

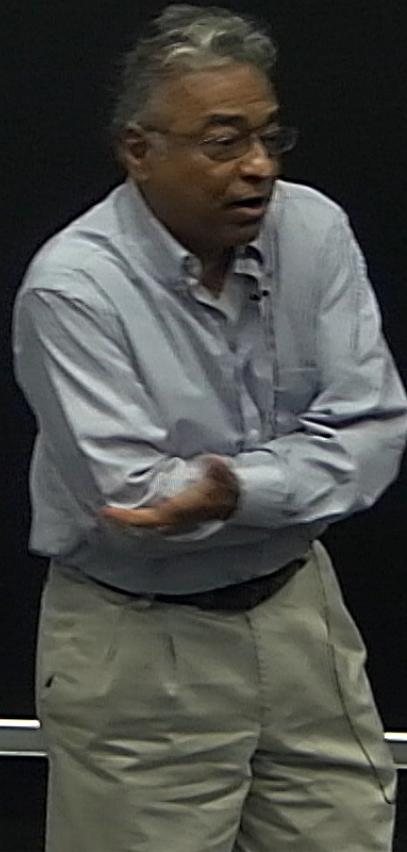
Title: Bounce in Loop Quantum Cosmology and its Implications

Date: Jun 26, 2017 10:05 AM

URL: <http://pirsa.org/17060095>

Abstract: There is a huge body of work in Loop Quantum Cosmology comprising of several thousand journal articles. I will provide an overview of conclusions, focusing on the difficult conceptual and mathematical issues that accompany the notion of a bounce and opening the way for phenomenological implications that will be discussed by Ivan Agullo.

Bounce in LQC & Its implications



Bounce in LQC & Its implications

1.82: New Physics: Quantum Riemannian Geometry

## Bounce in LQC & Its implications

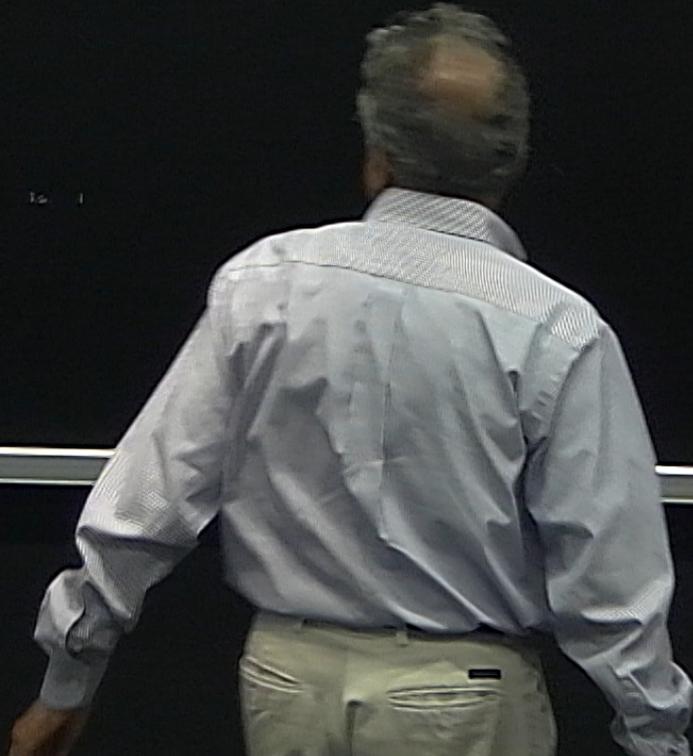
1 & 2: New Physics: Quantum Riemannian Geometry, sharp results on upper bounds of observables,

WDW Theory  $\Psi(a, \phi)$ , WDW eqn: Quantum Hamiltonian constraint:  $\hat{C}_{H+} \Psi = 0$

Source in LQC & its implications

1 & 2: New Physics: Quantum Riemannian Geometry; sharp results on upper bounds of observables.

WDW Theory  $\Psi(a, \phi)$ , WDW eqn: Quantum Hamiltonian Constraint:  $\hat{C}_H \Psi = 0$   
Hilbert space  $\mathcal{H}_{phys}$  on solutions:



1.22: New Physics: Quantum Riemannian Geometry; sharp results on upper bounds of observables;

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$\Psi(a, \phi) \in \mathcal{H}_{\text{phys}}$  st  $\hat{C}_H$  is SA on it;  $\Psi_{\text{phys}} = \int dx e^{ix\hat{C}_H} \Psi_{\text{kin}}$

182. Novel Physics: Quantum Riemannian Geometry; sharp results on upper bounds of observables;

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142: Nova Physics: Quantum Riemannian Geometry; sharp results on upper bounds of observables;

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$U(\lambda) = e^{i\lambda \hat{C}}$  compact;  $\hat{C} = \frac{1}{2} = 0$

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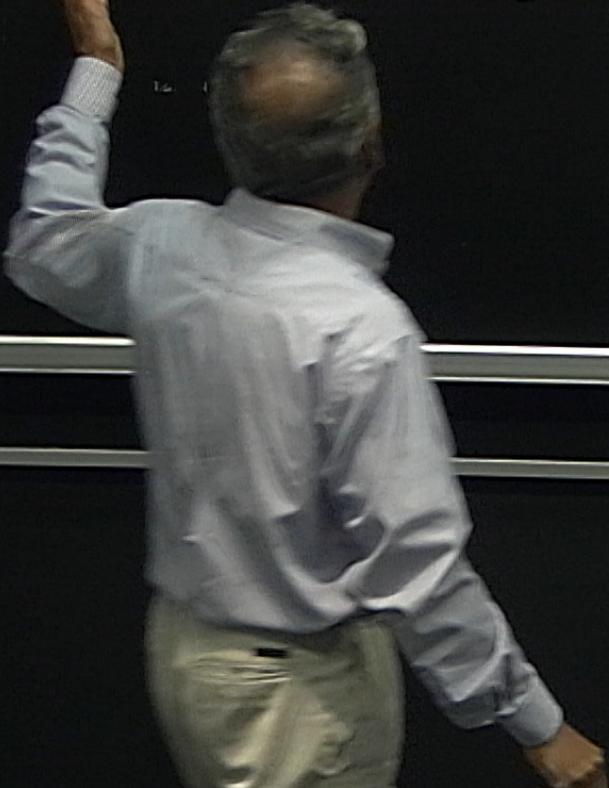
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for bounds of observables,

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$\psi_{phys} = \int dx_i e^{i\lambda \hat{C}_H} \psi_{kin}$   
"delta function"  
Not not normalizable in  $\mathcal{H}_{kin}$



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↳ "Group averaging"

... kinds of observables,

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$U(\chi) = e^{i\chi \hat{C}}$  compact;  $\hat{C} = \hat{C}_1 = 0$

$$\psi_{phys} = \int d\chi e^{i\chi \hat{C}_H} \psi_{kin}$$

"delta function"

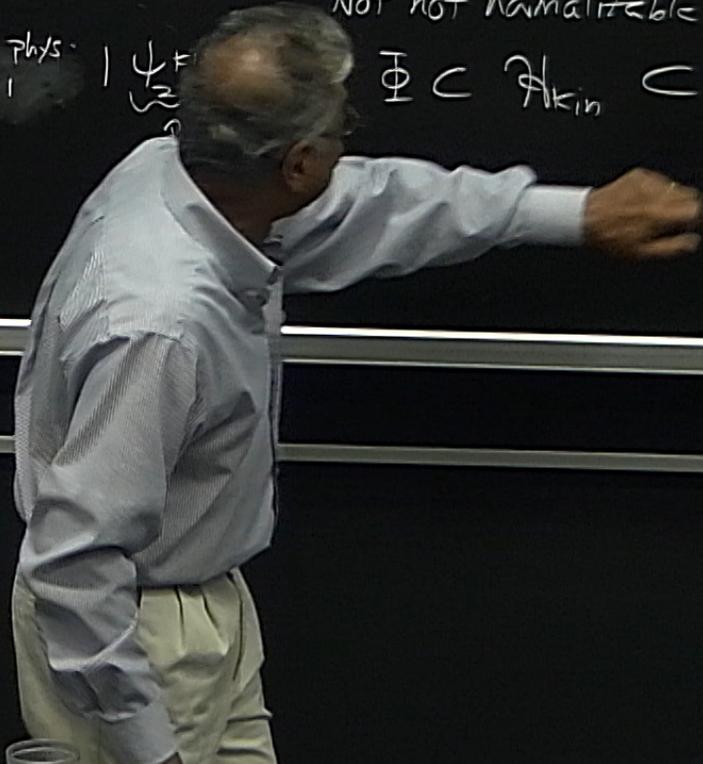
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"Group averaging"

Gelfand  
Riesz HS

$$\langle \psi_1^{phys} | \psi_2^{phys} \rangle = \langle \psi_1^{phys} | \psi_2^{phys} \rangle$$

$$\Phi \subset \mathcal{H}_{kin} \subset \Phi^*$$



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$$\Psi_{phys} = \int dx \underbrace{e^{i\hat{x}\hat{c}_H}}_{\delta(\hat{c}_H)} \Psi_{kin}$$

Not normalizable in  $\mathcal{H}_{kin}$

"Group averaging"

Gelfand Rigged HS

$$\langle \Psi_1^{Thy} | \Psi_2^{Thy} \rangle = \langle \Psi_1^{phys} | \underbrace{\Psi_2^{kin}}_{\Phi} \rangle, \quad \Phi \subset \mathcal{H}_{kin} \subset \Phi^*$$

NDE:

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"δ(C)"

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$$\Phi \subset \mathcal{H}_{kin} \subset \Phi^\times$$

NDE: matter  
scalar field  $\phi$  ( $m=0$ ).

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 Hilbert space  $\mathcal{H}_{phys}$  on solutions:

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 $U(\lambda) = e^{i\lambda \hat{C}}$  compact;  $\hat{C} = \frac{1}{2} = 0$       "  $\delta(C^2)$  "

Not normalizable in  $\mathcal{H}_{kin}$

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 matter  
 NDE: scalar field  $\phi$  ( $m=0$ ).  
 $\Psi_{phys}(a, \phi)$

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Not normalizable in  $\mathcal{H}_{\text{kin}}$

"Group averaging"

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$\langle \Psi_1^{\text{phys}} | \Psi_2^{\text{phys}} \rangle = \langle \Psi_1^{\text{phys}} | \underbrace{\Psi_2^{\text{kin}}}_{\in \Phi} \rangle$ ,  $\Phi \subset \mathcal{H}_{\text{kin}} \subset \Phi^*$

NDE: scalar field  $\phi$  ( $m=0$ ).

$\Psi_{\text{thy}}(a, \phi)$ ,  $\hat{p}$

WDW Theory  $\Psi(a, \phi)$ , WDW eqn: Quantum Hamiltonian Constraint:  $\hat{C}_H \Psi = 0$   
 Hilbert space  $\mathcal{H}_{phys}$  on solutions:

$\Psi(a, \phi) \in \mathcal{H}_{kin}$  st  $\hat{C}_H$  is SA-on it:  $\Psi_{phys} = \int dx_1 \dots dx_n \underbrace{e^{i\lambda \hat{C}_H}}_{\delta(C_H)} \Psi_{kin}$   
 $U(\lambda) = e^{i\lambda \hat{C}}$  compact;  $\hat{C} = \frac{1}{2} = 0$

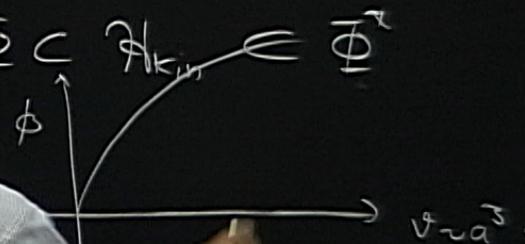
"Group averaging" not normalizable in  $\mathcal{H}_{kin}$

Gelfand  
Riesz HS

$\langle \Psi_1^{phys} | \Psi_2^{phys} \rangle = \langle \Psi_1^{phys} | \Psi_2^{kin} \rangle$   
 $\mathbb{R} \subset \mathcal{H}_{kin} \in \mathbb{R}^2$

NDE: scalar field  $\phi$  ( $m=0$ )

$\Psi_{phys}(a, \phi)$ ,  $\hat{p}$ , Unbo

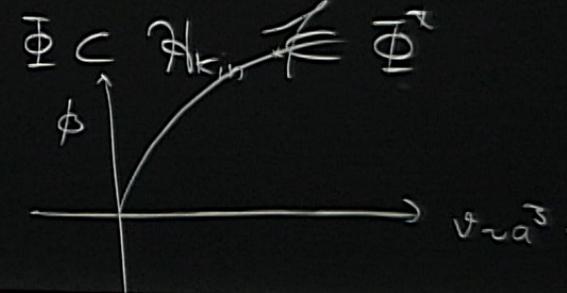


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"Group averaging"  
 Gelfand-Risned HS  
 $\langle \Psi_1^{phys} | \Psi_2^{phys} \rangle = \langle \Psi_1^{kin} | \Psi_2^{kin} \rangle$   
 matter scalar field  
 NDE:  $\Psi_{phys}(a, \phi)$



L&C:

1. New kinematic setup.

2. New Dynamics.

LQG:

1. New kinematic setup.

2. New Dynamics.

LQG:

~~( $G_{ab}$ ,  $F_{ab}$ ,  $K_{ab}$ )~~

$A_a^i$   
Spinconnection



L&C:

1. New kinematic setup.

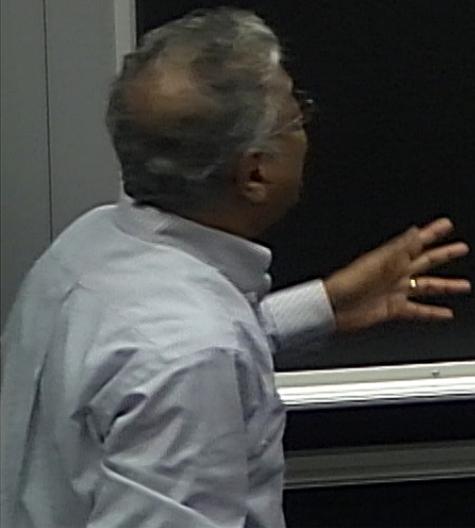
2. New Dynamics.

L&G:

~~( $g_{ab}$ ,  $\Gamma^a_b$ ,  $f^a_b$ )~~

$A^i_a$   
Spinconnection

$E^a_i$  ~ orthonormal Triad



LQC:

1. New kinematic setup.

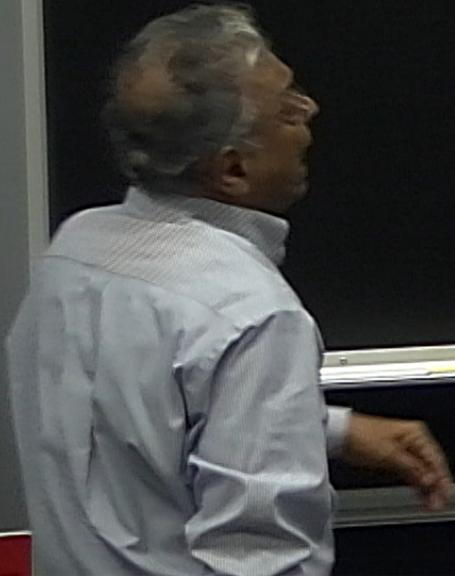
2. New Dynamics.

LQG:

~~( $g_{ab}$ ,  $\Gamma^a_{bc}$ )~~

( $A^i_a$ ,  
Spinconnection

$E^a_i \sim$  orthonormal Triad



CAUTION

LQC:

1. New kinematic setup

2. New

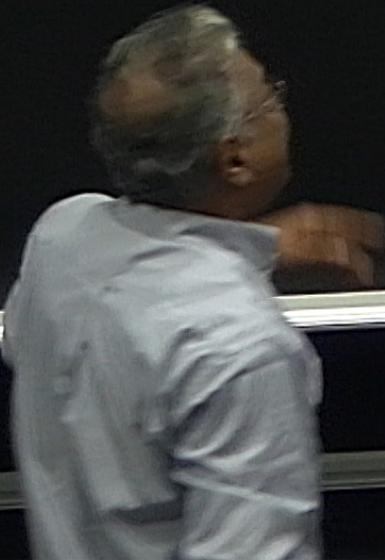
LQG: ~~( $g_{ab}, \gamma_{ab} \sim \text{Kob } \rho$ )~~

$(A_a^i, \text{Spin connection})$

$E_i^a \sim \text{orthonormal triad } ( )$

Background Ind Ansatz:  $(h_{\gamma}, F_S) \text{ Fluxes}$

$P \xrightarrow{h_{\gamma}} \rightarrow$   
 $= P \exp \int_{\gamma} A_{\mu} dx^{\mu}$



CAUTION

LGC:

1. New kinematic...

2. ...

LQG:  $(g_{ab}, \gamma^a_b \sim \text{Kob } \omega)$

$(A^i_a)$   
spinconnection

$E^a_i \sim \text{orthonormal triad ()}$

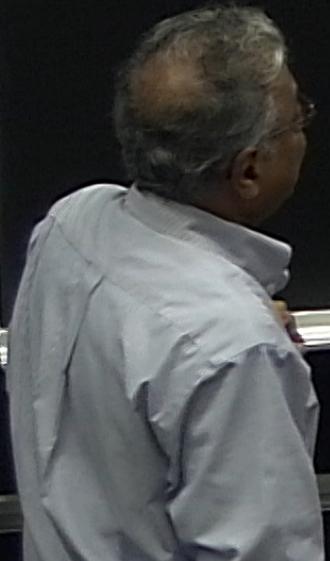
$P \xrightarrow{h\nu} \sqrt{\phantom{x}}$   
 $= P \exp \int_{\gamma} A^i_a dx^a$

Background Ind Algebra:  $(h_{ij}, F^i_j)$  Fluxes

Reps of this VHP,

LOST Theorem,

UNIQUE rep if you ask diff inv.



CAUTION

LQC:

LQG:  $(g_{ab}, \overset{\circ}{A}^a_b)$

$(\overset{\circ}{A}^a_b)$   
spinconnection

$E^a_i \sim$  orthonormal triad  $(\epsilon)$

Background Ind Algebra:  $(h_{ij}, F^i_j)$  Fluxes

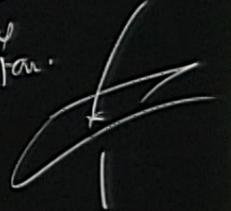
Reps of this VHP,

LOST Theorem, UNIQUE rep if

→ Riem Geo observables: Discrete eigenvalues.

if you ask Dirac  
elementary excitations.

$P = \exp(\int \overset{\circ}{A}^a_b dx^b)$



LQC:

1. New kinematic setup.

2. New Dynamics.

LQG:

$$\left( \begin{matrix} S_{ab} \\ S_a \end{matrix} \right) \sim \left( \begin{matrix} p_{ab} \\ p_a \end{matrix} \right)$$

$(A_a^i, \text{spinconnection})$

$E_i^a \sim \text{orthonormal triad}$

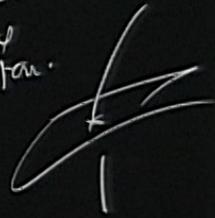
Background Ind Algebra:  $(h_{ij}, F_S)$  Fluxes

{ Reps of this AHP,

LOST Theorem, UNIQUE rep if

→ Riem Geo observables: Discrete eigenvalues.

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elementary excitations.



LQC:

$$A_a^i \sim C \underline{e}_a^i, \quad E_i^a \sim p_{(i}^a)$$

with an eye to the full theory.

LQC:

1. New kinematic setup.

2. New Dynamics.

LQG:

$(\begin{matrix} S_{ab} \\ S_a \end{matrix}, \begin{matrix} p_{ab} \\ p_a \end{matrix})$

$(A_a^i, \text{Spinconnection})$

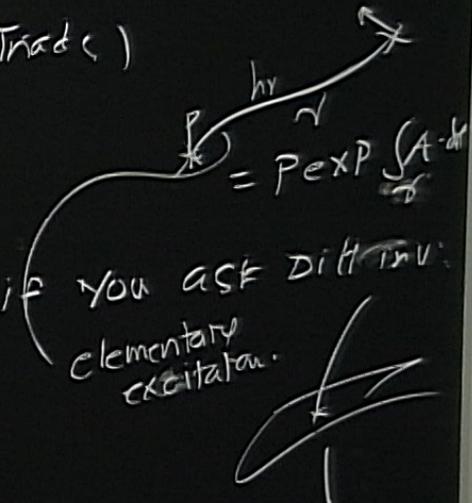
$E_i^a \sim \text{orthonormal Triad}$

Background Ind Algebra:  $(h_{ij}, F_S)$  Fluxes

Reps of this AHF,

LOST Theorem, UNIQUE rep

→ Riem Geo observables: Discrete eigenvalues.



LQC:

$A_a^i \sim c \tilde{e}_a^i$

$E_i^a \sim p_{(i}$

with an eye to the full theory,

observables:  $p_{ij} \sim e^{iAC} |_{\mathbb{R}}$

Fluxes  $\sim p$

LQG:  $(S_{ab}, \gamma_{ab}, K_{ab})$ ,  $(A_a^i, \text{spinconnection})$ ,  $E_i^a$

Background Ind Algebra:  $\hbar, \int S$  Fluxes

{ Reps of this AHF, LOST Theorem, UNIQUE rep if you ask Dirac elementary excitations.

→ Riem Geo observables: Discrete eigenvalues.

LQC:  $A_a^i \sim C \hat{e}_a^i$ ,  $E_a^i \sim P_{(i)}$  Fluxes  $\sim P$

with an eye to the full theory, observables:  $\hbar \sim e \int_{\mathbb{R}}$

unique rep:  $\psi(C) \text{ not } L^2(\mathbb{R})$

LQC:

LQG:  $(g_{ab}, \overset{ab}{p}_{ab})$   
 $\left\{ \begin{matrix} g_{ab} \\ p_{ab} \end{matrix} \right.$

$(A_a^i, \text{spin connection})$

$E_i^a \sim \text{orthonormal triad}$

Background Ind Algebra:  $(h_{ij}, F_S)$  Fluxes

{ Reps of this VHP

LOST Theorem

UNIQUE rep

if you ask Dirac elementary excitations.

→ Riem Geo observables: Discrete eigenvalues.

LQC:  $A_a^i \sim C \overset{a}{e}_a^i$ ,  $E_i^a \sim p_{(i}$

with an eye to the full theory

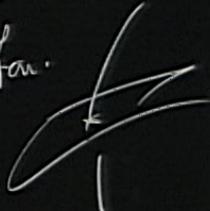
observables:

$p_{ij} \sim e^{i\alpha C} |_{\mathbb{R}}$

Fluxes  $\sim p$

Unique rep.  $\psi(C)$  not  $L^2(\mathbb{R})$   
(Inequivalent to std Schrödinger rep)

$$\psi(C) = \sum_{j=1}^N \alpha_j e^{i\alpha_j C}$$



... be classified in a useful way.

# Bounce in LQC & Its implications

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$$\Psi_{phys} = \int dx_i e^{i\hat{C}_H} \Psi_{kin}$$

"delta C hat"

Not not normalizable in  $\mathcal{H}_{kin}$

"Group averaging"

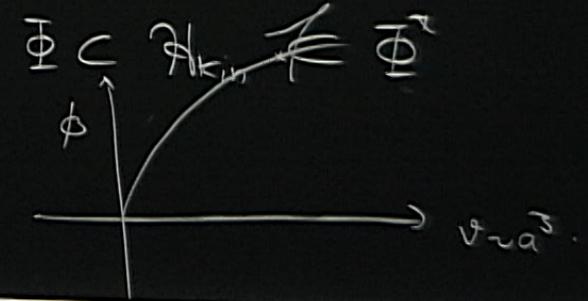
Gelfand  
Riesz HS

$$\langle \Psi_1^{phys} | \Psi_2^{phys} \rangle = \langle \Psi_1^{phys} | \Psi_2^{kin} \rangle$$

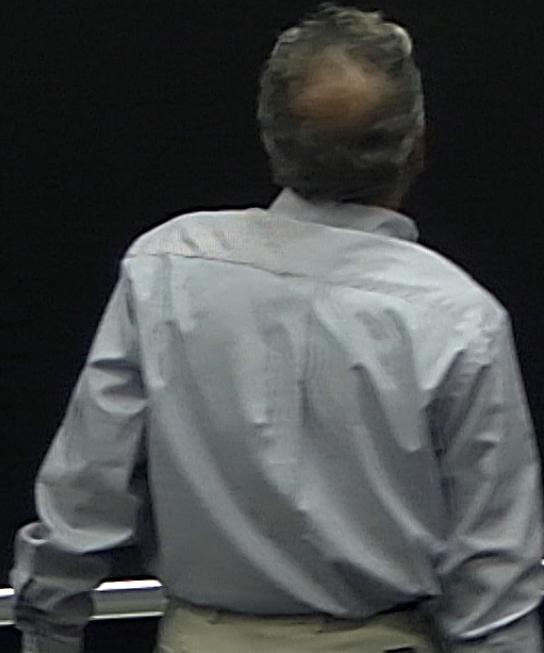
matter  $\hat{\Phi}$

NDE: scalar field  $\phi$  ( $m=0$ ).

$\Psi_{phys}(a, \phi)$ ,  $\hat{p}$ , unbounded;



Dynamics:  $\hat{C}_H$  on  $\mathcal{A}_{kin}$  (of almost periodic fn)  
Need new tools:  $\rightarrow F_{ab}^i = \lim_{A \rightarrow \infty} \frac{h_a}{Area}$



Dynamics:  $\hat{C}_H$  on  $\mathcal{P}_{kin}$  (of almost periodic fn)

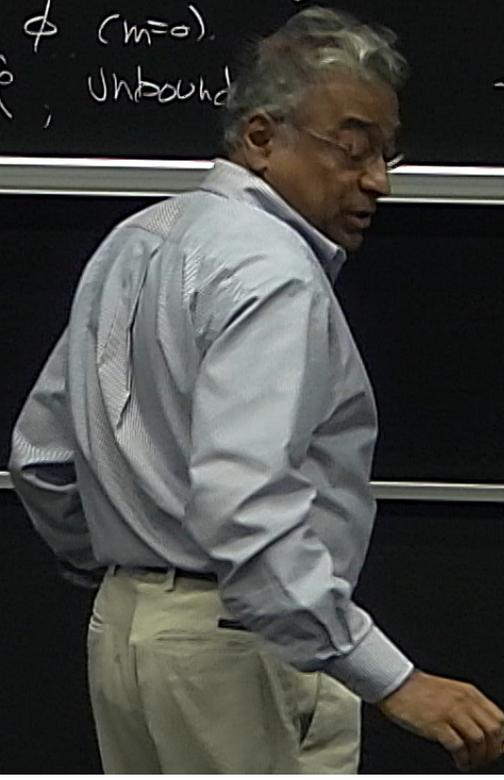
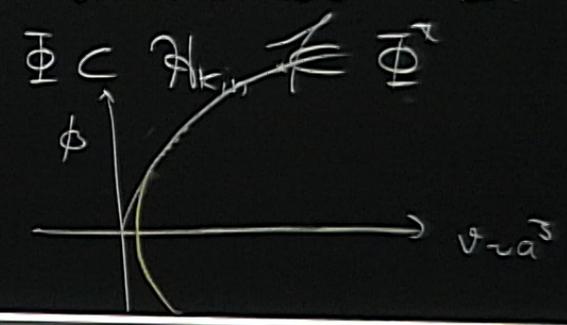
Need new tools:  $\rightarrow \int_{ab}^i = \lim_{A \rightarrow \Delta} \frac{h_a}{Area}$

$\Delta_p^2$ : min. non zero area eigenvalue

$\Psi(a, \phi) \in \mathcal{H}_{\text{kin}}$  st  $\hat{C}_H$  is SA-on it  
 $U(x) = e^{i\hat{x}\hat{c}}$  compact;  $\hat{C} = \frac{1}{2} = 0$

$\Psi_{\text{phys}} = \int dx_1 e^{i\hat{C}_H} \Psi_{\text{kin}}$   
 "delta C H"  
 Not normalizable in  $\mathcal{H}_{\text{kin}}$

"Group averaging"  
 Gelfand Rigged HS.  
 $\langle \Psi_1^{\text{phys}} | \Psi_2^{\text{phys}} \rangle = \langle \Psi_1^{\text{phys}} | \Psi_2^{\text{kin}} \rangle$   
 matter  $\hat{p}$   
 NDE: scalar field  $\phi$  ( $m=0$ )  
 $\Psi_{\text{thy}}(a, \phi)$ ,  $\hat{p}$ , Unbound



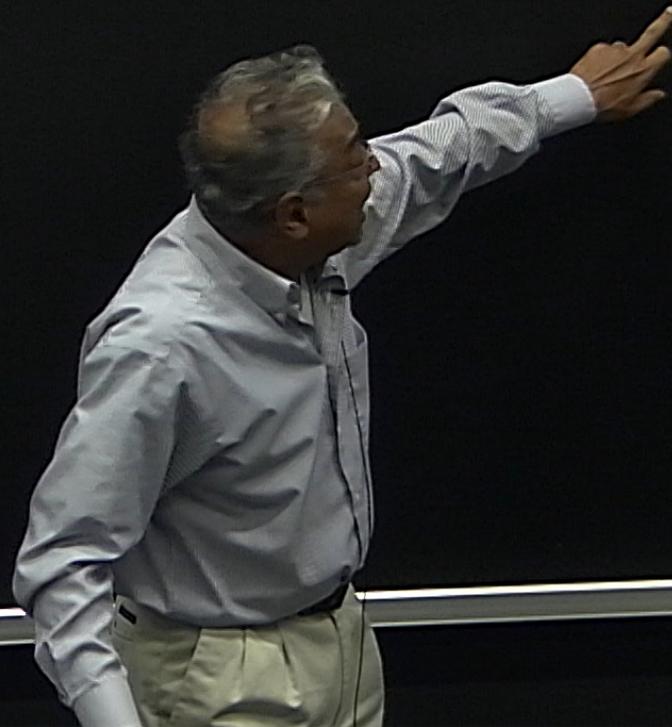
Dynamics:  $\hat{C}_H$  on  $\mathcal{H}_{kin}$  (of almost periodic fn)

Need new tools:  $\rightarrow \hat{F}_{ab}^i = \lim_{A \rightarrow \Delta} \frac{h_a}{h_{ab}}$

$\hat{P}$ : eigenvalues: Have an upper bound on  $\mathcal{H}_{phys}$ .

$\Delta_P^2$ : min. non zero area eigenvalue

$$P_{SOP} = \frac{18\pi}{\Delta^3} P_T$$



Dynamics:  $\hat{C}_H$  on  $\mathcal{H}_{kin}$  (of almost periodic fn)

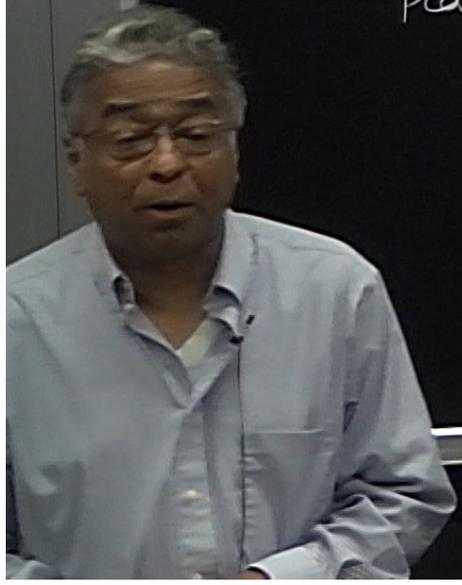
Need new tools:  $\rightarrow \hat{F}_{ab}^i = \lim_{A \rightarrow \Delta} \frac{h_a}{A \epsilon_a}$

$(\Delta)_P^2$ : min non zero area eigenvalue

$\hat{P}$ : eigenvalues: have an upper bound on  $\mathcal{H}_{phys}$

$\rho_{sub} = \frac{18\pi}{\Delta^3} \rho_P$

3. Quantum physics essential. BUT for states that are sharply peaked @ late times,



Dynamics:  $\hat{C}_H$  on  $\mathcal{H}_{kin}$  (of almost periodic fn)

Need new tools:  $\rightarrow \hat{F}_{ab}^i = \lim_{A \rightarrow \Delta} \frac{\overline{h a}}{Area}$

$(\Delta)_p^2$ : min non zero area eigenvalue

$\hat{P}$ : eigenvalues: have an upper bound on  $\mathcal{H}_{phys}$ .

$$P_{sub} = \frac{18\pi}{\Delta^3} P_A$$

3. Quantum physics essential. BUT for states that are sharply peaked @ late times, Effective eqns.

$$\left(\frac{\dot{a}}{a}\right)^2 = H^2 = \frac{8\pi G}{3} \rho \left(1 - \frac{\rho}{\rho_{sub}}\right), \text{ Raichadhuri also modified}$$

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Some continuity Eq<sup>n</sup>

SOME CONTINUITY ETC

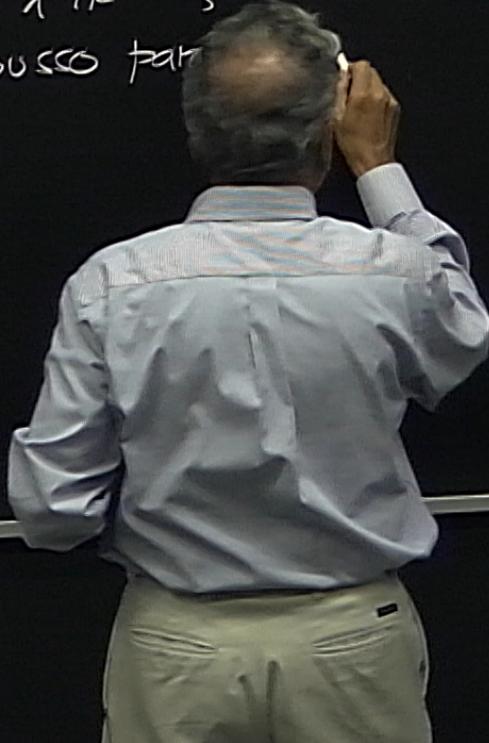
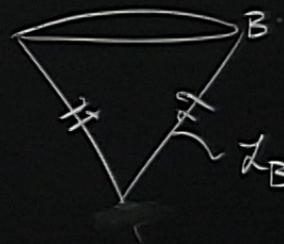
4. Examples:  
Potential Tension bet. UV & IR { Extreme Care

CAUTION

4. Examples:

- Potential Tension bet
- Entropy bounds

UV & IR } Extreme Care  
Bousso Par



CAUTION

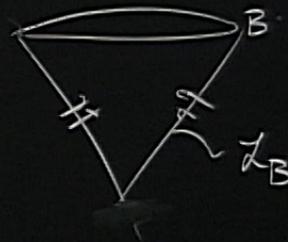
#### 4. Examples:

- Potential Tension bet
- Entropy bounds

UV & IR { Extreme Care

Bousso Paradigm  $S_B = \int_{\mathcal{L}_B} dA_a S^a$

$$S_B \leq \frac{A_B}{4 l_p^2}$$



#### 4. Examples:

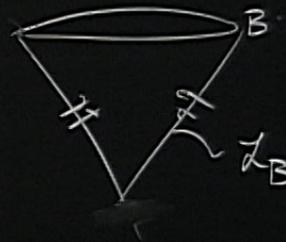
- Potential Tension bet
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- 1) Entropy current
- 2) QG
- 3) sharp geometry.

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$\rho_{\text{crit}}$   
Jobs

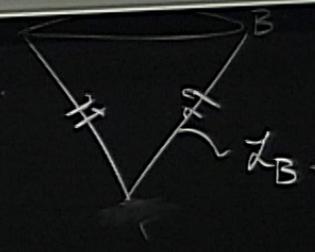
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(a)  
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Same continuity Eq<sup>n</sup>

- Potential Tension bet UV & IR } Extreme Care
- Entropy bounds - Bousso paradigm  $S_B = \int_{\mathcal{L}_B} dA \cdot S^a$
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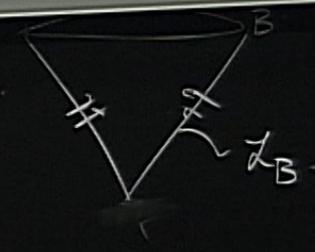


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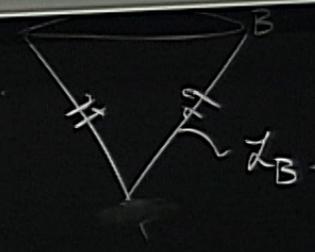
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$$S_B \leq \frac{A_B}{4 \ell_p^2}$$

$$S_B^{LCC} \leq 0.97 \frac{A_B}{4 \ell_p^2}$$

$$S_B^{GR}$$

Bound violated completely in the Planck regime



CAUTION

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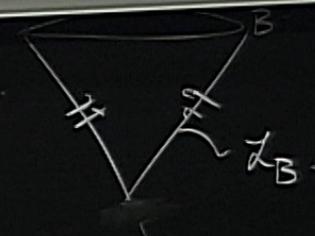
- Potential Tension bet UV & IR
- Entropy bounds
- Entropy current ✓
- G ✓
- sharp geometry ✓

Extreme Case  
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→ Result emerges just in the effective description automatically.

g<sup>eff</sup>  
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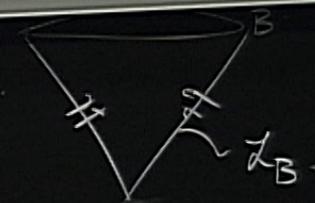
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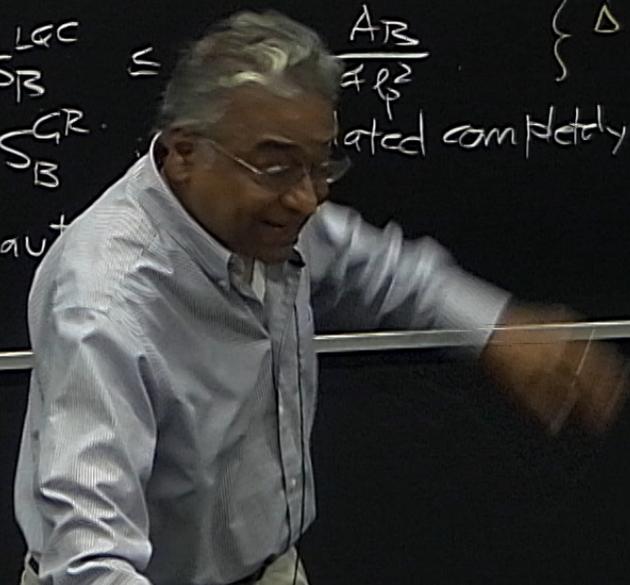
$$S_B^{GR}$$



fixed by BH entropy formula

related completely in the Planck regime

→ Result emerges just in the effective description



CAUTION

observations:

$$\Psi(a, \phi)$$

1. QFT on FLRW  $\rightarrow$  QFT on QFLRW geometries.

1. What is the new physics responsible for the bounce?
2. Might this new physics resolve the singularity for the perturbations as well as the background?
3. Does this new physics have any observational signature?
4. What general principles underlie the theory, beyond wanting to resolve the singularity?
5. Does a consistent picture for cosmology require that both the background and the perturbations are quantized?
6. Does the bounce or pre-bounce phase help in setting initial conditions?

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If  $\lambda_{\text{phys}} \ll R_{\text{curv}} = R_{\text{init}} \ln \mathcal{L}$

$$\partial_{\eta}^2 \chi_k + a^2(\eta) \left( \frac{k^2}{a^2(\eta)} - \frac{\ddot{R}(\eta)}{R(\eta)} \right) \chi_k = 0$$

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If  $\lambda_{\text{phys}} \ll R_{\text{curv}} = \text{finite in QFT}$

$$\partial_t^2 \chi_k + a^2(r) \left( \frac{k^2}{a^2(r)} - \frac{\ddot{a}(r)}{a(r)} \right) \chi_k = 0$$

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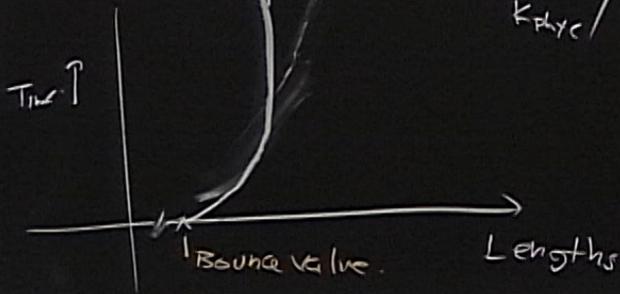
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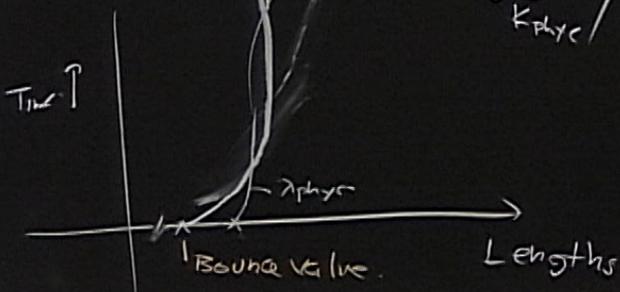
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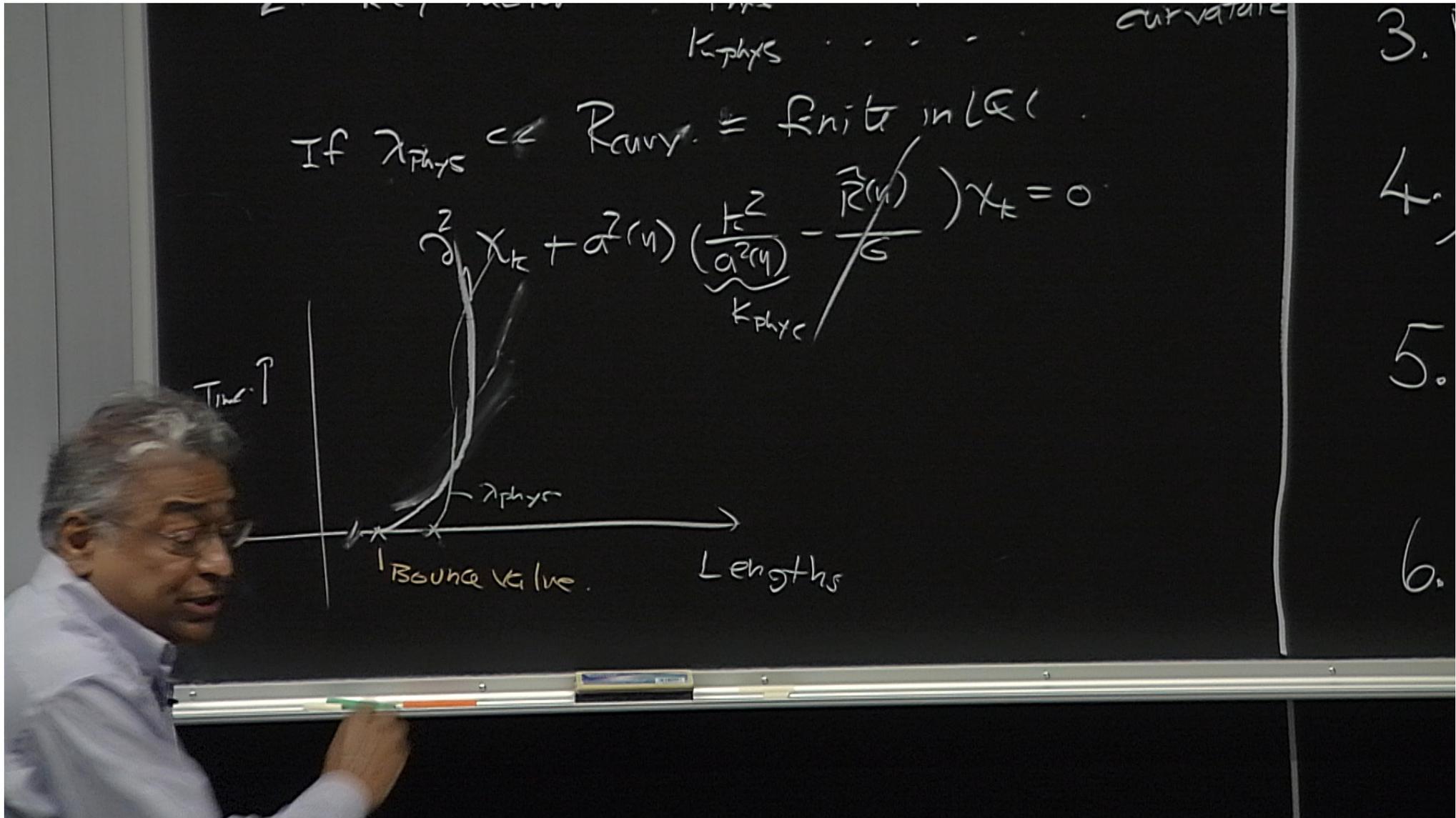
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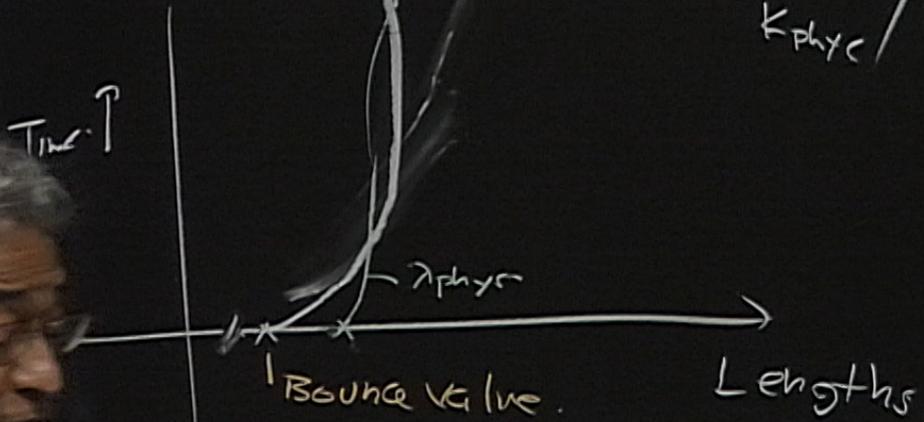


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If  $\lambda_{\text{phys}} \ll R_{\text{curv}} = \text{finite in } \mathbb{Q}(\dots)$

$$\frac{\partial^2}{\partial u^2} \chi_k + a^2(u) \left( \underbrace{\frac{k^2}{a^2(u)}}_{k_{\text{phys}}} - \frac{\bar{r}(u)}{G} \right) \chi_k = 0$$



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$\hat{G}_{ab}^{eff}$

same continuity Eq<sup>n</sup>

#### 4. Examples:

- Potential Tension bet
- Entropy bounds

- 1) Entropy current ✓
- 2) QG ✓
- 3) Sharp Geometry ✓

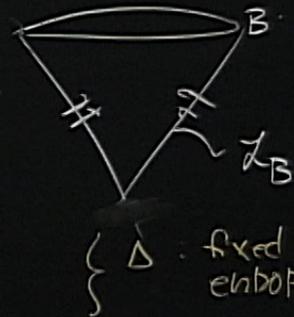
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