

Title: Remarks on cosmic censorship and its possible violations

Date: Nov 10, 2017 02:00 PM

URL: <http://pirsa.org/17110098>

Abstract:

On Cosmic Censorship and its potential violations

Roberto Emparan

ICREA & U Barcelona

Quantum Black Holes in the Sky?

Perimeter Institute – Nov 2017



Quantum Gravity in the sky?

How does one get there?

Cosmic censorship

A conjecture about *highest-energy physics*

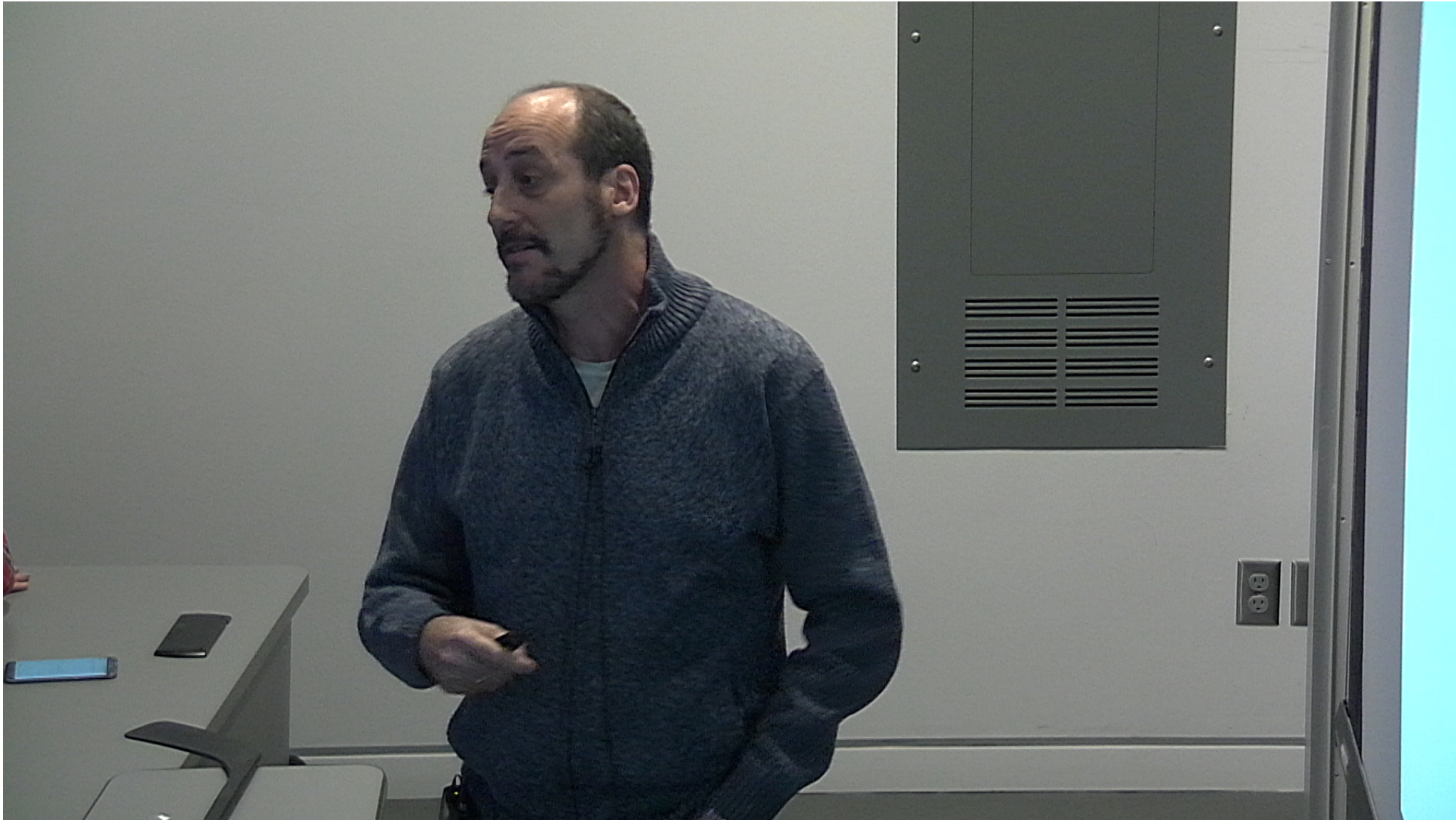
Starting from low-energy densities,
can we, *within a **quick, classical** time-scale*,
get to explore Quantum Gravity?

If cosmic censorship is violated,
predictability (within GR) is lost

May learn new physics

Naked Planck-scale energy densities
easily form

but in long, quantum timescales



Classical timescale

$$t_M^{(cl)} \sim \frac{GM}{c^3} = M$$

Collapse time

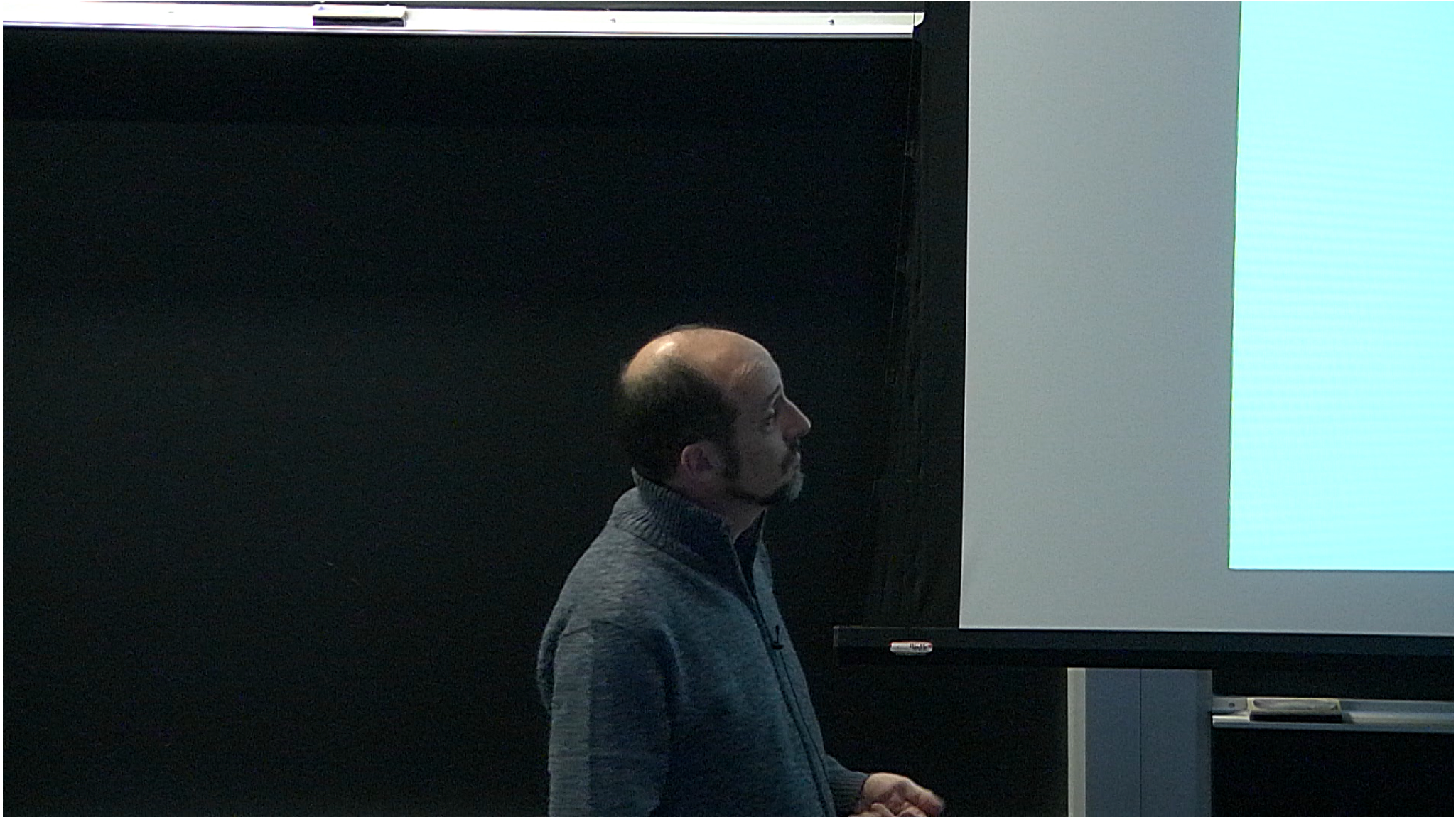
Quantum timescale

$$t_M^{(qu)} \sim t_{Pl} \left(\frac{M}{M_{Pl}} \right)^3 \sim M \left(\frac{M}{M_{Pl}} \right)^2 \gg t_M^{(cl)}$$

Hawking evaporation time & Page time

Firewall formation time (?)

If this quantum time-scale is required to reach
quantum gravity, then forget it...



Quantum timescale?

$$t_M^{(sc)} \sim M \log \frac{M}{M_{Pl}} \gtrsim t_M^{(cl)}$$

Scrambling time

Fuzzball tunneling time

Echo time

Does not appear to involve Planckian curvatures

Seems to require large non-locality

There exist a few well-established violations of
cosmic censorship

Naked singularities of arbitrarily high curvature
form within the classical time scale of the
system

Critical collapse

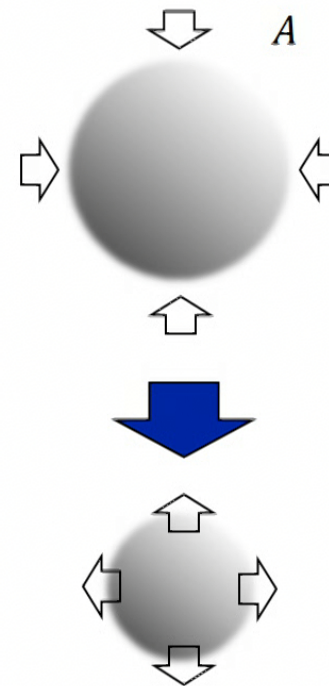
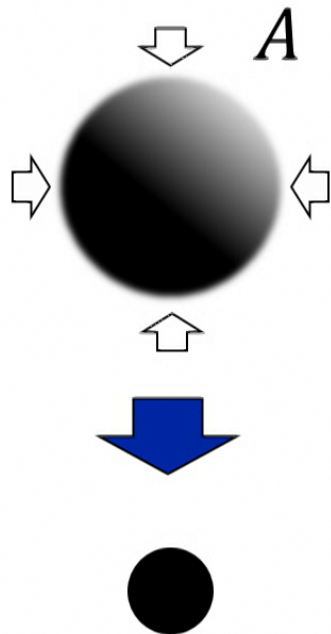
Black string instability

Do these examples suggest a useful
window for exploring quantum gravity?
(even at a very theoretical level)

Critical collapse

Choptuik 1993

Change initial amplitude A



Physically, for a small range of values around A_{crit} there forms a Planck-sized object, with naked Planckian curvature

Quantum gravity becomes observable

Is this a useful violation of CC?

Often disregarded since it requires fine-tuning

Fine-tuning of initial conditions

⇒ reaching Planckian scales requires
purposeful action

- It *is* possible—the job of experimentalists is to do fine-tuning
- It won't happen spontaneously in Nature –
not in the sky
(or only extremely rarely)

But anyway it is not very useful for studying
quantum gravity

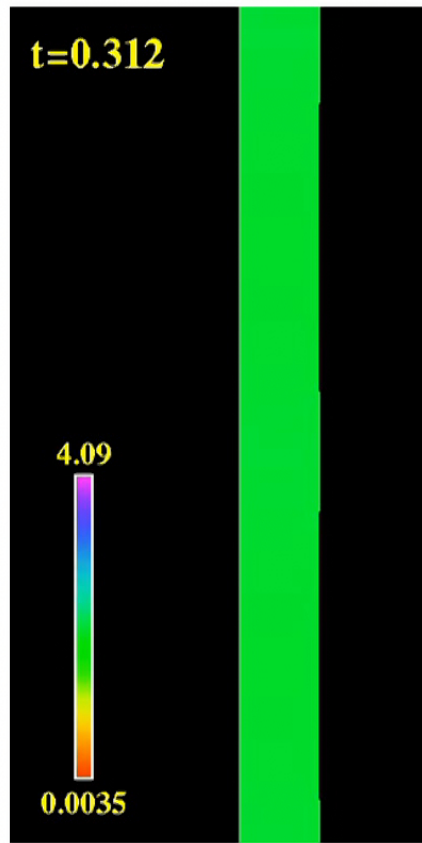
Too small fraction of initial energy goes into
singularity

Evolution across singularity is almost entirely
predictable

Planck-size “black hole” is likely to evaporate
into a few Planckian-energy quanta

Similar to endpoint of Hawking evaporation
(but without info-loss issues)

Black string instability

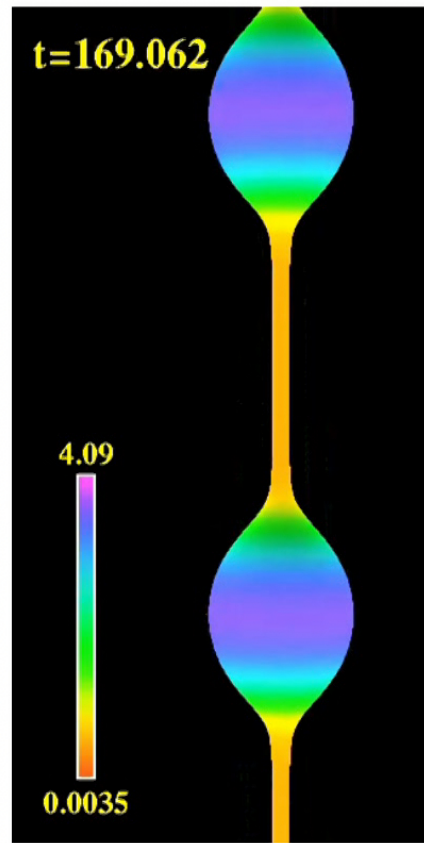


Gregory+Laflamme 1993

Lehner+Pretorius 2010

Evolution of unstable
5D black string

Black string instability

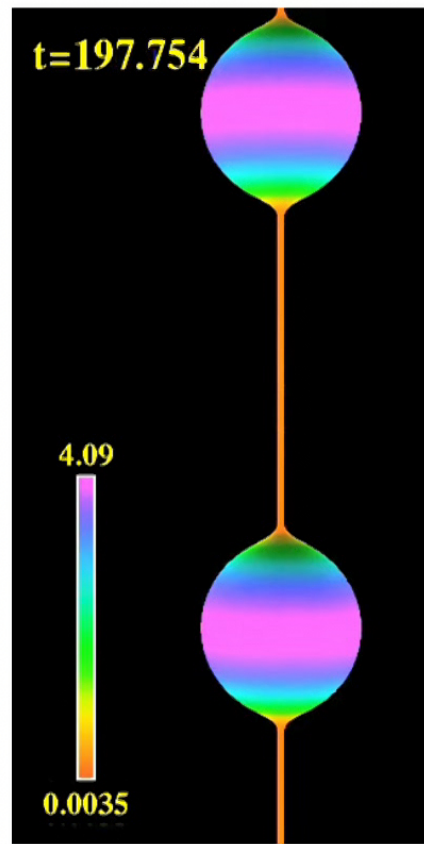


Gregory+Laflamme 1993

Lehner+Pretorius 2010

Evolution of unstable
5D black string

Black string instability

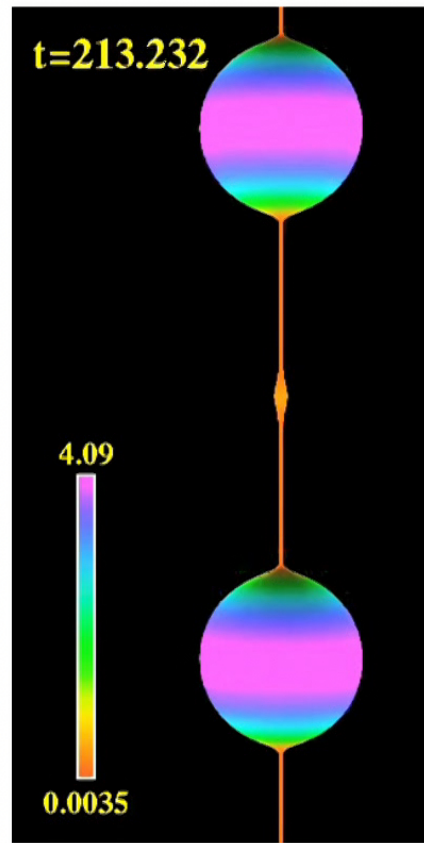


Gregory+Laflamme 1993

Lehner+Pretorius 2010

Evolution of unstable
5D black string

Black string instability

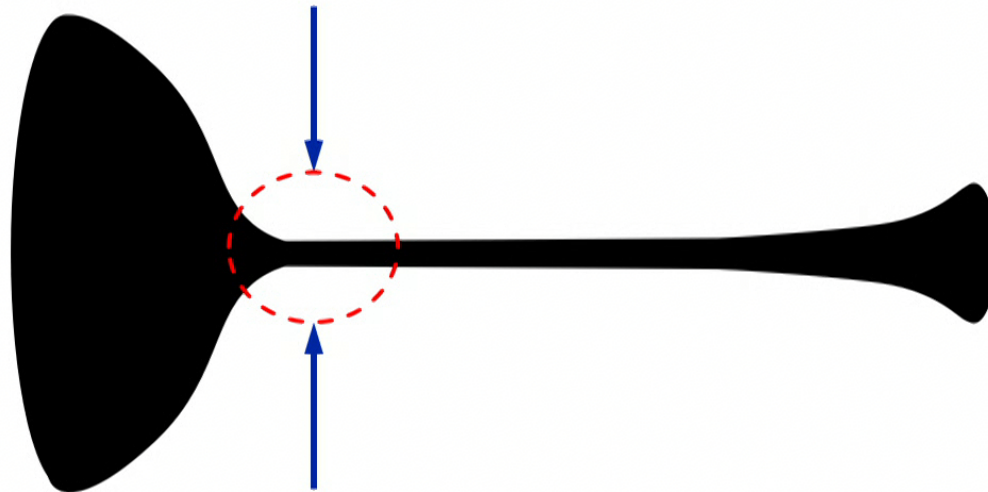


Gregory+Laflamme 1993

Lehner+Pretorius 2010

Evolution of unstable
5D black string

Naked arbitrarily large curvatures form within
the classical timescale of the system



No fine-tuning is needed

Is this useful?

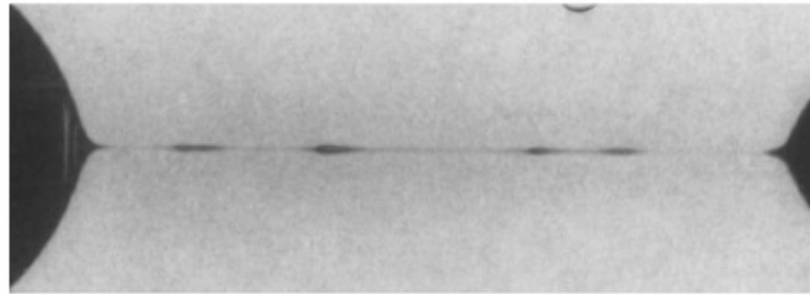
How strong is the loss of
predictability?

A model for black string breakup

RE w/ J Rocha & U Miyamoto

Breakup of fluid jets

Eggers 1993, 1995



Glycerol

FLUID JET BREAKUP

When surface tension becomes too large
molecular dynamics enters
thin neck evaporates

Subsequent evolution is quickly described by
hydrodynamics

BLACK STRING BREAKUP

When surface gravity becomes too large
quantum gravity/string theory enters
thin neck evaporates

Subsequent evolution is quickly described by
classical gravity

Proposal

Breakup controlled by same physics
as at
endpoint of Hawking evaporation

Singularity only occupies a
Planck-sized region
of space and time

Only a **Planck-size mass** is
radiated

Quantum gravity is just a little
pixie dust to effect breakup

Almost all is described
(*predictable*) by classical GR
attractors

If this is correct,
Cosmic Censorship is violated in the
mildest way
during black string break up

A strength-index s for CC violations

$$M_{sing} \sim M \left(\frac{M_{Pl}}{M} \right)^{1-s}$$

$s = 1$: very strong, maximal violation

$s = 0$: very mild, minimal violation

A strength-index s for CC violations

$$M_{sing} \sim M \left(\frac{M_{Pl}}{M} \right)^{1-s}$$

$s = 1$: very strong, maximal violation

$s = 0$: very mild, minimal violation

Forming **naked singularities**

with $s > 0$

could violate **entropy bounds**

$s = 0$ is harmless / useless
(black string, critical collapse)

Cosmic Censorship violations

with $s > 0$

could be forbidden by **Thermodynamics**

There seems to be little chance of
reaching observable quantum gravity
within classical timescales $t_M^{(cl)} \sim M$

Fuzzballs may form in $t_M^{(sc)} \sim M \log \frac{M}{M_{Pl}}$

But fuzzballs don't seem to reveal much
quantum gravity

Cosmic censorship of quantum gravity

A perfect conspiracy?