

Title: Simulation tools for neutrino experiments

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Abstract:



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Simulation tools for neutrino experiments

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Radiative Corrections at the Intensity Frontier of Particle Physics

June 13th, 2017

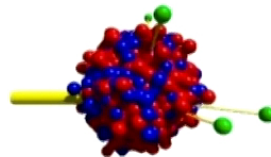
Overview

- The (accelerator beam) neutrino software stack
- What is GENIE?
 - How does GENIE work?
 - Recent developments
 - Future plans
- (Briefly) Other generators and tools
 - NEUT is the big omission - not an expert, so I will refrain from comment.
 - Also won't say much about Geant.



Geant 4

(won't say much about Geant here...)



GiBUU

The Giessen Boltzmann-Uehling-Uhlenbeck Project



First, some history...

- Common wisdom: The most natural theory community partner for this work is the nuclear theory community.
 - Some debate about this - there is a lot of room for HEP theory to contribute. But the NP theory community is hugely important and we've largely been walled off from each other - very different case than at the "Energy Frontier."
- HEP/NP separation (especially in the US) has meant that neutrino event generators in heavy use at experiments (especially GENIE and NEUT) were largely developed and maintained by experimentalists.
 - The inmates have control of the asylum...
 - More theory oriented generators came from Europe (even GENIE is UK via MINOS), and often lacked critical tools for use in experiments (flux, geometry, tools for estimating uncertainties, etc.).
 - The two generator "types" are growing towards each other...
- We are definitely seeking to remedy this situation.
 - There is a track record of success (e.g., A. Meyer et al and the z-expansion in GENIE), but the collaboration model probably needs some work in order to best serve all parties.
 - We need better mechanisms for giving credit to theorists who work with us.
 - Theorists should join generator groups.

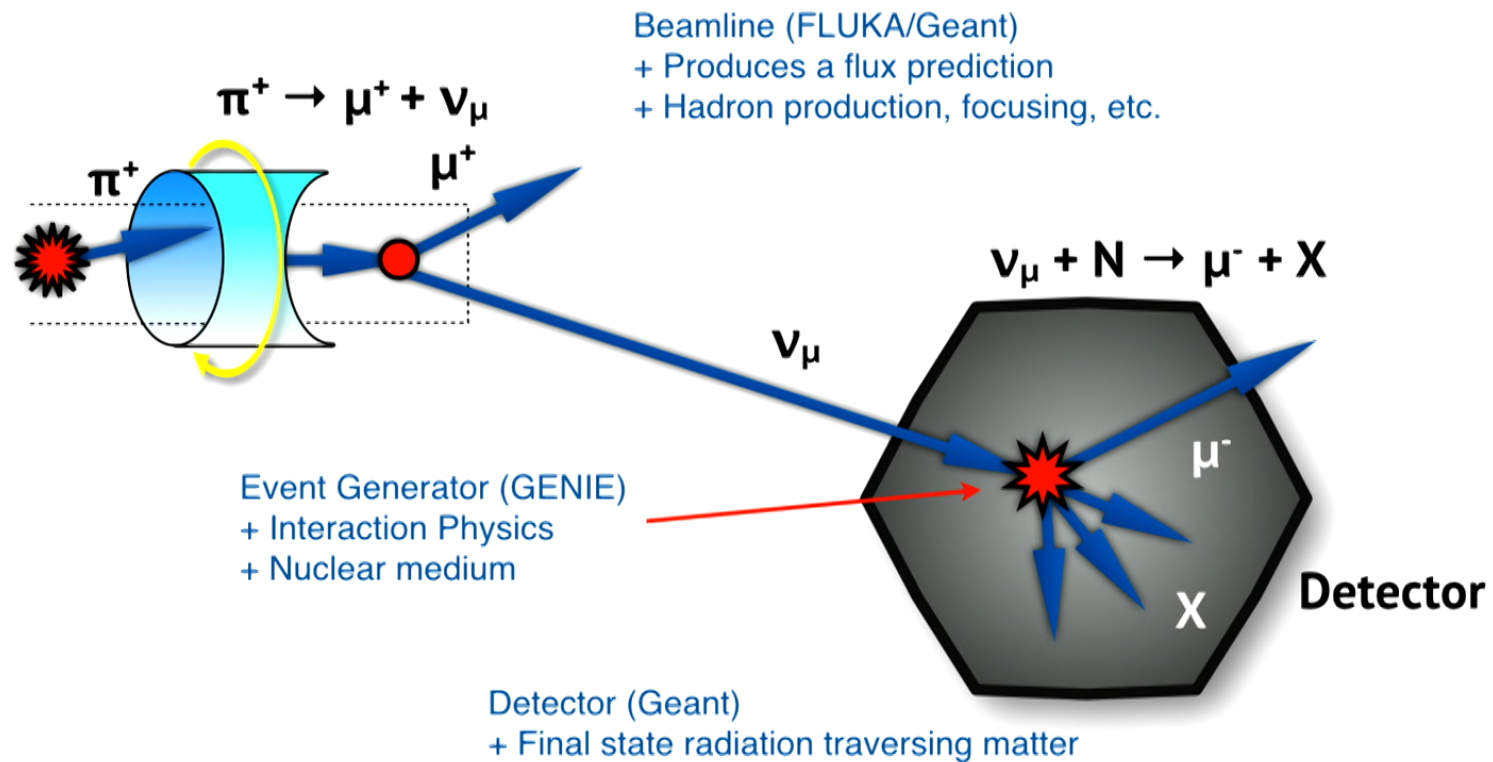
Perspectives*

- Theorists: The model doesn't need to match the data, it just needs to be correct.
- Experimentalist: The model doesn't need to be correct, it just needs to match the data.
 - *(Both camps are quite pleased with their positions.)*

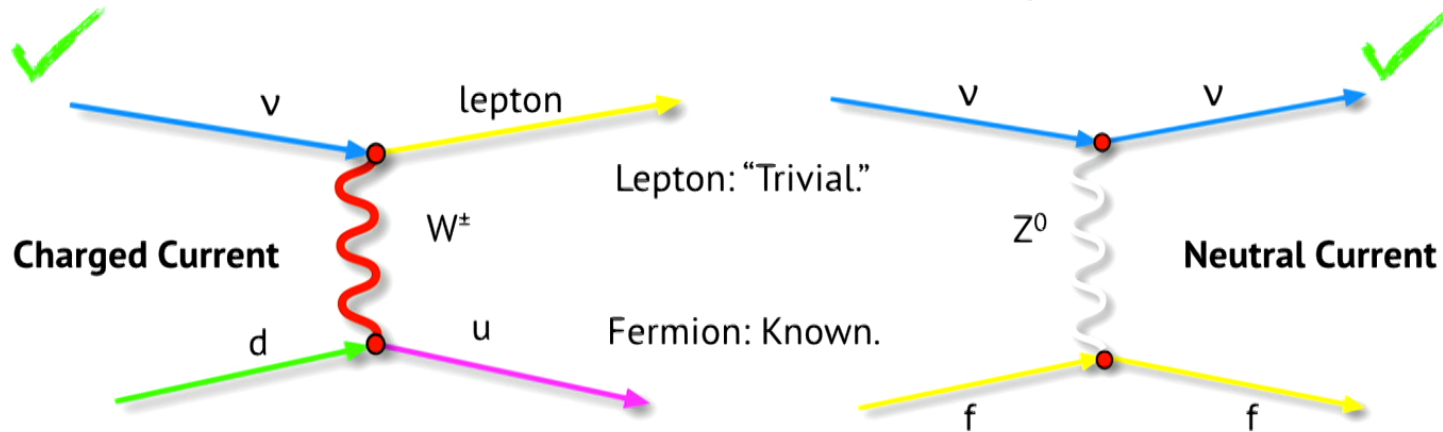
*Attributed to U. Mosel



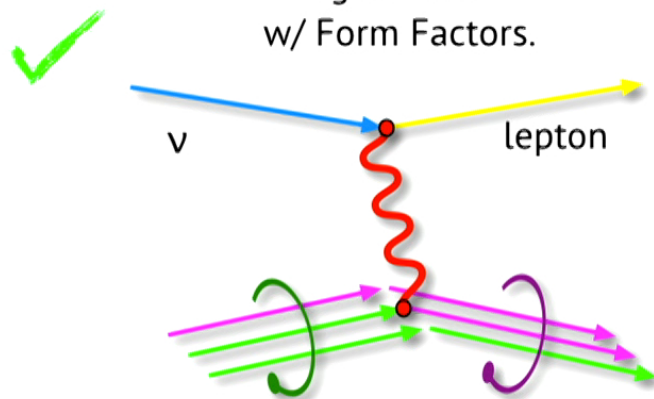
Neutrino Simulations: A Three-Part Software Stack



Neutrino Interactions - Weak Force Only!

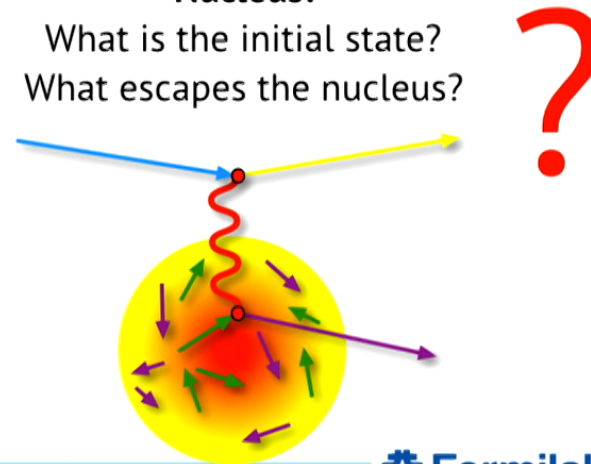


Free Nucleon: Parameterize ignorance w/ Form Factors.

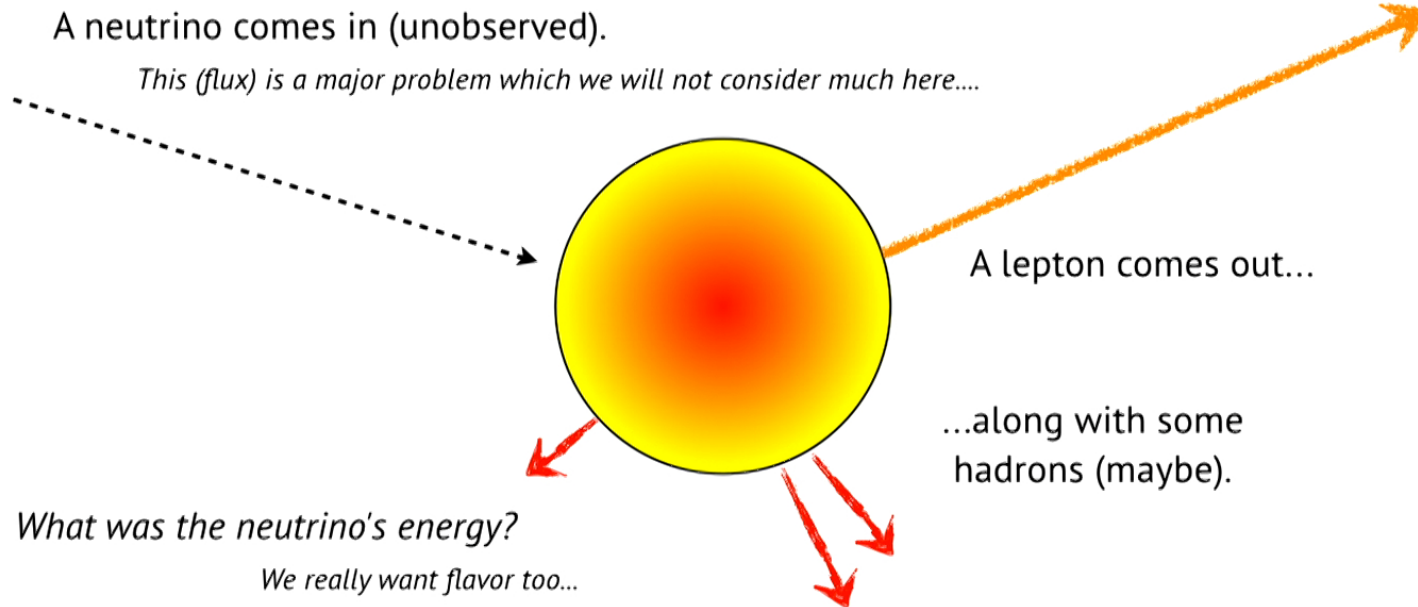


Nucleus:

What is the initial state?
What escapes the nucleus?

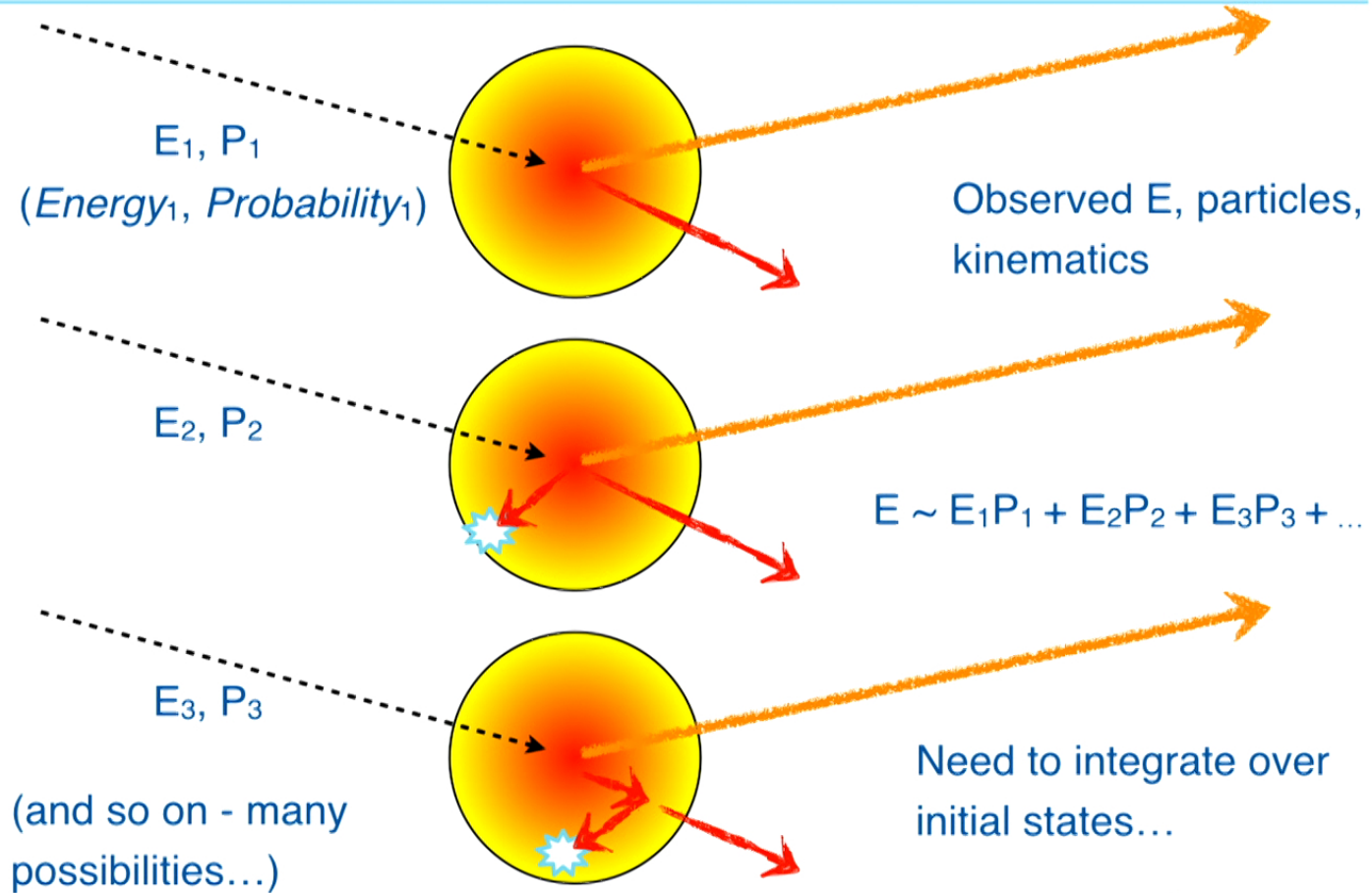


The Basic Problem



We have an unknown incoming energy and “missing” energy in the final state (neutral current reactions, neutrons in the final state, nuclear rescattering, etc.). We must infer the energy from incomplete final state information.

The Basic Problem: The Best We Can Do



The Basic Problem: The Best We Can Do

- The best we can do is build a map, weighted by probability, that provides all the possible initial states for an observed final state.
- With this map and a sample of events, we may infer a neutrino energy distribution (or some other kinematic distribution).
- How do we make any progress without an initial energy to begin with?
- For measurements, we use an *event generator* to predict backgrounds and the efficiency.
 - We may constrain the background prediction with data.
 - We must impose systematic uncertainties on our efficiency based on model estimates.
 - The more measurements we have, the better we may constrain these uncertainties and the better is our probability map.

`std::map<observed_topology, std::list<std::pair<probability, physics>>> = ?`

Neutrino MC Event Generators



- The generator must simulate all the types and momenta of every particle that appears in the final state.
- Some generators (MadGraph, Pythia, etc.) are computation aids for theorists, but GENIE is not.
- This is because we lack a theoretical framework that is both *complete* and *consistent*.
- The ideal input theory would be internally consistent and provide fully-differential cross sections in the kinematics of every final state particle over all reaction mechanisms, energies, and targets.
- Modern theory typically provides final state kinematics for the lepton only, and only over limited ranges in energy or momentum transfer, and may be fully exclusive or fully inclusive with no guidance on how to merge the regimes.
 - But the experiments must go on! So we must *stitch together* an ensemble that is consistent with all the data.

What else do neutrino event generators provide?



- Interfaces to geometry engines for modeling complex detectors.
- Flux drivers for computing exposure (atmospheric/solar sources) or normalizing responses to accelerator beams.
- Event re-weighting engines for studying systematic uncertainties and performing error propagation.
- Databases of electron, hadron, and neutrino scattering experiments with applications for comparing simulation and data.
 - Electron and hadron scattering event generator functionality.
- Nucleon decay generators.
- Libraries of pre-computed cross sections.

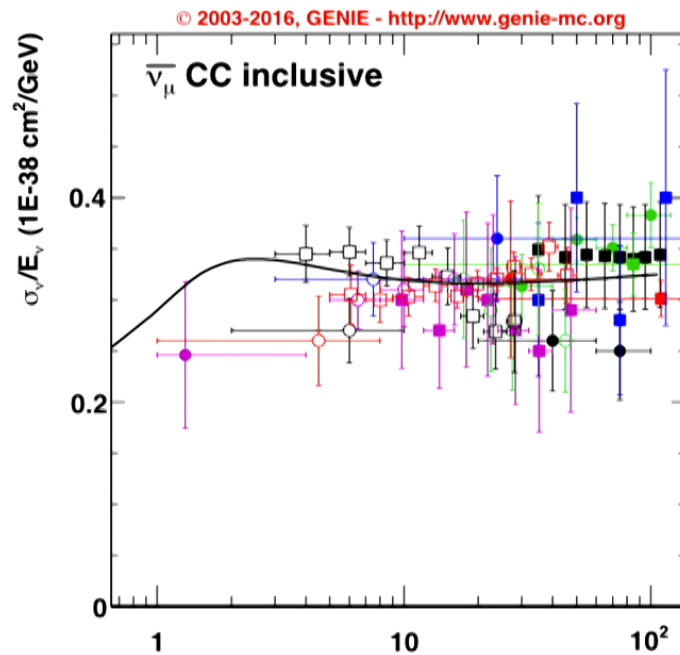
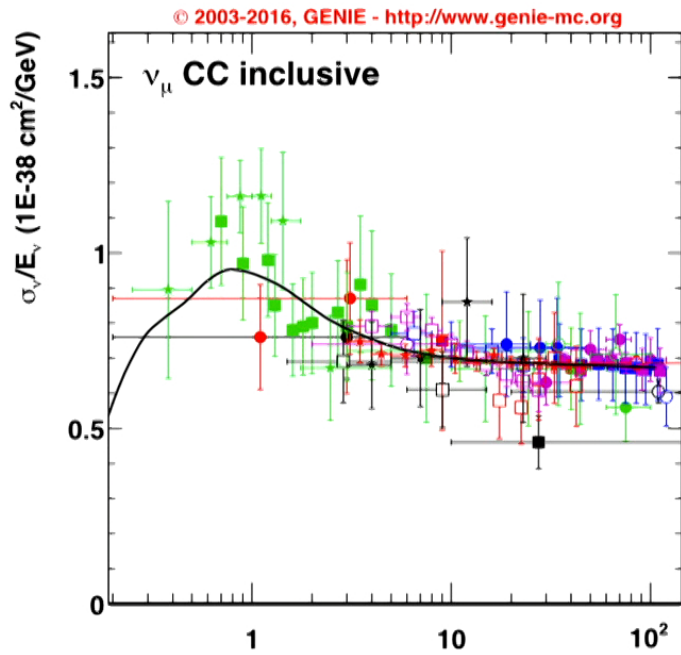
GENIE



- <https://genie.hepforge.org>
- The software:
 - Created to be a “universal event generator”.
 - Additionally run in electron and hadron scattering modes.
 - Many tools for studying systematics, comparison to data, etc.
 - Event handling is decoupled from physics routines, easy to create arbitrary algorithm stacks.
- The collaboration:
 - International collaboration with about a dozen collaborators (essentially all experimentalists) and many more contributors.
 - Collaborators do service work (validation, distribution, user support, developer support, etc.)
 - Contributors (many theorists) offer individual models or pieces of validation software, sometimes consulting, etc.

What is GENIE?

- We build a global physics model from a collection of exclusive state models (e.g., Llewellyn Smith QE, Rein-Sehgal resonant pion production, Bodek-Yang DIS, etc.).
 - (Many of these are *wrong but useful*.)
- When we add a new process (e.g., Nieves group MEC), we need to retune the total cross section by controlling the strength of the exclusive processes or subtracting processes.
- We try very hard to be consistent with data for the total cross section, so inclusive cross section calculations are very valuable as an additional constraint.
- We try to agree with a many other measured distributions as possible, but there are always tensions that are difficult to understand/reconcile.



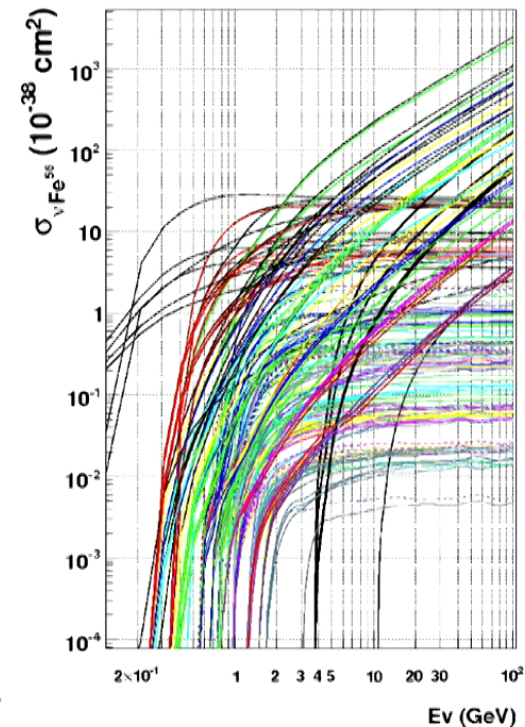
- | | | |
|--|--|----------------------|
| ● ANL_1FT.2 [Barish et al., Phys.Lett.B86:291 (1977)] | ○ FNAL_1SFT.1 [Kikugaki et al., Phys.Rev.Lett.49:99 (1982)] | E _ν (GeV) |
| ● ANL_1FT.4 [Barish et al., Phys.Rev.D19:2521 (1978)] | ○ FNAL_1SFT.2 [Baker et al., Phys.Rev.Lett.51:735 (1983)] | |
| ● BEBC.0 [Boselli et al., Phys.Lett.B70:273 (1977)] | ○ Gargamelle.0 [Eichten et al., Phys.Lett.B44:274 (1973)] | |
| ● BEBC.2 [Colby et al., Zeit.Phys.C2:187 (1978)] | ○ Gargamelle.10 [Ciancopollo et al., Phys.Lett.B84:381 (1979)] | |
| ● BEBC.3 [Boselli et al., Phys.Lett.B110:167 (1982)] | ○ Gargamelle.12 [Morfin et al., Phys.Lett.B104:235 (1981)] | |
| ● BEBC.6 [Parker et al., Nucl.Phys.B232:1 (1984)] | ○ IHEP_ITEP.0 [Asratyan et al., Phys.Lett.B76:239 (1978)] | |
| ● BNL_7FT.0 [Baltay et al., Phys.Rev.Lett.44:916 (1980)] | ○ IHEP_ITEP.2 [Yuzenko et al., Sov.J.Nucl.Phys.30:528 (1979)] | |
| ● BNL_7FT.4 [Baker et al., Phys.Rev.D25:817 (1982)] | ○ IHEP_JINR.0 [Anikeev et al., Zeit.Phys.C70:39 (1995)] | |
| ● CCFR.2 [Seligman et al., Nevis Report 292 (1996)] | ○ SKAT.0 [Baranov et al., Phys.Rev.B81:255 (1978)] | |
| ● CCFRR.0 [MacFarlane et al., Zeit.Phys.C26:1 (1984)] | ○ MINOS.0 [Adamson et al., Phys.Rev.D81:072002 (2010)] | |
| ○ CHARM.0 [Jonker et al., Phys.Lett.B99:265 (1981)] | ○ SciBusNE.0 [Nakajima et al., Phys.Rev.D63:G12001 (2001)] | |
| ○ CHARM.4 [Atyaby et al., Zeit.Phys.C38:403 (1988)] | ○ G00_00a | |

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|--|---|----------------------|
| ● BEBC.1 [Boselli et al., Phys.Lett.B70:273 (1977)] | ○ FNAL_1SFT.5 [Asratyan et al., Phys.Lett.B137:122 (1984)] | E _ν (GeV) |
| ● BEBC.3 [Colby et al., Zeit.Phys.C2:187 (1978)] | ○ Gargamelle.1 [Eichten et al., Phys.Lett.B44:274 (1973)] | |
| ● BEBC.6 [Boselli et al., Phys.Lett.B110:167 (1982)] | ○ Gargamelle.11 [Ertiquiza et al., Phys.Lett.B80:309 (1979)] | |
| ● BEBC.7 [Parker et al., Nucl.Phys.B232:1 (1984)] | ○ Gargamelle.13 [Morfin et al., Phys.Lett.B104:235 (1981)] | |
| ● BNL_7FT.1 [Favourakis et al., Phys.Rev.D21:562 (1980)] | ○ IHEP_ITEP.1 [Asratyan et al., Phys.Lett.B76:239 (1978)] | |
| ● CCFR.3 [Seligman et al., Nevis Report 292 (1996)] | ○ IHEP_ITEP.3 [Yuzenko et al., Sov.J.Nucl.Phys.30:528 (1979)] | |
| ● CHARM.1 [Jonker et al., Phys.Lett.B99:265 (1981)] | ○ IHEP_JINR.1 [Anikeev et al., Zeit.Phys.C70:39 (1995)] | |
| ● CHARM.5 [Atyaby et al., Zeit.Phys.C38:403 (1988)] | ○ MINOS.1 [Adamson et al., Phys.Rev.D81:072002 (2010)] | |
| ● FNAL_1SFT.4 [Taylor et al., Phys.Rev.Lett.51:739 (1983)] | ○ G00_00a | |

How does GENIE work?

- The first step is to compute the total cross section for the input energy, flavor, helicity, and target isotope.
- Perform a sum over exclusive channels (square then sum, sigh).
- Numerical integration of the corresponding differential cross section expression:
 - Computationally intensive procedure (100's of millions of differential cross section evaluations), but only needs to be run once per release.

$\nu_\mu, \bar{\nu}_\mu + Fe$, all processes



<https://www.hepforge.org/archive/genie/data/>



The default model:

```
<param_set name="Default">
  <param type="int" name="NGenerators"> 13 </param>
  <param type="alg" name="Generator-0"> genie::EventGenerator/QEL-CC </param>
  <param type="alg" name="Generator-1"> genie::EventGenerator/QEL-NC </param>
  <param type="alg" name="Generator-2"> genie::EventGenerator/RES-CC </param>
  <param type="alg" name="Generator-3"> genie::EventGenerator/RES-NC </param>
  <param type="alg" name="Generator-4"> genie::EventGenerator/DIS-CC </param>
  <param type="alg" name="Generator-5"> genie::EventGenerator/DIS-NC </param>
  <param type="alg" name="Generator-6"> genie::EventGenerator/COH-CC </param>
  <param type="alg" name="Generator-7"> genie::EventGenerator/COH-NC </param>
  <param type="alg" name="Generator-8"> genie::EventGenerator/DIS-CC-CHARM </param>
  <param type="alg" name="Generator-9"> genie::EventGenerator/QEL-CC-CHARM </param>
  <param type="alg" name="Generator-10"> genie::EventGenerator/NUE-EL </param>
  <param type="alg" name="Generator-11"> genie::EventGenerator/IMD </param>
  <param type="alg" name="Generator-12"> genie::EventGenerator/IMD-ANH </param>
</param_set>
```

Interesting additions / alternatives:



```
<param_set name="DFR">
  <param type="int" name="NGenerators"> 2 </param>
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</param_set>
```

How does GENIE work?

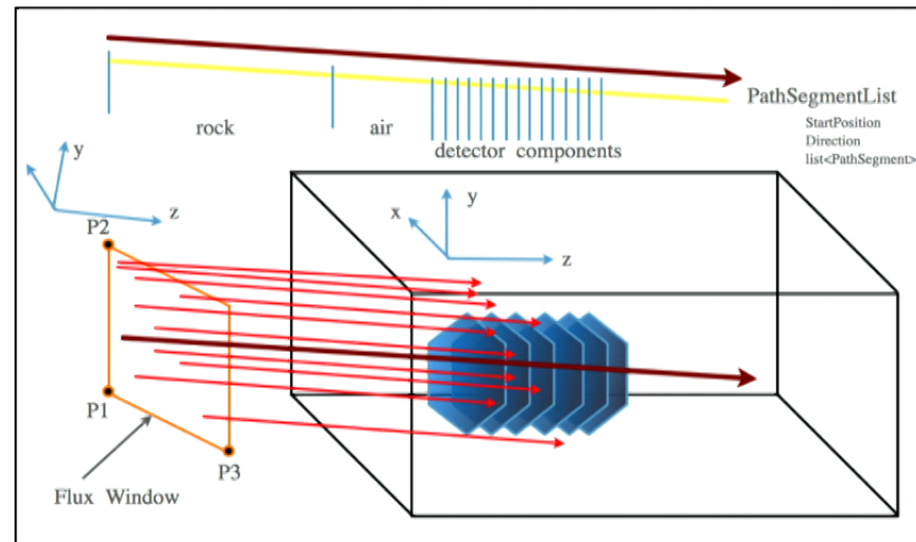
- With the total cross sections in hand, event generation proceeds by projecting rays through the detector geometry and computing the total path length of all the materials along a trajectory.
- At the start of a run, we find the longest path length through the detector and normalize the interaction probability to 1 on that path, scaling the interaction probabilities appropriately, and incorporating this information into the flux driver.
 - Necessary to keep running times reasonable.
- Then for any given path, events are chosen randomly by channel according to their contribution to the total cross section in an accept-reject loop.

Why does GENIE need Geometry?

- Real fluxes and geometries are never uniform.
 - Experiments need to generate interaction vertices in the correct locations.
 - Fluxes vary in intensity and energy profile across the detector.
 - Detector structures (and the surrounding area!) have specific structures and boundaries.



"... so complicated!"



Fluxes

- Many choices (including making your own):
 - User-specified histograms (no spatial variation, only energy and flavor)
 - Encapsulations of common parameterizations (e.g., atmospheric)
 - Simple, generic ntuple format (`GSimpleNtpFlux`*`)
 - Experiment (NuMI, T2K) or institution specific.
- Wrap any of the above in a "flavor blender" adapter (`GFlavorMixerI``) - this is how you handle far detectors in an oscillation experiment.
- Some drivers have exposure counters (e.g., time, protons on target).

*FNAL beamlines committed to migrating to this common ntuple format (dk2nu).



How does GENIE work?

- Currently implemented GENIE physic models rely heavily on a factorization assumption.
- Some cases blend boxes together a bit (but for the most part they do not).



"Is that safe?"

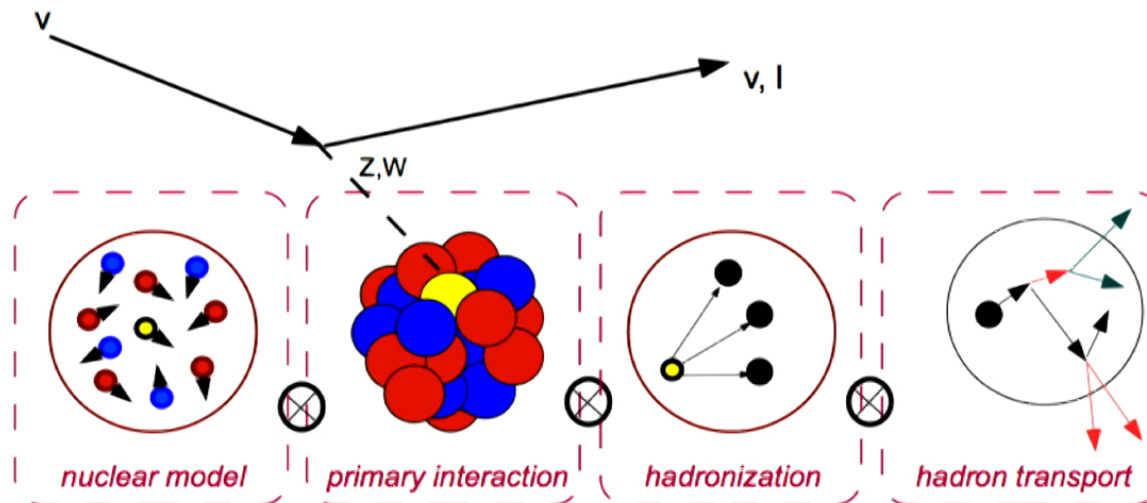


Figure by C. Andreopoulos



GENIE Physics Models



- GENIE 2.0 (~2007) used identical physics models as NEUGEN, a Fortran generator that was developed over a number of years by a succession of physicists, and used by MINOS. GENIE has evolved with each subsequent release.
- There are currently dozens of different physics models.
- The default nuclear model is the relativistic Fermi gas with Bodek and Ritchie high-momentum tails. GENIE also implements the Effective Spectral Function, and the Local Fermi Gas. Other spectral function implementations exist in development branches and need a bit more effort to become public.
- The quasielastic process defaults to Llewellyn-Smith, but we also have the Nieves et al model. The axial form factor model is the dipole but we offer (and are preparing to default to) the z-expansion model as well.
- Excitation of nucleon resonances (decaying by meson emission) and coherent pion production are both described by models by Rein and Sehgal, but we offer a number of alternatives (Berger and Sehgal, different form factor models, etc.).
 - We also offer a diffractive pion production model (Rein).
- Models for neutrino-electron scattering and inverse muon decay are included and mostly complete (additional radiative corrections required for neutrino-electron scattering).

GENIE Physics Models



- We offer (non-default) a custom built and the Valencia 2p2h models.
- Bodek and Yang (2003) is used for nonresonant inelastic scattering.
- Other interesting exclusive states (QEL hyperon production, single Kaon production, etc.) are optional (making them default would lead to double counting in the hadronization model).
- The custom "AGKY" hadronization model, developed internally, covers the transition between PYTHIA at high ($W > 3\text{GeV}/c^2$) invariant masses and an empirical model based on KNO-scaling at lower invariant masses.
- GENIE has two* internally developed models for final-state interactions; one is a cascade model and the other (the default) parameterizes the cascade a single effective interaction for easy re-weighting.
 - Actually many more than two - we are snap-shotting major changes with dated timestamps as we make improvements. Users can choose from our long-standing default and the bleeding edge, with a variety of options in between.
- GENIE uses the SKAT parametrization of formation zones (the effective distance over which a quark hadronizes).
- More detail in the back-ups...

Pieces (Usually)

- Vertex selection
 - Simple nuclear density model
 - Initial state nuclear model
 - Removal energy and momentum
 - RFG with Bodek-Ritchie tails.
 - New: Local Fermi Gas
 - New: Effective Spectral Function
 - Almost there: "Benhar" spectral function
 - Just started: Correlated Fermi Gas (MIT)
 - Hard scattering process
 - Differential cross section formula to get event kinematics (x, y, Q², W, t, etc.)
 - Lepton kinematics
 - Hadronic system
 - Propagation/transport (default is an "effective cascade")
 - Fast and re-weightable
- GROUND STATE
- INITIAL STATE
- FINAL STATE

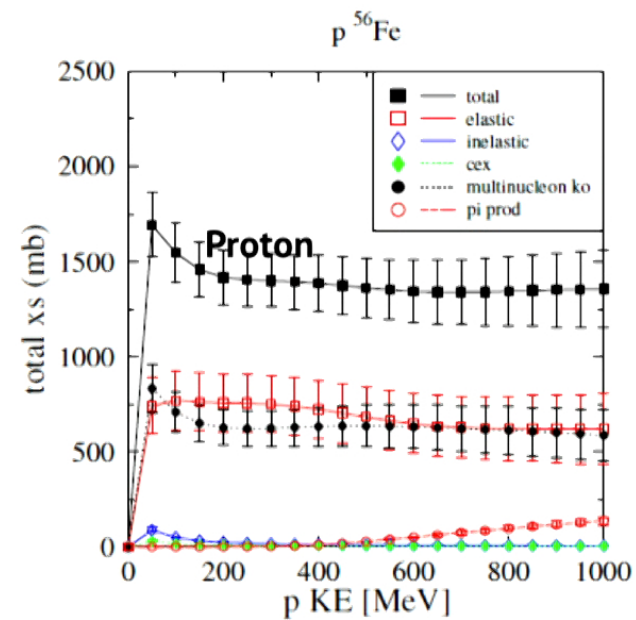
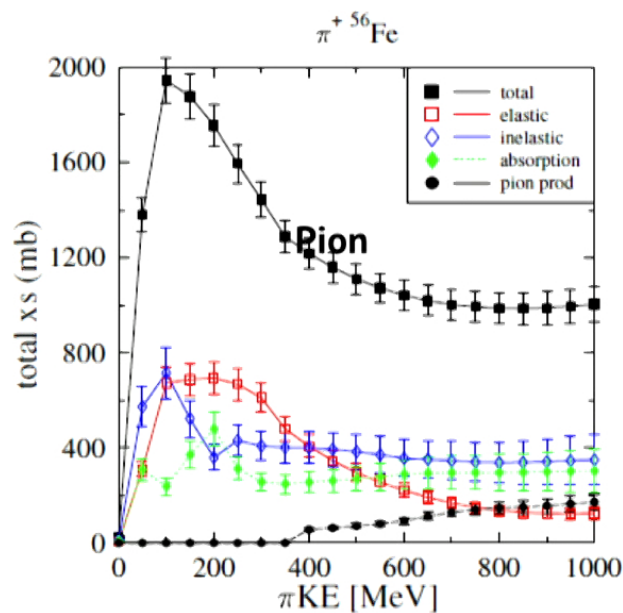


Usual Pieces

- Decays before and after propagation
 - Remnant decay
 - Just started caring about this, really...
 - Current model is very simple
 - Working on adopting other codes (Geant4, INCL++, possibly GiBUU) to handle clustering, de-excitation, evaporation
 - May be a bridge to more sophisticated transport codes
- REMNANT STATE
- Sometimes models can't work this way - e.g., discovering we can't separate lepton and hadron kinematics into separate modules for QE events (can't compute cross section in Q2 and then compute lepton and hadron kinematics, need to flip the procedure and then accept-reject based on Q2), etc.
 - (Actually, we should do all events this way - but the code runs much slower and so we're working on ways to make that process fast enough to be more widely used.)

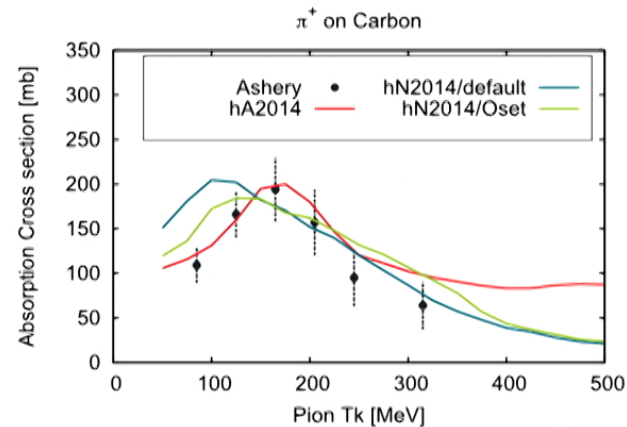
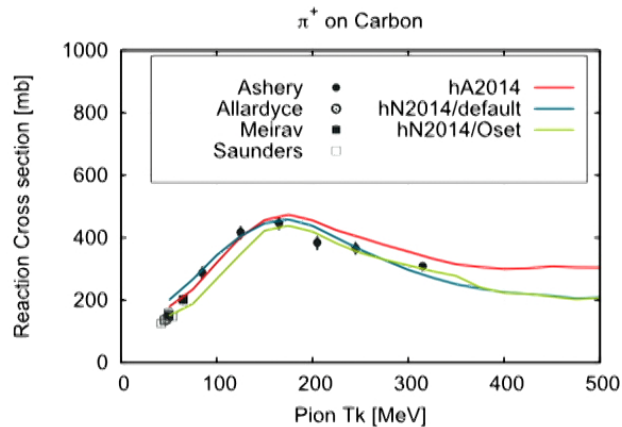
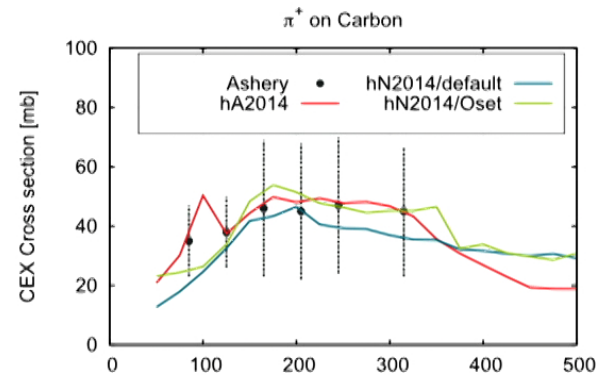
FSI Models

- GENIE: "hA" (default) - use iron reaction cross section data, isospin symmetry, and $A^{2/3}$ scaling to predict the FSI reaction rates.
- Individual particle energies and angles use data templates or sample from the allowed phase space.



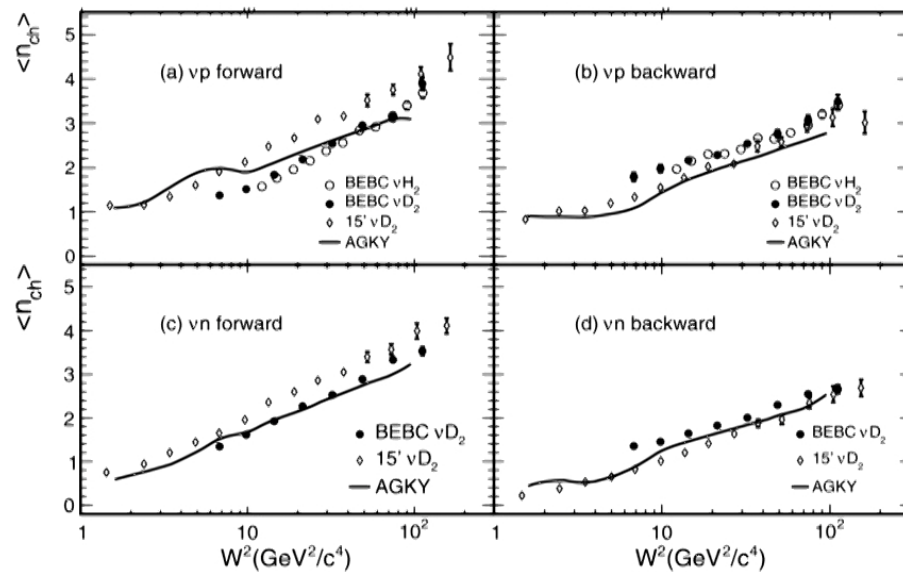
FSI Models

- GENIE "hN" is our cascade model.
- New to hN are: Oset et al, Nucl. Phys. A468 (1987), Oset et al, Nucl. Phys. A484 (1998)
- Model describes low energy (kinetic E around Delta peak, 85 MeV - 350 MeV) pion interactions inside nuclear matter.
 - Nuclear effects are implemented as modifications of the Delta width.
- Introduced here as a modification of the GENIE cascade model (hN). Modifications not yet filtered down into the parameterized (hA, default) model.



Modeling Nuclear Effects

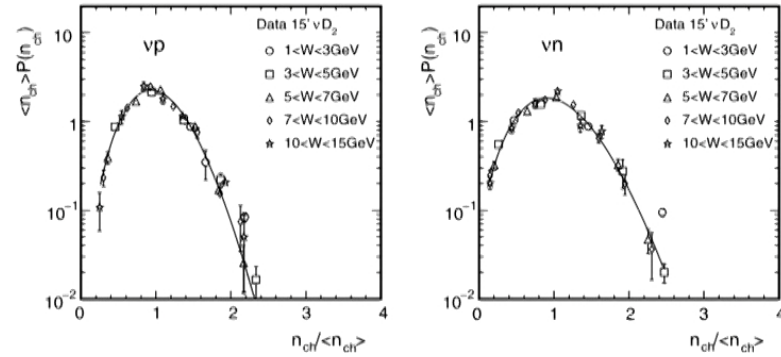
- What about hadronization in the nuclear medium?
- We use Pythia (currently version 6, migration to 8 is on-going).
- GENIE does reasonably well, but the validation uses deuterium or hydrogen - little influence from nuclear effects.



T. Yang et al, Eur. Phys. J C (2009) 63:1-10

AGKY Hadronization

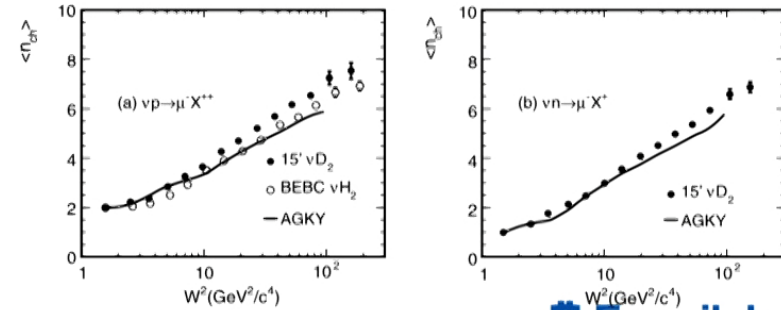
Fig. 1 KNO scaling distributions for νp (left) and νn interactions. The curve represents a fit to the Levy function. Data points are taken from [7]



The AGKY model, which is now the default hadronization model in the neutrino Monte Carlo generators NEUGEN [9] and GENIE-2.0.0 [10], includes a phenomenological description of the low invariant mass region based on Koba–Nielsen–Olesen (KNO) scaling [11], while at higher masses it gradually switches over to the PYTHIA/JETSET model. The transition from the KNO-based model to the PYTHIA/JETSET model takes place gradually, at an intermediate invariant mass region, ensuring the continuity of all simulated observables as a function of the invariant mass. This is accomplished by using a transition window $[W_{\min}^{\text{tr}}, W_{\max}^{\text{tr}}]$ over which we linearly increase the fraction of neutrino events for which the hadronization is performed by the PYTHIA/JETSET model from 0% at W_{\min}^{tr} to 100% at W_{\max}^{tr} . The default values used in the AGKY model are

$$W_{\min}^{\text{tr}} = 2.3 \text{ GeV}/c^2, \quad W_{\max}^{\text{tr}} = 3.0 \text{ GeV}/c^2.$$

Fig. 3 Average charged-hadron multiplicity (n_{ch}) as a function of W^2 . (a) νp events. (b) νn events. Data points are taken from [7, 20]



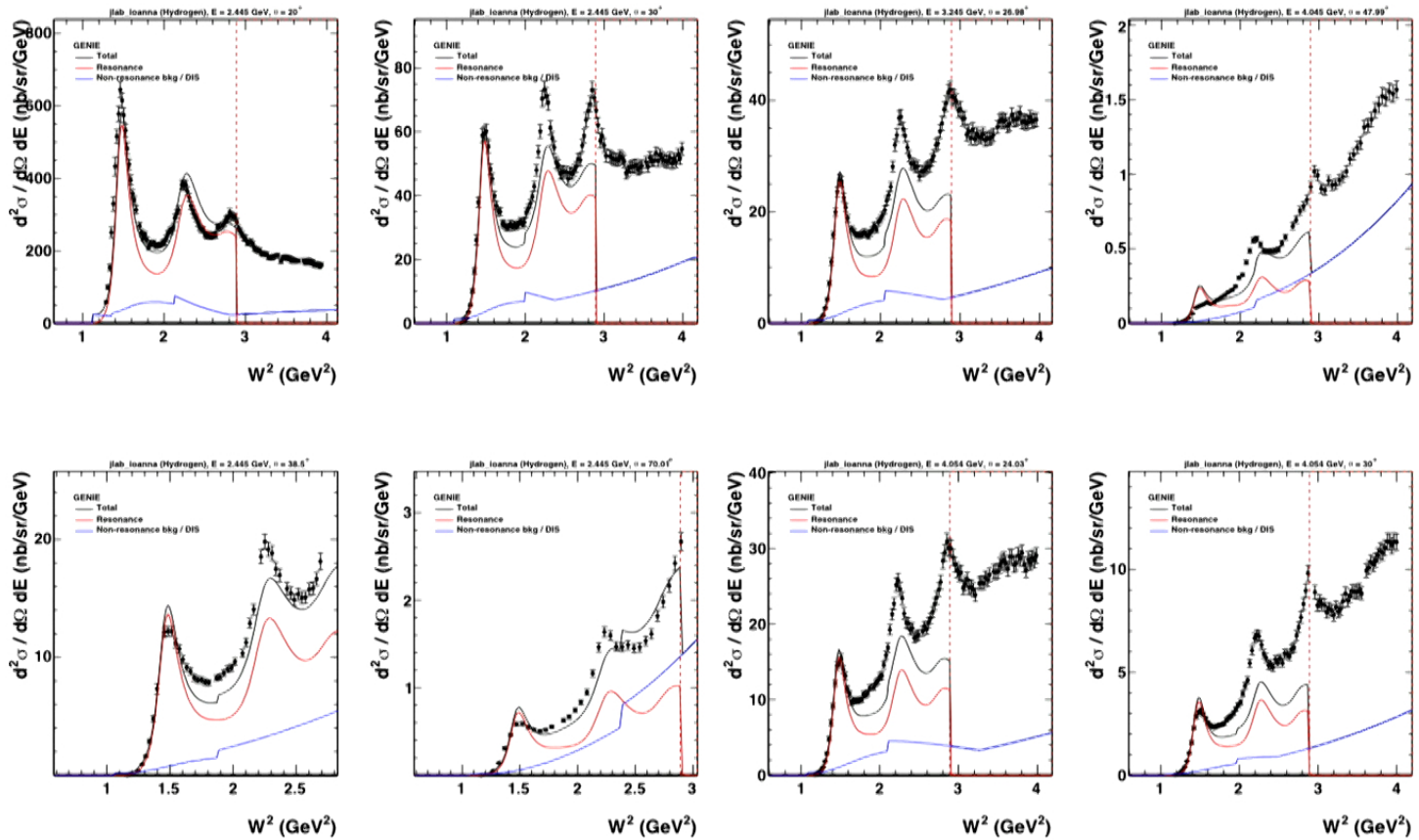
T. Yang et al, Eur. Phys. J C (2009) 63:1-10



Electrons



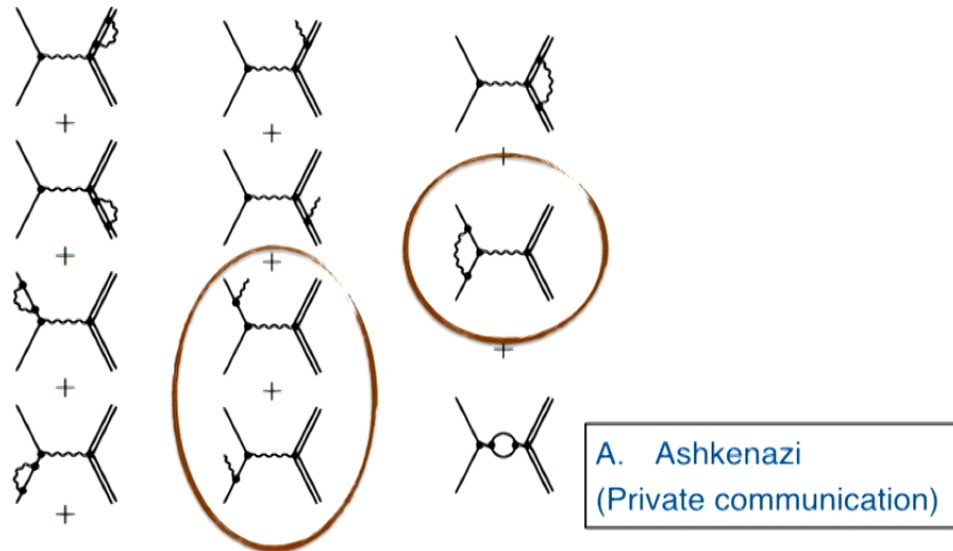
- Some distributions look good. others are more challenging.



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Radiative corrections in GENIE

- To first order, we don't do them.
- Some on-going work (O. Hen's group at MIT) to include some basic effects.
 - Corrections for nuclear effects are not part of the current effort.



initial state radiation

For the ISR, relevant only for the electron scattering case.

We aim to subtract energy and momentum from the initial state probe according to a given distribution.

final state radiation

Final state radiation should be applicable for both electron and neutrino mode.

We aim to add a final state photon and keep record of the decay. We think of the possibility of doing this process via GEANT and not GENIE.

internal

We're currently looking at two approaches:

1. Based on PHYSICAL REVIEW C, VOLUME 62, 025501
The energy loss, ΔE , can be sampled using the distribution:

$$I_{int}(E, \Delta E, a) = \frac{a}{\Delta E} \left(\frac{\Delta E}{E}\right)^a$$

where E is the incoming electron energy

$$a = \frac{\alpha_{EM}}{\pi} \left[\ln\left(\frac{Q^2}{m^2}\right) - 1 \right] \quad \text{good approximation using } Q^2 \text{ from elastic scattering}$$

For Nuclei a should be multiply by Z

2. Based on L. W. Mo and Y. S. Tsai, SLAC-PUB-380.

$$I(E, \Delta E, t) = \frac{1}{E} \frac{\ln\left(\frac{E}{E-\Delta E}\right)^{\frac{t}{\ln 2} - 1}}{\Gamma\left(\frac{t}{\ln 2}\right)}$$

where t is target thickness in radiation lengths

A. Ashkenazi
(Private communication)

New Releases: GENIE 3.0

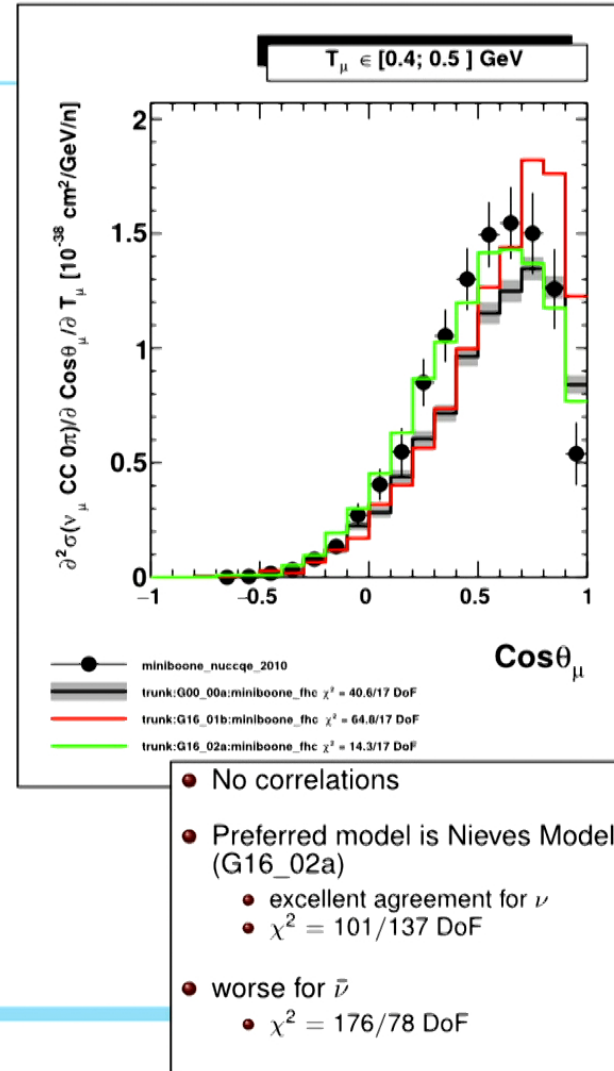
- The basic idea behind GENIE 3.0 is to vastly improve the configuration mechanisms to make it easier to change between model “constellations” and provide a mechanism for storing new constellations.
 - It should be a command line option to switch,
 - e.g., `--constellation minerva2017``, or
 - e.g., `--constellation best_theory``
- Of course, there will be *some* default and we will likely do some tuning/updates to improve it.
 - We are currently testing four basic variations.



GENIE 3.0 - versions

- Default - G00_00a
 - No MEC
 - CCQE process is LwlynSmith Model
 - Dipole Axial Form Factor - Depending on $M_A = 0.99 \text{ GeV}$
 - Nuclear model: Fermi Gas Model - Bodek, Ritchie
- Default + MEC - G16_01b
 - with **Empirical MEC**
 - CCQE process is LwlynSmith Model
 - Dipole Axial Form Factor - Depending on $M_A = 0.99 \text{ GeV}$
 - Nuclear model: Fermi Gas Model - Bodek, Ritchie
- Nieves, Simo, Vacas Model - G16_02a
 - **Theory motivated MEC**
 - CCQE process is Nieves
 - Dipole Axial Form Factor - Depending on $M_A = 0.99 \text{ GeV}$
 - Nuclear model: Local Fermi Gas Model
- G17_02a (not presented in this talk) - G17_02a
 - with Z-Expansion for Axial form factor
 - Get rid of M_A

Marco Roda - mroda@liverpool.ac.uk
on behalf of GENIE collaboration



GENIE Comparisons

The GENIE suite contains a package devoted to comparing GENIE predictions against publicly released datasets.

- Crucial technology for **new GENIE global fit** to neutrino scattering data
- Provides the opportunity to improve and develop GENIE models
- All sorts of data
 - **Modern Neutrino Cross Section measurement**
 - nuclear targets
 - typically flux-integrated differential cross-sections
 - MiniBooNE, T2K, MINERvA
 - **Historical Neutrino Cross Section Measurement**
 - Bubble chamber experiment
 - Measurements of neutrino-induced **hadronic system characteristics**

Marco Roda - mroda@liverpool.ac.uk
on behalf of GENIE collaboration



GENIE 4.0 - GENIE + Professor



<https://professor.hepforge.org>

Current authors:

- Andy Buckley (Glasgow)
- Holger Schulz (IPPP)

Former members:

- Hendrik Hoeth
- Heiko Lacker
- Jan Eike von Seggern
- Daniel Weyh
- Simone Amoroso

Professor is a **tuning tool for Monte Carlo event generators**, based on the ideas described in "Tuning and Test of Fragmentation Models Based on Identified Particles and Precision Event Shape Data" (Z. Phys., C73 (1996) 11-60).

Professor has been successfully used to produce **most of the established "tunings"** of the general purpose MC event generators.

A collaboration between Professor and GENIE authors to produce a **Professor/GENIE interface** and **Professor-based GENIE tunes** was supported by Inst. of Particle Physics Phenomenology via an IPPP Associateship Award.
→ **Active ongoing work!**

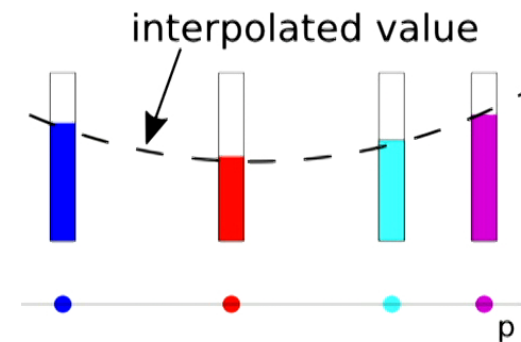
Energy Frontier

Slide by C. Andreopoulos



GENIE + Professor

- <http://professor.hepforge.org>
- Numerical assistant
- Developed for ATLAS experiment
- $I(p)$ used instead of a full MC
 - ① MC runs subset of param space
 - ② sample bin's behaviour
 - ③ Parametrization $I(p)$
 - Polynomial interpolation
 - Repeat for each bin
- a parameterization $I_j(p)$ for each bin
- Minimize according to $\vec{I}(p)$
- ~ 15 parameters

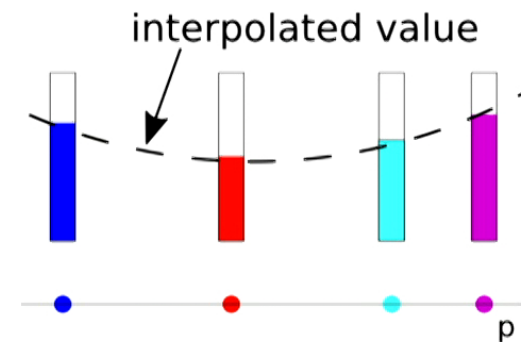


Marco Roda - mroda@liverpool.ac.uk
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GENIE + Professor

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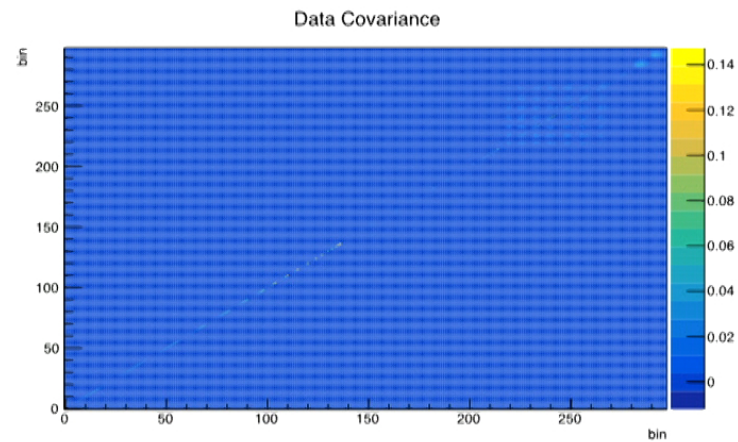
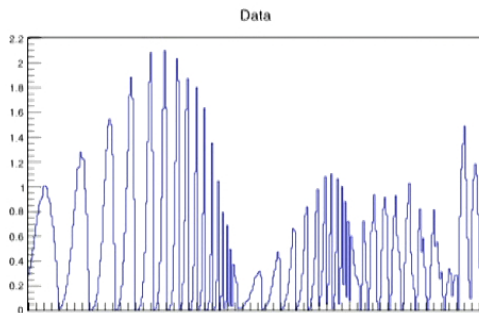


GENIE + Professor

- Highly parallelizable
 - independent from the minimization
- All kind of parameters can be tuned
 - Not only reweight-able
- Advanced system
 - Take into account correlations
 - weights specific for each bin and/or dataset
 - Proper treatment while handling multiple datasets
 - Restrict the fit to particular subsets
 - Nuisance parameters can be inserted
 - proper treatment for datasets without correlations (MiniB)
- Reliable minimization algorithm
 - based on Minuit

Marco Roda - mroda@liverpool.ac.uk
on behalf of GENIE collaboration

- MiniBooNE ν_μ CCQE
 - 2D histogram
 - 137 points
- MiniBooNE $\bar{\nu}_\mu$ CCQE
 - 2D histogram
 - 78 points
- T2K ND280 0π (2015)
 - irregular 2D histogram
 - 67 points
- MINERvA ν_μ CCQE
 - 1D histogram
 - 8 points
- MINERvA $\bar{\nu}_\mu$ CCQE
 - 1D histogram
 - 8 points

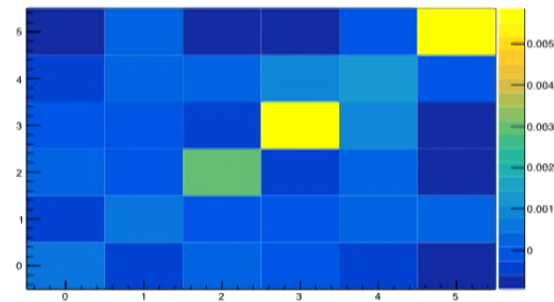


GENIE + Professor

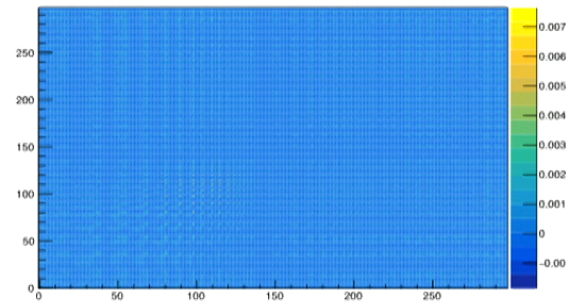
Marco Roda - mroda@liverpool.ac.uk
on behalf of GENIE collaboration

- Default + Empirical MEC
- G16_01b in the new naming scheme
- Parameters:
 - QEL- $M_A \in [0.7; 1.8]$ GeV - Default value is 0.99 GeV
 - QEL-CC-XSecScale $\in [0.8; 1.2]$ - Default value is 1
 - RES-CC-XSecScale $\in [0.5; 1.5]$ - Default value is 1
 - MEC-FracCCQE $\in [0; 1]$ - Default value is 0.45
 - FSI-PionMFP-Scale $\in [0.6; 1.4]$ - Default value is 1
 - FSI-PionAbs-Scale $\in [0.4; 1.6]$ - Default value is 1
- Parameters best fit
 - 0 M_A
 - 1 QEL-CC-XSecScale
 - 2 RES-CC-XSecScale
 - 3 MEC-FracCCQE
 - 4 FSI-PionMFP-Scale
 - 5 FSI-PionAbs-Scale
- Prediction covariance
 - due to the propagation of the param. covariance
 - So far not used
 - Tool to propagate systematics parameters

Parameter Covariance



Prediction Covariance



Datasets were fitted separately (and together)

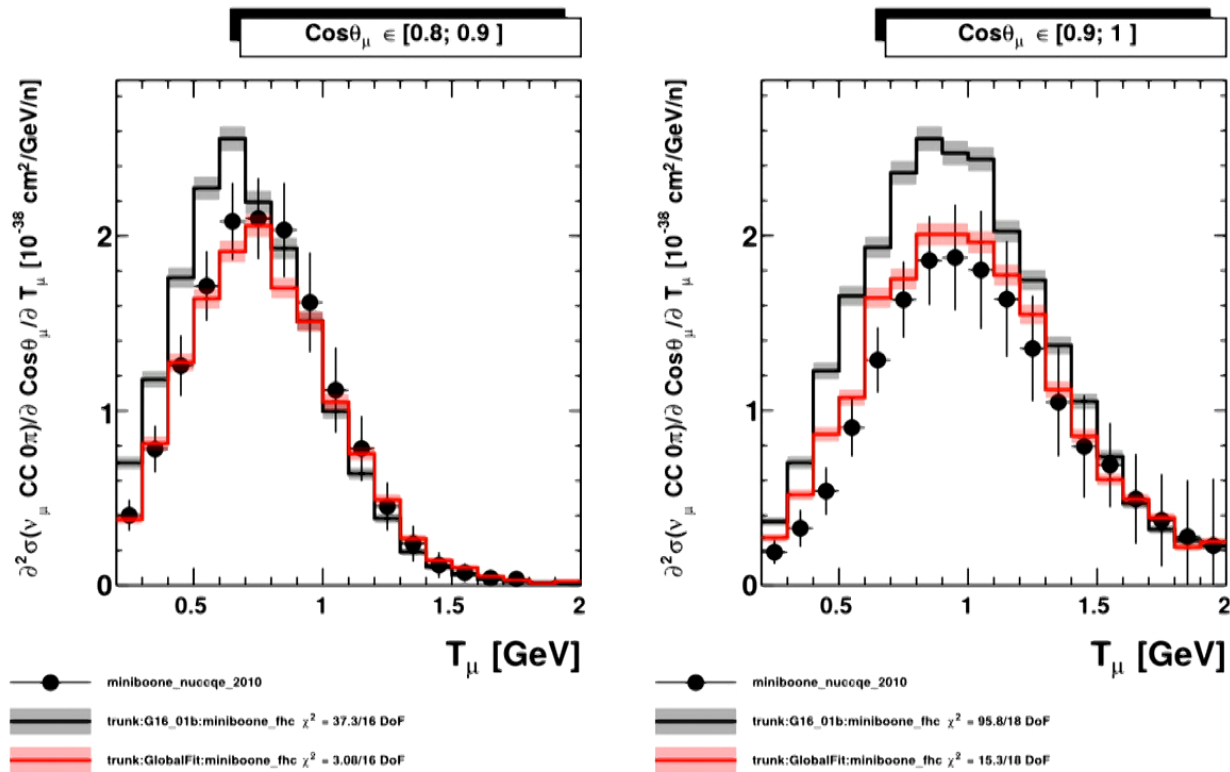
Parameter	Neutrino fit	Anti-neutrino fit	Global fit
M_A (GeV/c ²)	1.17 ± 0.02	1.26 ± 0.03	1.21 ± 0.02
QEL-CC-XSecScale	0.93 ± 0.01	0.97 ± 0.02	0.95 ± 0.02
RES-CC-XSecScale	0.86 ± 0.05	0.98 ± 0.09	1.02 ± 0.05
MEC-FracCCQE	0.85 ± 0.03	0.7 ± 0.1	0.53 ± 0.08
FSI-PionMFP-Scale	0.87 ± 0.02	1.39 ± 0.03	0.75 ± 0.04
FSI-PionAbs-Scale	1.51 ± 0.03	0.7 ± 0.1	0.87 ± 0.07

Fit Results	Neutrino fit	Anti-neutrino fit	Global fit	Nominal Values
Miniboone ν_μ χ^2	152 / 137	171 / 137	138 / 137	441 / 137
MiniBooNE $\bar{\nu}_\mu$ χ^2	60 / 78	32.4 / 78	36.2 / 78	50.4 / 78
T2K χ^2	237 / 67	276 / 67	252 / 67	135 / 67
MINERvA ν_μ χ^2	6.11 / 8	8.07 / 8	7.79 / 8	17.5 / 8
MINERvA $\bar{\nu}_\mu$ χ^2	8.19 / 8	11.5 / 8	5.7 / 8	6.23 / 8
Global dataset χ^2	463 / 292	499 / 292	440 / 292	650 / 298

- M_A and cross section scale factors are in good agreement
- FSI parameters are not
- The agreement with data is reasonable
 - Better than original model

- Default + MEC - G16_01b
 - with Empirical MEC
 - CCQE process is LwlynSmith Model
 - Dipole Axial Form Factor - Depending on $M_A = 0.99$ GeV
 - Nuclear model: Fermi Gas Model - Bodek, Ritchie

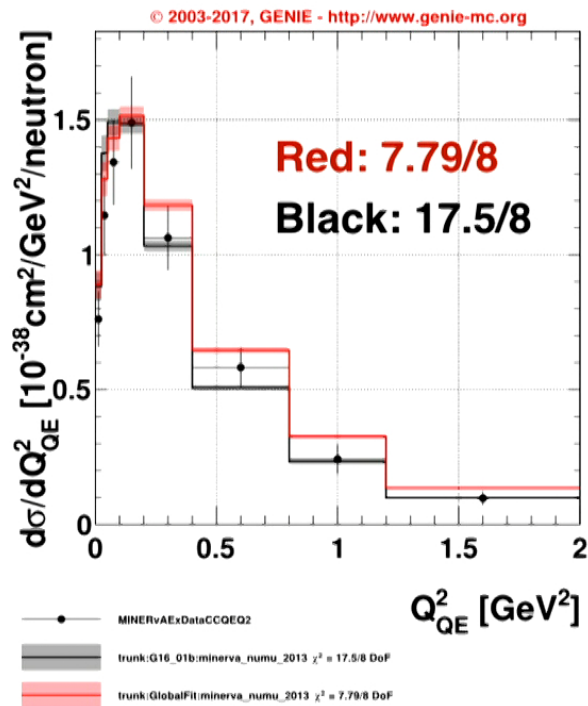
Best fit - MiniBooNE ν_μ CCQE



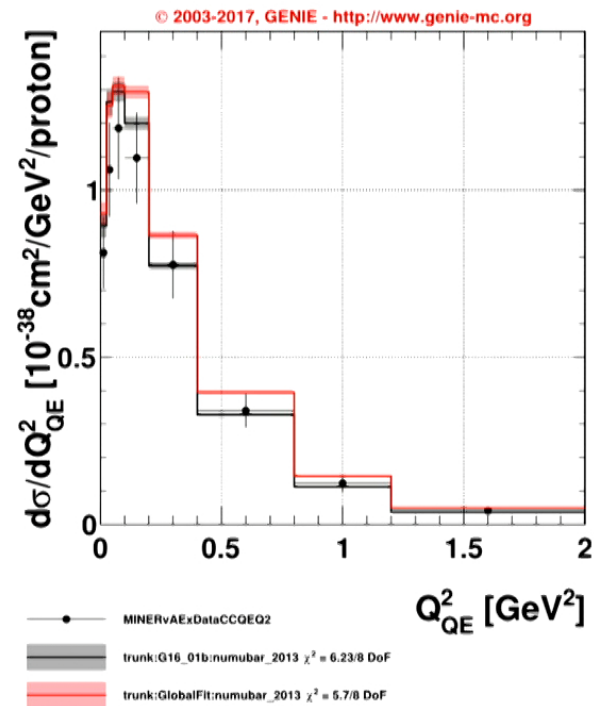


Best fit - MINERvA

Neutrinos



Antineutrinos



⇒ "Eye evaluation" would prefer default model

Peelle's Pertinent Puzzle

Datasets were fitted separately (and together)

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Schedule

- GENIE 3.0 "soon"
 - finalizing comparisons
 - deciding on technical changes
- GENIE 4.0 is on an aggressive schedule also.
 - the goal is this calendar year
 - lots of technical progress already at Liverpool
 - but many thorny physics issues (dataset tensions, experiment-theory tensions, opinions about model selection, etc.) to sort out...

NuWro

- ▶ NuWro is not an official MC in any experiment and serves as a laboratory for new developments.
- ▶ New (or relatively new) ingredients:
 - ▶ Berger-Sehgal coherent pion production
 - ▶ π momentum distribution from Δ decay
 - ▶ effective density and momentum dependent potential for CCQE (C. Juszczak, J. Nowak, J. Sobczyk)
- ▶ eWro - electron scattering module (a work in progress) C. Juszczak, K. Graczyk, JTS, J. Zmuda

Jarek Nowak, Lancaster University

IPPP/NuSTEC topical meeting on
Neutrino-Nucleus scattering

- <http://school.genie-mc.org> (lecture by T. Golan)
- <https://github.com/NuWro/nuwro>
- <https://nuwro.github.io/user-guide/>





- ▶ All major interaction channels are implemented, for charged and neutral current, covering neutrino energy region from a few hundreds MeV (Impulse Approximation limit) to several TeV:
- ▶ QEL (quasi-)elastic scattering
- ▶ RES pion production through a Δ resonance excitation
- ▶ DIS more inelastic processes
- ▶ COH coherent pion production
- ▶ np-nh two body current contribution
- ▶ **Transition region treatment:** smooth transition from full RES(Δ) to full DIS starting from $W=1.3-1.6 \text{ GeV}/c^2$

The main idea: to test NuWro nuclear model using electron scattering data

- ▶ Fermi gas and local Fermi gas
- ▶ QE and Δ regions only
- ▶ for Δ non-resonant background after E. Hernandez, J. Nieves, M. Valverde, Phys. Rev. D76 033005 (2007)
- ▶ EM form factors from J. Zmuda, K.M. Graczyk, arXiv:1501.03086v4
- ▶ Δ self-energy following E. Oset, L.L. Salcedo, Nucl. Phys. A468 631 (1987)

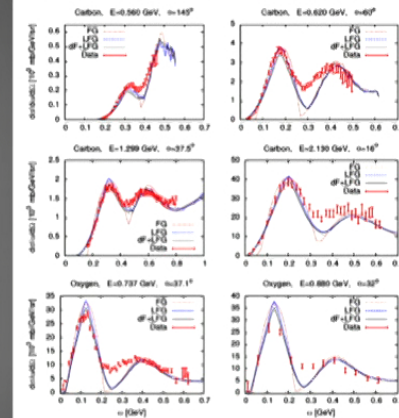
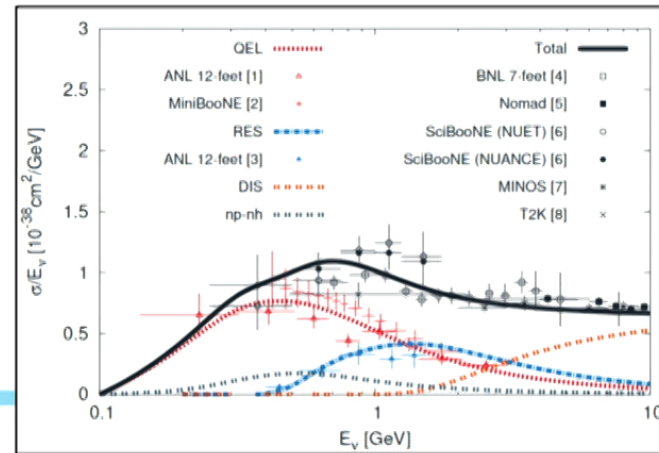
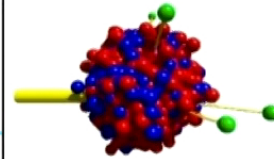


Fig. 1. Differential cross sections for electron scattering off carbon and oxygen obtained within eWro (for various beam energies, E' , and scattering angle, θ)

K. Graczyk, C. Juszczak, JTS, J. Zmuda, arXiv:1510.03268

- Re-weighting utilities are new.





- <http://gibuu.hepforge.org>
- Strives to use the “best possible theory” in all cases.

■ *Initial interactions:*

- Mean field potential with local Fermigas momentum distribution, nucleons are bound (not so in generators!)
- Initial interactions calculated by summing over interactions with all bound, Fermi-moving nucleons
- 2p2h from electron phenomenology

■ *Final state interaction:*

- propagates outgoing particles through the nucleus using *quantum-kinetic transport theory*, fully relativistic (off-shell transport possible).
Initial and final interactions come from the same Hamiltonian.
CONSISTENCY of inclusive and semi-inclusive X-sections

Ulrich Mosel

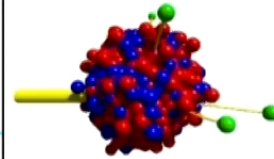
New in 2016:

IPPP/NuSTEC (Durham) 2017

- Stable groundstate implemented -> improved hole spectral functions
- 2p2h structure function for all kinematics, fitted to e-scattering, is used for neutrinos as well



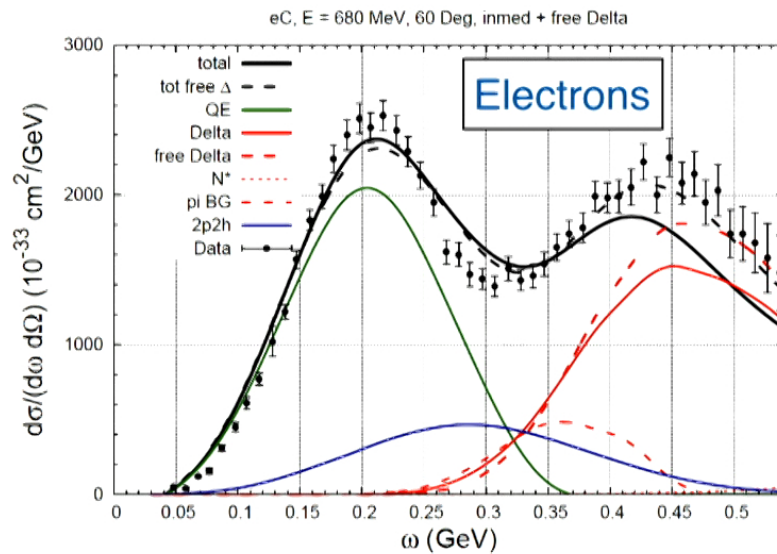
“Nature”



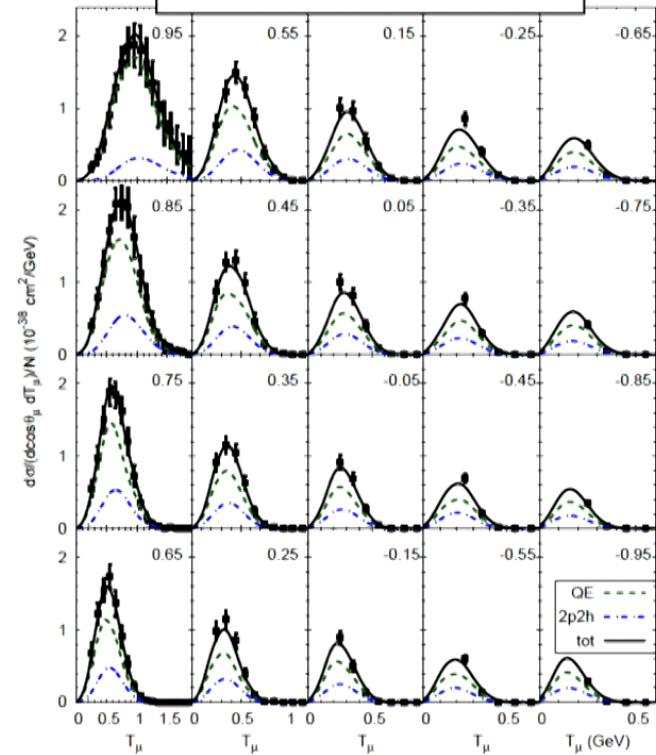
GiBUU

The Giessen Boltzmann-Uehling-Uhlenbeck Project

- Compares well to many electron and neutrino data sets.
- Typically not re-weightable, no geometry/flux



MiniBooNE Neutrinos

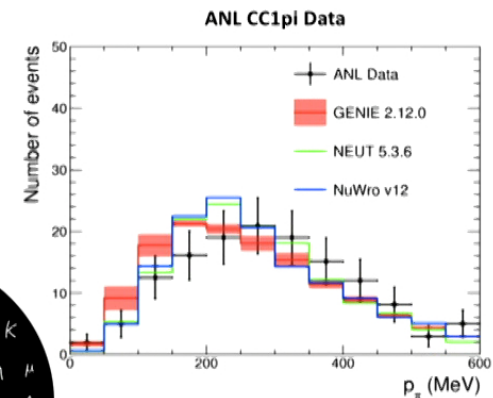
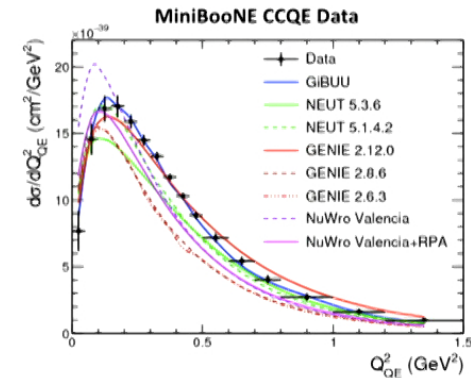


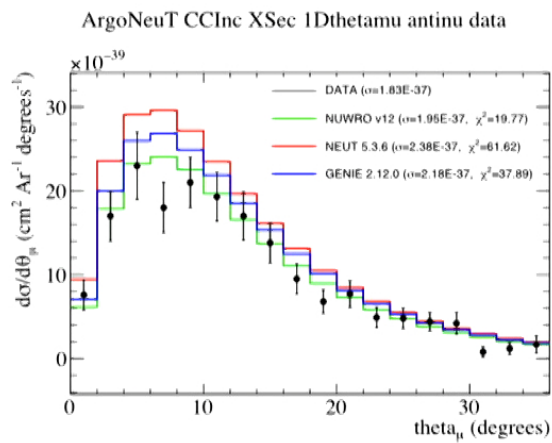
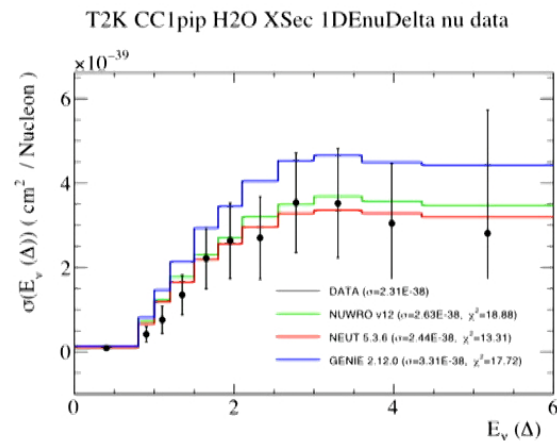
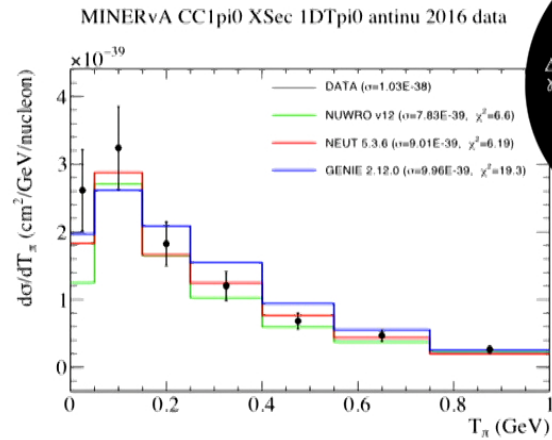
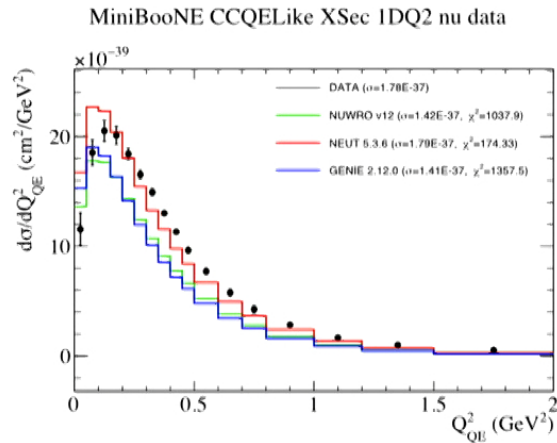
Ulrich Mosel

IPPP/NuSTEC (Durham) 2017

NUISANCE

- **Open source** generator tuning framework.
 - Tools for comparison plots, re-weighting (for NEUT, NuWro, and GENIE), fitting.
 - Interfaces to re-weighting tools in generators; can add ad-hoc weights as well.
 - Tuning mechanisms include support for priors via penalties in the likelihood.
 - Migrad & Bayesian tuning.
 - Reproducible results via job cards.
- <http://nuisance.hepforge.org>





http://nuisance.hepforge.org/files/validation/nuisancevalidation_v1r0_280217/nuisance_v1r0_validation_280217.pdf



NUISANCE

P. Stowel,
Communication to MINERvA

• Tuning mechanisms:

Frequentist Method (Migrad)

1. Form a joint likelihood for all samples included in a fit.
2. Use ROOT's GSL minimizer libraries to find a best fit.
3. Use MINOS to evaluate errors and parameter contours.

Bayesian Method

1. Throw 1-sigma prior uncertainties for all free params.
2. Bin prior distribution with no weights for each param.
3. Bin distribution again weighting each throw by the likelihood

NUISANCE can throw parameters according to arbitrary prior distributions and saves the Data-MC χ^2 value for each dataset.



Method	Pros	Cons
'Migrad' Tuning	<ul style="list-style-type: none">• Can quickly see if parameters are pulled to crazy values.	<ul style="list-style-type: none">• Need to run many different fits when looking at subsets of the data.• Risk of getting stuck in local minima.
Bayesian Tuning	<ul style="list-style-type: none">• One single large set of throws should cover all possible fits to subsets of the data.• Additional prior penalty terms (BC tuning) can be included after throws are made.	<ul style="list-style-type: none">• Requires many throws when looking at many dimensional fits.• Need to make sure prior covers all datasets.

Thanks!



Luis Alvarez-Ruso [8], Costas Andreopoulos [2,5], Christopher Barry [2],
Francis Bench [2], Steve Dennis [2], Steve Dytman [3], Hugh Gallagher [7], Tomasz Golan [1,4,*],
Robert Hatcher [1], Rhiannon Jones [2], Libo Jiang [3], Anselmo Mereaglia [6],
Donna Naples [3], Gabriel Perdue [1], Marco Roda [2],
Jeremy Wolcott [7], Julia Yarba [1]

(The GENIE Collaboration)

- (1) Fermi National Accelerator Laboratory, USA
- (2) University of Liverpool, UK
- (3) University of Pittsburgh, USA
- (4) University of Rochester, USA
- (5) STFC Rutherford Appleton Laboratory UK
- (6) Strasbourg IPHC, France
- (7) Tufts University, USA
- (8) University of Valencia, Spain
- * Currently at the University of Wrocław, Poland

