

Title: Emergent State-Dependence in Holographic Models of Black Holes

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Abstract:

State Dependence in Holographic Models of Black Holes



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Motivation

- Information transfer through weak interactions
[Giddings et al]
- Coupling between entangled CFTs dual to AdS - Schwarzschild wormhole allows to glean information from behind the horizon
[Gao, Jafferis & Wall; Stanford & Maldacena;..]
- Interaction (in addition to entanglement) between the dual CFT and its thermofield double (through heat bath) may avoid firewalls/information paradoxes [Chowdhury,...]
- State-dependent operators
[Papadodimas & Raju]

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Bulk Model

Consistent truncation of N=8 SUGRA to gravity in AdS₄
coupled to an $m^2 = -2$ scalar with potential

$$V(\phi) = -2 - \cosh(\sqrt{2}\phi)$$

In all asymptotically AdS solutions, the scalar falls off like

$$\phi = \frac{\alpha}{r} + \frac{\beta}{r^2}$$

Consider designer gravity boundary conditions:
[TH & Horowitz]

$$\beta(\alpha)$$

$$C_{ij} = \hat{Y}_{ij} \frac{\hat{r}_i + \hat{r}_j}{2}$$

$$= \frac{(\vec{Y}_i - \vec{Y}_j) \cdot \vec{r}_{\text{het}}}{2}$$

$$s z = \int dz n_e(x,y,z) V_e(x,y,z)$$

$$P \sim -(\delta T_i - \delta T_j) C_{ij}$$

$$z \rightarrow -z$$

$$\delta T_{kz} \rightarrow -\delta T_{kz}$$

$$AdS_3 \times S^2$$

$$H_2$$

$$SE \cup \dots = e$$

$$N \rightarrow 0$$

$$\langle \delta x^2 \rangle$$

$$H_2$$

$$H_1$$

$$r = \frac{r_0}{n} \rightarrow \text{loc}$$

Dual Description

The 3D dual is ABJM theory, which has 8 scalars. With $\beta = 0$ boundary conditions the bulk scalar is dual to the $\Delta = 1$ operator

$$\mathcal{O} = \text{Tr}(\phi_1^2 - \phi_2^2)$$

Designer gravity boundary conditions correspond to adding a multi-trace interaction $W(\mathcal{O})$ such that

$$\beta = \frac{\delta W}{\delta \alpha}$$

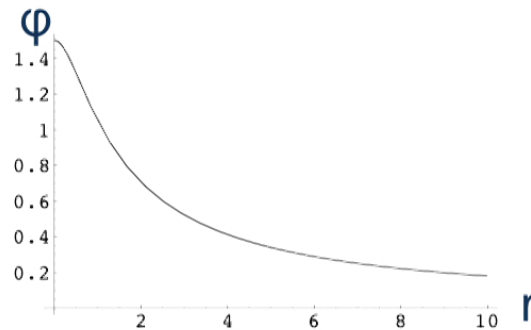
Vacuum structure: bulk

$$\beta_{bc}(\alpha) = -c_1\alpha^2 + c_2\alpha^3$$

Solitons?

$$ds^2 = -h(r)e^{-2\chi(r)}dt^2 + h^{-1}(r)dr^2 + r^2d\Omega$$

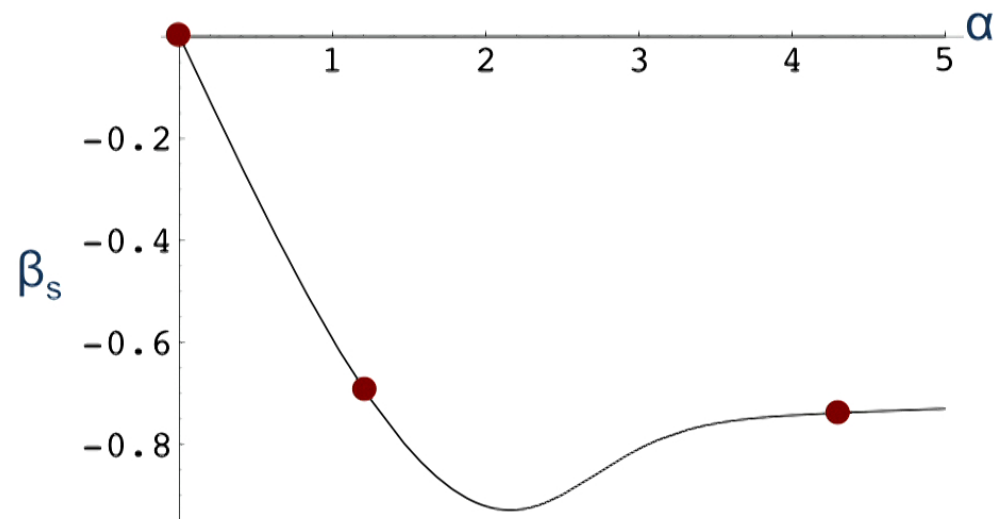
Regularity at origin: $h=1, h'=\varphi'=\chi'=0$



$\varphi(0) \rightarrow (\alpha, \beta)$

Vacuum structure: bulk

$$\beta_{bc}(\alpha) = -c_1\alpha^2 + c_2\alpha^3$$

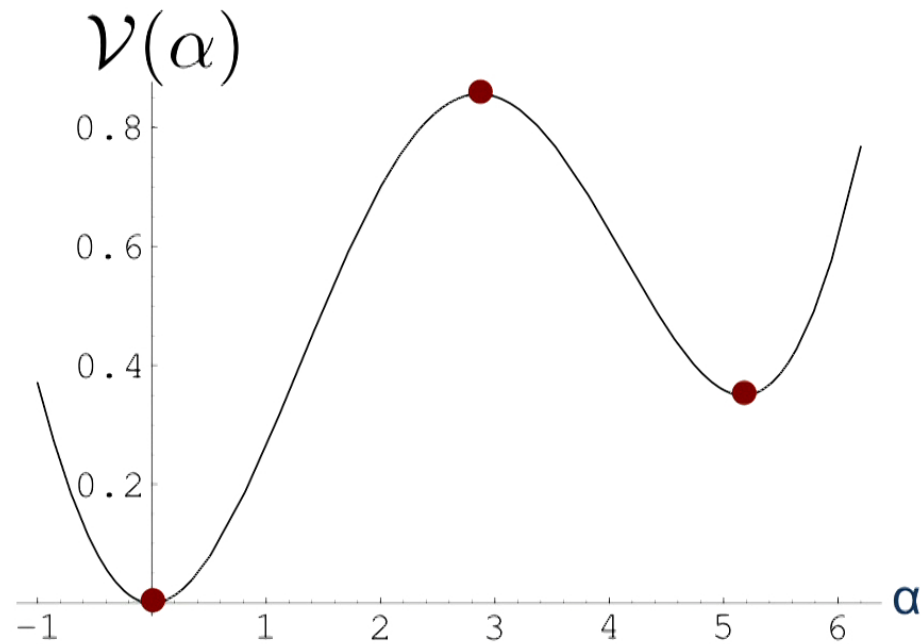


Solitons: $\beta_s(\alpha) = \beta_{bc}(\alpha)$

$$M = \text{Vol}(S^2) [M_0 + \alpha\beta + W]$$

Vacuum structure: dual

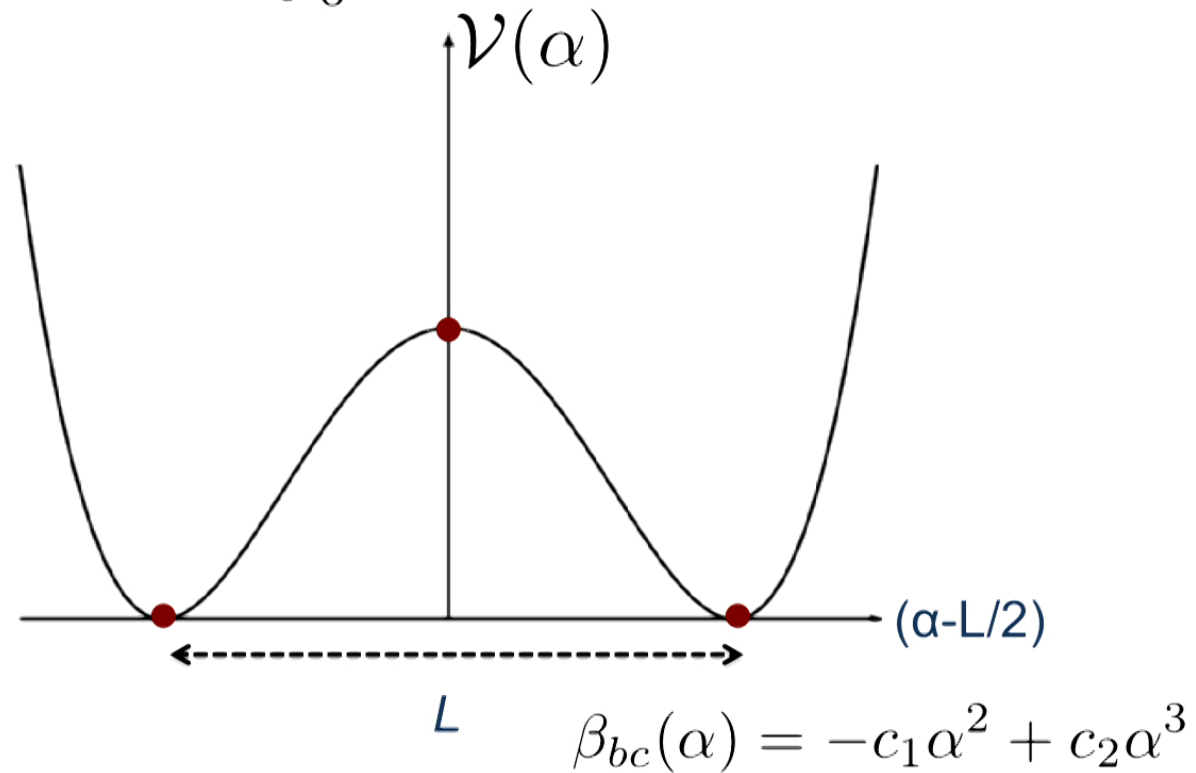
$$\mathcal{V}(\alpha) = - \int_0^\alpha \beta_s(\tilde{\alpha}) d\tilde{\alpha} + W(\alpha)$$



$$\beta_{bc}(\alpha) = -c_1\alpha^2 + c_2\alpha^3$$

Vacuum structure: dual

$$\mathcal{V}(\alpha) = - \int_0^\alpha \beta_s(\tilde{\alpha}) d\tilde{\alpha} + W(\alpha)$$



Black Holes

$$ds^2 = -h(r)e^{-2\chi(r)}dt^2 + h^{-1}(r)dr^2 + r^2d\Omega$$

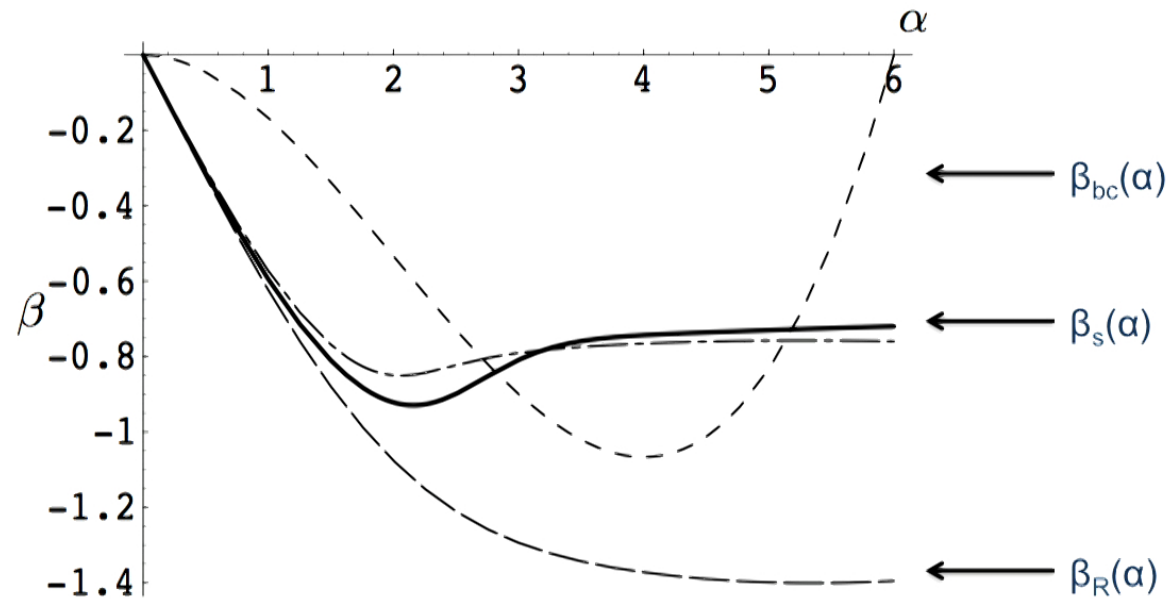
Regularity at horizon:

$$\phi'(R_e) = \frac{V_{,\phi}}{h_{,r}} = \frac{R_e V_{,\phi_e}}{1 - R_e^2 V(\phi_e)}$$

$$\varphi(R_e) \rightarrow (\alpha, \beta)$$

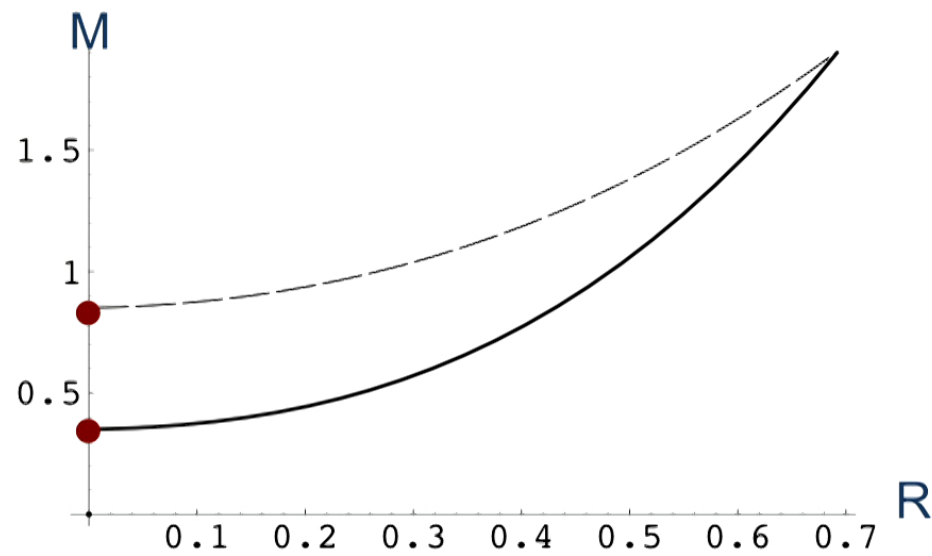
1. Schwarzschild-AdS black holes: $\varphi = 0$ outside horizon
2. Black holes with scalar hair

Black Holes



Black holes with hair: $\beta_R(\alpha) = \beta_{bc}(\alpha)$

Black Holes

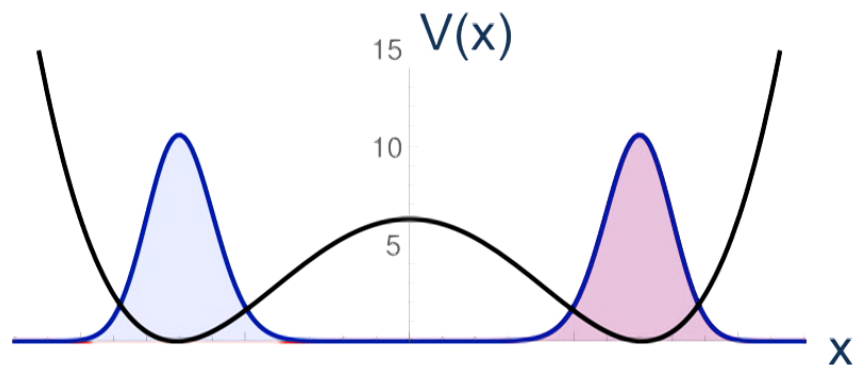


Black holes with hair: $\beta_R(\alpha) = \beta_{bc}(\alpha)$

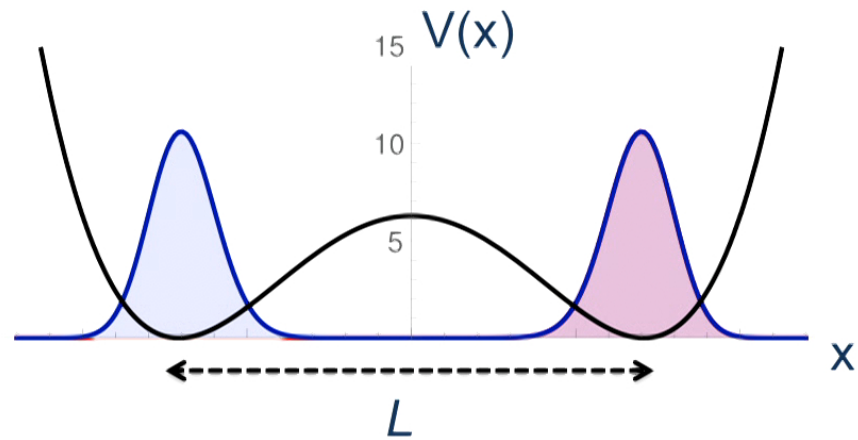
Black Holes: dual description

Black holes with hair: Schwarzschild - AdS inside horizon smoothly glued onto soliton vacuum outside.

Dual toy model description:

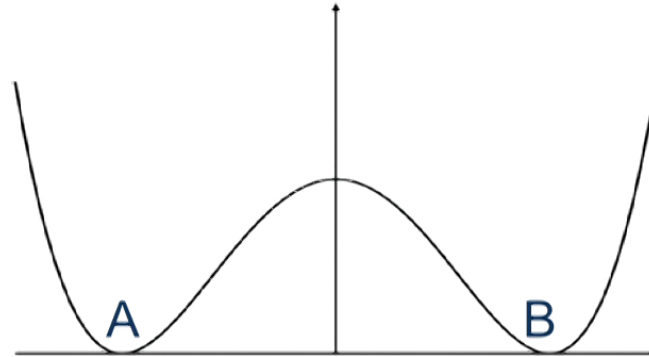


Dual toy model



- Two semi-classical vacua, without introducing a second boundary.
- Non-perturbative interaction $\lambda \sim 1/N^2$
- As the separation $L \rightarrow \infty$: decoupling/cosmic censorship violation

Dual toy model



Perturbatively: Two harmonic oscillators

$$b_R|0\rangle_R = 0, \quad |n\rangle_R = \frac{1}{\sqrt{n!}}(b_R^+)^n|0\rangle_R$$

No tensor product: $\mathcal{H} \cong \mathcal{F}_R \cong \mathcal{F}_L$

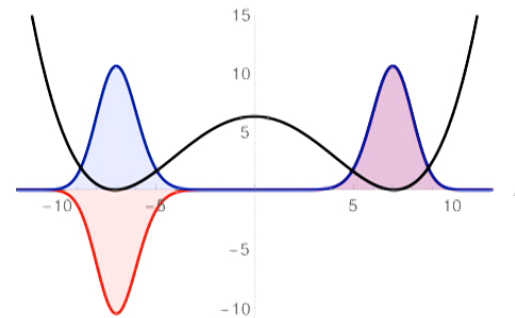
Naively there is a firewall, since vacuum A is a highly excited state from the viewpoint of vacuum B.

Black hole 'microstates'

More natural to consider parity based decomposition: $\mathcal{H} = \mathcal{H}^+ \oplus \mathcal{H}^-$

Energy eigenstates:

$$H\Psi_n^\pm = E_n^\pm \Psi_n^\pm$$



Black hole microstates:

$$\mu = \alpha_+ \Psi_0^+ + \alpha_- \Psi_0^- :$$

Contain both coarse and fine-grained information, e.g.

$$E_0^- - E_0^+ = \frac{2}{\sqrt{\pi\lambda}} e^{-\frac{1}{6\lambda}} [1 + O(\lambda)]$$

State-dependence emerges

Many different black hole microstates,

$$\mu = \alpha_+ \Psi_0^+ + \alpha_- \Psi_0^- = \alpha_R \Psi_0^R + \alpha_L \Psi_0^L$$

are mapped to the same doubled state

$$\Psi_{00}^{\text{dbl}} = \begin{cases} \Psi_0^L \otimes \Psi_0^R & \text{if } \alpha_L, \alpha_R \neq 0, \\ 0 & \text{if } \alpha_L = 0 \text{ or } \alpha_R = 0. \end{cases}$$

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- Fine-grained information about original microstate lost.
- Vice versa, since no correlation in the effective tensor product theory distinguishes fine-grained features of microstates, this information must be supplied to match predictions in the full theory

—————→ emergent state-dependence

- In terms of operators, state-dependence can be viewed as the problem to define operators on an overcomplete basis of states [Jafferis]