

Title: PSI 17/18 - Relativity - Lecture 4

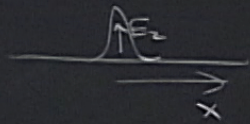
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URL: <http://pirsa.org/17090023>

Abstract:

$$\left. \begin{array}{l} \partial_\mu F^{\mu\nu} \\ \epsilon^{\alpha\beta\gamma\delta} \partial_\beta F_{\gamma\delta} \end{array} \right\} \Rightarrow \underline{\underline{\eta^{\mu\nu} \partial_\mu \partial_\nu F_{\alpha\beta} = 0}}$$

Recap: M.E.s  $\rightarrow F_{\mu\nu}$ ,  $J_\nu$  invariant



$\eta^{\mu\nu}$

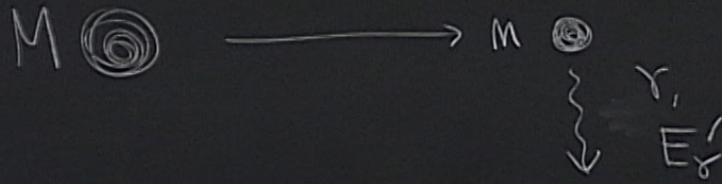
$$P^\mu = \left( \frac{E}{c}, P^x, 0, 0 \right)$$

$$P^x = \frac{E}{c} \quad |P| = \frac{E}{c} \text{ in general}$$

$$x'^\mu = \Lambda^\mu_\nu x^\nu$$
$$\frac{\partial}{\partial x'^\mu} = \tilde{\Lambda}^\alpha_\mu \frac{\partial}{\partial x^\alpha}$$

Einstein (1905, improved 1947)

① work in frame where  $M, m$  at rest



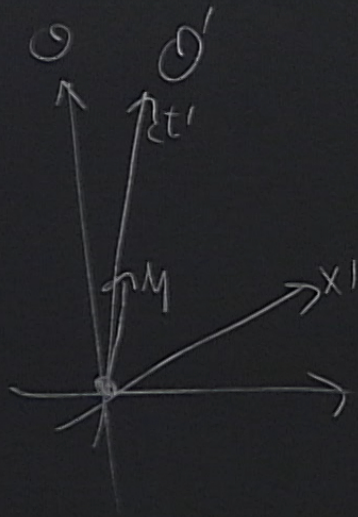
$$E'_M = E'_m + 2E'_y$$

Einstein (1905, improved 1947)

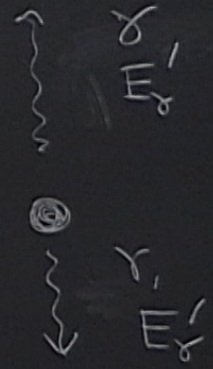
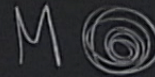
invariant

$$x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu}$$

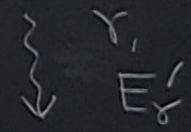
$$\frac{\partial}{\partial x'^{\mu}} = \tilde{\Lambda}^{\alpha}_{\mu} \frac{\partial}{\partial x^{\alpha}}$$



① work in frame where  $M, m$  at rest

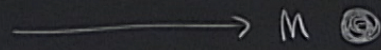
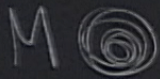


② go to frame where  $M$  is moving with velocity  $v$



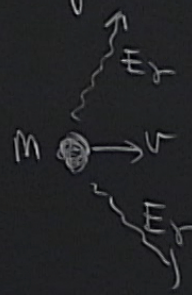
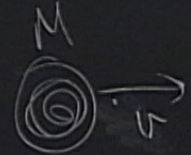
Einstein (1905, improved 1947)

① work in frame where  $M, m$  at rest



② go to frame where  $M$  is moving with velocity  $v$

Momentum  
CMS in ①



$$E'_M = E'_m + 2E'_y$$

$$p^M = \begin{pmatrix} ct' \\ \frac{E'_y}{c}, 0, 0, \frac{E'_y}{c} \end{pmatrix}$$

$$p^M = \left( \gamma \frac{E'_y}{c}, \dots \right) \Rightarrow E_y = E'_y \gamma$$

$$ct = \gamma(ct' + vx')$$

$$E'_M = E'_m + 2E'_y$$

$$p'_M = \begin{pmatrix} ct' & x' & y' & z' \\ \frac{E'_y}{c} & 0 & 0 & \frac{E'_y}{c} \end{pmatrix}$$

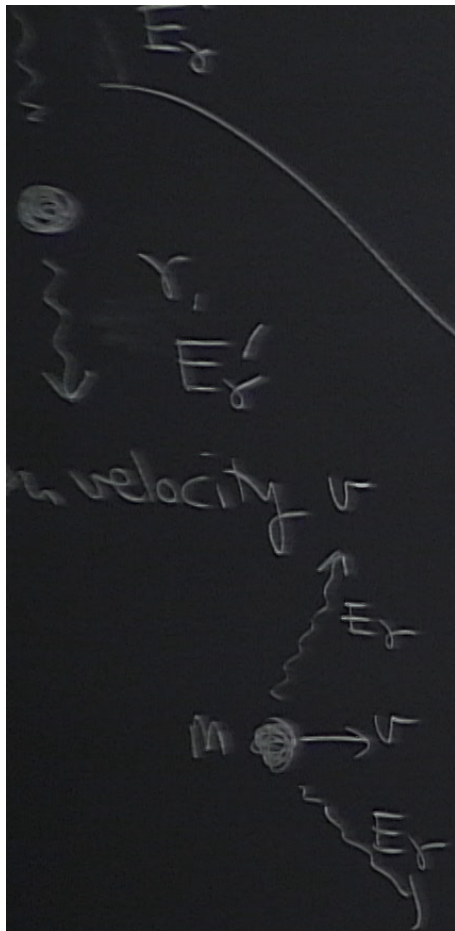
$$ct = \gamma(ct' + vx')$$

$$p_M = \begin{pmatrix} \gamma \frac{E'_y}{c} & \dots \end{pmatrix}$$

$$\Rightarrow E_y = E'_y \gamma$$

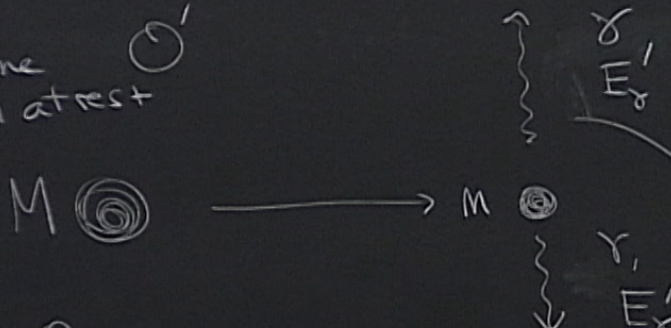
$E_y \approx E'_y$  to order  $v^2$ .

$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}} \approx 1 + \frac{1}{2} \frac{v^2}{c^2}$$



Einstein (1905, improved 1947)

① work in frame where  $M, m$  at rest

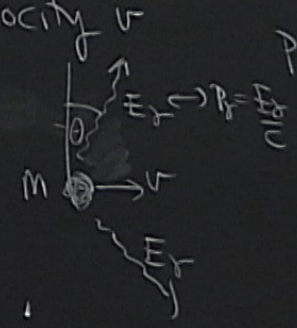
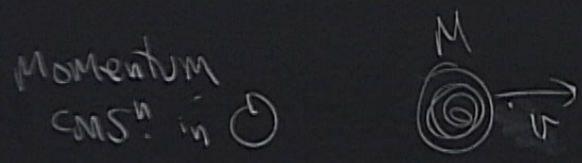


$$E'_M = E'_m + 2E'_y$$

$$ct = \gamma(ct + vx)$$

$$p'^M = \left( \frac{E'_y}{c}, 0, 0, \frac{E'_y}{c} \right)$$

② go to frame where  $M$  is moving with velocity  $v$



$$p^M = \left( \gamma \frac{E'_y}{c}, \dots \right) \Rightarrow E_y = E'_y \gamma$$

$$E_y \approx E'_y \text{ to order } \frac{v}{c}$$

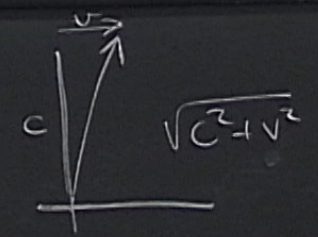
$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}} \approx 1 + \frac{1}{2} \frac{v^2}{c^2}$$

(x-comp!)

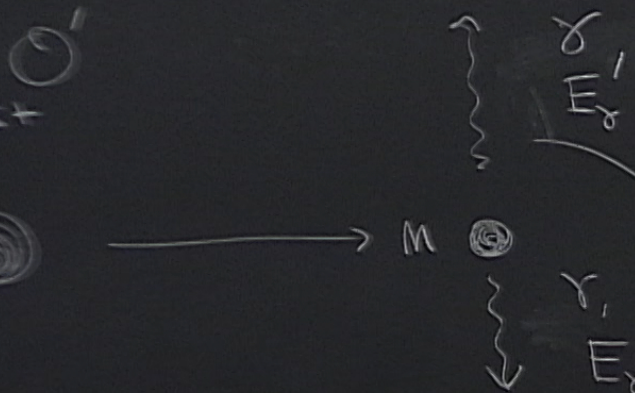
$$Mv = Mv + 2 \frac{E'_y}{c} \sin \theta$$



Einstein (1905, improved 1947)



to frame M, in rest

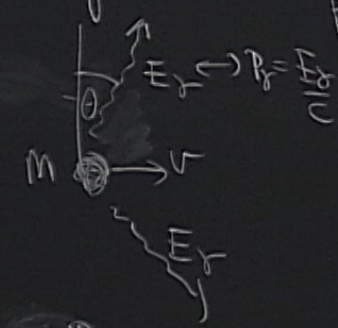


$$E'_M = E'_m + 2E'_y$$

$$p'_M = \left( \frac{E'_y}{c}, 0, 0, \frac{E'_y}{c} \right)$$

$$ct = \gamma(ct' + vx')$$

to frame where M is moving with velocity v



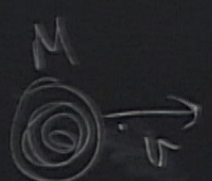
$$p_M = \left( \frac{E_y}{c}, \dots \right) \Rightarrow E_y = E'_y \gamma$$

$E_x \approx E'_x$  to order  $v^2$ .

$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}} \approx 1 + \frac{1}{2} \frac{v^2}{c^2}$$

But  $\sin \theta \approx \frac{v}{c}$  if  $v \ll c$

consider  $\sin^2 \theta$



$$Mv = Mv + 2 \frac{E_y \sin \theta}{c}$$

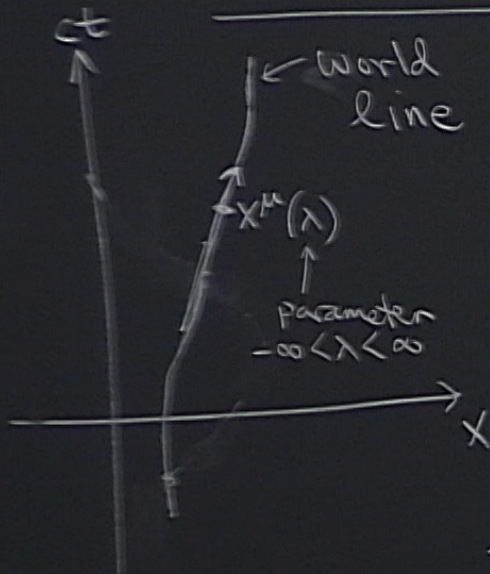
$$\Rightarrow Mv = mv + 2 \frac{E_r}{c} \cdot \frac{v}{c} \quad \text{for } v \ll c$$

$$\Rightarrow (M-m)c^2 = 2E_r \Rightarrow \Delta M c^2 = \text{energy released.}$$

in photons.

in particular, if  $m=0$  then  $E_{\text{released}} = M c^2$ .  
(e.g.  $\pi \rightarrow 2\gamma$ )

# Particles and their action



Lorentz invariant properties

$$\eta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = \eta_{\mu\nu} \frac{dx'^\mu}{d\lambda} \frac{dx'^\nu}{d\lambda} \quad \text{under L.T.}$$

$$< 0 \Rightarrow \text{time like}$$

$$-c^2 dt^2 + dx^2 = 0 \Rightarrow \text{null} \quad |dx| = c dt$$

$$> 0 \Rightarrow \text{spacelike. (tachyon } \cancel{E})$$

Massive particles follow timelike trajectories

can define "proper time" along world line  $d\tau = \sqrt{dt^2 - d\mathbf{x}^2/c^2}$

- Lorentz invariant.

So can define  $u^M = \frac{dx^M}{d\tau}$  - the 4-velocity.

$$p^M = \left( \frac{E}{c}, \mathbf{p} \right)$$

$$m_0 u^M = p^M$$

↑  
rest mass  
(same in every  
frame)

the 4-momentum

$$\Rightarrow E = p^0 c = \sqrt{p^2 c^2 + m_0^2 c^4}$$

$$\Rightarrow \eta_{\mu\nu} p^\mu p^\nu = -m_0^2 c^2$$
$$= -\left(\frac{E}{c}\right)^2 + |\mathbf{p}|^2$$

td line  $d\tau = \sqrt{dt^2 - \frac{dx^2}{c^2}} = dt \sqrt{1 - \left(\frac{dx}{dt}\right)^2 \frac{1}{c^2}} = dt \sqrt{1 - \frac{v^2}{c^2}}$

$$-\frac{d(t)^2}{d\tau^2} + \frac{(dx)^2}{c^2} = -1$$

- the 4-velocity.

$$\eta_{\mu\nu} \frac{U^\mu U^\nu}{c^2} = -1$$

the 4-momentum

$$\eta_{\mu\nu} P^\mu P^\nu = -m_0^2 c^2$$

$$= -\left(\frac{E}{c}\right)^2 + |\mathbf{p}|^2$$

$$\sqrt{E^2 - m_0^2 c^4}$$

$$= -\left(\frac{E}{c}\right)^2 + |\mathbf{p}|^2$$

$$\Rightarrow \text{If } p^2 c^2 \ll m_0^2 c^4 \rightarrow E \approx \underbrace{m_0 c^2}_{\text{same old}} + \frac{1}{2} \frac{p^2}{m_0} + \dots$$

$\searrow \frac{1}{2} m_0 v^2$

$$p = m_0 v$$

$$\underline{p} = m_0 \underline{u} = m_0 \frac{dx}{dt} \cdot \frac{dt}{d\tau} = \gamma m_0 \frac{dx}{dt} = \gamma m_0 (\underline{v}).$$

$$v \ll c \Rightarrow \underline{p} \approx \underline{m_0 v}$$

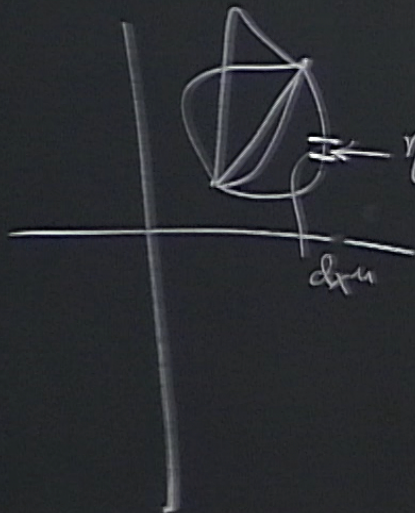
$$\underline{v} = \frac{dx}{dt}$$

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \approx 1 + \frac{1}{2} \frac{v^2}{c^2} + \dots$$

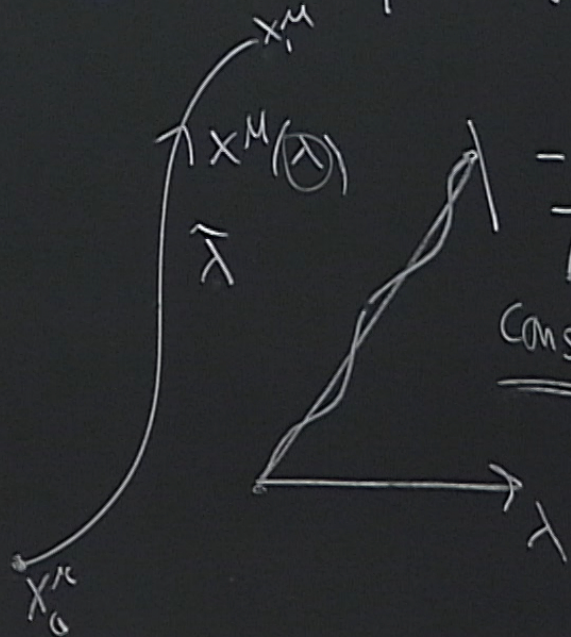
# Action for a relativistic

$$\delta S = 0 \Rightarrow \text{eqns of motion}$$

making



$$\eta_{\mu\nu} dx^\mu dx^\nu$$



$$-m_0 c \int_{\lambda_0}^{\lambda_1} d\lambda \sqrt{-\eta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}}$$

$\equiv$   
 $\text{const}$   
 $d\tau$

$$S = \int d\tau$$



massive particle

making  $S$  invariant under some symmetry is the best way we know to ensure

+

coupling to electromagnetic fields

$$F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$$

$$A_\mu \frac{dx^\mu}{d\lambda}$$

free  
particle

# on for a relativistic particle

of motion, making  $S'$  invariant under some symmetry is the

$$S = \int_{\lambda_0}^{\lambda_1} d\lambda \left[ \underbrace{-m_0 c \sqrt{-\eta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}}}_{d\tau} + q A_\mu(x(\lambda)) \frac{dx^\mu}{d\lambda} \right]$$

$\underbrace{-m_0 c}_{\text{const}}$   
 $\underbrace{d\tau}_{\text{free particle}}$   
 $\underbrace{q A_\mu(x(\lambda)) \frac{dx^\mu}{d\lambda}}_{\text{coupling to electromagnetic fields}}$   
 $F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$        $A_{\mu\lambda} \frac{dx^\mu}{d\lambda}$

$$m \frac{d^2 \underline{x}}{dt^2} = q (\underline{E} + \underline{v} \wedge \underline{B})$$

↙ Lorentz force.

or some symmetry is the best way we know to ensure eq's of motion are invariant under that symm.

$$1) \left. \frac{dx^M}{d\lambda} \right]$$

electromagnetic fields

$$\partial_\nu \frac{dx^M}{d\lambda} \frac{dx^\nu}{d\lambda} = 0 \quad A_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}$$

eqs of motion, making  $S$  invariant under some symmetry is

$$-M_0 c \int_{\lambda_0}^{\lambda_1} d\lambda \left[ -\eta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} \right] + \int d\lambda q [A_\mu(x(\lambda)) \frac{dx^\mu}{d\lambda}]$$

$$S = \int d\tau$$

(with  $\lambda$  axis and  $\tau$  axis indicated)

Coupling to electromagnetic fields:  $F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$

Free particle

$$V = \eta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = c^2 \dot{\tau}^2$$

$$\int d\tau m_0 \eta_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d}{d\tau} (\delta x^\mu)$$

$$\left[ \frac{dx^M dx^N}{\lambda d\lambda} \right]$$

$$+ \int d\lambda q A_\mu(x(\lambda)) \frac{dx^\mu(\lambda)}{d\lambda} \Big|_{x^M + \delta x^M}$$

coupling to electromagnetic fields

$$x \sim \lambda$$

$$F \sim \lambda^2$$

$$A \sim \frac{1}{\lambda}$$

free  
particle

$$F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0 \quad A \frac{dx^\mu}{d\lambda}$$

$$\eta_{\mu\nu} dx^\mu d\delta x^\nu$$

$$\sqrt{-g_{\mu\nu} dx^\mu dx^\nu} = c d\tau$$

$$\delta(A_\mu(x))$$

$$m_0 \eta_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d}{d\tau} (\delta x^\nu)$$

$$+ q \partial_\nu A_\mu \delta x^\nu \frac{dx^\mu}{d\tau} d\tau + q A_\mu \frac{d\delta x^\mu}{d\tau} d\tau$$

$$S = \int d\tau$$

free  
particle

coupling to electromagnetic fields

$$A \sim \frac{1}{r}$$

$$F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$$

$$A_{\mu} \frac{dx^\mu}{d\lambda}$$

$$\int \frac{m_0 c \eta_{\mu\nu} dx^\mu d\delta x^\nu}{\sqrt{-\eta_{\mu\nu} dx^\mu dx^\nu}} = c d\tau$$

$$\int d\tau \underbrace{m_0 \eta_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau}}_{-M_0 \frac{d^2 x^\mu}{d\tau^2} \delta x_\mu}$$

$$- M_0 \frac{d^2 x^\mu}{d\tau^2} \delta x_\mu$$

$$+ q \overbrace{\partial_\nu A_\mu \delta x^\nu}^{\delta(A_\mu(x))} \frac{dx^\mu}{d\tau} d\tau + q \underbrace{A_\mu \frac{d\delta x^\mu}{d\tau}}_{}$$

$$S = \int d\tau$$

free particle

coupling to electromagnetic fields

$$A \sim \frac{1}{r}$$

$$F_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$$

$$A_{\mu} \frac{dx^\mu}{d\lambda}$$

$$\int \frac{m_0 c \eta_{\mu\nu} dx^\mu d\delta x^\nu}{\sqrt{-\eta_{\mu\nu} dx^\mu dx^\nu}} = c d\tau$$

$$\int d\tau \underbrace{m_0 \eta_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau}}_{-m_0 \frac{d^2 x^\mu}{d\tau^2} \delta x_\mu} + q \underbrace{\partial_\nu A_\mu \delta x^\nu}_{\delta(A_\mu(x))} \frac{dx^\mu}{d\tau} d\tau + q \underbrace{A_\mu \frac{d\delta x^\mu}{d\tau} d\tau}_{-q \frac{d}{d\tau} A_\mu \delta x^\mu}$$

$$\begin{aligned}
 & \delta(A_\mu(x)) \\
 & + q \partial_\nu A_\mu \delta x^\nu \frac{dx^\mu}{d\tau} d\tau + q A_\mu \frac{d\delta x^\mu}{d\tau} d\tau \\
 & \quad \downarrow \\
 & + q (\partial_\nu A_\mu - \frac{d}{d\tau} A_\mu \delta x^\mu) \frac{dx^\nu}{d\tau} \delta x^\mu
 \end{aligned}$$



$$\begin{aligned}
 & \delta(A_\mu(x)) \\
 & + \underbrace{q \partial_\nu A_\mu \delta x^\nu \frac{dx^\mu}{d\tau} d\tau}_{\downarrow} + q A_\mu \frac{d\delta x^\mu}{d\tau} d\tau \\
 & + q (\partial_\nu A_\mu - \partial_\mu A_\nu) \delta x^\nu \frac{dx^\mu}{d\tau} \quad - q \frac{d}{d\tau} A_\mu \delta x^\mu \rightarrow \partial_\nu A_\mu \frac{dx^\nu}{d\tau} \delta x^\mu
 \end{aligned}$$

$$\int \frac{m_0 c \eta_{\mu\nu} dx^\mu d\delta x^\mu}{\sqrt{-\eta_{\mu\nu} dx^\mu dx^\nu}} = c d\tau$$

$$\int d\tau \underbrace{m_0 \eta_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d}{d\tau} (\delta x^\mu)}_{-m_0 \frac{d^2 x^\nu}{d\tau^2} \delta x^\nu} + q \overbrace{\partial_\nu A_\mu \delta x^\nu}^{\delta(A_\mu(x))} \frac{dx^\mu}{d\tau} d\tau$$

$$- m_0 \frac{d^2 x^\nu}{d\tau^2} \delta x^\nu$$

$$+ q (\partial_\nu A_\mu - \partial_\mu A_\nu) \delta x^\nu \frac{dx^\mu}{d\tau} d\tau$$

$$E \approx \underbrace{m_0 c^2}_{\text{same old}} + \frac{1}{2} \frac{p^2}{m_0} + \dots$$

$\searrow$   
 $\frac{1}{2} m_0 v^2$

$$p = m_0 \underline{u} = m_0 \frac{dx}{dt} \cdot \frac{dt}{dt}$$

$v \ll c$

$$-m_0 \frac{d^2 x^\nu}{dt^2} + q \left( \partial_\nu A_\mu - \partial_\mu A_\nu \right) \frac{dx^\mu}{dt} = 0 \quad \forall \delta x^\nu$$

$$\Rightarrow m_0 \frac{d^2 x^\nu}{dt^2} = q F_{\nu\mu} \frac{dx^\mu}{dt}$$

Lorentz Force Law