

Title: PSI 17/18 - Relativity - Lecture 3

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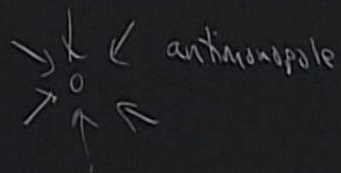
Abstract:

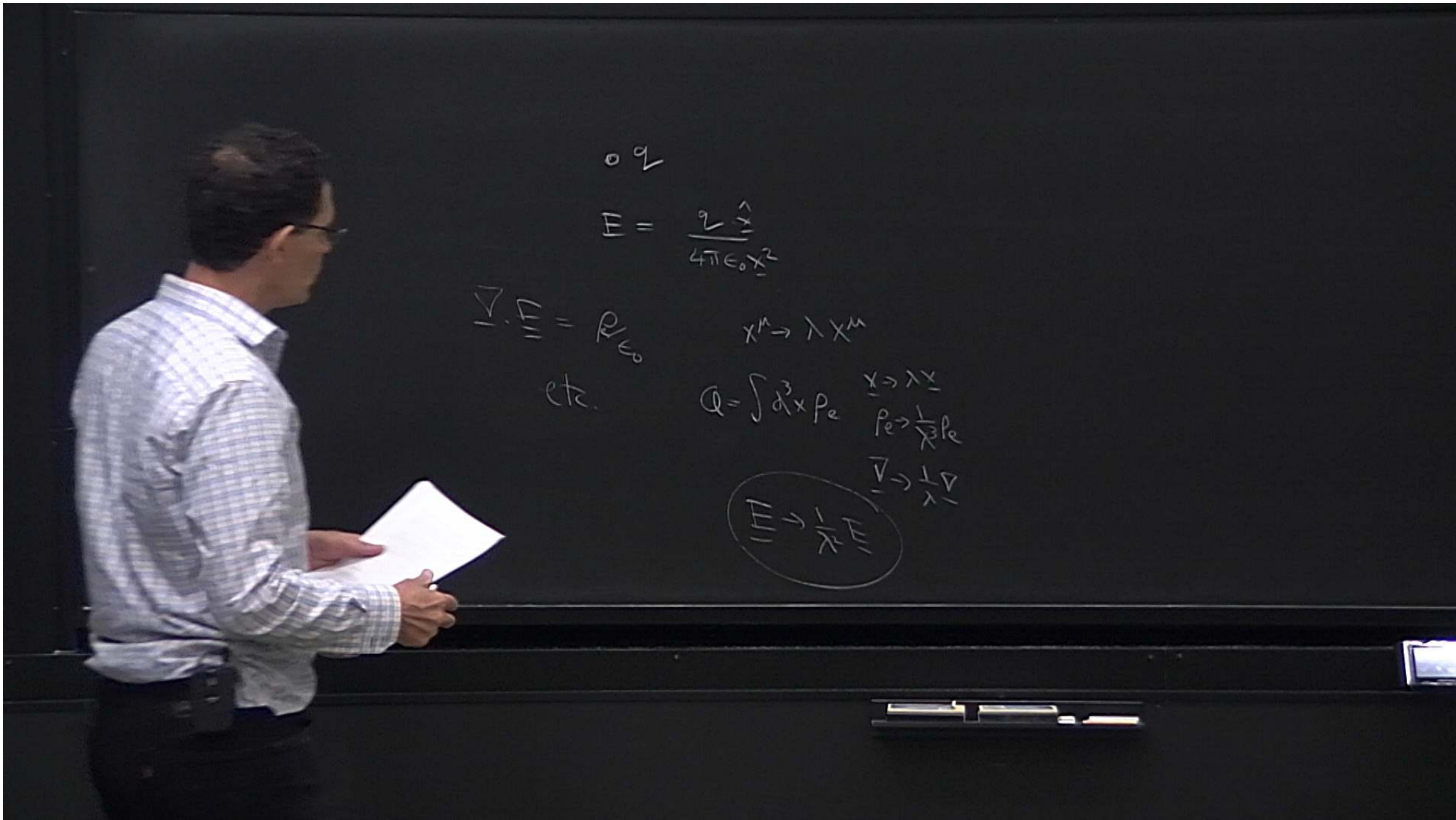
magnetic monopole



$$\underline{\nabla} \cdot \underline{E} = \frac{\partial B}{\partial t}$$

$$B_k = \frac{1}{2} \epsilon_{kij} F_{ij}$$





o q

$$E = \frac{q \hat{r}}{4\pi\epsilon_0 r^2}$$

$$\nabla \cdot \underline{E} = \rho/\epsilon_0$$

etc.

$$x^M \rightarrow \lambda x^M$$

$$Q = \int d^3x \rho_e \quad \begin{array}{l} x \rightarrow \lambda x \\ \rho_e \rightarrow \frac{1}{\lambda^3} \rho_e \end{array}$$

$$\nabla \rightarrow \frac{1}{\lambda} \nabla$$

$$\underline{E} \rightarrow \frac{1}{\lambda^2} \underline{E}$$

$$\epsilon_{\mu\nu\rho\sigma}$$

$$\mathbb{E} = \frac{q}{4\pi\epsilon_0 x^2} \begin{pmatrix} 1 \\ \dots \end{pmatrix}$$

$$\nabla \cdot \mathbb{E} = \rho_{\text{cl}}$$

etc.

$$Q = \int d^3x \rho_e \quad \begin{matrix} x \rightarrow \lambda x \\ \rho_e \rightarrow \frac{1}{\lambda^3} \rho_e \\ \nabla \rightarrow \frac{1}{\lambda} \nabla \end{matrix}$$

$$\mathbb{E} \rightarrow \frac{1}{\lambda} \mathbb{E}$$

$$x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu} + a^{\mu}$$

0 9

$$\underline{E} = \frac{q}{4\pi\epsilon_0 x^2} \begin{pmatrix} \hat{x} \\ \hat{y} \\ \hat{z} \end{pmatrix}$$

$$\underline{\nabla} \cdot \underline{E} = \rho/\epsilon_0$$

etc.

$$\sum_{\alpha} \sum_{\beta} \eta^{\alpha\beta} = \eta^{\mu\nu}$$

$$x^{\mu} \rightarrow \lambda x^{\mu}$$

$$Q = \int d^3x p_e \quad \begin{matrix} x \rightarrow \lambda x \\ p_e \rightarrow \frac{1}{\lambda^3} p_e \end{matrix}$$

$$\underline{E} \rightarrow \frac{1}{\lambda^2} \underline{E}$$

$$\underline{\nabla} \rightarrow \frac{1}{\lambda} \underline{\nabla}$$

$$x'^{\mu} = \underbrace{\sum_{\nu} \Lambda^{\mu}_{\nu} x^{\nu}}_{\text{Lorentz}} + \underbrace{a^{\nu}}_{\text{translations}}$$

- Poincaré symm  
10-dim Lie group.

examples

$$\Lambda_{\beta}^{\alpha} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \\ & & 1 \end{pmatrix}$$

i.e. orthogonal

$$\Lambda^T \eta \Lambda = \eta$$

-rotation

$$\Lambda_{\beta}^{\alpha} = \begin{pmatrix} \cosh\theta & -\sinh\theta \\ -\sinh\theta & \cosh\theta \\ & & 1 \\ & & & 1 \end{pmatrix} \Rightarrow$$

$$ct' = \cosh\theta ct - \sinh\theta x$$

$$x' = \cosh\theta x - \sinh\theta ct$$

$$(ct')^2 - x'^2 = (ct)^2 - x^2$$

$$ct' = ct$$

$$x' = \cos\theta x + \sin\theta y$$

$$y' = \cos\theta y - \sin\theta x$$

$$z' = z$$

$$x'^2 + y'^2 = x^2 + y^2$$

examples

$$\Lambda_{\beta}^{\alpha} = \begin{pmatrix} \cos\theta & \sin\theta & \\ -\sin\theta & \cos\theta & \\ & & 1 \end{pmatrix}$$

i.e. orthogonal

$$\Lambda^T \eta \Lambda = \eta$$

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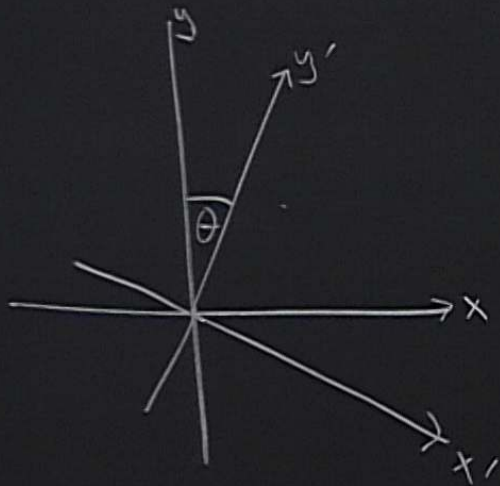
$$\begin{aligned} ct' &= \cosh\theta ct - \sinh\theta x \\ x' &= \cosh\theta x - \sinh\theta ct \end{aligned}$$

$$(ct')^2 - x'^2 = (ct)^2 - x^2 \quad \text{using } \cosh^2\theta - \sinh^2\theta = 1$$

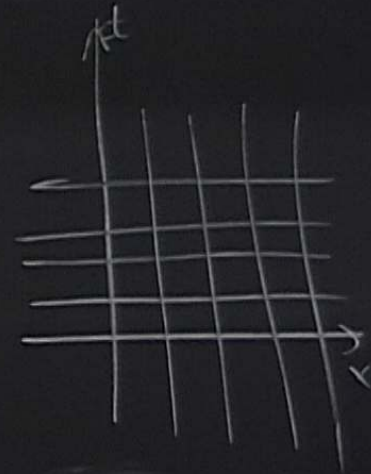
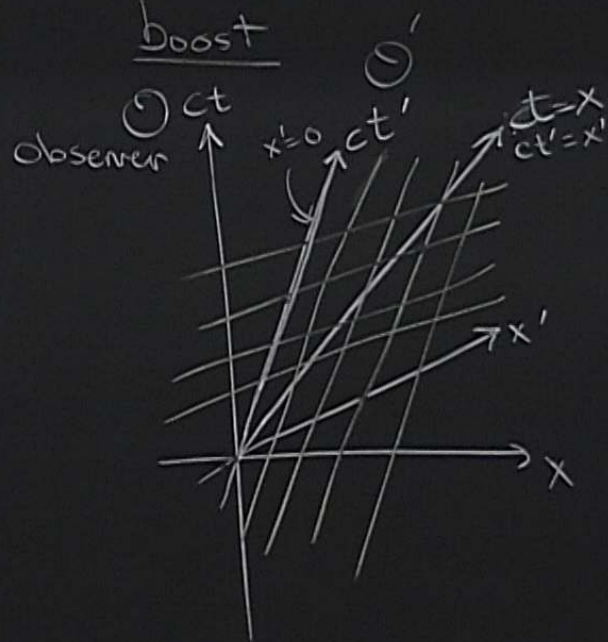
$$\begin{aligned} ct' &= ct \\ x' &= \cos\theta x + \sin\theta y \\ y' &= \cos\theta y - \sin\theta x \\ z' &= z \end{aligned}$$

$$x'^2 + y'^2 = x^2 + y^2$$

rotation



boost



$$(ct)^2 - X^2 = (ct')^2 - X'^2 \quad (\cosh^2\theta - \sinh^2\theta = 1)$$



$$x' = 0 \Rightarrow x = \tanh\theta ct \equiv vt$$

$$\Rightarrow \frac{v}{c} = \tanh\theta < 1$$

$$\cosh\theta = \frac{1}{\sqrt{1 - \tanh^2\theta}} = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \equiv \gamma$$

$\Rightarrow$  Lorentz transformations only allow you to transform to observers moving slower than  $c$ .

$$\sinh\theta = \tanh\theta \cosh\theta = \gamma \frac{v}{c}$$



$$ct = \gamma(ct' + \frac{v}{c}x')$$

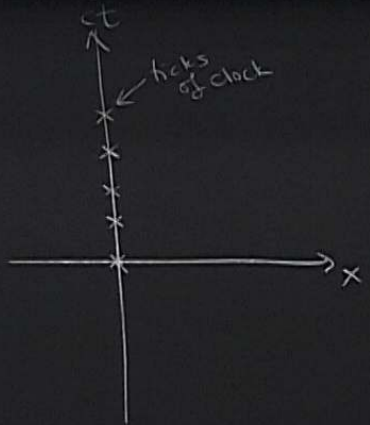
$$x = \gamma(x' + vt')$$

$$\Leftrightarrow \begin{cases} ct' = \gamma(ct - \frac{v}{c}x) \\ x' = \gamma(x - vt) \end{cases}$$

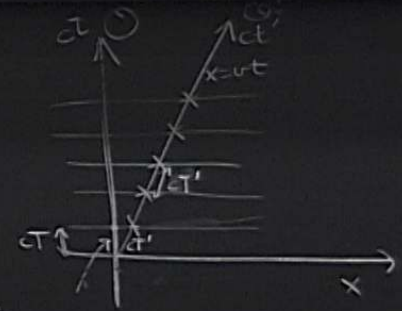
$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{1}{\sqrt{1 - \tanh^2 \theta}}$$

$$\sinh \theta = \frac{v}{c} \cosh \theta$$

# Time Dilation



Imagine a moving clock

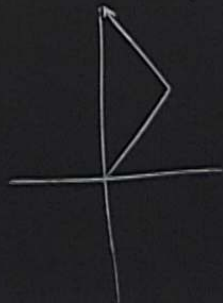


$$x' = \gamma(x - vt) = 0$$

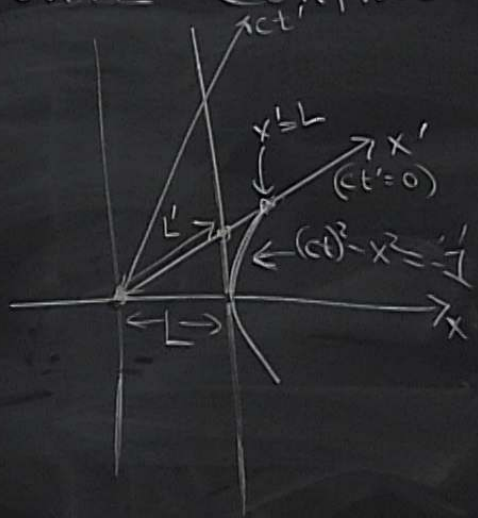
$$ct = \gamma(ct' + \frac{v}{c}x')$$

$$= \gamma ct'$$

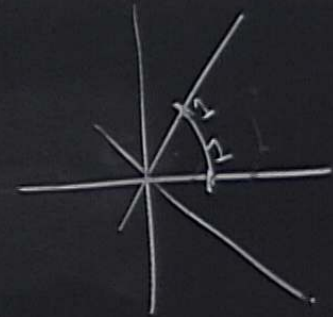
$$\Rightarrow T = \gamma T'$$



# Lorentz Contraction



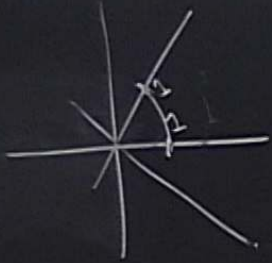
$$(ct')^2 - x'^2 = (ct)^2 - x^2 = -1$$



collection of atoms



consider some Lorentz scalar e.g.  $\underline{P}_{(x)} = F_{\mu\nu} F^{\mu\nu}(x) = 2(-\underline{E}^2 + \underline{B}^2)$



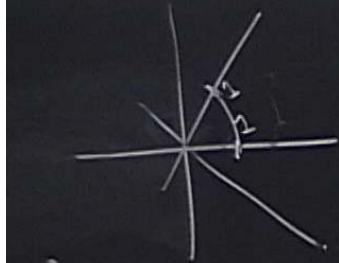
$$(ct)^2 - x^2$$
$$-1$$

collection of atoms



consider some Lorentz scalar e.g.  $\varphi_{(s)} = F_{\mu\nu} F^{\mu\nu}(x) = 2(-\underline{E}^2 + \underline{B}^2)$   
 $= \varphi_s(x)$

$$\varphi'(t', \underline{x}') = \varphi(t, \underline{x}) = \varphi_s(x)$$



$$(ct)^2 - x^2$$
$$-1$$

collection of atoms



consider some Lorentz scalar e.g.  $\varphi_{(x)} = F_{\mu\nu} F^{\mu\nu}(x) = 2(-\underline{E}^2 + \underline{B}^2)$   
 $= \varphi_S(x)$

$$\begin{aligned}\varphi'(t', x') &= \varphi(t, \underline{x}) = \varphi_S(\underline{x}) \\ &= \varphi_S(x, y, z) \\ &= \varphi_S(\gamma(x + vt'), y', z')\end{aligned}$$

If  $\varphi_S$  has peaks separated by  $\Delta x = L$   
then  $\varphi'$  " " " "  $\Delta x' = L/\gamma$

transform  
laws

from  $\phi$  to  $\psi$

$$\underline{E = mc^2}$$

Einstein (1905) knew about Maxwell eq's and that they have a conserved energy

transform  
laws

know  $\phi$  ...

$$E = mc^2$$

Einstein (1905) knew about Maxwell eq's and that they have a conserved energy

$\rho_{em}$   
↑  
energy-momentum

$$\int_V d^3x \rho_{em} = \text{Energy in } V$$

$$\underline{E = mc^2}$$

Einstein (1905) knew about Maxwell eq's and that they have a conserved energy

$$\rho_{em} = \frac{1}{2} \epsilon_0 E^2 + \frac{B^2}{2\mu_0}$$

↑  
energy-momentum

flux  $\underline{J}_{em} = \frac{1}{\mu_0} (\underline{E} \wedge \underline{B})$

(Poynting)  $\Rightarrow \frac{dE_{em}}{dt} = - \int_{S'} \underline{J}_{em} \cdot d\underline{S}$

$\Rightarrow$  2-index symmetric tensor, energy-momentum tensor  $T^{\mu\nu} = \frac{1}{\mu_0} \left( F^{\mu\alpha} F^{\nu\beta} - \frac{1}{4} \eta^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta} \right)$

$$\int_V d^3x \rho_{em} = \text{Energy in } V$$



Energy  
in Poy

$\Rightarrow$  2-index symmetric tensor, energy-momentum tensor.

$$T_{em}^{\mu\nu} = \frac{1}{\mu_0} \left( F^{\alpha\mu} F_{\alpha}^{\nu} \right)$$

$$\partial_{\mu} F^{\alpha\mu}$$

If  $J_c^M = 0$ , then M.E.s  $\Rightarrow \partial_{\mu} T_{em}^{\mu\nu} = 0$

(using ①)

$$\partial_{\mu} T_{em}^{\mu\nu} = \frac{1}{\mu_0} \left( F^{\alpha\mu} \partial_{\mu} F_{\alpha}^{\nu} - \frac{1}{2} F^{\alpha\beta} \partial^{\nu} F_{\alpha\beta} \right)$$

Energy-momentum

Energy  
in  $T^{\mu\nu}$

$\Rightarrow$  2-index symmetric tensor, energy-momentum tensor  $T_{em}^{\mu\nu} = \frac{1}{\mu_0} \left( F^{\mu\alpha} F_{\alpha}^{\nu} - \frac{1}{4} \eta^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta} \right) : T^{\mu\nu} = T^{\nu\mu} = J^{\mu\nu}$

$$\partial_{\alpha} F^{\alpha\mu} = 0$$

If  $J_c^M = 0$ , then M.E.s  $\Rightarrow \boxed{\partial_{\alpha} T_{em}^{\mu\nu} = 0}$

$$\begin{aligned} \partial_{\alpha} T_{em}^{\mu\nu} &= \frac{1}{\mu_0} \left( F^{\alpha\mu} \partial_{\alpha} F_{\alpha}^{\nu} - \frac{1}{2} F^{\alpha\beta} \partial^{\nu} F_{\alpha\beta} \right) \\ &\quad \text{(using ①)} \\ &= \frac{1}{2\mu_0} F^{\alpha\mu} \left( \partial_{\alpha} F_{\alpha}^{\nu} + \partial_{\alpha} F_{\mu}^{\nu} + \partial^{\nu} F_{\mu\alpha} \right) = 0 \\ &\quad \text{②} \end{aligned}$$

index symmetric tensor, energy-momentum tensor  $T_{em} = \frac{1}{\mu_0} (F_{\alpha\beta} F^{\alpha\beta} - 4 \dots)$

$$\partial_\mu F^{\alpha\mu} = 0$$

0, then M.E.s  $\Rightarrow \boxed{\partial_\mu T_{em}^{\mu\nu} = 0}$   
 (using 0)

$$= \frac{1}{\mu_0} \left( F^{\alpha\mu} \partial_\mu F_\alpha^\nu - \frac{1}{2} F^{\alpha\beta} \partial^\nu F_{\alpha\beta} \right)$$

$$= \frac{1}{2\mu_0} F^{\alpha\mu} (\partial_\mu F_\alpha^\nu + \partial_\alpha F_\mu^\nu + \partial^\nu F_{\alpha\mu}) = 0$$

②

$$\Rightarrow \partial_0 T_{em}^{00} = -\partial_i T_{em}^{i0}$$

$$\parallel \frac{1}{c} \partial_t$$

and  $\partial_0 T_{em}^{0i} = -\partial_j T_{em}^{ji}$

$$P^i = \int d^3x T^{0i}$$

obeys a similar law.

$\Rightarrow$  if Energy  $\equiv \int d^3x T_{em}^{00}(t,x)$

$$\text{then } \frac{d}{dt} \text{Energy} = \int d^3x \partial_t T_{em}^{00}(t,x)$$

$$= -c \int d^3x \partial_j T_{em}^{0j}$$

$$= -c \int d^3x J^i \delta^i$$

$$J^i = \frac{T^{0i}}{c}$$



$$\begin{aligned}
 ct &= \gamma(ct' + \frac{v}{c}x') \\
 x &= \gamma(x' + vt') \\
 \Leftrightarrow \\
 ct' &= \gamma(ct - \frac{v}{c}x) \\
 x' &= \gamma(x - vt)
 \end{aligned}$$

$$\begin{aligned}
 \cosh\theta &= \frac{1}{\sqrt{1 - \tanh^2\theta}} \\
 &= \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \equiv \gamma \\
 \sinh\theta &= \tanh\theta \cosh\theta \\
 &= \frac{v}{c}
 \end{aligned}$$

$\Rightarrow$  Lorentz transformations only allow you to transform to observers moving slower than  $c$ .

These can be combined into  $P^M$  (4-vector)

$$P^M = \frac{1}{6c} \int T^{0M} E_{\text{wave}} dx^1 dx^2 dx^3$$

Energy-Momentum 4-vector in EM field.

Consider plane wave

$$\begin{aligned}
 \underline{E} &= (0, 0, f(x-ct)) \\
 \underline{B} &= (0, -\frac{f(x-ct)}{c}, 0)
 \end{aligned}$$

$$\Rightarrow T^{00} = \frac{f^2}{\underbrace{Mc^2}_{\frac{1}{Mc^2} = c_0}}$$

$$T^{0x} = \frac{f^2}{Mc^2}$$

$$\Rightarrow E = \text{Area} \int dx f^2 E_0, \quad P^x = \frac{\text{Area}}{c} \int dx f^2$$

$$\Rightarrow P^r = (E/c, P, 0, 0) \text{ with } E = Pc$$

tensor, energy-momentum tensor

$$T_{em}^{\mu\nu} = \frac{1}{\mu_0} \left( F^{\mu\alpha} F_{\alpha}^{\nu} - \frac{1}{4} \eta^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta} \right)$$

$\partial_{\mu} F^{\mu\lambda} = 0$

$$\frac{1}{\mu_0} \left( F^{i0} F_{i0} + \frac{1}{2} (F^{0i} F_{0i} + F^{ij} F_{ij}) \right)$$

$T^{00} = \rho_{em}$   
 $T^{0i} = T^{i0} = J_{em}^i$

$\partial_{\mu} T_{em}^{\mu\nu} = 0$  (conserved)

$$\partial_0 T_{em}^{00} = -\partial_i T_{em}^{i0}$$

$$\frac{1}{c} \partial_t$$

if Energy  $\equiv \int d^3x T_{em}^{00}(t, \underline{x})$

then  $\frac{d}{dt} \text{Energy} = \int \partial_t T^{00}(t, \underline{x})$

$$= -c \int \partial_i T^{0i} d^3x$$

$$= -c \int d^3x J^i \partial_i$$

$$J^i = \frac{T^{0i}}{c}$$

$$\partial_{\mu} F^{\mu\nu} = \partial_{\alpha} F^{\alpha\nu} + \partial_{\beta} F^{\beta\nu} = 0$$

and  $\partial_0 T_{em}^{0i} = -\partial_j T_{em}^{ji}$

$P^i = \int d^3x T^{0i}$

obeys a similar law.