

Title: PSI 2017/2018 - Relativistic Quantum Information - Lecture 11

Date: Apr 03, 2018 09:00 AM

URL: <http://pirsa.org/18040004>

Abstract:

Communication through massless fields

Communication through a massless fields in vacuum

- Information propagates arbitrarily slow even for massless field.
- Recover the message even if the beam is missed.
- Information flow not supported by real quanta (photons) flow.
- Information flow in absence of energy flow.

The screenshot shows the Phys.org website header with navigation links for Nanotechnology, Physics, Earth, Astronomy & Space, Technology, Chemistry, Biology, and Other Sciences. Below the header are social media icons for Facebook, Twitter, RSS, Email, and a mobile phone icon, along with a search bar. The article breadcrumb is 'Home > Physics > Quantum Physics > March 31, 2015'. The main title is 'Photon 'afterglow' could transmit information without transmitting energy' and the author is 'March 31, 2015 by Lisa Zyga' with a 'feature' tag.

(Phys.org)—Physicists have theoretically shown that it is possible to transmit information from one location to another without transmitting energy. Instead of using real photons, which always carry energy, the technique uses a small, newly predicted quantum afterglow of virtual photons that do not need to carry energy. Although no energy is transmitted, the receiver must provide the energy needed to detect the incoming signal—similar to the

Antenna states

$$|\psi_A\rangle = \alpha_A |e_A\rangle + \beta_A |g_A\rangle$$

$$|\psi_B\rangle = \alpha_B |e_B\rangle + \beta_B |g_B\rangle$$

$\int_{t=0}^{\infty}$

$$\int dt \int dt' \int dx \int dx' \chi_A(t) \chi_B(t') \frac{1}{r(\vec{x}-\vec{x}_A)} \frac{1}{r(\vec{x}'-\vec{x}_B)} \left\langle \left[\phi(t, \vec{x}), \phi(t', \vec{x}') \right] \right\rangle_{\psi_A, \psi_B}$$

State independent

ONLY NON-ZERO IF
THERE ARE COHERENCES IN INITIAL STATE

Mathematical Methods: Beyond the Strong Huygens Principle

Subtleties in the behaviour of the solutions of certain PDEs:
The strong Huygens principle

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The Green's function of the (massless) wave equation in 3+1D Minkowski space has support only on the light cone. Hence, any disturbances propagate strictly along null geodesics (at the speed of light)

Mathematical Methods: Beyond the Strong Huygens Principle

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The strong Huygens principle

The Green's function of the (massless) wave equation in 3+1D Minkowski space has support only on the light cone. Hence, any disturbances propagate strictly along null geodesics (at the speed of light)

Exploitable when emitters are quantum!

TECHNICAL DETAILS

R. H. Jonsson, E. Martin-Martinez, A. Kempf, Phys. Rev. Lett. 114, 110505 (2015)

A. Blasco, L. J. Garay, M. Martin-Benito, E. Martin-Martinez, Phys. Rev. Lett. 114, 141103 (2015)

A. Blasco, L. J. Garay, M. Martin-Benito, E. Martin-Martinez, Phys. Rev. D 93, 024055 (2016)

P. Simidzija, E. Martin-Martinez, Phys. Rev. D 95, 025002 (2017)

See also:

R. H. Jonsson, J. of Phys. A, 44, 445402 (2016)

STRONG HUYGENS PRINCIPLE

The radiation Green's function (or equivalently the commutator) of a massless field has support only on the light-cone

$$\square G(x, x') = -4\pi\delta_4(x, x') \quad [\Phi(x), \Phi(x')] = \frac{i}{4\pi}G(x, x')$$

—————> **Communication** has support only on the **light-cone**

True in 3+1 Flat spacetime

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BEYOND THE STRONG HUYGENS PRINCIPLE

In general: if there is curvature (unless there is conformal invariance)

→ In curved spacetimes, **communication through massless fields** is not confined to the light-cone, but there can be a leakage of information towards the **inside of the light-cone decoupled from energy propagation.**

SPATIALLY **FLAT**, **OPEN FRW** SPACETIME 3+1D:

$$ds^2 = a(\eta)^2(-d\eta^2 + dr^2 + r^2 d\Omega^2)$$

η : conformal time
 $a(\eta)$: scale factor
 t : cosmological time,
 $dt = a(\eta)d\eta$
 units: $\hbar = c = 1$

This geometry will be generated by:

a **perfect fluid** with a constant density-to-pressure ratio $p = w\rho$

→ the **scale factor** evolves as $a \propto \eta^{\alpha + \frac{1}{2}} \propto t^{\frac{2\alpha + 1}{2\alpha + 3}}$ with $\alpha = \frac{3 - 3w}{6w + 2}$

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A TEST SCALAR FIELD QUANTIZED IN THE **BUNCH-DAVIS VACUUM** WILL BE COUPLED TO THE BACKGROUND GEOMETRY.

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TRANSMISSION OF INFORMATION

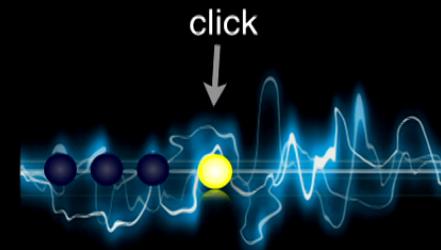
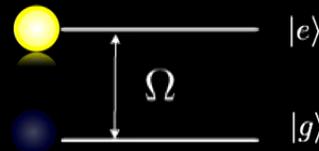
Influence of the presence of A on B \longrightarrow **SIGNALING ESTIMATOR, S**

how much information can be sent? \longrightarrow **CHANNEL CAPACITY, C**

ALICE & BOB'S DETECTOR MODEL

Unruh-DeWitt DETECTOR

-Two-level system



-Energy gap ground-excited states:

Ω

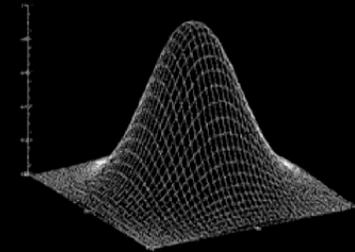
-Monopole moment operator:

$$\nu = \{A, B\}$$

$$\mu_\nu(t) = |e_\nu\rangle\langle g_\nu|e^{i\Omega_\nu t} + |g_\nu\rangle\langle e_\nu|e^{-i\Omega_\nu t}$$

-Spatially smeared:

$$F(\vec{x}, t) = \frac{1}{\sigma^3\sqrt{\pi^3}}e^{-a(t)^2\vec{x}^2/\sigma^2}$$



Detectors: $|\psi_\nu\rangle = \alpha_\nu |e_\nu\rangle + \beta_\nu |g_\nu\rangle$

DETECTOR-FIELD
INTERACTION
HAMILTONIAN

$$H_{I,\nu} = \lambda_\nu \chi_\nu(t) \mu_\nu(t) \int d^3\mathbf{x} a(t)^3 F[\mathbf{x} - \mathbf{x}_\nu(t), t] \Phi[\mathbf{x}, \eta(t)]$$

Coupling strength

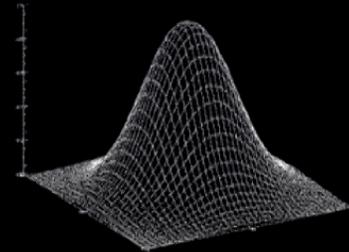
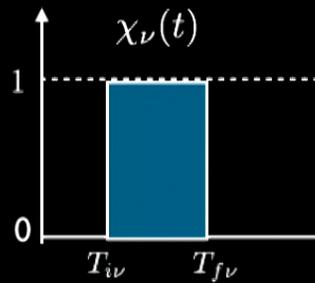
Monopole moment

Scale factor

Smearing function

Detector's trajectory

Switching function



Total Interaction
Hamiltonian:

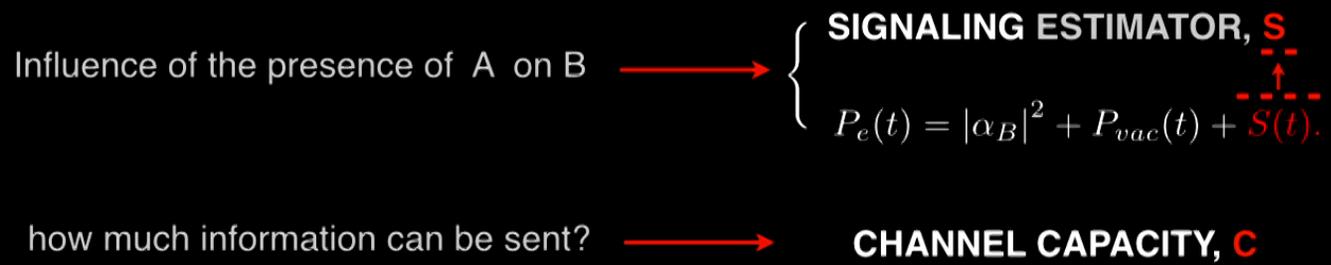
$$H_I = H_{I,A} + H_{I,B}$$

TRANSMISSION OF INFORMATION

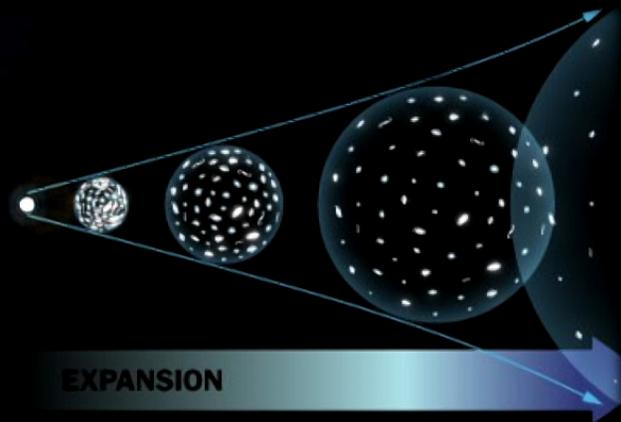
Influence of the presence of A on B \longrightarrow **SIGNALING ESTIMATOR, S**

how much information can be sent? \longrightarrow **CHANNEL CAPACITY, C**

TRANSMISSION OF INFORMATION



THE BIG BANG Setting



BIG BANG CASE, ST. COSMOLOGICAL MODEL: GENERAL RELATIVITY

**SCALAR FIELD:
COUPLING TO GRAVITY**

KLEIN-GORDON EQUATION

$$(\square - m^2 + \xi R)\phi = 0 \quad \square = \frac{1}{\sqrt{|g|}} \partial_\mu (\sqrt{|g|} g^{\mu\nu} \partial_\nu)$$

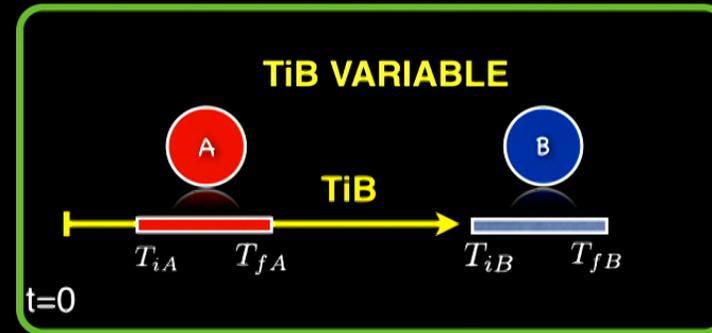
CONFORMAL COUPLING

$$\xi = \frac{1}{6} \quad \text{Yields Conformally Invariant Action}$$

MINIMAL COUPLING

$$\xi = 0 \quad \text{Gives good predictions (Cosmology, etc..)}$$

CASE : Variation of **temporal** separation



CASE: Variation of **spatial** separation



IS INFORMATION
TRANSMITTED?

Influence of the presence of A on B \longrightarrow { **SIGNALING ESTIMATOR, S**
 $P_e(t) = |\alpha_B|^2 + P_{vac}(t) + \overset{\uparrow}{S(t)}$.

**SIGNALING
ESTIMATOR, S**

$$S = \lambda_A \lambda_B S_2 + \mathcal{O}(\lambda_\nu^4)$$

$$S_2 = 4 \int a(t)^3 d^3 \mathbf{x} dt \int a(t')^3 d^3 \mathbf{x}' dt' \chi_A(t) \chi_B(t') \operatorname{Re}(\alpha_A^* \beta_A) F(\mathbf{x} - \mathbf{x}_A, t) \\ \times F(\mathbf{x}' - \mathbf{x}_B, t) \operatorname{Re}(\alpha_B^* \beta_B [\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')])$$

CONFORMAL COUPLING

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$$[\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')] = \frac{i}{4\pi} \left[\frac{\delta(\Delta\eta + |\mathbf{x} - \mathbf{x}'|) - \delta(\Delta\eta - |\mathbf{x} - \mathbf{x}'|)}{a(t)a(t')|\mathbf{x} - \mathbf{x}'|} \right]$$

$$\Delta\eta = \eta(t) - \eta(t')$$

$$|\psi_{0,\nu}\rangle = \alpha_\nu |e_\nu\rangle + \beta_\nu |g_\nu\rangle$$

CHANNEL CAPACITY

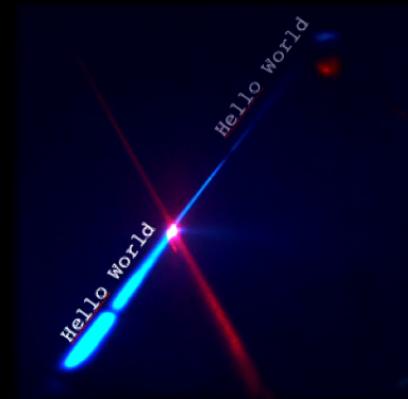
To obtain a lower bound to the channel capacity, we use a simple
COMMUNICATION PROTOCOL:

- **Alice** encodes “1” by coupling her detector A to the field, and “0” by not coupling it.
- Later **Bob** switches on B and measures its energy. If B is excited, Bob interprets a “1”, and a “0” otherwise.

$$C \simeq \lambda_A^2 \lambda_B^2 \frac{2}{\ln 2} \left(\frac{S_2}{4|\alpha_B||\beta_B|} \right)^2 + \mathcal{O}(\lambda_\nu^6)$$

(noisy asymmetric binary channel)

Robert H. Jonsson, Eduardo Martín-Martínez, and Achim Kempf.
Quantum Collect Calling.
Phys. Rev. Lett. 114, 110505 (2015).



CONFORMAL COUPLING

$$[\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')] = \frac{i}{4\pi} \left[\frac{\delta(\Delta\eta + |\mathbf{x} - \mathbf{x}'|) - \delta(\Delta\eta - |\mathbf{x} - \mathbf{x}'|)}{a(t)a(t')|\mathbf{x} - \mathbf{x}'|} \right]$$

support on the
light cone

Decay with Spatial
separation

MINIMAL COUPLING

$$S = \lambda_A \lambda_B S_2 + \mathcal{O}(\lambda_\nu^4)$$

$$S_2 = 4 \int a(t)^3 d^3 \mathbf{x} dt \int a(t')^3 d^3 \mathbf{x}' dt' \chi_A(t) \chi_B(t') \text{Re}(\alpha_A^* \beta_A) F(\mathbf{x} - \mathbf{x}_A, t) \\ \times F(\mathbf{x}' - \mathbf{x}_B, t) \text{Re}(\alpha_B^* \beta_B [\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')])$$

$$[\phi(\mathbf{x}_A, t_A), \phi(\mathbf{x}_B, t_B)] = i \frac{\theta(\eta(t_B) - \eta(t_A)) - \theta(\eta(t_A) - \eta(t_B))}{(2\pi)^3 |\mathbf{x} - \mathbf{x}'| a(\eta(t_A)) a(\eta(t_B))} \int_0^\infty dk k \sin(k|\mathbf{x} - \mathbf{x}'|) \hat{g}(\eta(t_A), \eta(t_B), k)$$

$$\Delta\eta = \eta(t) - \eta(t')$$

$$|\psi_{0,\nu}\rangle = \alpha_\nu |e_\nu\rangle + \beta_\nu |g_\nu\rangle$$

MINIMAL COUPLING

SIGNALING ESTIMATOR, S

$$[\phi(\mathbf{x}_A, t_A), \phi(\mathbf{x}_B, t_B)] = i \frac{\theta(\eta(t_B) - \eta(t_A)) - \theta(\eta(t_A) - \eta(t_B))}{(2\pi)^3 |\mathbf{x} - \mathbf{x}'| a(\eta(t_A)) a(\eta(t_B))} \int_0^\infty dk k \sin(k|\mathbf{x} - \mathbf{x}'|) \hat{g}(\eta(t_A), \eta(t_B), k)$$

$$\hat{g}(\eta, \eta', k) = \frac{8\pi}{k} \sqrt{\left| \frac{\eta}{\eta'} \right|} \frac{\text{sgn}(\eta') [J_{\alpha-1/2}(k|\eta|) Y_{\alpha-1/2}(k|\eta'|) - Y_{\alpha-1/2}(k|\eta|) J_{\alpha-1/2}(k|\eta')]}{Y_{\alpha-1/2}(k|\eta'|) [J_{\alpha-3/2}(k|\eta'|) - J_{\alpha+1/2}(k|\eta')]} - J_{\alpha-1/2}(k|\eta') [Y_{\alpha-3/2}(k|\eta') - Y_{\alpha+1/2}(k|\eta')]$$

J_α, Y_α BESSEL FUNCTIONS

**SIGNALING
ESTIMATOR, S**

$$[\phi(\mathbf{x}_A, t_A), \phi(\mathbf{x}_B, t_B)] = i \frac{\theta(\eta(t_B) - \eta(t_A)) - \theta(\eta(t_A) - \eta(t_B))}{(2\pi)^3 |\mathbf{x} - \mathbf{x}'| a(\eta(t_A)) a(\eta(t_B))} \int_0^\infty dk k \sin(k|\mathbf{x} - \mathbf{x}'|) \hat{g}(\eta(t_A), \eta(t_B), k)$$

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J_α Y_α BESSEL FUNCTIONS

**MATTER DOMINATED
UNIVERSE** \longrightarrow $\alpha = 2$ \longrightarrow $a \propto \eta^2 \propto t^{2/3}$

$$J_{2-1/2}(k|\eta|) = \sqrt{2/\pi} \frac{1}{\sqrt{k|\eta|}} \left[-\cos(k\eta) + \frac{\sin(k\eta)}{k\eta} \right]$$

$$Y_{2-1/2}(k|\eta|) = \sqrt{2/\pi} \frac{\text{sgn}(\eta)}{\sqrt{k|\eta|}} \left[-\sin(k\eta) + \frac{\cos(k\eta)}{k\eta} \right]$$

MINIMAL COUPLING

SIGNALING
ESTIMATOR, S

MATTER DOMINATED
UNIVERSE $\longrightarrow \alpha = 2 \longrightarrow a \propto \eta^2 \propto t^{2/3}$

$$[\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')] = \frac{i}{4\pi} \left[\frac{\delta(\Delta\eta + |\mathbf{x} - \mathbf{x}'|) - \delta(\Delta\eta - |\mathbf{x} - \mathbf{x}'|)}{a(t)a(t')|\mathbf{x} - \mathbf{x}'|} + \frac{\theta(-\Delta\eta - |\mathbf{x} - \mathbf{x}'|) - \theta(\Delta\eta - |\mathbf{x} - \mathbf{x}'|)}{a(t)a(t')\eta(t)\eta(t')} \right]$$

CONFORMAL COUPLING

SIGNALING ESTIMATOR, S

$$[\phi(\mathbf{x}, t), \phi(\mathbf{x}', t')] = \frac{i}{4\pi} \left[\frac{\delta(\Delta\eta + |\mathbf{x} - \mathbf{x}'|) - \delta(\Delta\eta - |\mathbf{x} - \mathbf{x}'|)}{a(t)a(t')|\mathbf{x} - \mathbf{x}'|} \right]$$

support on the light cone

Decay with Spatial separation

MINIMAL COUPLING

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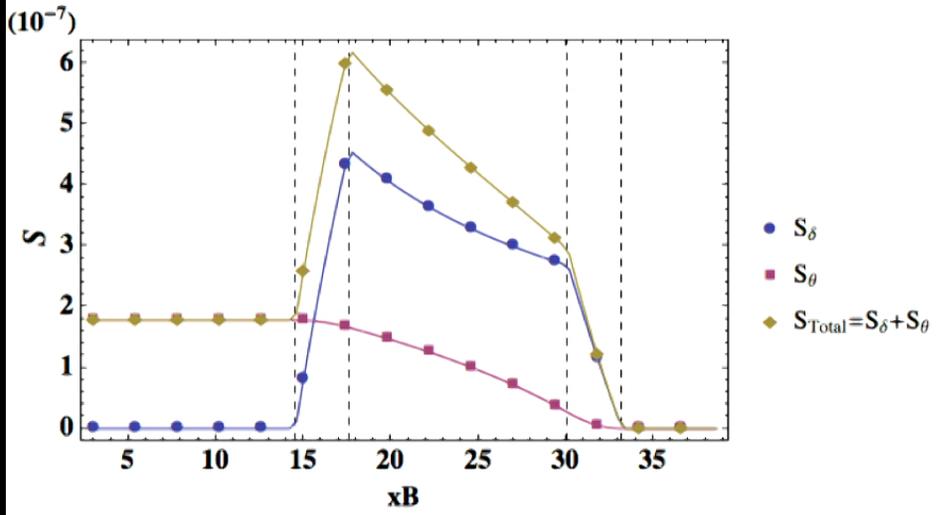
VIOLATION OF STRONG HUYGENS PRINCIPLE !!!!

Does NOT decay with Spatial separation

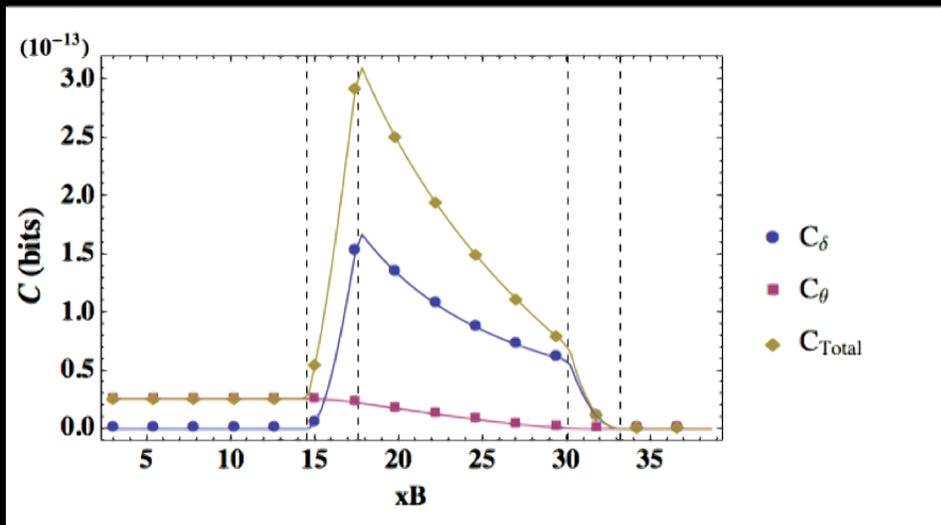
Timelike-leakage

Case:
Variation of
spatial
separation

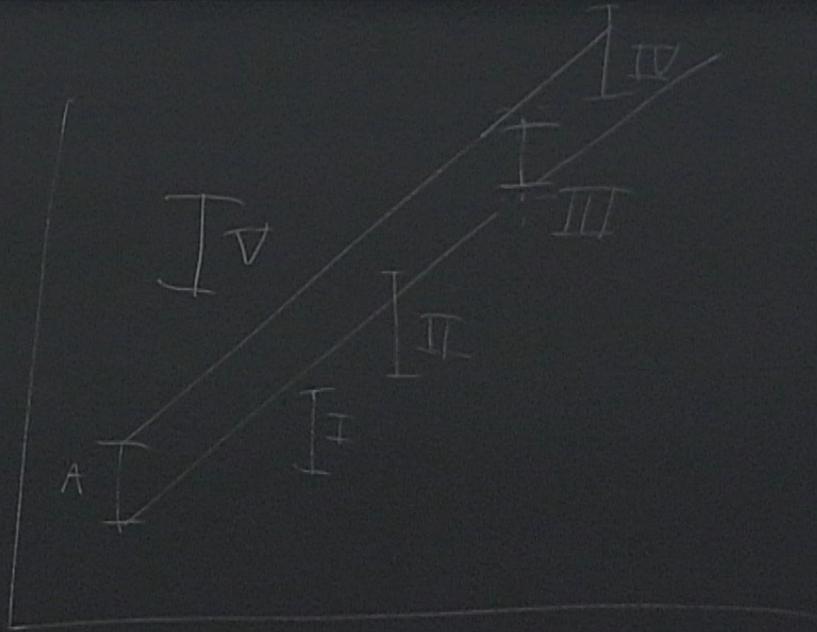
MINIMAL COUPLING



**SIGNALING
ESTIMATOR, S**

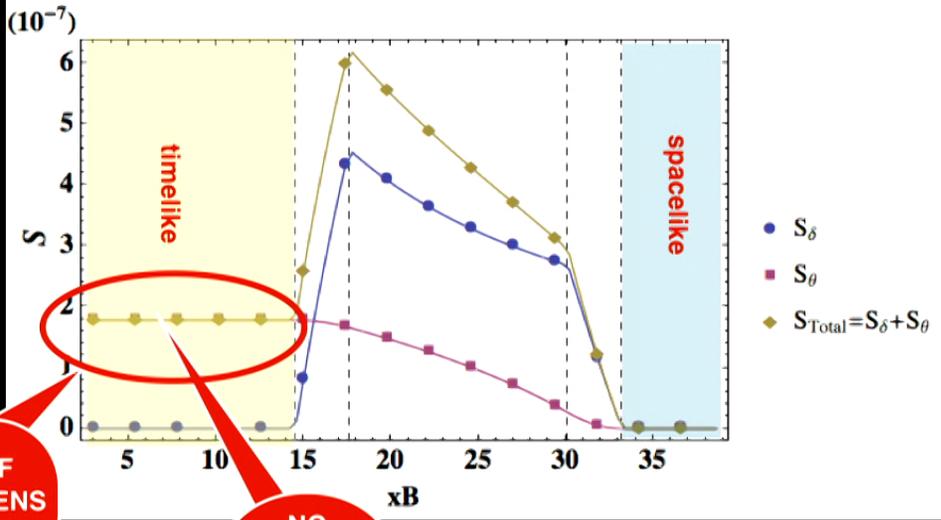


**CHANNEL
CAPACITY**



Case:
Variation of
spatial
separation

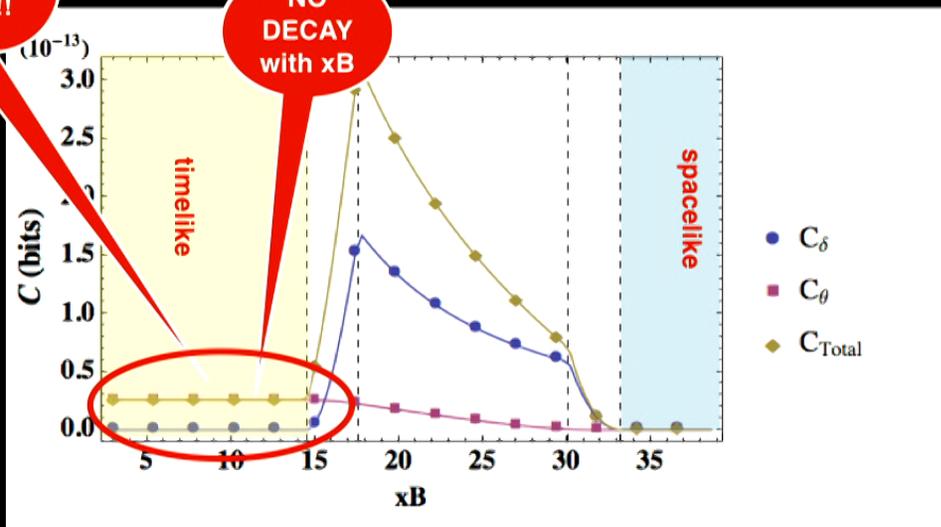
MINIMAL COUPLING



**SIGNALING
ESTIMATOR, S**

**VIOLATION OF
STRONG HUYGENS
PRINCIPLE !!!!**

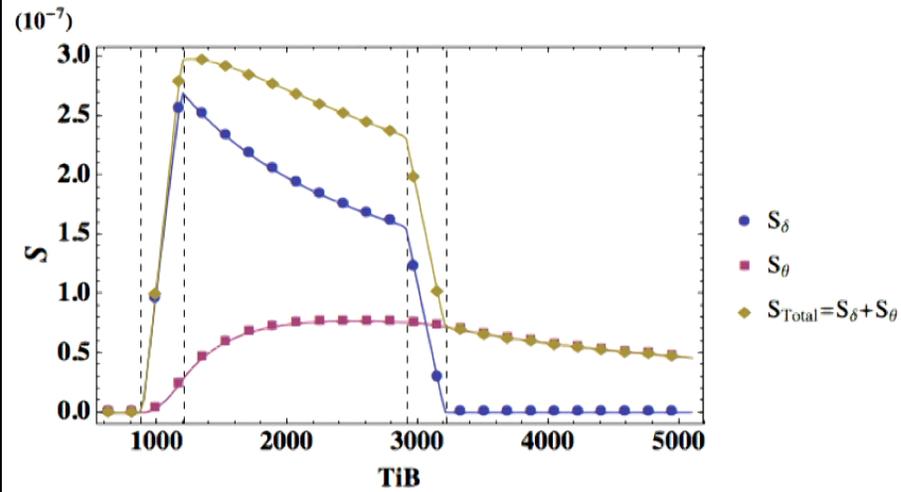
**NO
DECAY
with x_B**



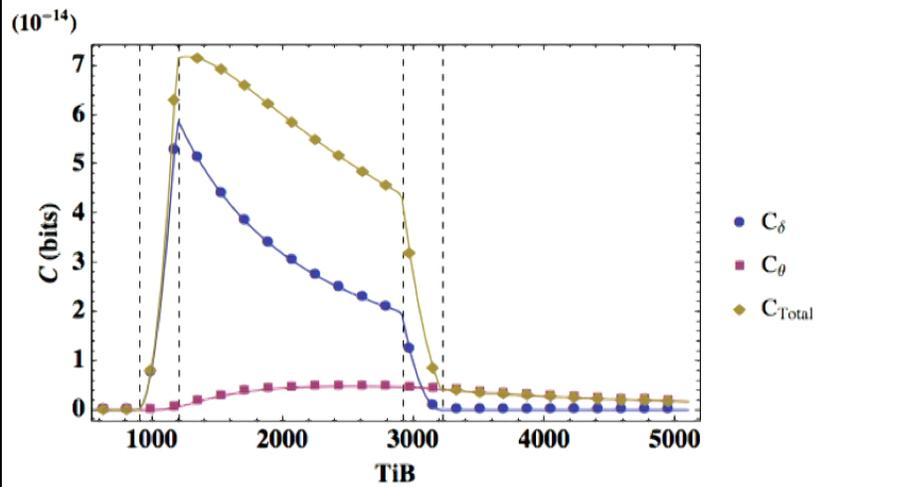
**CHANNEL
CAPACITY**

Case:
Variation of
temporal
separation

MINIMAL COUPLING



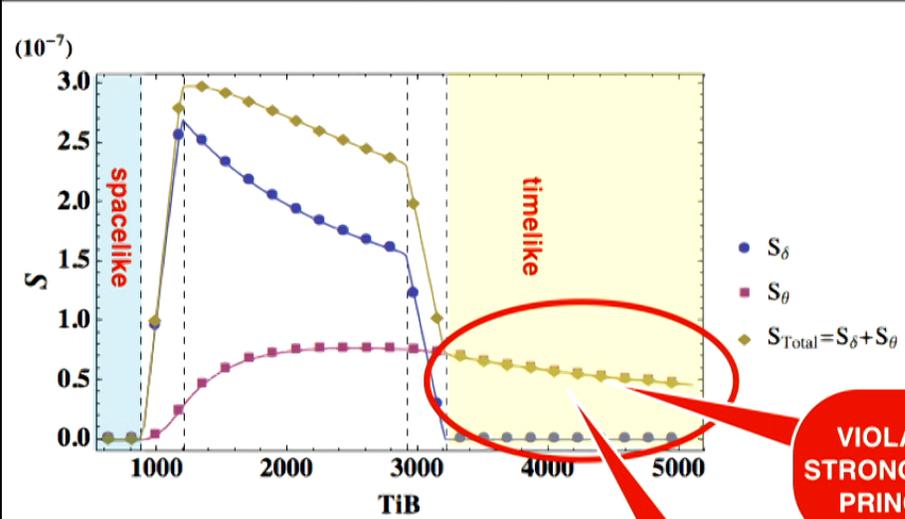
**SIGNALING
ESTIMATOR, S**



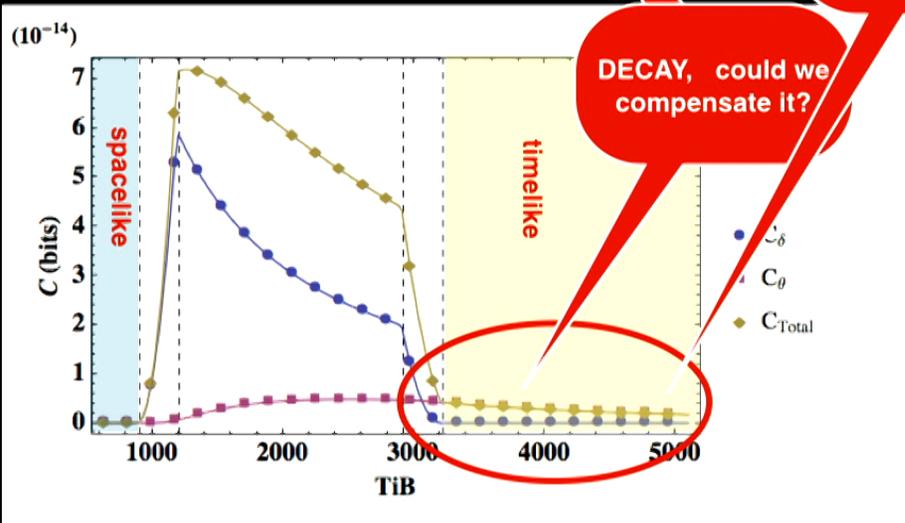
**CHANNEL
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Case:
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SIGNALING
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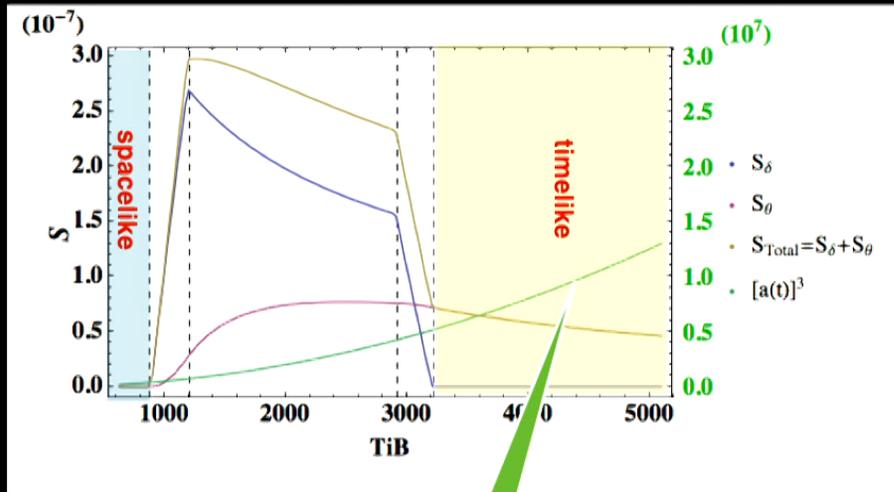
CHANNEL
CAPACITY

VIOLATION OF
STRONG HUYGENS
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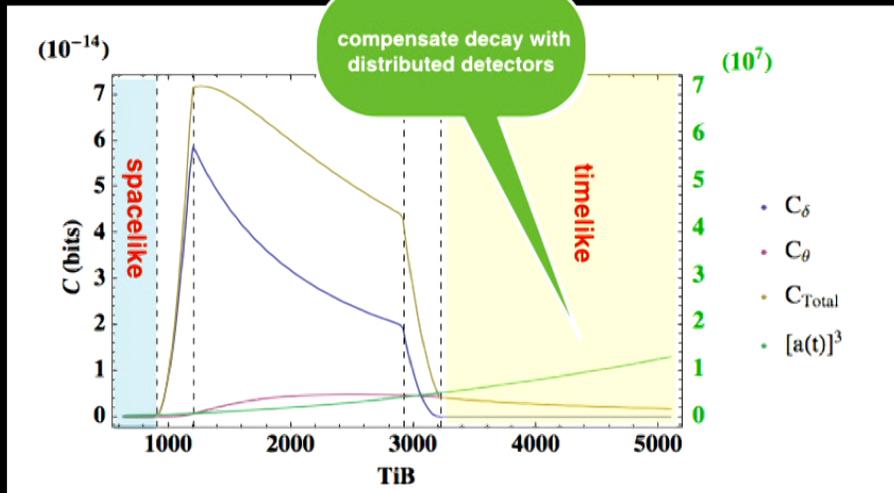
DECAY, could we
compensate it?

Case:
Variation of
temporal
separation

MINIMAL COUPLING



SIGNALING
ESTIMATOR, S



CHANNEL
CAPACITY

Exponential Expansion (deSitter): No decay in time!

P. Simidzija, E. Martin-Martinez, *Phys. Rev. D* 95, 025002 (2017)

AND...WHAT HAPPENS in the
case of **QUANTUM BOUNCE**
Setting ?

How much information survives a Cosmological
cataclysm!!!!



Cosmological
cataclysm!!!!



information