

Title: PSI 17/18 - Quantum Field Theory II - Lecture 3

Date: Nov 08, 2017 09:00 AM

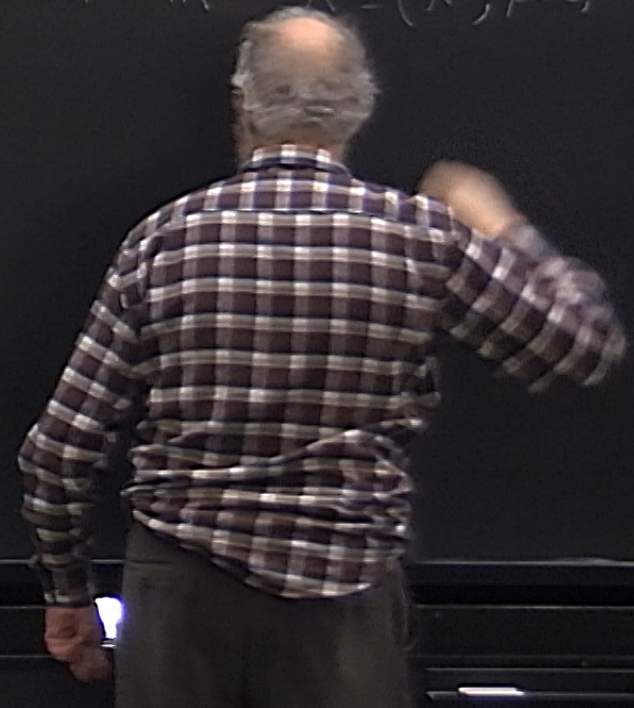
URL: <http://pirsa.org/17110013>

Abstract:

Scalar Free Field: Continued

Space Time: Euclidean \mathbb{R}^d $X = (x^\mu, \mu = 1, \dots, d)$

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Space Time: Euclidean \mathbb{R}^d $X = (x^\mu, \mu=1, \dots, d)$
Real $\phi(x)$

Space Time: Euclidean \mathbb{R}^d $X = (X^\mu, \mu=1, \dots, d)$ mass of the quanta
Real $\phi(x)$ Action $S_E[\phi] = \int d^d X \left[\frac{1}{2} (\partial_\mu \phi)^2 + \frac{m^2}{2} \phi^2 \right]$

Scalar Free Field: Continued

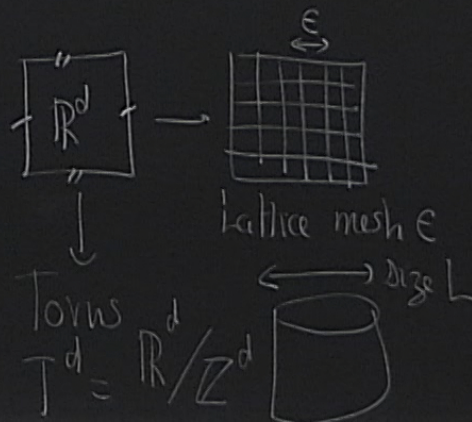
Recipe $\int D[\phi] \exp(-\frac{1}{\hbar} S_E[\phi])$

Euclidean
Funct Integral

• UV regularization (small distances)

• IR regularization (large distances)

e.g. periodic b.c



Space Time: Euclidean \mathbb{R}^d

Real $\phi(x)$ Action

Mathematical recipe $\rightarrow \epsilon$

Physical p

Space Time: Euclidean \mathbb{R}^d $X = (X^\mu, \mu = t, \dots, d)$ mass of the quan.
↓

Real $\phi(x)$ Action $S_E[\phi] = \int d^d X \left[\frac{1}{2} (\partial_\nu \phi)^2 + \frac{m^2}{2} \phi^2 \right]$

Mathematical
recipe $\xrightarrow{\text{continuum limit}}$ $\epsilon \rightarrow 0$

mesh ϵ
→ size L

Physical
 p.b.c in Euclidean Time \leftrightarrow Finite temperature
 p.b.c in Spacelike directions \leftrightarrow ϕ Field in a finite Box

Scalar Free Field: Continued

Recipe $Z = \int D[\phi] \exp\left(-\frac{1}{\hbar} S_E[\phi]\right)$

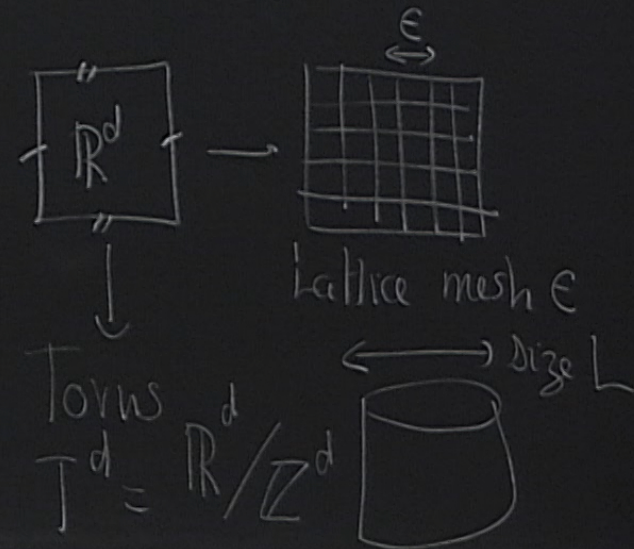
Euclidean

Funel Integral

• UV regularization (small distances)

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e.g. periodic b.c



Space

Real

Mathem
recipe

Physic

mass of the quanta



$$\left[\dots + \frac{m^2}{2} \phi^2 \right]$$

Free Field

$S_E[\phi]$ is quadratic

p.b.c integrals by part

$$S_E = - \int d^d x \frac{1}{2} \phi(x) [-\Delta_x \phi(x)]$$

mass of the quanta

Free Field

$$\left[\left(\frac{\partial \phi}{\partial x^\mu} \right)^2 + \frac{m^2}{2} \phi^2 \right]$$

$S_E[\phi]$ is quadratic
p.b.c integranda by part

$$S_E = - \int d^d x \frac{1}{2} \phi(x) [-\Delta_x + m^2] \phi(x)$$

$$\Delta_x = \sum_{\mu=1}^d \frac{\partial^2}{\partial x^\mu \partial x^\mu} \quad \begin{array}{l} \text{Euclidean Laplace} \\ \text{Beltrami diff operator} \end{array}$$

$$S_E i = \frac{1}{2} \phi \cdot (-\Delta + m^2) \cdot \phi$$

mass of the quanta

Free Field

$$\left[\left(\frac{\partial \phi}{\partial x^\mu} \right)^2 + \frac{m^2}{2} \phi^2 \right]$$

$S_E[\phi]$ is quadratic

p.b.c integrals by part

$$S_E = - \int d^d x \frac{1}{2} \phi(x) [-\Delta_x + m^2] \phi(x)$$

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$$S_E i = \frac{1}{2} \phi \cdot (-\Delta + m^2) \cdot \phi$$

$$Z = \int d\phi \exp\left(-\frac{1}{2\hbar} \phi (-\Delta + m^2) \phi\right)$$

② Gaussian integration

$$Z = \left(\det \left[\frac{-\Delta + m^2}{2\pi \cdot \hbar} \right] \right)^{-1/2} \leftarrow \phi \text{ is real}$$

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$-\Delta + m^2$ diff. operator

$$\det = \infty$$

② Gaussian integration

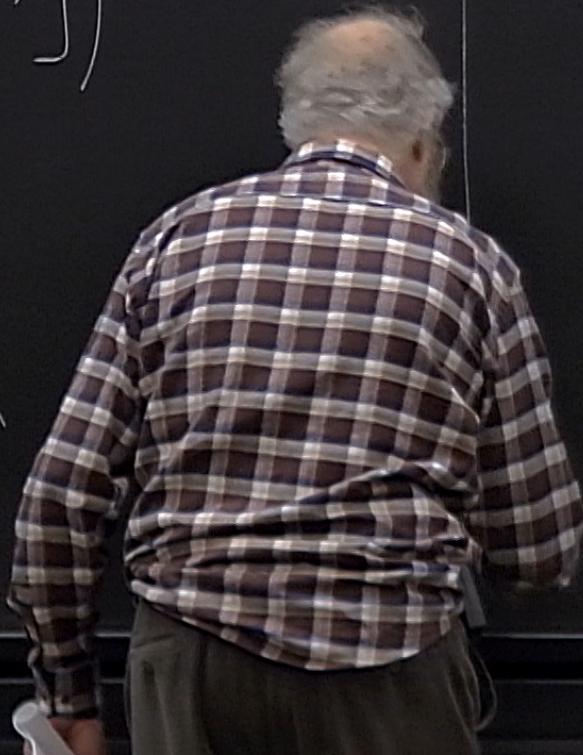
$$Z = \left(\det \left[\frac{-\Delta + m^2}{2\pi \cdot \hbar} \right] \right)^{-1/2} \leftarrow \phi \text{ is real}$$

$-\Delta + m^2$ diff. operator

$\det = \infty$ if space
continuous

$UV + IR \rightarrow$ finite

this ∞ is related to the problem
of the v.e. density in QFT



Correlation functions:

al Path Integral $\rightarrow \langle 0 | T(A_1 A_2) | 0 \rangle$

$q(t)$ in path integral

Functional integral $\rightarrow \langle 0 | T(\underbrace{\Phi(x_1)}_{\text{Field Operator}} \underbrace{\bar{\Phi}(x_2)}_{\text{Field Operator}}) | \underbrace{0}_{\text{quantum vacuum}} \rangle$

$q(t)$ in path integral \longrightarrow $Q(t)$ position operator
in quantum th.
(Heisenberg Picture)

$|0\rangle$
 \uparrow
quantum vacuum

$\phi(x)$ \longrightarrow $\bar{\Phi}(x)$ Field operator

Correlation functions.

real

Path Integral $\rightarrow \langle 0 | T(A_1 A_2) | 0 \rangle$

$q(t)$ in path integral $\rightarrow Q(t)$ position operator in quantum th. (Heisenberg Picture)

Functional integral $\rightarrow \langle 0 | T(\underbrace{\Phi(x_1)}_{\text{Field operator}} \underbrace{\Phi(x_2)}_{\text{Field operator}}) | \underbrace{0}_{\text{quantum vacuum}} \rangle$

$\phi(x) \rightarrow \Phi(x)$ Field operator

Now compute in Euclidean

$$\langle \phi(x_1) \phi(x_2) \rangle = \frac{1}{Z} \times \int D[\phi] \exp\left(-\frac{1}{\hbar} S_E[\phi]\right) \phi(x_1) \phi(x_2)$$

x_1, x_2 2 points in Euclidean space

exp. value of product of random variables

ϕ in Spacetime \leftrightarrow Field in a finite Box

Same for the Free Field

$$\langle \phi(x_1) \phi(x_2) \rangle = G(x_1, x_2) = \frac{1}{\hbar} \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

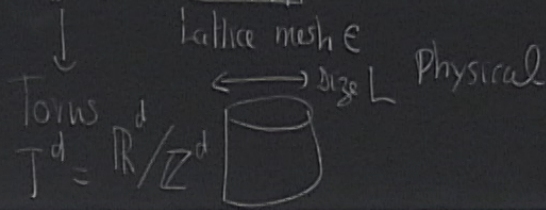
$\phi \in \mathbb{C}$ in Spacetime \leftrightarrow \mathbb{Q} Field in a finite Box

Same for the Free Field

$$\langle \phi(x_1) \phi(x_2) \rangle = \hbar \cdot G_0(x_1, x_2) = \hbar \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

↙ Free Field Case

IR regularization (large distance)
e.g. periodic b.c



ϕ, b, c in Euclidean Time \leftrightarrow Finite temperature
 ϕ, b, c in Space directions \leftrightarrow Q Field in a finite Box

③ Use result of Gaussian integration theory

ϕ is a gaussian random variable

$$\phi(x) \rightarrow \phi_a^i \quad a=1, \dots, d$$

Finite dimensional
Gaussian integrals

$$-\Delta + m^2 \rightarrow \text{Kernel symm. } K = (K_{ab})_{a,b=1, \dots, d}$$

$d \times d \text{ matrix}$

$$\int \prod d\phi^a \exp\left(-\frac{1}{2} \phi^a K_{ab} \phi^b\right) = \left[\det\left(\frac{K}{2\pi}\right)\right]^{-1/2}$$

$$\langle \phi^a \phi^b \rangle = (K^{-1})_{ab} \quad \text{Variance matrix}$$

general result

$$\langle \phi^a \rangle = 0$$

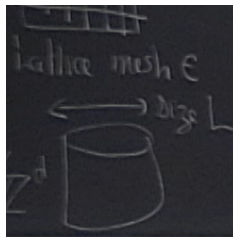
$$K_{ab} (K^{-1})_{bc} = \delta_{ac}$$

Same for the Free Field

$$\langle \phi(x_1) \phi(x_2) \rangle = \hbar G_0(x_1, x_2) = \hbar \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

Free Field case

$$(-\Delta_{x_1} + m^2) \cdot G_0(x_1, x_2) = \delta(x_1 - x_2)$$



Physical

pb.c in Euclidean Time \leftrightarrow Finite Temperature
 pb.c in Spacelike directions \leftrightarrow QField in a finite Box

$$S_E = \frac{i}{2} \phi \cdot (-\Delta + m^2) \cdot \phi$$

$$Z = \int d\phi \exp\left(-\frac{i}{2} \phi \cdot (-\Delta + m^2) \cdot \phi\right)$$

ory

Same for the Free Field

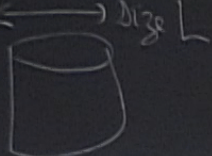
$$\langle \phi(x_1) \phi(x_2) \rangle = \frac{1}{i} G_0(x_1, x_2) = \frac{1}{i} \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

\leftarrow Free Field Case

$$(-\Delta_{x_1} + m^2) \cdot G_0(x_1, x_2) = \delta(x_1 - x_2)$$

From End: Functional Calculus
 Kernels, etc...

δ_{ac}

\mathbb{Z}^d  Physical problem in Euclidean time \leftrightarrow Time temperature
 p.b.c in Spacetime \leftrightarrow QField in a finite Box

$Z =$

theory

Same for the Free Field

$$\langle \phi(x_1) \phi(x_2) \rangle = \hbar \cdot G_0(x_1, x_2) = \hbar \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

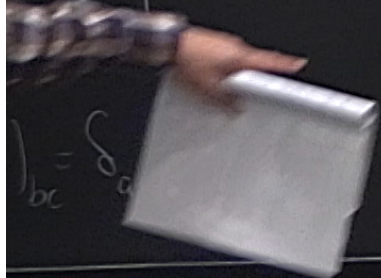
Free Field Case

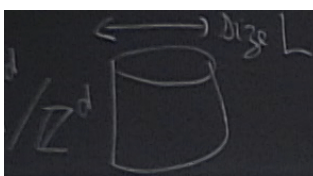
From End. Functional Calculus
Kernels, etc...

$$(-\Delta_{x_1} + m^2) \cdot G_0(x_1, x_2) = \delta(x_1 - x_2)$$

Klein Gordon equation
with source term

Solve this: $G_0(x_1, x_2) = G_0(x_1 - x_2)$ Translation Invariance





Physical picture in Euclidean time \leftrightarrow Time temperature
 picture in Spacetime \leftrightarrow QField in a finite Box

$Z =$

theory

Same for the Free Field

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Fourier Analysis $\hat{G}_0(k) = \int d^D x e^{-i k \cdot x} G_0(x)$ F.T $K^2 = \sum_{\mu=1}^D K_\mu K_\mu$

$X = (x^2, x^d)$
 $k = (k_1, k_d)$

the eqn becomes $(K^2 + m^2) \cdot \hat{G}_0(k) = 1$

Physical picture in Euclidean time \leftrightarrow Time temperature
 p.b.c in Spacetime \leftrightarrow QField in a finite Box

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$$\langle \phi(x_1) \phi(x_2) \rangle = \hbar \cdot G_0(x_1, x_2) = \hbar \left(\frac{1}{-\Delta + m^2} \right)_{x_1, x_2}$$

Free Field Case

From End. Functional Calculus
 Kernels, etc...

$$\hat{G}_0(k) = \frac{1}{k^2 + m^2}$$

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Fourier Analysis $\hat{G}_0(k) = \int dx e^{-i k \cdot x} G_0(x)$ F.T

$$k^2 = \sum_{\mu=1}^d k_\mu k_\mu$$

Euclidean

$x = (x^1, \dots, x^d)$
 $k = (k_1, \dots, k_d)$ the eqn becomes $(k^2 + m^2) \cdot \hat{G}_0(k) = 1$

From End: Functional Calculus
Kernels, etc...

$$\hat{G}_0(k) = \frac{1}{k^2 + m^2}$$

$$G_0(x) = \int_{\mathbb{R}^d} \frac{d^d k}{(2\pi)^d} e^{i k \cdot x} \frac{1}{k^2 + m^2}$$

Euclidean
Propagator

$(k^2 + m^2)^{-1} x_1, x_2$

denominator
source term

$$k^2 = \sum_{p=1}^d k_p k_p$$

Euclidean

From End. Functional Calculus
Kernels, etc.

$$\hat{G}_0(k) = \frac{1}{k^2 + m^2}$$

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Euclidean
Propagator

Explicit

Frona End: Functional Calculus
Kernels, etc.

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Euclidean
Propagator

Explicit: $G_0(x) = G_0(|x|)$

rotation invariance

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Kernels, etc.

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$$G_0(x) = \int \frac{d^d k}{(2\pi)^d} e^{i k \cdot x} \frac{1}{k^2 + m^2}$$

Euclidean
Propagator

Explicitly: $G_0(x) = G_0(|x|)$

rotation invariance

radial coordinate $\Delta = \frac{\partial^2}{\partial r^2} + \frac{d-1}{r} \frac{\partial}{\partial r}$

Bessel Function

$$\frac{1}{2\pi} \left(\frac{2\pi}{m} |x| \right)^{\frac{2-d}{2}} K_{\frac{d-2}{2}}(|x| \cdot m)$$

At large distances $|x| \gg m$ $G_0(x) \simeq \exp(-m|x|)$

Euclidean Exponential decay if $m > 0$

- If mass is zero, makes sense only if $d > 2$

If $d \leq 2$, massless free field has IR divergences
singularities at $k=0$ (large distances)

- $G_0(|x|) \simeq |x|^{\frac{2-d}{2}}$ at large distances

At large distances $|x| \gg m$ $G_0(x) \simeq \exp(-m|x|)$

Euclidean (2) Exponential decay if $m > 0$

(2) If mass is zero, makes sense only if $d > 2$

If $d \leq 2$: massless free field has IR divergences
singularities at $k=0$ (large distances)

- $G_0(|x|) \simeq |x|^{-\frac{2-d}{2}}$ at large distances

if $m=0$

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Euclidean (2) Exponential decay if $m > 0$

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If $d \leq 2$, massless free field has IR divergences
singularities at $k=0$ (large distances)

$$- G_0(|x|) \simeq |x|^{2-d} \text{ at large distances}$$
$$\text{if } m=0$$

Dimensional
if $m=0$

$$G_0(x) = \int d^d k$$

$k \simeq \text{dimension}$

Dimensional analysis

if $m=0$

$$G_0(x) = \int d^d k \frac{e^{i k \cdot x}}{|k|^2} \sim [x]^{-d} [x]^2$$

$k \simeq \text{dimension of } \vec{x}$

convergence

$$\sim [x]^{-d} [x]^2$$

At small distances $|x| \ll m$

the mass m is unimportant, so the expression for the massless propagator is exact

$$G_0(x) \simeq |x|^{2-d}$$

$|x| \rightarrow 0$

$$\sim [x]^{-d} [x]^2$$

At small distances $|x| \cdot m \ll 1$

the mass m is unimportant, so the expression for the massless propagator is exact

$$G_0(x) \approx |x|^{2-d}$$

$|x| \rightarrow 0$

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if

if $d = 2$

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$$G_0(x) \approx |x|^{2-d}$$

$|x| \rightarrow 0$

if $d > 2$ this diverges

$$G_0(x) \approx -\frac{1}{2\pi} \log(|x|m) \quad \text{if } d = 2$$

At small distances

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this diverges too

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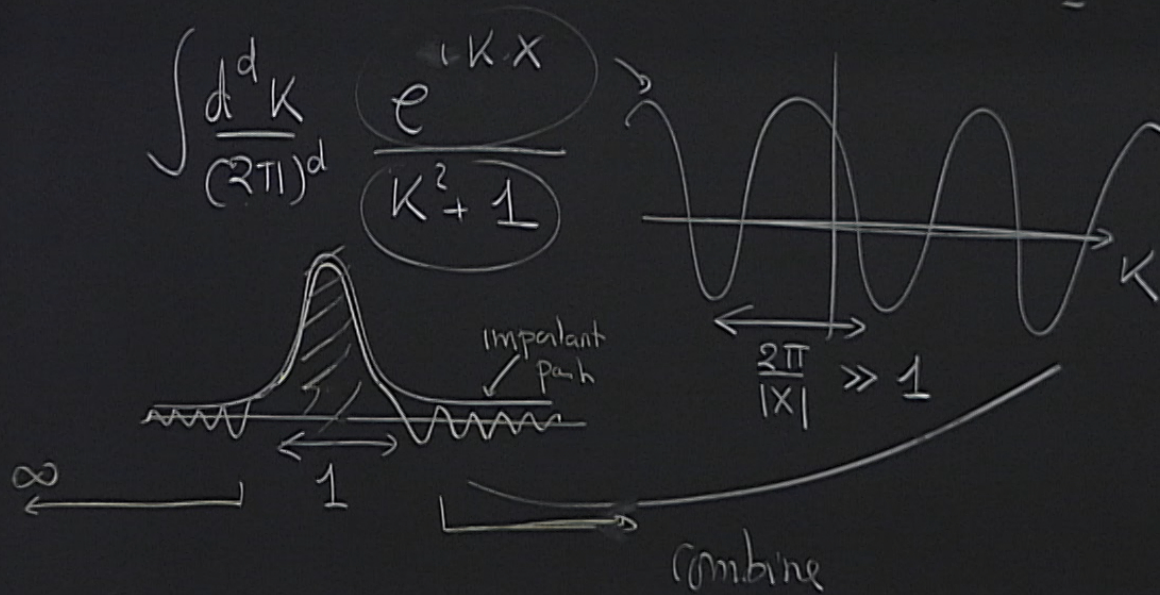
$$G_0(x) \approx -\frac{1}{2\pi} \log(|x|m)$$

if $d = 2$ this diverges too
logarithmic divergence

$$\langle \phi^a \rangle = 0$$

general model $ab^2 + bc + ac$ $k = (k_1 - k_d)$ the eqn becomes

$$|x| m \ll 1 \quad \text{set } m = 1 \quad |x| \ll 1$$

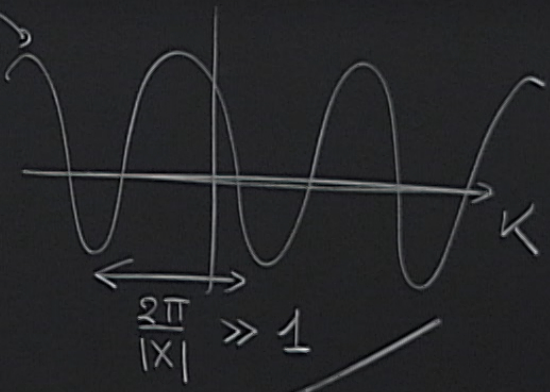
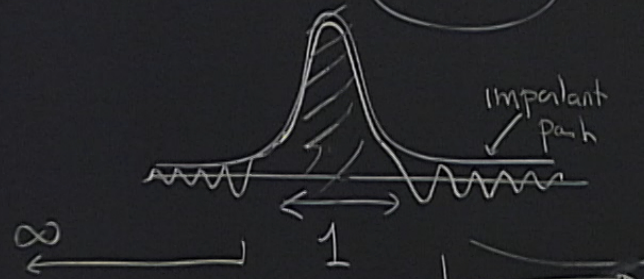


$$\langle \phi^a \rangle = 0$$

general vector $\vec{k} = (k_1, \dots, k_d)$ the eqn becomes

$$|x|m \ll 1 \quad \text{set } m=1 \quad |x| \ll 1$$

$$\int \frac{d^d k}{(2\pi)^d} \frac{e^{i\vec{k}\cdot\vec{x}}}{k^2 + 1}$$



$$\vec{k} = H \cdot \vec{u} \quad \int dH$$

vector $\vec{H} = \dots$

combine

Big part when $|x| \ll 1$

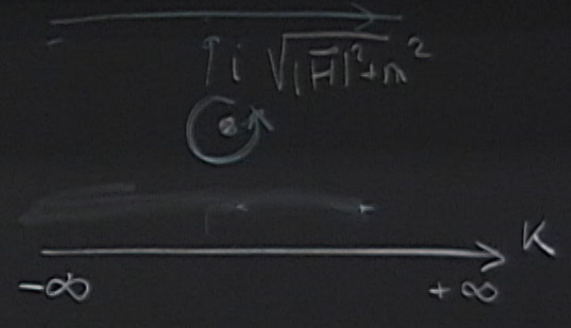


$k = (k_1, \dots, k_d)$ The eqn becomes $(K + M) \cdot G_0(K) = 1$

Bessel function

$$\vec{x} = \begin{pmatrix} x \\ 0 \\ \vdots \end{pmatrix} \quad \vec{k} = (K, \vec{H})$$

$$\int dK d\vec{H} \frac{e^{iK \cdot X}}{K^2 + \vec{H}^2 + m^2} \rightarrow \int dK \frac{e^{iKX}}{K^2 + (\vec{H}^2 + m^2)}$$



in beta

dominant residue of the pole at $e^{-X \cdot \sqrt{H^2 + m^2}}$

$$K = i \sqrt{H^2 + m^2}$$

pole

$k = (k_1, \dots, k_d)$ The eqn becomes $(K + M) \cdot G_0(k) = 1$

Bessel function

$$\vec{x} = \begin{pmatrix} x \\ 0 \\ \vdots \end{pmatrix} \quad \vec{k} = (K, \vec{H})$$

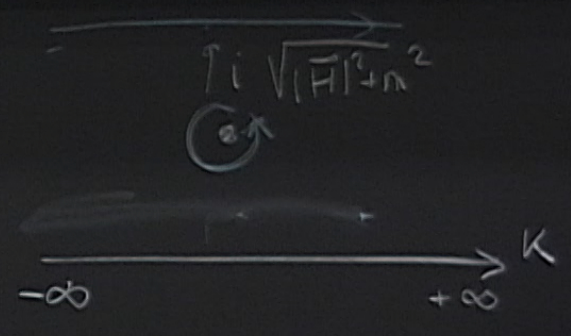
$$\int d\vec{k} d\vec{H} \frac{e^{i\vec{k} \cdot \vec{x}}}{K^2 + \vec{H}^2 + m^2} \rightarrow \int dK \frac{e^{iKx}}{K^2 + (\vec{H}^2 + m^2)}$$

$x > 0$

dominant

residue of the pole at

$$e^{-x \cdot \sqrt{\vec{H}^2 + m^2}}$$



$$K = i \sqrt{\vec{H}^2 + m^2}$$

pole

intra slow

Pessey function

$$i \sqrt{|H|^2 + m^2}$$



$$-i \sqrt{|H|^2 + m^2}$$

pole at

$$K = i \sqrt{|H|^2 + m^2}$$

pole

integrate over \vec{H}
slowest decaying term is for $H=0$
 $\exp(-X \sqrt{m^2})$

$$\int_{-\infty}^{+\infty} \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{H^2 + m^2}}$$

integrate over \vec{H}
 slowest decaying term is for $H=0$
 $\exp(-X \sqrt{m^2})$

Another derivation in radial coordinates

$$\left[-\left(\frac{\partial^2}{\partial r^2} + \frac{d-1}{r} \frac{\partial}{\partial r} \right) + m^2 \right] \psi(r) = 0$$

$r \cdot m \gg 1$ negligible

$$\left(-\frac{\partial^2}{\partial r^2} + m^2 \right) \psi(r) = 0$$

$$\vec{k} = \sqrt{H^2 + m^2}$$



$$= -i \sqrt{H^2 + m^2}$$

$$= i \sqrt{H^2 + m^2}$$

pole

integrate over \vec{H}
 slowest decaying term is for $H=0$
 $\exp(-X \sqrt{m^2})$

Another derivation in radial coordinates
 $\left[-\left(\frac{\partial^2}{\partial r^2} + \frac{d-1}{r} \frac{\partial}{\partial r} \right) + m^2 \right] G(r) = \delta$
 $r \cdot m \gg 1$ negligible

$$\left(-\frac{\partial^2}{\partial r^2} + m^2 \right) G(r) = 0 \quad e^{-mr} \quad e^{+mr}$$

$$\int_{-\infty}^{+\infty} \frac{dH}{\sqrt{|H|^2 + m^2}}$$



$$= -i \sqrt{|H| + m^2}$$

at

$$K = i \sqrt{|H|^2 + m^2}$$

pole

integrate over \vec{H}
 slowest decaying term is for $H=0$
 $\exp(-X \sqrt{m^2})$

Another derivation in radial coordinates
 $\left[-\left(\frac{\partial^2}{\partial r^2} + \frac{d-1}{r} \frac{\partial}{\partial r} \right) + m^2 \right] G(r) = \delta$

$r \cdot m \gg 1$ negligible

grows at ∞

$$\left(-\frac{\partial^2}{\partial r^2} + m^2 \right) G(r) = 0$$

$$\boxed{e^{-mr}}$$

~~$$e^{+mr}$$~~

$$\text{if } m=0$$

Short distance singularity : or UV.

$$\text{if } d \geq 2$$

A feature of QFT \neq NR Q.M

short time & distances
large Energy & momentum

no short time singularities

$$\text{if } m=0$$

short distance singularity : or UV.

$$\text{if } d \geq 2$$

A feature of QFT \neq NR Q.M

short time & distances
large Energy & momentum

no short time singularities

$$G_0(x) \approx -\frac{1}{2\pi} \log(|x|m) \quad \left| \begin{array}{l} d=2 \\ \text{logarithmic decay} \end{array} \right.$$

Distribution of $\phi(x)$ Random Variable (Gaussian)
 Field in space (Euclidean)

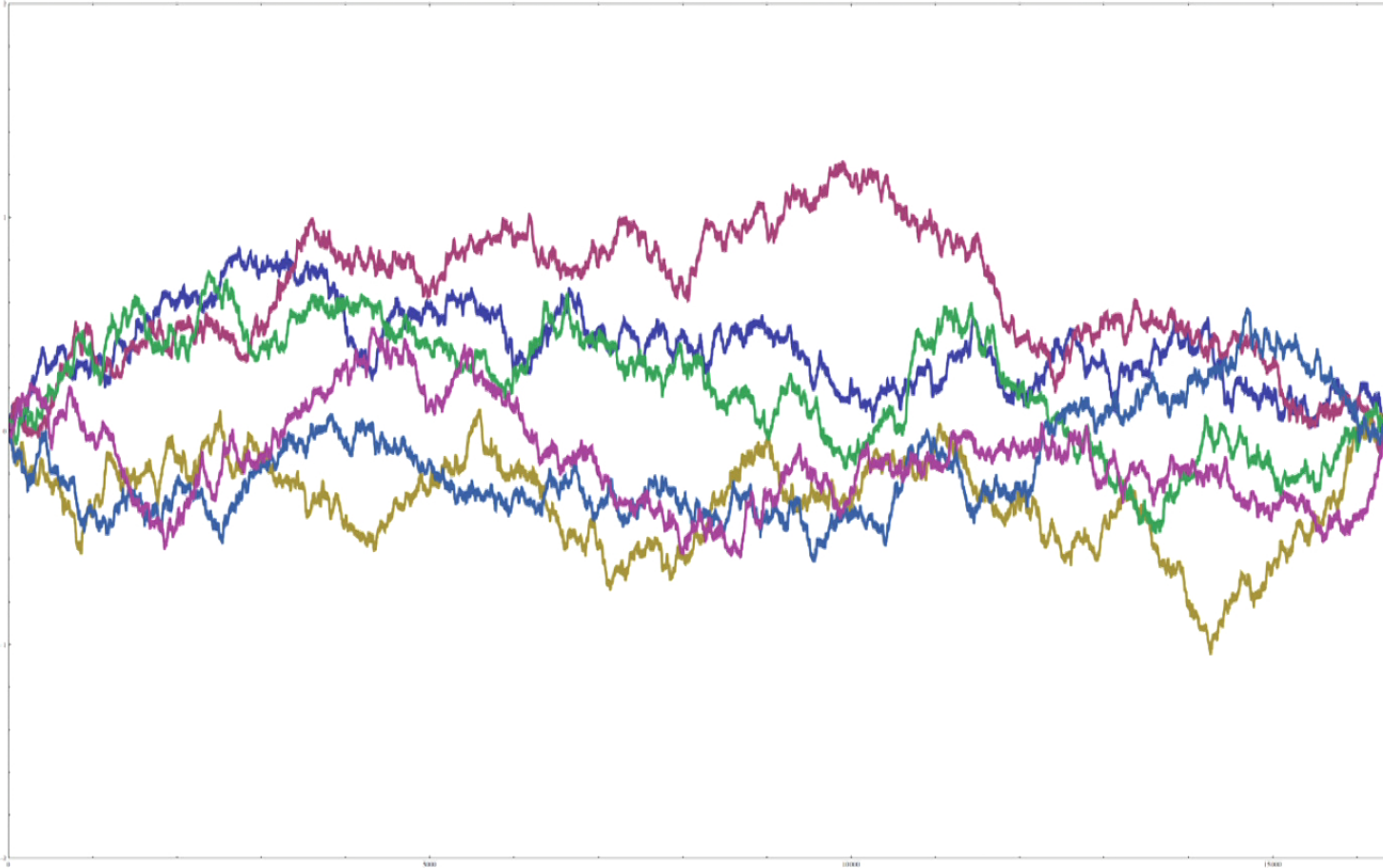
$$P[\phi] \propto \exp(-S_E[\phi])$$



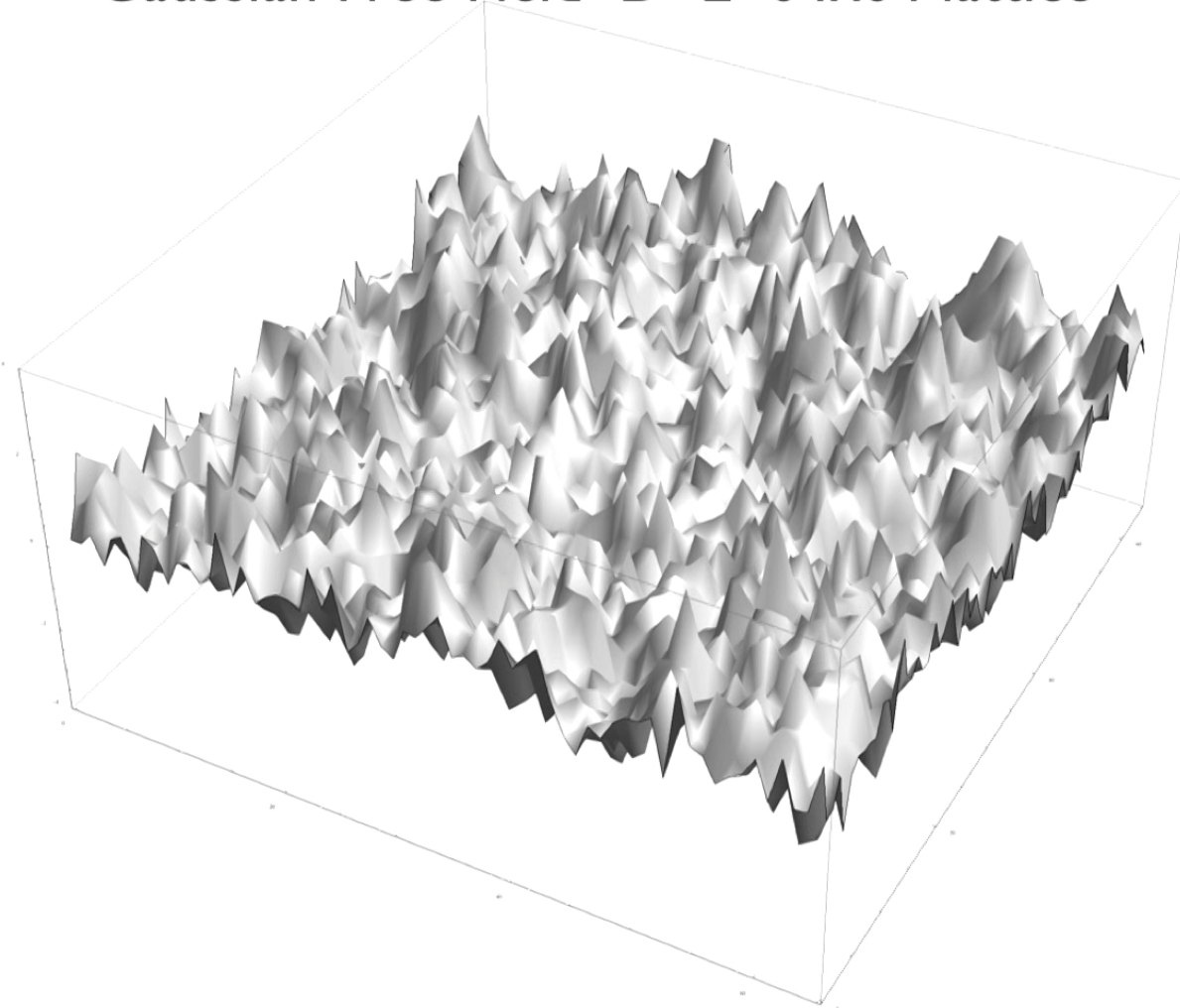
→ finer lattice

Samples of Gaussian free field configurations (Euclidean, periodic b.c.)

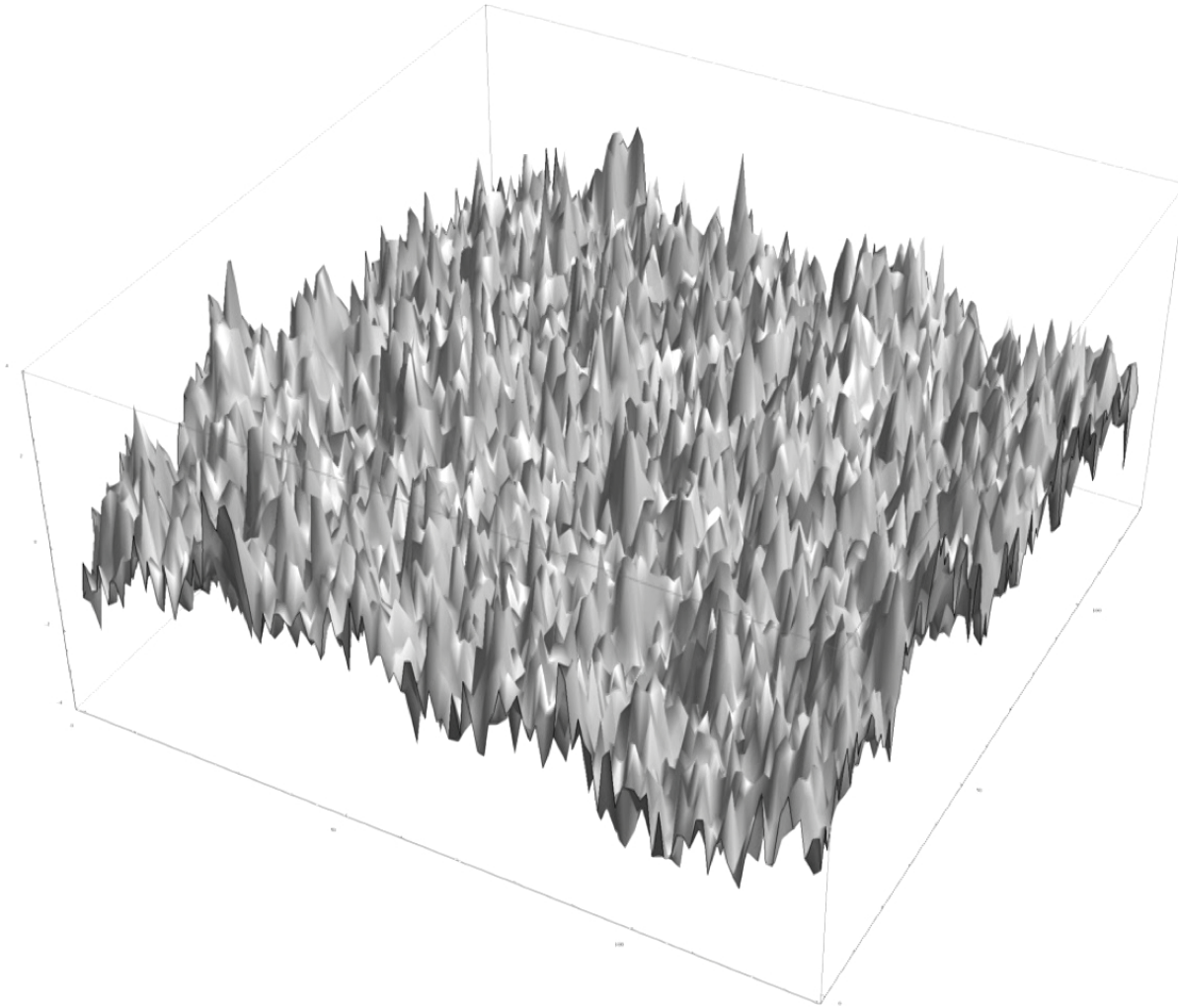
$D=1$: Gaussian Free Fields = Random Walk
(i.e. Brownian or Wiener Process)



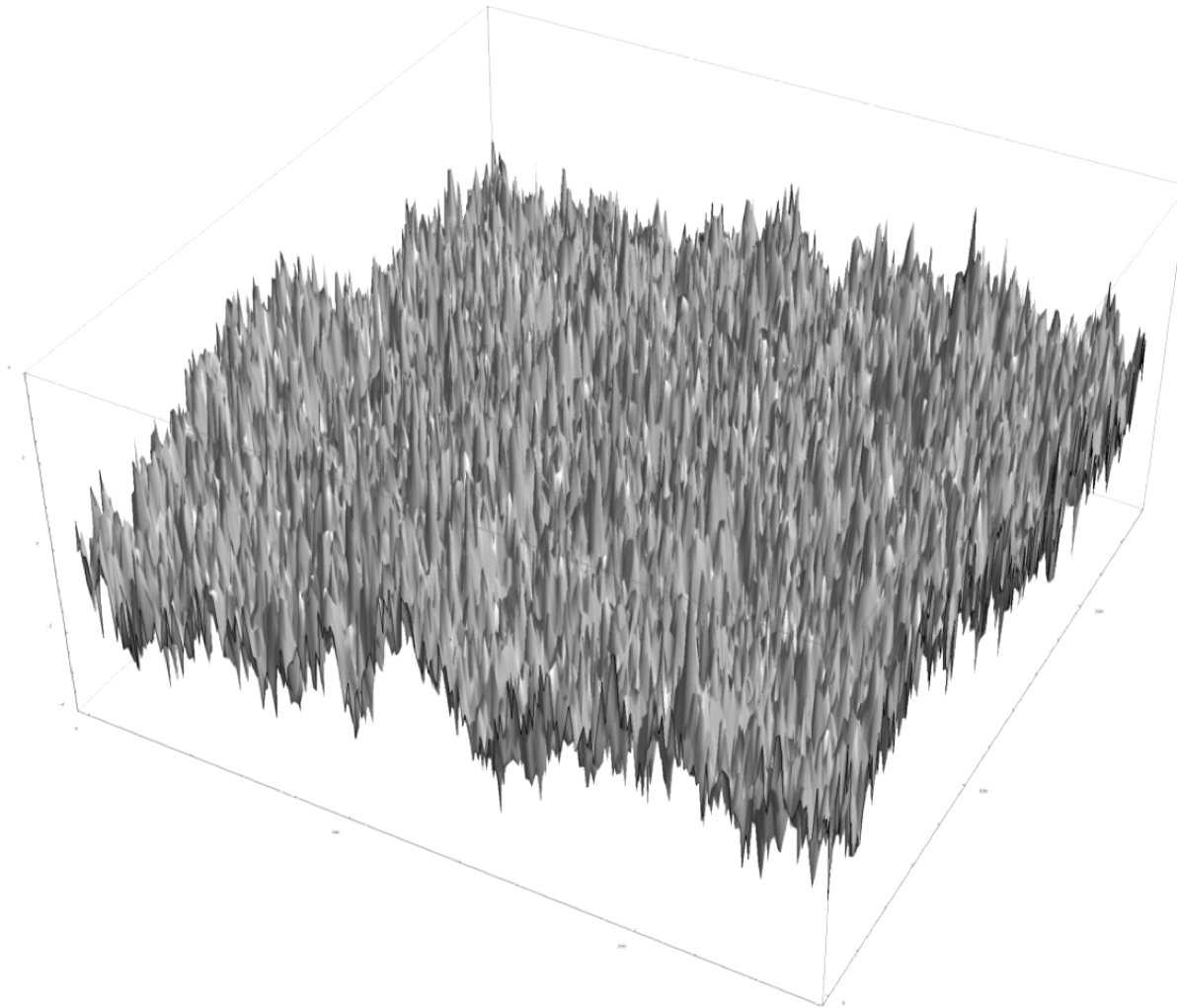
Gaussian Free Field $D=2$ 64x64 lattice



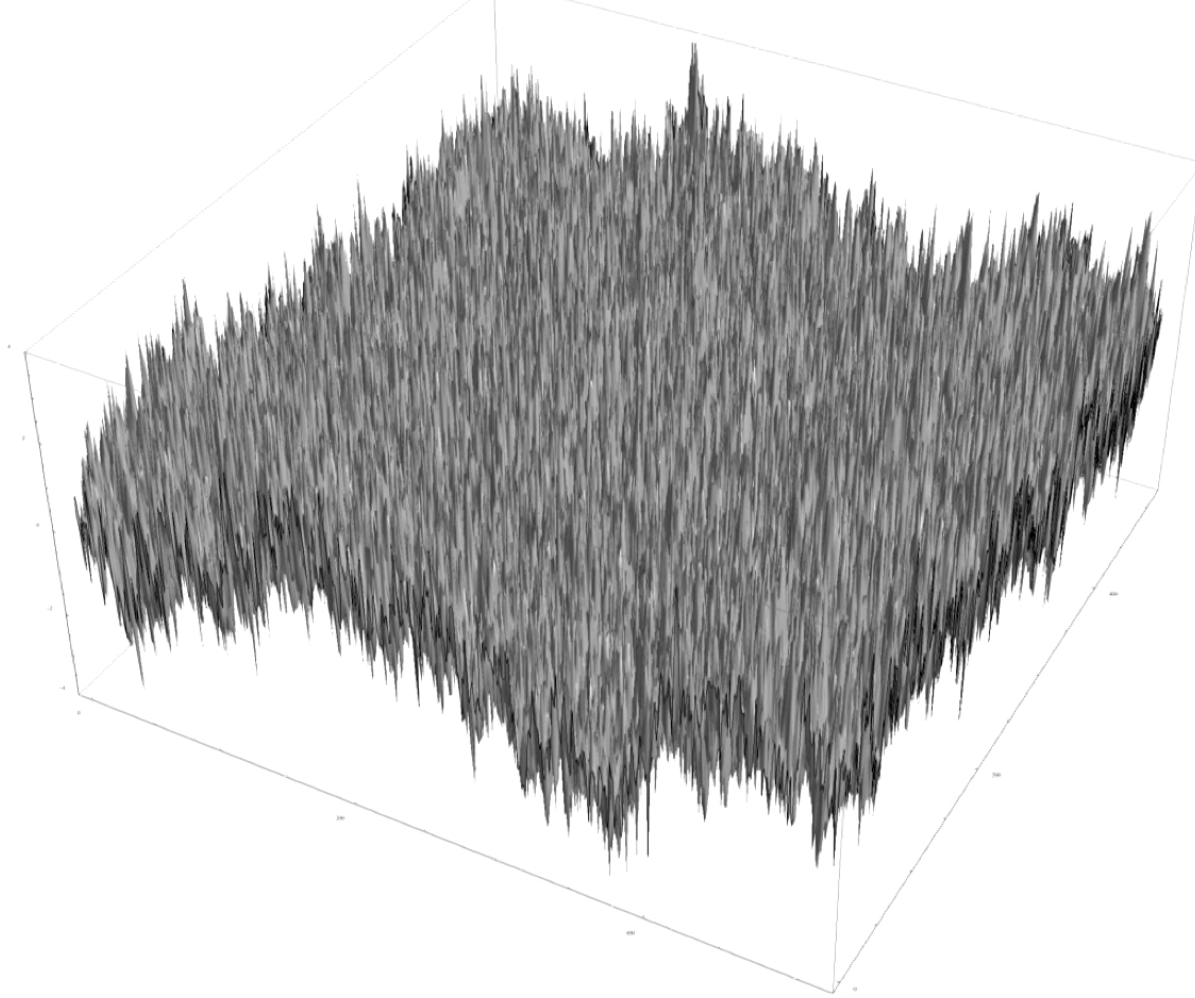
Gaussian Free Field $D=2$ 128×128 lattice



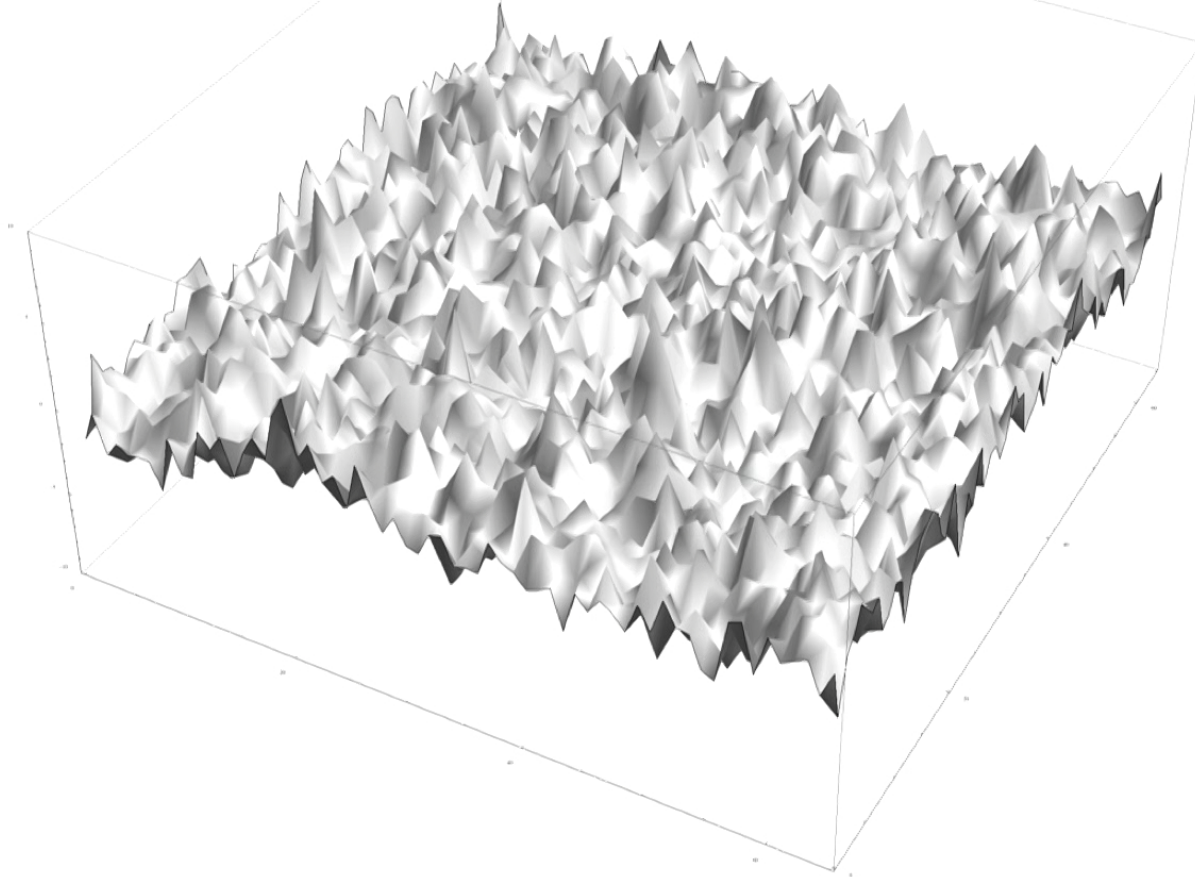
Gaussian Free Field $D=2$ 256x256 lattice



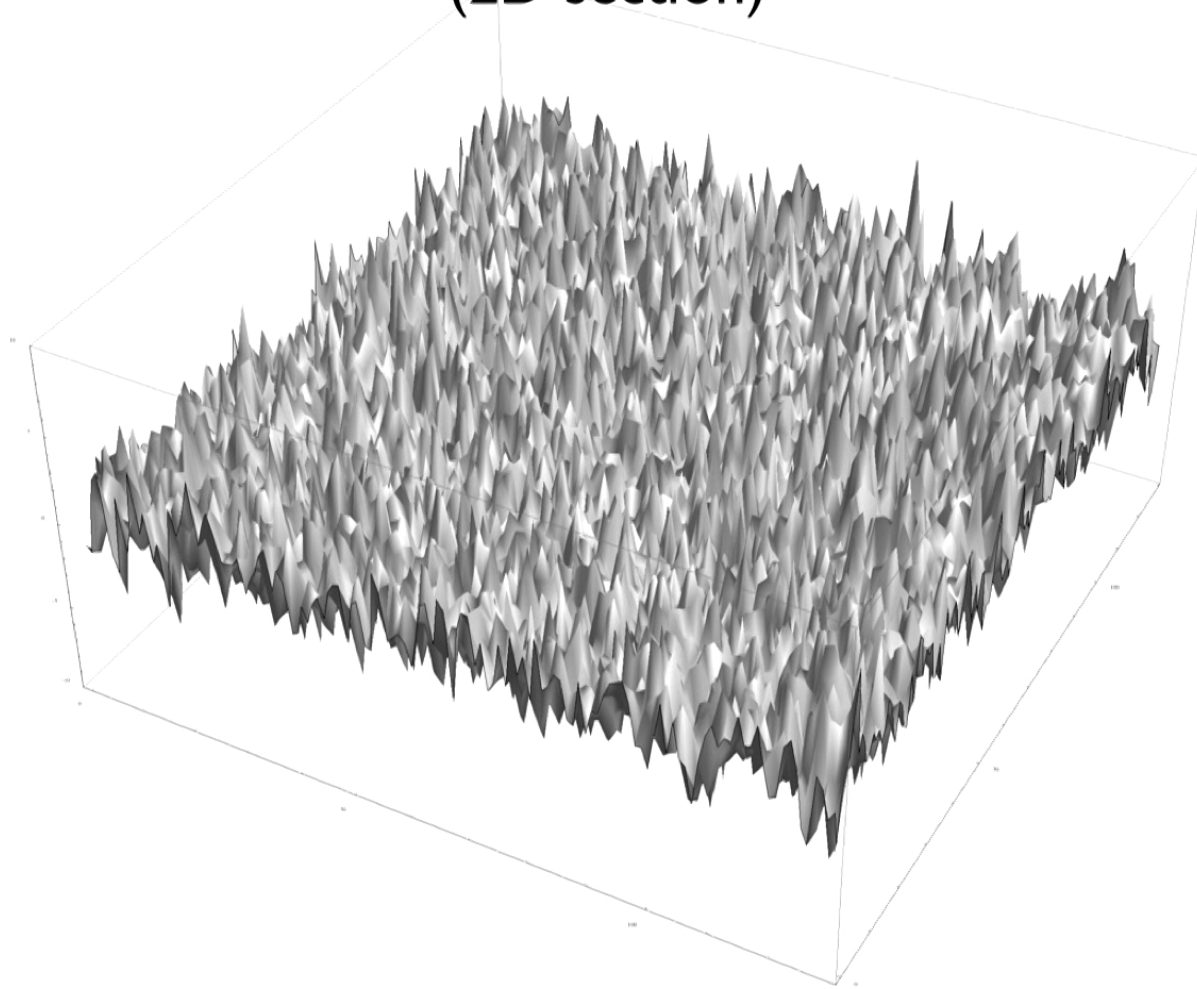
Gaussian Free Field $D=2$ 512x512 lattice



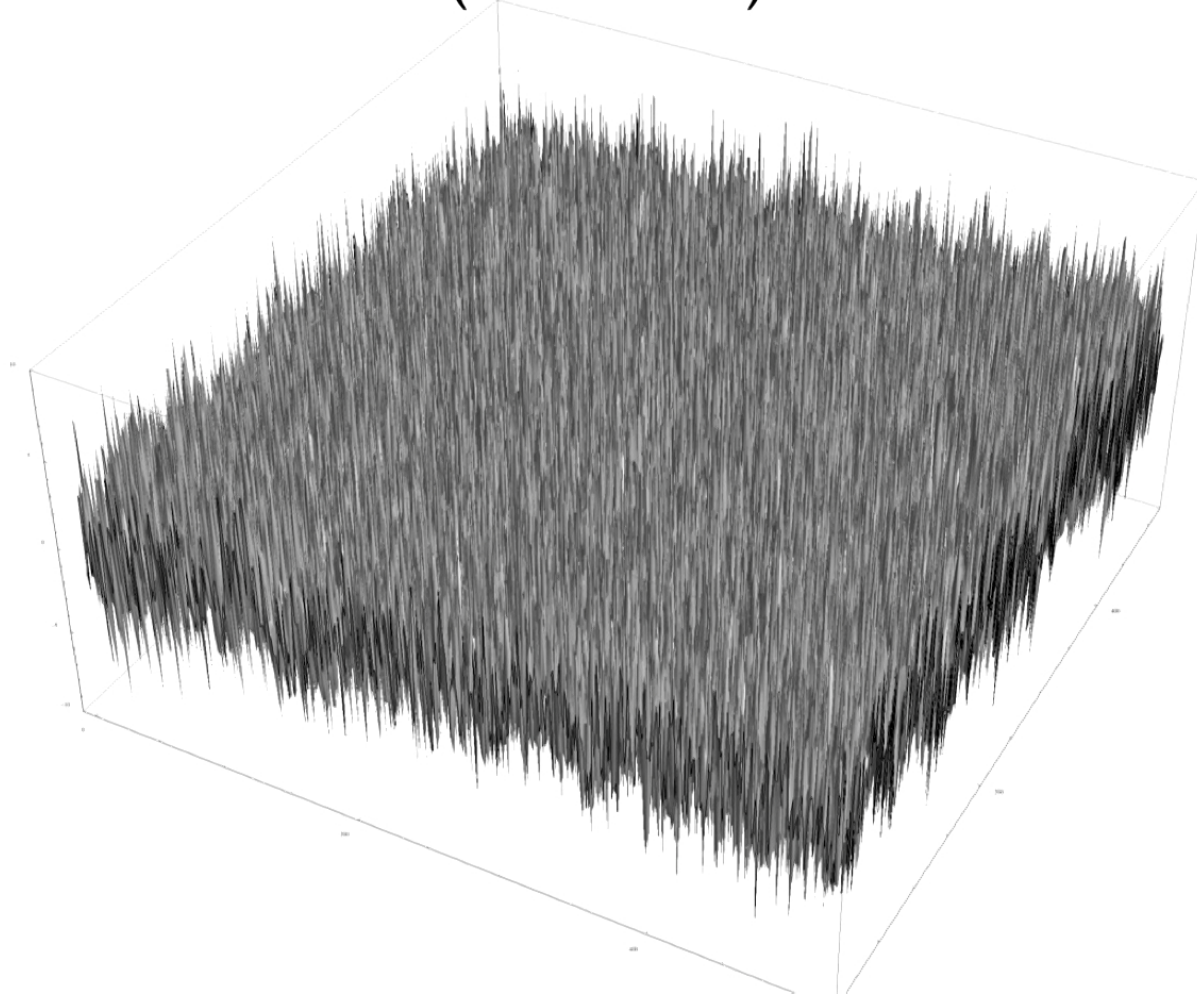
Gaussian Free Field $D=3$ $64 \times 64 \times 64$ lattice
(2D section)



Gaussian Free Field $D=3$ $128 \times 128 \times 128$ lattice
(2D section)



Gaussian Free Field $D=3$ $512 \times 512 \times 512$ lattice
(2D section)



Free field mass m , dimension $D=2$ propagator in Euclidean space

