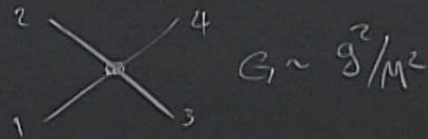
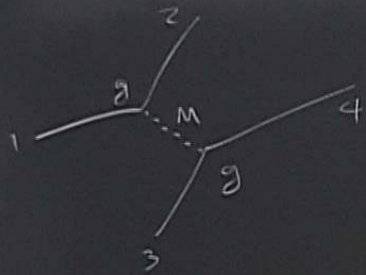


Title: PSI 17/18 - Beyond Standard Model - lecture 7

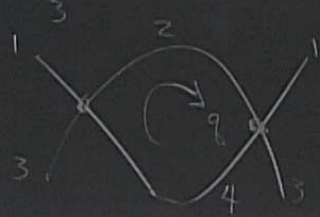
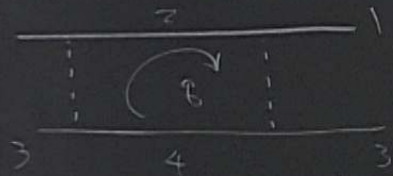
Date: Feb 27, 2018 09:00 AM

URL: <http://pirsa.org/18020107>

Abstract:



$$G \sim g^2/M^2$$



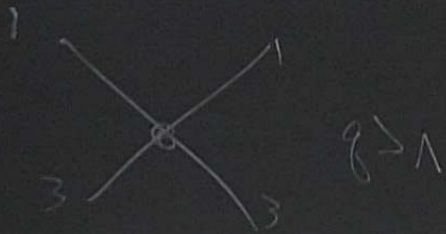
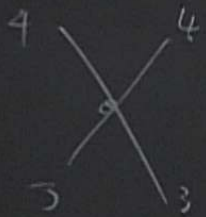
$$g \ll 1$$



$$g \gg 1$$

$$m \ll \Lambda \ll M$$

$$G \sim g^2/M^2$$



$$G^2 \int \frac{d^4 q}{(2\pi)^4} \frac{1}{q^2} \sim G^2 \Lambda^2$$

$$G^2 \int \frac{d^4 q}{(2\pi)^4} \frac{1}{q^2} \sim G^2 \Lambda^2$$

→ low-high energy split.

the effective couplings  
describing low-energy implications  
of high-energy states must be  
 $\Lambda$ -dependent (depend on details of

Formally:

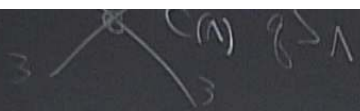
$l$ : light degrees of freedom (SM)

$h$ : heavy " " "

Only can access observables involving  $l$ .

$\langle O_1(x) \dots O_n(x) \rangle$

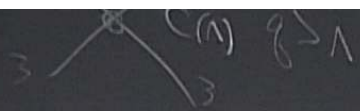
$D_l D_h$



$\hbar$  dependent (depend on details of)

$$\langle T^* [O_1(x) \dots O_n(x)] \rangle = \int \mathcal{D}\ell \mathcal{D}h e^{iS(\ell, h)} O_1(x) \dots O_n(x)$$

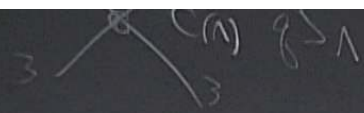
involving  $\ell$ .



$\hbar$  dependent (depend on details of)

$$\langle \mathbb{T}^* [O_1(x) \dots O_n(x)] \rangle = \int \mathcal{D}x \left[ \mathcal{D}h e^{iS(x,h)} \right] O_1(x) \dots O_n(x)$$

involving  $h$ .



$\Lambda$  dependent (depend on details of)

$$\langle \prod_{i=1}^n O_i(x_i) \rangle = \int \mathcal{D}\phi \left[ \mathcal{D}h e^{iS(\phi, h)} \right] O_1(x_1) \dots O_n(x_n)$$

$$\underbrace{e^{iS_W(\phi)}}_{\text{Wilson action}} = \int \mathcal{D}h e^{iS(\phi, h)}$$

involving  $h$ .

independent (depend on details of)

$$|Q_n(t)\rangle = \int \mathcal{D}x \left[ \mathcal{D}h e^{iS(x,h)} \right] Q_1(x) \dots Q_n(x) = \int \mathcal{D}x e^{iS_n(x)} Q_1(x) \dots Q_n(x)$$

$\underbrace{iS_n(x)}_{\text{on action}} = \int \mathcal{D}h e^{iS(x,h)}$

independent (depend on details of)

$$\langle O_n(t) \rangle = \int \mathcal{D}x_1 \mathcal{D}x_n \left[ \mathcal{D}h e^{iS(x,h)} \right] O_1(t) \dots O_n(t) = \int \mathcal{D}x e^{iS_n(x)} O_1(t) \dots O_n(t)$$

$\underbrace{iS_n(x)}_{\text{on action}} = \int \mathcal{D}h e^{iS(x,h)}$

(a)  $\lambda \rightarrow \Lambda$

energy states must be  $\Lambda$ -dependent (depend on details of)

$$\langle T [O_1(x) \dots O_n(x)] \rangle = \int \mathcal{D}\phi_1 \left[ \mathcal{D}\psi e^{iS(\phi, \psi)} \right] O_1(x) \dots O_n(x) = \int \mathcal{D}\phi_1 e^{iS_w(\phi; \Lambda)} O_1(x) \dots O_n(x)$$

$$\underbrace{e^{iS_w(\phi; \Lambda)}}_{\text{Wilson action}} = \int \mathcal{D}\psi e^{iS(\phi, \psi)}$$

energy states must be  
 $\Lambda$ -dependent (depend on details of)

$$\langle T^* [O_1(\ell) \dots O_n(\ell)] \rangle = \int \mathcal{D}q_1 \mathcal{D}q_n e^{iS(q,h)} O_1(\ell) \dots O_n(\ell) = \int \mathcal{D}q_1 e^{iS_w(\ell;N)} O_1(\ell) \dots O_n(\ell)$$

$e^{iS_w}$   
 Wilson  
 $\mathcal{D}q_n e^{iS(q,h)}$

$$S_w(\ell;N) = \int dt dx \mathcal{L}_w$$

$$\mathcal{L}_w = \sum_n \frac{1}{\Lambda^2} O_n(\ell)$$

independent (depend on details of)

$$|O_n(\ell)\rangle = \int \mathcal{D}\phi_n \mathcal{D}\psi_n e^{iS(\phi_n, \psi_n)} |O_1(\ell) \dots O_n(\ell)\rangle = \int \mathcal{D}\phi_n e^{iS_n(\ell; \Lambda)} |O_1(\ell) \dots O_n(\ell)\rangle$$

$$iS_n(\ell; \Lambda) = \int \mathcal{D}\phi_n e^{iS/\hbar}$$

an action

$$S_n(\ell; \Lambda) = \int d^4x \mathcal{L}_n$$

$$\mathcal{L}_n = \sum_n \frac{1}{\Lambda^{d_n}} O_n(\ell)$$

$[O_n]$  = mass dimension of  $O_n$   
 $= d_n$

independent (depend on details of)

$$\langle \mathcal{O}_n \rangle = \int \mathcal{D}\ell_1 \mathcal{D}\ell_n e^{iS(\ell_1, \ell_n)} \mathcal{O}_1(\ell) \dots \mathcal{O}_n(\ell) = \int \mathcal{D}\ell_1 e^{iS_n(\ell; N)} \mathcal{O}_1(\ell) \dots \mathcal{O}_n(\ell)$$

$$\mathcal{D}\ell_n e^{iS(\ell_1, \ell_n)}$$

$$S_n(\ell; N) = \int d^4x \mathcal{L}_W$$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n}} \mathcal{O}_n(\ell)$$

$[\mathcal{O}_n]$  = mass dimension  
of  $\mathcal{O}_n$   
=  $d_n$

What usually doesn't  
happen:

$$c_n(\lambda) = c_n(\lambda_0) Z_n(\lambda/\lambda_0)$$

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$$c_n(\lambda) = c_n(\lambda_0) Z_n(\lambda/\lambda_0)$$

$$\begin{pmatrix} c_1(\lambda) \\ \vdots \\ c_n(\lambda) \end{pmatrix} = Z_{ij}(\lambda/\lambda_0) \begin{pmatrix} c_1(\lambda_0) \\ \vdots \\ c_n(\lambda_0) \end{pmatrix}$$

$$Z_{ij} = \delta_{ij} + b_{ij} \frac{\alpha}{4\pi} \ln\left(\frac{\lambda}{\lambda_0}\right) + \dots$$

$$= e^{b \frac{\alpha}{4\pi} \ln(\lambda/\lambda_0)}$$

$$= \left(\frac{\lambda}{\lambda_0}\right)^{\frac{\alpha}{4\pi} b}$$

what usually doesn't  
happen:

$$c_n(\lambda) = c_n(\lambda_0) Z_n(\lambda)$$

$$\begin{pmatrix} q(\omega) \\ \vdots \\ s(\omega) \end{pmatrix} = Z_j(\lambda/\lambda_0)$$

$$Z_j = \delta_{ij} + b_{ij} \frac{\alpha}{4\pi} \ln\left(\frac{\lambda}{\lambda_0}\right) + \dots$$

$$= e^{b \frac{\alpha}{4\pi} \ln(\lambda/\lambda_0)}$$

$$= \left(\frac{\lambda}{\lambda_0}\right)^{\frac{\alpha}{4\pi} b}$$

$$\alpha \ll 1$$

$$\alpha \ln(\lambda/\lambda_0) \approx b(i)$$

what usually doesn't  
happen:

$$c_n(\lambda) = c_n(\lambda_0) Z_n(\lambda/\lambda_0)$$

$$\begin{pmatrix} c_1(\lambda) \\ \vdots \\ c_n(\lambda) \end{pmatrix} = Z_{ij}(\lambda/\lambda_0) \begin{pmatrix} c_1(\lambda_0) \\ \vdots \\ c_n(\lambda_0) \end{pmatrix}$$

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anomalous  
dimension

$$\alpha \ll 1$$

$$\alpha \ln(\lambda/\lambda_0) \approx b(i)$$

what usually doesn't  
happen:

$$c_n(\lambda) = c_n(\lambda_0) Z_{ij}(\lambda/\lambda_0)$$

$$\begin{pmatrix} c_1(\lambda) \\ \vdots \\ c_n(\lambda) \end{pmatrix} = Z_{ij}(\lambda/\lambda_0) \begin{pmatrix} c_1(\lambda_0) \\ \vdots \\ c_n(\lambda_0) \end{pmatrix}$$

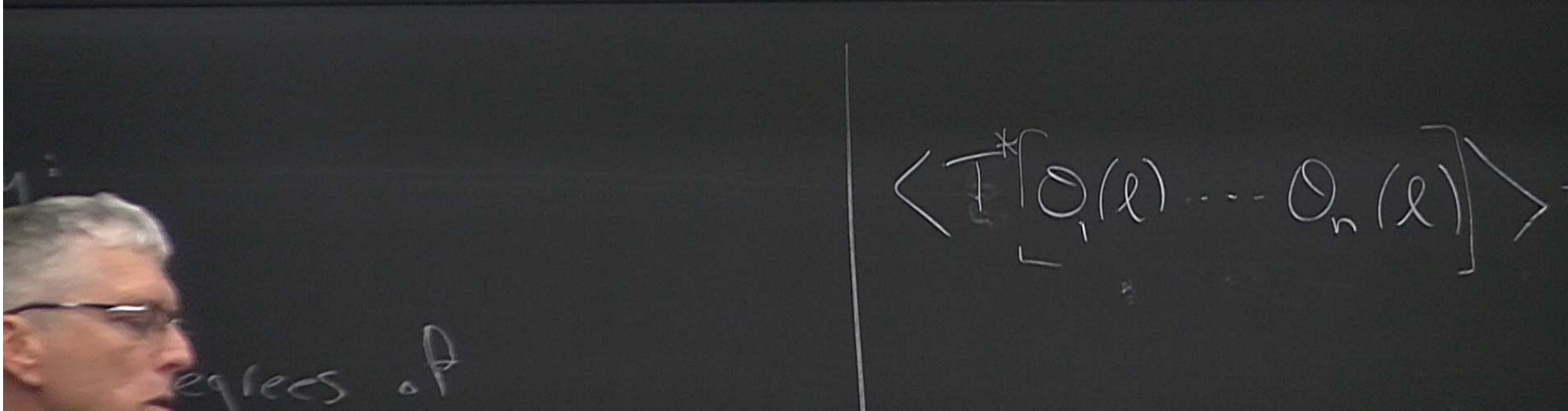
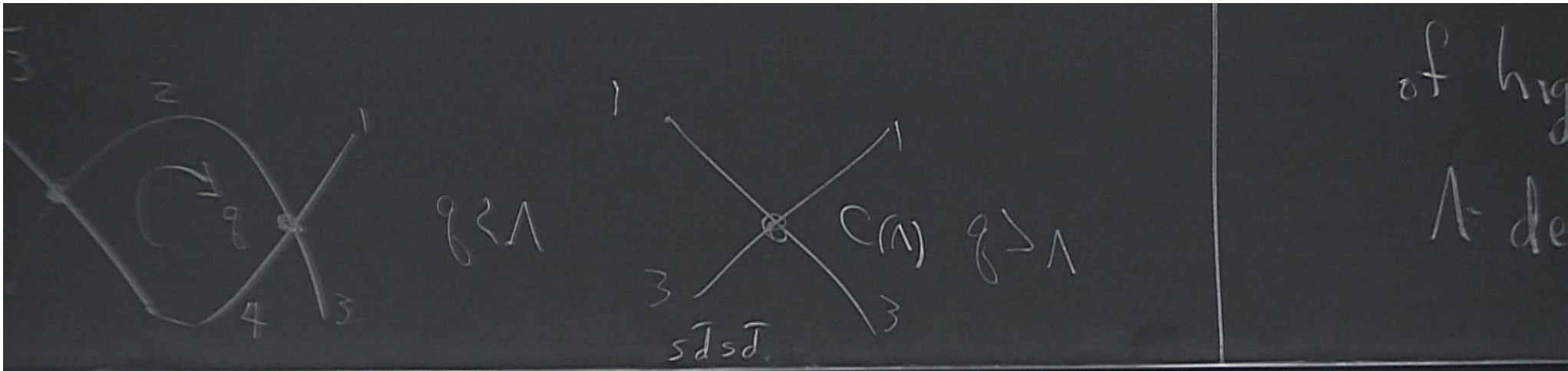
$$Z_{ij} = \delta_{ij} + b_{ij} \frac{\alpha}{4\pi} \ln\left(\frac{\lambda}{\lambda_0}\right) + \dots$$

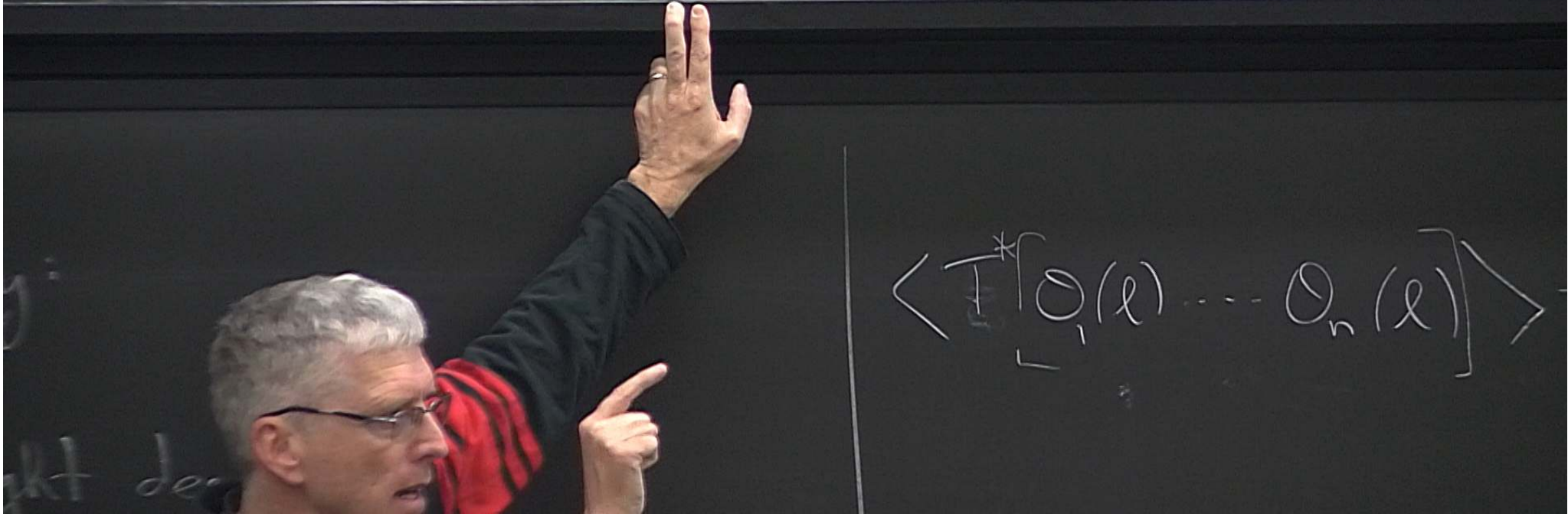
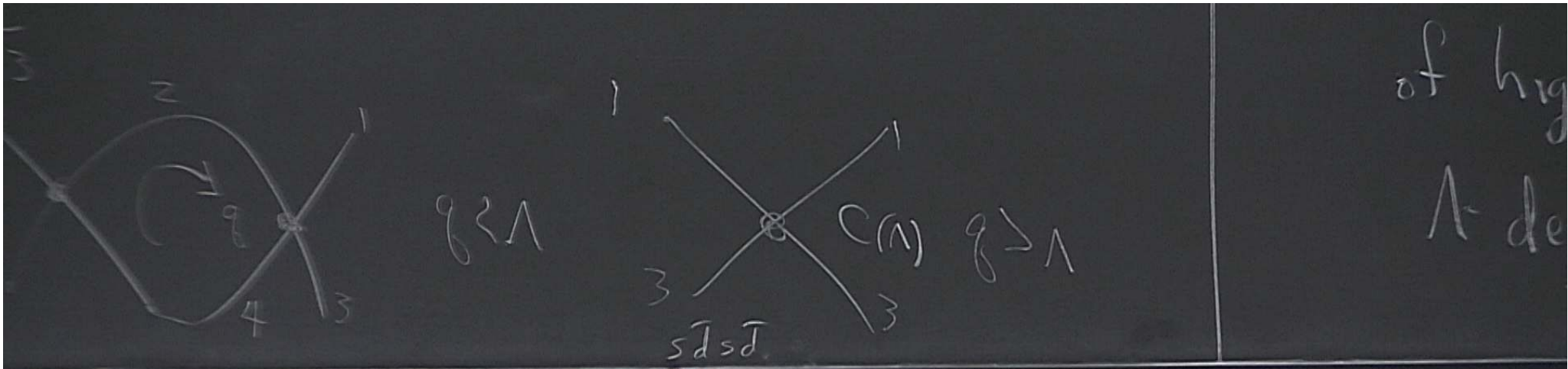
$$= \rho_{ij} + b_{ij} \frac{\alpha}{4\pi} \ln(\lambda/\lambda_0)$$

anomalous  
dimension

$$\alpha \ll 1$$

$$\alpha \ln(\lambda/\lambda_0) \approx \epsilon$$





$$\alpha \ll 1$$

$$\alpha \ln(M_{\text{UV}}) = \mathcal{O}(1)$$

Picture: SM is the most general renormalizable\*  
interaction built from only SM field content  
+ gauge symmetries:

$$* \sum_n \frac{c_n}{\Lambda^{4-d_n}} \mathcal{O}_n(\ell) \quad \text{with } d_n \leq 4$$

$$G \sim \frac{1}{M^2} \text{ (diagram)} G^{n/p}$$



+ gauge symmetries:

$$\sum_n \frac{c_n}{\Lambda^{4-d_n}} O_n(x) \quad \text{with } d_n \leq 4$$

→ Should regard SM as that part of the Wilson action not suppressed by  $\Lambda$ .

Only can access observables involving  $l$ .

Plan: write down most general  $\mathcal{L}$  involving SM fields.

- do not stop at  $\dim \leq 4$

- check when coefficients are not of generic size  $\rightarrow$  selection rules  
+ approx. symm.

$\mathcal{L} = \frac{c_2}{\Lambda^2} \mathcal{O}_2$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\phi)$$

$d_n$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

dim 0:  $1$  ( $\sqrt{-g}$ ) cosmological constant (Dark Energy)  $\approx c.f. \Lambda^4$  (cosmology const prob)  
 If coeff  $\approx (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \approx 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$  (electroweak hierarchy)

dim 4:  $(\phi^\dagger \phi)^2$   $ff\phi$   $F_{\mu\nu}^2$   $f\cancel{D}f$   $(D\phi)^\dagger(D\phi)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\phi)$$

$d_n$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

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dim 4:  $(\phi^\dagger \phi)^2$   $ff\phi$   $F_{\mu\nu}^2$   $f\cancel{D}f$   $(D\phi)^\dagger(D\phi)$   $\left| \theta \in \Lambda^4 \rho F_{\mu\nu} F_{\mu\nu} \right.$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\psi)$$

$d_n$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

dim 0:  $1$  ( $\sqrt{-g}$ ) cosmological constant (Dark Energy)  $\leftarrow$  c.f.  $\Lambda^4$  (cosmology const prob)  
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dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \approx 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$  (electroweak hierarchy)

dim 4:  $(\phi^\dagger \phi)^2$   $f\bar{f}\phi$   $F_{\mu\nu}^2$   $f\bar{f}F$   $(D\psi)^\dagger(D\psi)$   $\left| \theta \in \frac{\text{GeV}^4}{\Lambda^4} \frac{F_{\mu\nu} F_{\mu\nu}}{\Lambda^4} \leftarrow \text{CP violating} \right.$  (Strong CP problem)  
 $\theta_{\text{QCD}} \lesssim 10^{-9} \rightarrow d_H$

$$\langle T^* [O_1(x) \dots O_n(x)] \rangle = \int \mathcal{D}\phi_1 \mathcal{D}\phi_2 \dots e^{iS(\phi, h)} [O_1(x) \dots O_n(x)] = \int \mathcal{D}\phi_1 e^{iS(\phi, h)}$$

$$e^{iS_W(\phi; \Lambda)} = \int \mathcal{D}\phi_1 e^{iS(\phi, h)}$$

↑  
Wilson action

$$S_W(\phi; \Lambda) = \int d^4x$$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} O_n$$

involving  $\phi$ .

$$V_{ij} \bar{u}_i \gamma^m \alpha_L d_j W$$

general fields.  
p et  
coeff

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} O_n$$

dim 0: 1 ( $\sqrt{-g}$ ) cosmological constant (Dark Energy)  
if coeff  $\approx (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \approx 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$

$$e^{iS_W(\ell; \Lambda)} = \int \mathcal{D}h_1 e^{iS(\ell, h)}$$

↑  
Wilson action

$$S_W(\ell; \Lambda) = \int d^4x \mathcal{L}_W$$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n}} \mathcal{O}_n(\ell)$$

$$V_{ij} \bar{u}_i \gamma^m \alpha_L d_j W \quad S \in V_{ij}$$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n}} \mathcal{O}_n$$

dim 0:  $1 (\sqrt{-g})$  cosmological constant (Dark Energy) c.f.  $\Lambda^4$   
 If  $\omega_{eff} = (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \sim 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$

dim 4:  $(\phi^\dagger \phi)^2$   $f\bar{f}\phi$   $F_{\mu\nu}^2$   $f\bar{f}F$   $(D\phi)^\dagger (D\phi)$   $|g \in \text{MCP} \frac{F_{\mu\nu} F_{\mu\nu}}{2} \leftarrow \text{CP violating}$   
 $\theta_{QCD} \lesssim 10^{-9} \rightarrow c_h$

selection rules  
 X. Symm.

$$e^{iS_W(\ell; \Lambda)} = \int \mathcal{D}h_\mu e^{iS(\ell, h)}$$

↑  
Wilson action

$$S_W(\ell; \Lambda) = \int d^4x \mathcal{L}_W$$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\ell)$$

$$V_{ij} \bar{u}_i \gamma^m \alpha_L d_j W \quad S \in V_{ij} \sim O(1)$$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

dim 0:  $1$  ( $\sqrt{-g}$ ) cosmological constant (Dark Energy) c.f.  $\Lambda^4$   
 If  $\omega_{eff} = (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \sim 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$

dim 4:  $(\phi^\dagger \phi)^2$   $f\bar{f}\phi$   $F_{\mu\nu}^2$   $f\bar{f}f$   $(D\phi)^\dagger(D\phi)$   $|g \in \text{NRP } F_{\mu\nu} F_{\mu\nu} \leftarrow \text{CP violating}$   
 $\partial_{\text{acc}} \lesssim 10^{-7} \rightarrow d_n$

$d_j W$

$\delta E V_{ij} \sim \delta(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\phi)$$

$d_n$

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

dim 0:  $1$  ( $\sqrt{-g}$ ) cosmological constant (Dark Energy)  $\propto c.f. \Lambda^4$  (cosmolog const prob)  
if coeff  $\sim (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \sim 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$  (electroweak hierarchy)

dim 4:  $(\phi^\dagger \phi)^2$   $ff\phi$   $F_{\mu\nu}^2$   $f\cancel{D}f$   $(D\phi)^\dagger(D\phi)$   $\theta \in \text{NLP } F_{\mu\nu} F_{\mu\nu} \leftarrow \text{CP violating}$  (Strong CP problem)  
 $\theta_{\text{QCD}} \lesssim 10^{-7} \rightarrow d_H \lesssim 10^{-26} \text{ e-cm}$

$\mathcal{L}_W$

$\delta E_{ij} \sim \mathcal{O}(1)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\phi)$$

edn.

$$\mathcal{L} = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

dim 0:  $1$  ( $\sqrt{-g}$ ) cosmological constant (Dark Energy)  $\leftarrow$  c.f.  $\Lambda^4$  (cosmolog const prob)  
If coeff  $\sim (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \sim 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$  (electroweak hierarchy)

dim 4:  $(\phi^\dagger \phi)^2$   $f\bar{f}\phi$   $F_{\mu\nu}^2$   $f\bar{f}f$   $(D\phi)^\dagger(D\phi)$   $\left| \theta \in \frac{h}{2\pi} \frac{F_{\mu\nu} F_{\mu\nu}}{\Lambda^4} \leftarrow \text{CP violating (Strong CP problem)} \right.$   
 $\theta_{\text{dec}} \lesssim 10^{-7} \rightarrow d_n \lesssim 10^{-26} \text{ e-cm} \sim \frac{e}{(10^{19} \text{ GeV})^2}$

$$\delta E V_{ij} \sim \mathcal{O}(1)$$

$$d_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(\phi)$$

$$\mathcal{L}_\phi = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n$$

$$T_{\mu\nu} = \lambda g_{\mu\nu} \leftarrow p = -p = -\lambda$$

dim 0:  $\lambda \sqrt{-g}$  cosmological constant (Dark Energy) c.f.  $\Lambda^4$  (cosmology const prob)  
 if coeff  $\approx (10^{-3} \text{ eV})^4$

dim 2:  $\mu^2 \phi^\dagger \phi$   $\mu \approx 100 \text{ GeV}$  (Higgs mass,  $G_F$ )  $\Lambda^2$  (electroweak hierarchy)

dim 4:  $(\phi^\dagger \phi)^2$   $f \bar{f} \phi$   $F_{\mu\nu}^2$   $f \not{D} f$   $(D\phi)^\dagger (D\phi)$   $\theta \in \text{mix } F_{\mu\nu} F_{\mu\nu}$   $\leftarrow$  CP violating (Strong CP problem)  
 $\theta_{QCD} \lesssim 10^{-9} \rightarrow d_n \lesssim 10^{-26} \text{ e-cm} \sim \frac{e}{(10^{12} \text{ GeV})^{-1}}$

Only can access observables involving  $l$ .

$$V_{ij} = \bar{u}_i \gamma^m \alpha_{ij} d_j W$$

$$\delta E_{V_{ij}} \sim v(l)$$

Other small things amongst the dim-4 ints:

$$\rightarrow M_W = M_Z \cos \theta \left( 1 + \frac{\alpha}{\pi} + \dots \right) \left[ \begin{array}{l} \text{ruined by: higgs non} \\ \text{doublets} \\ \text{(custodial approx sym)} \end{array} \right]$$

$$G^2 \int \frac{d^4 q}{(2\pi)^4} \frac{1}{q^2} \sim G^2 \Lambda^2$$

the effective couplings  
describing low-energy impl  
of high-energy states m  
 $\Lambda$  dependent (depend on

Only can access observables involving  $l$ .

$$V_{ij} = \bar{u}_i \gamma^m \alpha_{ij} d_j W$$

$$SEV_{ij} \sim v(i)$$

Other small things amongst the dim-4 ints:

$$\rightarrow M_W = M_Z \cos \theta \left( 1 + \frac{\alpha}{\pi} + \dots \right) \left[ \begin{array}{l} \text{ruined by: higgs non} \\ \text{doublets} \\ \text{(custodial approx sym)} \end{array} \right]$$

$$\rightarrow \text{FCNC} \quad K^0 \rightarrow \bar{K}^0 \quad \left( \begin{array}{l} \text{ruined by couplings like} \\ Z \bar{s}d, X \bar{s}d \rightarrow \text{almost} \\ \text{"minimal flav. violat" models} \end{array} \right) \left[ \begin{array}{l} \text{all flav.} \\ \text{models} \end{array} \right]$$

$\Lambda$  dependent

$\bar{u}_i \gamma^\mu \chi_j d_j W$

$\delta \in V_{ij} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(W)$$

$d_n$

(c)  $Z$ -coupling well known  
(LEP)

(d) accidental symmetries:

$u_i \gamma^{\mu} d_j W$

$\delta E V_{ij} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(W)$$

etc.

(c)  $Z$ -coupling well known  
(LEP)

(d) "accidental" symmetries:

although not an input, conservation of  
 $\mathbb{Z}_3 L_e, L_\mu, L_\tau$  are accidentally valid

$\bar{u}_i \gamma^\mu \alpha_j d_j W$

$\delta e V_{ij} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{4-n}} \mathcal{O}_n(W)$$

etc.

(c)  $Z$ -coupling well known  
(LEP)

(d) "accidental" symmetries:

although not an input, conservation of  
 $\mathbb{Z} L_e, L_\mu, L_\tau, \dots$  are accidentally valid

$\bar{u}_i \gamma^m \alpha_{ij} d_j W$

$\delta \in V_{ij} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{4-n}} \mathcal{O}_n(W)$$

etc.

(c)  $Z$ -coupling well known  
(LEP)

(d) "accidental" symmetries:

(ruined by supersymmetry)

although not an input, conservation of  
 $\mathbb{Z}_3 L_e, L_\mu, L_\tau, \dots$  are accidentally valid

$\bar{u}_i \gamma^\mu d_j W$

$\delta E V_{ij} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{4-n}} \mathcal{O}_n(\psi)$$

$\sim d_n$

(c)  $Z$ -coupling well known  
(LEP)

(f) presence of  
Lorentz-violat

(d) "accidental symmetries"

(ruined by supersymmetry)

although not an input, conservation of

$\mathbb{Z}_3 L_{e, \mu, \tau}$  are accidentally valid ( " " "

(e) CP violation is small + involves all  
generations

$\alpha_d; W$

$S \in V_{ii} \sim 0(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{d_n-4}} \mathcal{O}_n(l)$$

$d_n$

(c)  $Z$ -coupling well known  
(LEP)

(f) absence of  
Lorentz-violating  
interactions

(d) accidental symmetries:

(ruined by supersymmetry)

although not an input, conservation of

$L_e, L_\mu, L_\tau$  are accidentally valid ( " " " )

(e) CP violation is small + involves all  
generations

$\sum_j d_j W$

$\delta \in V_{ii} \sim v(i)$

$$\mathcal{L}_W = \sum_n \frac{c_n}{\Lambda^{4-n}} \mathcal{O}_n(\psi)$$

$\sim d_n$

(c)  $Z$ -coupling well known  
(LEP)

(f) absence of  
Lorentz-violating  
interactions

(d) accidental symmetries:

although not an input, conservation of  
 $\mathbb{E} L_e, L_\mu, L_\tau$  are accidentally valid

(g) (parity symmetry)

$$c_1 \partial_\epsilon \phi \partial_\epsilon \phi$$

$$c_2 \nabla^\mu \phi \nabla_\mu \phi$$

(e) CP violation is small + involves all  
generations

Only can access observables involving  $l$ .

$$V_{ij} = \bar{u}_i \gamma^m \chi_L d_j W$$

$$SEV_{ii} \sim v(i)$$

Other small things amongst the dim-4 int's:

$$\rightarrow M_W = M_Z \cos \theta \left( 1 + \frac{\alpha}{\pi} + \dots \right) \left[ \begin{array}{l} \text{ruined by: higgs non} \\ \text{doublets} \\ \text{(custodial approx sym)} \end{array} \right]$$

$$\rightarrow \text{FCNC} \quad K^0 \rightarrow \bar{K}^0 \quad \left( \begin{array}{l} \text{ruined by couplings like} \\ Z \bar{s}d, \chi \bar{s}d \rightarrow \text{almost} \\ \text{"minimal flav. violat"} \text{ models} \end{array} \right)$$

$$\mathcal{L}_5 \quad (1 \text{ oper.}) \quad \Delta L = 2$$

$$\mathcal{L}_6 \quad (B\text{-viol}) \quad \Delta B = \Delta L$$

(c)  $Z$ -coupling well known (LEP)

(d) "accidental symmetries" although not an input, by  $L_e, L_\mu, L_\tau$ .

(e) CP violation is small + general