma(\frac{d}{2}+i\nu)} \delta(\nu - \bar{\nu}) \frac{1}{|\vec{w}_1 - \vec{w}_2|^{2(\frac{d}{2}+i\nu)}}$$

$$\mu(\nu) = \frac{\Gamma(\frac{d}{2}+i\nu)\Gamma(\frac{d}{2}-i\nu)}{4\pi^{d+1}\Gamma(i\nu)\Gamma(-i\nu)}$$

[\[arXiv:1404.5625\]](#)

❖ Imply we can go in the opposite direction **highest-weight** \Rightarrow **plane wave**

$$e^{\pm im\hat{p}\cdot X} = 2 \int_0^\infty d\nu \mu(\nu) \int d^d \vec{w} G_{\frac{d}{2}-i\nu}(\hat{p}; \vec{w}) \phi_{\frac{d}{2}+i\nu}^\pm(X^\mu; \vec{w})$$

$$\Delta \in \frac{d}{2} + i\mathbb{R}_{\geq 0}$$



$$\vec{w} \in \mathbb{R}^d$$

Constructing A Conformal-Primary Basis

- ❖ By forming the combination $\omega = \frac{m}{2y}$ we can further use the boundary behavior of G_Δ to explore the massless analog:

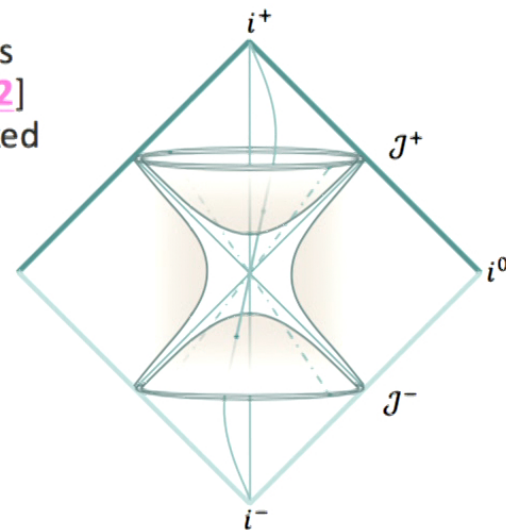
$$G_\Delta(y, \vec{z}; \vec{w}) \xrightarrow{m \rightarrow 0} \pi^{\frac{d}{2}} \frac{\Gamma(\Delta - \frac{d}{2})}{\Gamma(\Delta)} y^{d-\Delta} \delta^{(d)}(\vec{z} - \vec{w}) + \frac{y^\Delta}{|\vec{z} - \vec{w}|^{2\Delta}} + \dots$$

- ❖ The first term results in a Mellin transform of the energy, in which the reference direction is the same as the momentum, and satisfies the desired properties of a massless conformal primary.

$$\varphi_\Delta^\pm(X^\mu; \vec{w}) \equiv \int_0^\infty d\omega \omega^{\Delta-1} e^{\pm i\omega q \cdot X - \epsilon\omega} = \frac{(\mp i)^\Delta \Gamma(\Delta)}{(-q(\vec{w}) \cdot X \mp i\epsilon)^\Delta}$$

Amplitude Transforms

- ❖ Note that transforming momentum space amplitudes directly, is an alternative to previous approaches [[hep-th/0303006](#), [arXiv:1609.00732](#)] towards flat space holography, which have looked at a foliation of Minkowski space to reproduce AdS/CFT, dS/CFT on each slice.



Constructing A Conformal-Primary Basis


❖ Photon

$$\left(\frac{\partial}{\partial X^\sigma} \frac{\partial}{\partial X_\sigma} \delta^\mu_\nu - \frac{\partial}{\partial X^\nu} \frac{\partial}{\partial X_\mu} \right) A_{\mu a}^{\Delta \pm}(X^\rho; \vec{w}) = 0 \quad A_{\mu a}^{\Delta \pm}(\Lambda^\rho_\nu X^\nu; \vec{w}'(\vec{w})) = \frac{\partial w^b}{\partial w'^a} \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{-(\Delta-1)/d} \Lambda_\mu^\sigma A_{\sigma b}^{\Delta \pm}(X^\rho; \vec{w})$$

❖ Graviton

$$\partial_\sigma \partial_\nu h^\sigma_{\mu; a_1 a_2} + \partial_\sigma \partial_\mu h^\sigma_{\nu; a_1 a_2} - \partial_\mu \partial_\nu h^\sigma_{\sigma; a_1 a_2} - \partial^\rho \partial_\rho h_{\mu\nu; a_1 a_2} = 0 \quad \begin{aligned} h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} &= h_{\mu_2 \mu_1; a_1 a_2}^{\Delta, \pm}, \\ h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} &= h_{\mu_1 \mu_2; a_2 a_1}^{\Delta, \pm}, \quad \delta^{a_1 a_2} h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} = 0 \end{aligned}$$

$$h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm}(\Lambda^\rho_\nu X^\nu; \vec{w}'(\vec{w})) = \frac{\partial w^{b_1}}{\partial w'^{a_1}} \frac{\partial w^{b_2}}{\partial w'^{a_2}} \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{-(\Delta-2)/d} \Lambda_{\mu_1}^{\sigma_1} \Lambda_{\mu_2}^{\sigma_2} h_{\sigma_1 \sigma_2; b_1 b_2}^{\Delta, \pm}(X^\rho; \vec{w})$$

\longleftrightarrow $\Delta \in \frac{d}{2} + i\mathbb{R}$ \times  $\vec{w} \in \mathbb{R}^d$

Constructing A Conformal-Primary Basis


❖ Photon

$$\left(\frac{\partial}{\partial X^\sigma} \frac{\partial}{\partial X_\sigma} \delta^\mu_\nu - \frac{\partial}{\partial X^\nu} \frac{\partial}{\partial X_\mu} \right) A_{\mu a}^{\Delta \pm}(X^\rho; \vec{w}) = 0 \quad A_{\mu a}^{\Delta \pm}(\Lambda^\rho_\nu X^\nu; \vec{w}'(\vec{w})) = \frac{\partial w^b}{\partial w'^a} \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{-(\Delta-1)/d} \Lambda_\mu^\sigma A_{\sigma b}^{\Delta \pm}(X^\rho; \vec{w})$$


❖ Graviton

$$\partial_\sigma \partial_\nu h^\sigma_{\mu; a_1 a_2} + \partial_\sigma \partial_\mu h^\sigma_{\nu; a_1 a_2} - \partial_\mu \partial_\nu h^\sigma_{\sigma; a_1 a_2} - \partial^\rho \partial_\rho h_{\mu\nu; a_1 a_2} = 0 \quad \begin{aligned} h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} &= h_{\mu_2 \mu_1; a_1 a_2}^{\Delta, \pm}, \\ h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} &= h_{\mu_1 \mu_2; a_2 a_1}^{\Delta, \pm}, \quad \delta^{a_1 a_2} h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm} = 0 \end{aligned}$$

$$h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm}(\Lambda^\rho_\nu X^\nu; \vec{w}'(\vec{w})) = \frac{\partial w^{b_1}}{\partial w'^{a_1}} \frac{\partial w^{b_2}}{\partial w'^{a_2}} \left| \frac{\partial \vec{w}'}{\partial \vec{w}} \right|^{-(\Delta-2)/d} \Lambda_{\mu_1}^{\sigma_1} \Lambda_{\mu_2}^{\sigma_2} h_{\sigma_1 \sigma_2; b_1 b_2}^{\Delta, \pm}(X^\rho; \vec{w})$$



$\Delta \in \frac{d}{2} + i\mathbb{R}$



 $\vec{w} \in \mathbb{R}^d$

Constructing A Conformal-Primary Basis

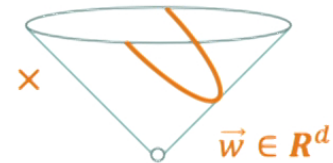
- ❖ The shadow is linearly independent.
- ❖ Demanding conformal profile fixes residual gauge transformations but within gauge equivalence class can return to Mellin representative.

$$A_{\mu a}^{\Delta, \pm}(X^\mu; \vec{w}) = \frac{\partial_a q_\mu}{(-q \cdot X \mp i\epsilon)^\Delta} + \frac{\partial_a q \cdot X}{(-q \cdot X \mp i\epsilon)^{\Delta+1}} q_\mu$$

$$- \text{const.} \frac{\partial}{\partial X^\mu} \left(\frac{\partial_a q \cdot X}{(-q \cdot X \mp i\epsilon)^\Delta} \right)$$



$$\Delta \in \frac{d}{2} + i\mathbb{R}$$



Constructing A Conformal-Primary Basis

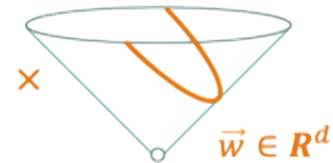
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$$\Delta \in \frac{d}{2} + i\mathbb{R}$$



Constructing A Conformal-Primary Basis

❖ The orthogonality conditions $\int_{-\infty}^{\infty} d\nu \mu(\nu) \int d^d \vec{w} G_{\frac{d}{2}+i\nu}(\hat{p}_1; \vec{w}) G_{\frac{d}{2}-i\nu}(\hat{p}_2; \vec{w}) = \delta^{(d+1)}(\hat{p}_1, \hat{p}_2)$

$$\int_{H_{d+1}} [d\hat{p}] G_{\frac{d}{2}+i\nu}(\hat{p}; \vec{w}_1) G_{\frac{d}{2}+i\bar{\nu}}(\hat{p}; \vec{w}_2) = \mu(\nu) = \frac{\Gamma(\frac{d}{2}+i\nu)\Gamma(\frac{d}{2}-i\nu)}{4\pi^{d+1}\Gamma(i\nu)\Gamma(-i\nu)}$$

$$2\pi^{d+1} \frac{\Gamma(i\nu)\Gamma(-i\nu)}{\Gamma(\frac{d}{2}+i\nu)\Gamma(\frac{d}{2}-i\nu)} \delta(\nu+\bar{\nu}) \delta^{(d)}(\vec{w}_1 - \vec{w}_2) + 2\pi^{\frac{d}{2}+1} \frac{\Gamma(i\nu)}{\Gamma(\frac{d}{2}+i\nu)} \delta(\nu-\bar{\nu}) \frac{1}{|\vec{w}_1 - \vec{w}_2|^{2(\frac{d}{2}+i\nu)}} \quad [\text{arXiv:1404.5625}]$$

❖ Imply we can go in the opposite direction **highest-weight** \Rightarrow **plane wave**

$$e^{\pm im\hat{p}\cdot X} = 2 \int_0^\infty d\nu \mu(\nu) \int d^d \vec{w} G_{\frac{d}{2}-i\nu}(\hat{p}; \vec{w}) \phi_{\frac{d}{2}+i\nu}^\pm(X^\mu; \vec{w})$$

$$\Delta \in \frac{d}{2} + i\mathbb{R}_{\geq 0}$$

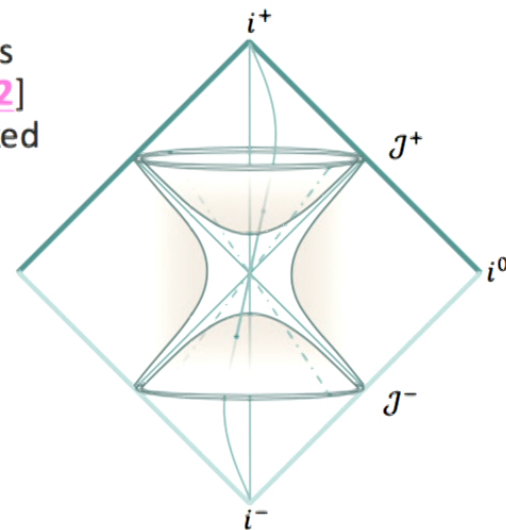
×



$$\vec{w} \in \mathbb{R}^d$$

Amplitude Transforms

- ❖ Note that transforming momentum space amplitudes directly, is an alternative to previous approaches [[hep-th/0303006](#), [arXiv:1609.00732](#)] towards flat space holography, which have looked at a foliation of Minkowski space to reproduce AdS/CFT, dS/CFT on each slice.



What Needs To Be Done

- ❖ The current map interpreting S-matrix elements as 2D CFT correlators seems to imply either an exotic CFT 2 or that the map needs to be finessed... **Options?**

➤ *Is there a better shadow-related basis?*

$$\begin{aligned}\mathcal{O}_{i\lambda}^+(w, \bar{w}) &= \phi_{i\lambda}^+(w, \bar{w}) + C_{+, \lambda} \int d^2 z \frac{1}{(z-w)^{2+i\lambda}(\bar{z}-\bar{w})^{i\lambda}} \phi_{-i\lambda}^-(z, \bar{z}) \\ \mathcal{O}_{i\lambda}^-(w, \bar{w}) &= \phi_{i\lambda}^-(w, \bar{w}) + C_{-, \lambda} \int d^2 z \frac{1}{(z-w)^{i\lambda}(\bar{z}-\bar{w})^{2+i\lambda}} \phi_{-i\lambda}^+(z, \bar{z})\end{aligned}$$

➤ The mode combination that decouples in the soft limit is precisely a linear combination of Mellin and Mellin+shadow in the limit where $\text{Im } \Delta = 0$:

$$\mathbf{a}_- \equiv a_-(\omega \hat{x}) - \frac{1}{2\pi} \int d^2 w \frac{1}{\bar{z} - \bar{w}} \partial_{\bar{w}} a_+(\omega \hat{y})$$

Understand the conformally soft limit!

What Has Been Done

- ❖ Beautiful expressions for full mellin transform (which inherently probe UV structure) of string amplitudes [[arXiv:1806.05688](#)]
- ❖ Systematic n-pt N^k MHV [[arXiv:1711.08435](#)]
- ❖ 3D example of CB decomposition [[arXiv:1711.06138](#)]
- ❖ Interesting statements about symplectic pairing of conformally soft modes [[arXiv:1810.05219](#)]