

Title: Evolution of Black Holes in Inflation

Date: Jul 25, 2017 11:00 AM

URL: <http://pirsa.org/17070002>

Abstract: <p>We present an analytic, perturbative solution that describes dynamical black holes in slow-roll inflation with a general potential. It is shown that the spacetime evolves quasi-statically through a sequence of quasi-Schwarzschild-deSitter metrics with time dependent cosmological constant and mass parameters, such that the cosmological constant is instantaneously equal to the value of the scalar potential.&nbsp; The areas of the black hole and cosmological horizons each increase in time as the effective cosmological constant decreases, and the the fractional area increase is proportional to the fractional change of the cosmological constant, times a geometrical factor. For black holes ranging in size from much smaller than to comparable to the cosmological horizon, the pre-factor varies from very small to order one. The change in the horizon areas happens in such a way that the first law of thermodynamics is obeyed between successive time steps.</p>

4. McInerny, Satishchandran, J.T.  
"Cosmography of KNdS  
BHs & Isentropic Phase  
Transitions" (ArXiv)

0. Chodburn & Gregory  
"Time Dep BHs & Scalar Hair"

1. Gregory, Kastor, J.T. "BH Thermo w/ dynamical  
Lambda" (ArXiv)

2. ——— " ——— " Evolution of BHs in Inflat.  
(WP)

3. Dolan, Kastor, Kubiznak, J.T. "Thermo. Vols &  
Isoperimetric Inequalities" (ArXiv)

main Refs



# "Evolution of BHs in Inflation"

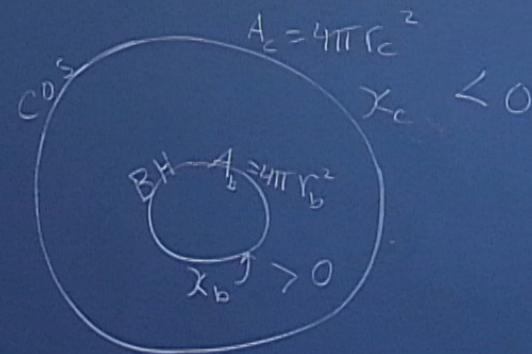
I. Static SdS

$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2$$

$$f(r) = 1 - \frac{2M}{r} - \frac{1}{3}r^2$$

$$= \frac{1}{3}(r_c - r)(r - r_b)(r + r_b + \epsilon)$$

$$(M, \Lambda) \leftrightarrow (1, r_b)$$



$$\chi_h = \frac{1}{2} f'(r_h)$$



# "Evolution of BHs in Inflation"

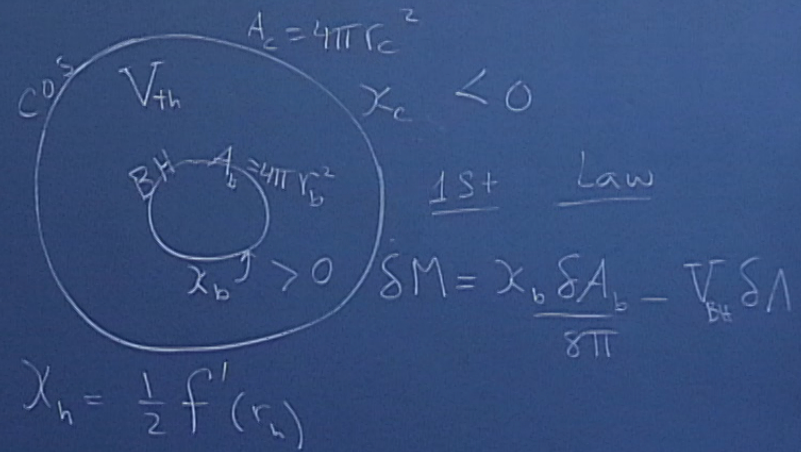
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$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2$$

$$f(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$$(M, \Lambda) \leftrightarrow (1, r_b)$$



$$\chi_h = \frac{1}{2} f'(r_h)$$



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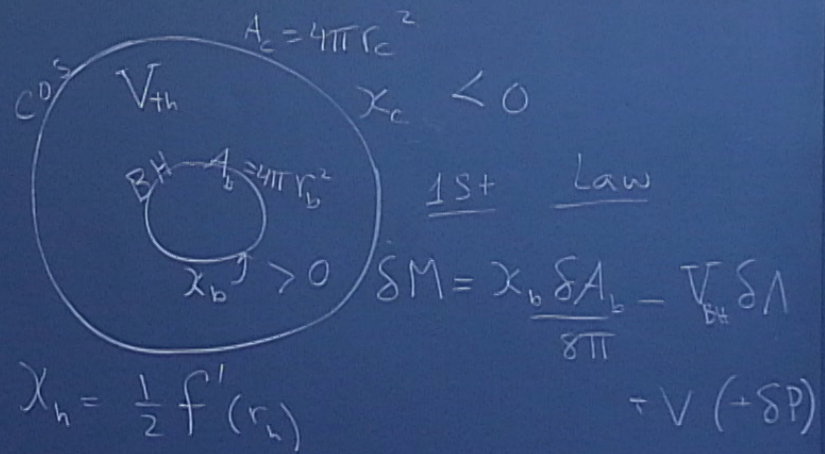
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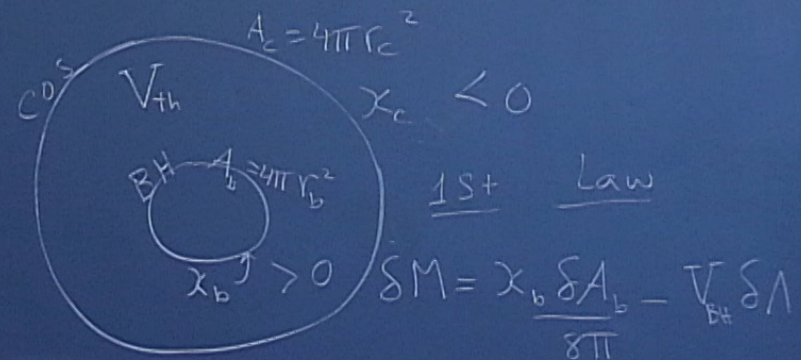
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$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$



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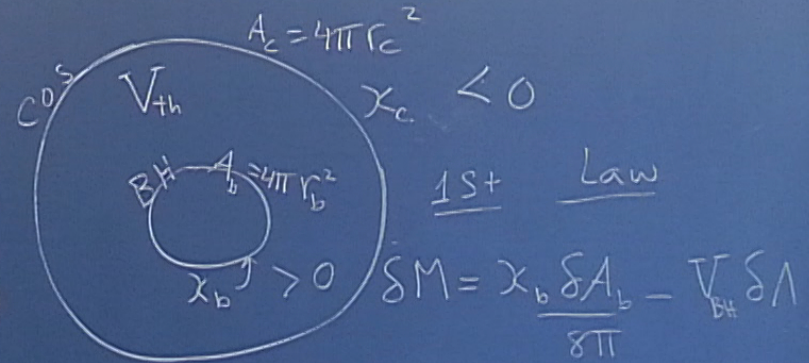
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$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2$$

$$f(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$$(M, \Lambda) \leftrightarrow (r_b)$$



$$\frac{1}{8\pi} \delta A_b - \frac{V_{BH}}{8\pi} \delta \Lambda$$

$$\delta M = \chi_b \frac{\delta A_b}{8\pi} - \frac{V_{BH}}{8\pi} \delta \Lambda$$

$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$

$$\text{for SdS } V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$$



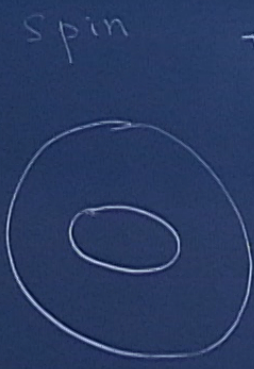
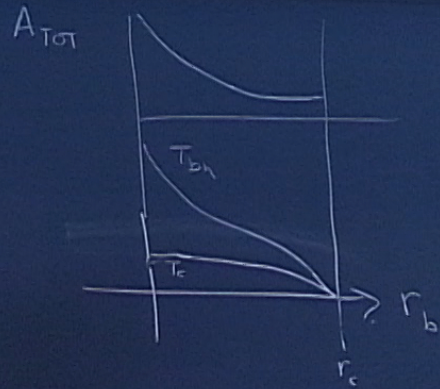
$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$(M, \Lambda) \leftrightarrow (r_b, r_c)$

$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$

for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$



$V, \Lambda, a$

$$A_{TOT} = A_c + A_b$$

fix  $\Lambda, a$



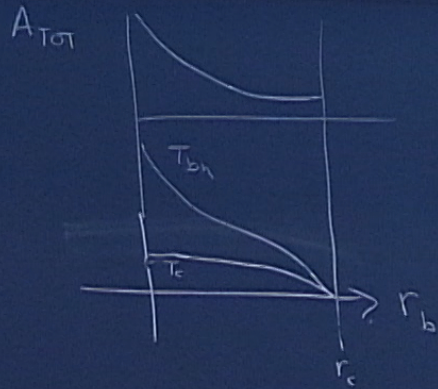
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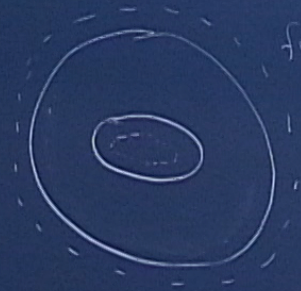
$(M, \Lambda) \leftrightarrow (r_b, r_c)$

$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$

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Spin



$V, \Lambda, a$

fixed  $= A_{TOT} = A_c + A_b$

fix  $\Lambda, a$



$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

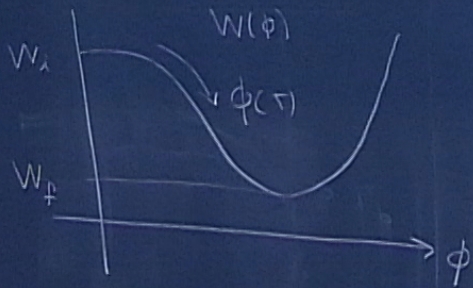
$(M, \Lambda) \leftrightarrow (r_b, r_c)$

$$\kappa_b \delta A_b + |\kappa_c| \delta A_c = -V_{th} \delta \Lambda$$

for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$

II

$ds \rightarrow ds$  w/ BH



a) Using techniques of (O.)



$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

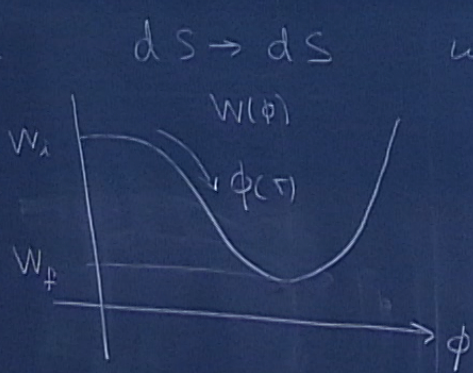
$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$(M, \Lambda) \leftrightarrow (r_b, r_c)$

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for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$

II



a) Using techniques of (0.)

$$\Delta A_b = A_b^f - A_b^i, \quad \Delta A_c \propto |\Delta \Lambda| = W_i - W_f$$

b) found that 1<sup>st</sup> Law Sat  $i \rightarrow f$

$$\chi_b \Delta A_b + |\chi_c| \Delta A_c =$$



$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

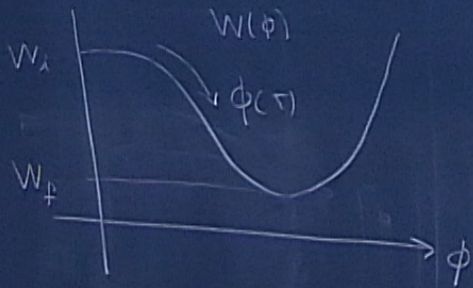
$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$(M, \Lambda) \leftrightarrow (\Lambda, r_b)$

$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$

for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$

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w/ BH

a) using techniques of (O.)

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$$\chi_b \Delta A_b + |\chi_c| \Delta A_c = \bar{V}_{th} |\Delta \Lambda|$$



$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

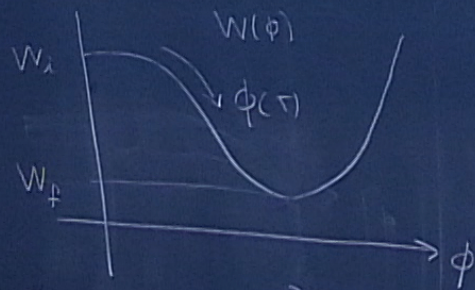
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$(M, \Lambda) \leftrightarrow (\Lambda, r_b)$

$$\chi_b \delta A_b + |\chi_c| \delta A_c = -V_{th} \delta \Lambda$$

for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$

II



$$P_\phi = W(\phi) + \frac{1}{2} \dot{\phi}^2 \text{ (metric dep)}$$

a) Using techniques of (0.) } perturb, slow roll

$$\Delta A_b = A_b^f - A_b^i, \quad \Delta A_c \propto |\Delta \Lambda| = W_i - W_f$$

b) found that 1<sup>st</sup> Law Sat  $i \rightarrow f$

$$\chi_b \Delta A_b + |\chi_c| \Delta A_c = \bar{V}_{th} |\Delta \Lambda|$$



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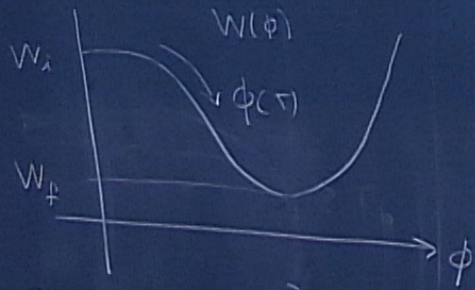
$$= \frac{\Lambda}{3} (r_c - r)(r - r_b)(r + r_b + r_c)$$

$(M, \Lambda) \leftrightarrow (r_b, r_c)$

$$\kappa_b \delta A_b + |\kappa_c| \delta A_c = -V_{th} \delta \Lambda$$

for SdS  $V_{th} = \frac{4\pi}{3} (r_c^3 - r_b^3)$

II  $ds \rightarrow dS$  choose  $W(\phi)$ ,  $r_b^i$ , slow roll  $\phi(\tau)$ , perturb Einst



$$\rho_\phi = W(\phi) + \frac{1}{2} \dot{\phi}^2 \text{ (metric dep)}$$

a) Infl, No BH,  $\psi(\tau)$ ,  $W(\psi)$

$$\Lambda(\tau) \simeq W(\psi(\tau)) ?$$

b) Infl, with BH  
And  $M(\tau)$  is true



BHs & Isentropic Phase Transitions" (ArXiv)

2. ——— " ——— " Evolution of BHs in Inflat. (WIP)

3. Dolan, Kastor, Kubiznak, JT, "Thermo. Vols & Isoperimetric Inequalities" (ArXiv)

III. Metric

a) whose time?

choose  $T = t + hcr$

s.t. (1) slow-roll ( $\dot{\phi} = h.o.t.$ )

$\exists$  sols  $\Psi(T)$  in SdS

(2)  $\Psi$  in-going at BH, out...cos



# "Evolution of BHs in Inflation"

## I. Static SdS

$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2 = -dT^2 f + 2\gamma dT dr + \frac{(1-\gamma^2)}{f} dr^2 + r^2 d\Omega^2$$

$$f_{SdS}(r) = 1 - \frac{2M}{r} - \frac{1}{3}r^2$$

$$\gamma_{SdS} = -\gamma r + \beta/r^2 \rightarrow \begin{cases} 1 & \text{BH} \\ -1 & \text{cosm.} \end{cases}$$

$$T \rightarrow \begin{cases} \text{in.} & \text{BH} \\ \text{out} & \text{cos} \end{cases}$$



# "Evolution of BHs in Inflation"

## I. Static SdS

$$T \rightarrow \begin{cases} \text{in.} & \text{BH} \\ \text{out} & \text{cos}^\circ, \psi = \psi(T) \end{cases}$$

$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2 = -dT^2 f + 2\gamma dT dr + \frac{(1-\gamma^2)}{f} dr^2 + r^2 d\Omega^2$$

$$f_{\text{SdS}}(r) = 1 - \frac{2M}{r} - \frac{1}{3}r^2$$

$$\gamma_{\text{SdS}} = -\gamma r + \beta/r^2 \rightarrow \begin{cases} 1 & \text{BH} \\ -1 & \text{cosm.} \end{cases}$$



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$\exists$  sols  $\psi(T)$  in SdS

(2)  $\psi$  in-going at BH, out ... cos

$$-3\gamma \dot{\psi} = \frac{2W}{2\psi}$$



# "Evolution of BHs in Inflation"

with  $\Lambda(T) \simeq W(T)$ ,  $\varphi = \varphi(T)$

## I. Static SdS

$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2 = -dT^2 f + 2\gamma dT dr + \frac{(1-\gamma^2)}{f} dr^2 + r^2 d\Omega^2$$

$$f_{SdS}(r) = 1 - \frac{2M}{r} - \frac{1}{3} r^2$$

$$\gamma_{SdS} = -\gamma r + \beta/r^2$$

$$f(T, r) = 1 - \frac{2M(T)}{r} - \frac{\Lambda(T)}{3} r^2 + \delta f(T, r)$$

$$= f_{qst} + \delta f$$

$= W(\varphi)$ ,  $r_b^i$ , slow roll  $\varphi(T)$ , perturb Einst

a) Infl, No BH,  $\varphi(T)$ ,  $W(\varphi)$



$$f_{\text{Sch}}(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$\gamma_{\text{Sch}} = -\gamma r + \beta/r^2$$

$$f(T, r) = 1 - \frac{2M(T)}{r} - \frac{\Lambda(T)}{3} r^2 + \delta f(T, r)$$
$$= f_{\text{Sch}} + \delta f$$

and  $\gamma(T, r) = \gamma_{\text{Sch}}(r) + \delta\gamma(T, r)$



$$T(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$\gamma_{SdS} = -\gamma r + \beta/r^2$$

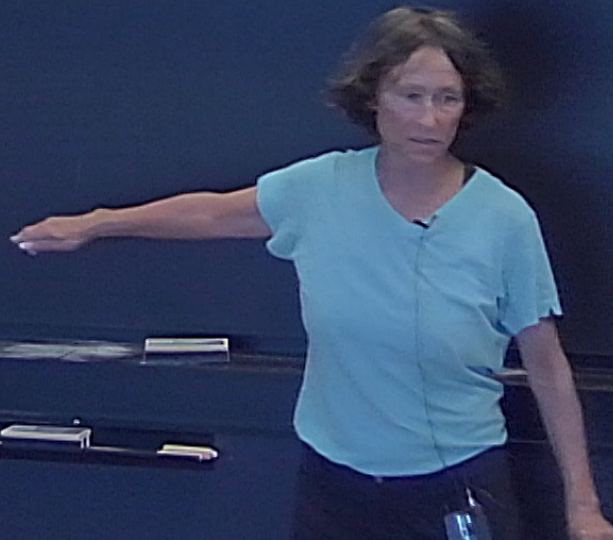
$$f(T, r) = 1 - \frac{2M(T)}{r} - \frac{\Lambda(T)}{3} r^2 + \delta f(T, r)$$

$$= f_{qst} + \delta f$$

and  $\gamma(T, r) = \gamma_{SdS}(r) + \delta\gamma(T, r)$

$$\delta\gamma = \frac{r(1-\gamma^2)}{2f} \int^T \dot{\phi}^2 dT'$$

$$\delta f = -\frac{\dot{\phi}^2}{2r} \int_{r_0}^r dr' \frac{r'^2(1-\gamma'^2)}{f}$$





$$\gamma(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$\gamma_{sds} = -\gamma r + \beta/r^2$$

$$f(\tau, r) = 1 - \frac{2M(\tau)}{r} - \frac{\Lambda(\tau)}{3} r^2 + \delta f(\tau, r)$$

$$= f_{qst} + \delta f$$

and  $\gamma(\tau, r) = \gamma_{sds}(r) + \delta\gamma(\tau, r)$

$$\delta\gamma = \frac{r(1-\gamma^2)}{2f} \int^{\tau} \dot{\phi}^2 d\tau'$$

$$\delta f = -\frac{\dot{\phi}^2}{2r} \int_{r_b}^r dr' \frac{r'^2(1-\gamma'^2)}{f}$$

$$2\dot{M} = \beta \dot{\phi}^2, \quad \dot{\Lambda} = -3\gamma \dot{\phi}$$

or  $\Lambda = W(\phi)$

$$\gamma = \gamma_{sds} + \delta\gamma \rightarrow 1 \quad \text{BH}$$

$$1 - \gamma \rightarrow C_b(r - r_b(\tau))$$

$$r_b(\tau), r_c(\tau) \leftrightarrow \Lambda(\tau), M(\tau)$$



"Evolution of BHs in Inflation"

I. Static SdS

with  $\Lambda(T) \approx W(T)$

$\varphi = \varphi(T)$

$$ds^2 = -f dt^2 + \frac{dr^2}{f} + r^2 d\Omega^2 = -dT^2 f + 2\eta dT dr + \frac{(1-\eta^2)}{f} dr^2 + r^2 d\Omega^2$$

$$f(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$\eta = -\gamma r + \beta/r^2$$

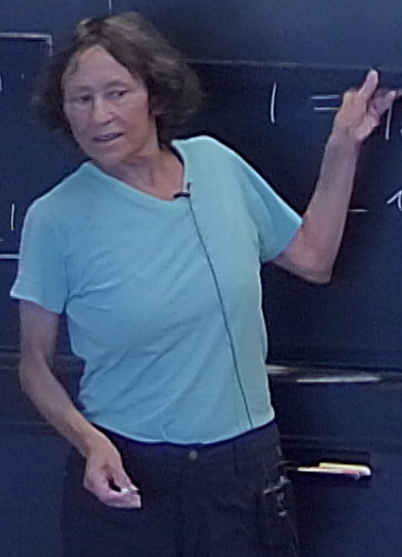
$$\rightarrow -dT^2 f + 2 dT dr + \frac{C_b}{2b} dr^2 + r^2 d\Omega^2$$

$$\delta f = -\frac{\dot{\phi}^2}{2r} \int_{r_b}^r dr' r'^2 (1 - \frac{2M}{r'})$$

$$1 = 1_{SdS} + \delta \eta \rightarrow 1 \quad BH$$

$$\eta \rightarrow C_b (r - r_b(T))$$

$$r_b(T), r_c(T) \leftrightarrow \Lambda(T), M(T)$$





$$\gamma_{sds}(r) = 1 - \frac{2M}{r} - \frac{\Lambda}{3} r^2$$

$$\dot{\gamma}_{sds} = -\dot{\gamma} r + \beta/r^2$$

$$\rightarrow + 2 dT dr + \frac{C_b}{2b} dr^2 + r^2 d\Omega^2$$

and  $\gamma(T, r) = \gamma_{sds}(r) + \delta\gamma(T, r)$

$$\delta\gamma = \frac{r(1-\gamma^2)}{2f} \int \dot{\phi}^2 dT'$$

$$\delta f = -\frac{\dot{\phi}^2}{2r} \int_{r_b}^r dr' \frac{r'^2(1-\gamma'^2)}{f}$$

$$2\dot{M} = \beta \dot{\phi}^2, \quad \dot{\Lambda} = -3\gamma \dot{\phi} \rightarrow$$

$$\Lambda = W(\phi)$$

$$\gamma = \gamma_{sds} + \delta\gamma \rightarrow 1 \quad \text{BH}$$

$$1 - \gamma \rightarrow C_b(r - r_b(T))$$

$$r_b(T), r_c(T) \leftrightarrow \Lambda(T), M(T)$$



BHs & Isentropic Phase Transitions" (ArXiv)

2. ——— " Evolution of BHs in Inflat. (WP)

3. Dolan, Kastor, Kubiznak, JT, "Thermo. Vols & Isoperimetric Inequalities" (ArXiv)

$$\Rightarrow \dot{r}_h = \frac{r_h}{2|\chi_h|} \dot{\phi}^2 ; \quad \delta A_h^{(\tau)} = \frac{A_h}{|\chi_h|} \frac{V_{Th}}{A_{Tot}} [W_i - W(\psi(\tau_1))]$$

$$\Rightarrow \chi_b \dot{A}_b + |\chi_c| \dot{A}_c = -V_{Th} \dot{A}$$



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$$\Rightarrow \dot{r}_h = \frac{r_h}{2|\chi_h|} \dot{\phi}^2 ; \quad \delta A_h^{(\tau)} = \frac{A_h}{|\chi_h|} \frac{V_{Th}}{A_{TOT}} [W_i - W(\psi(\tau_1))]$$

$$\Rightarrow \chi_b \dot{A}_b + |\chi_c| \dot{A}_c = -V_{Th} \dot{\Lambda} \left| \frac{\delta A_b}{A_b} \sim \frac{|\Delta \Lambda|}{\Lambda} \cdot \begin{cases} (r_b \sqrt{\Lambda}) & \text{small} \\ 1 & \text{large} \end{cases} \right.$$