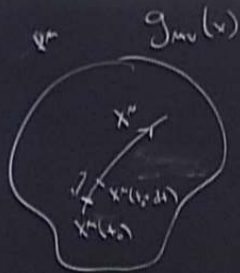


Title: PSI 16/17 Explorations in Cosmology (Kendrick Smith) - Lecture 8

Date: Apr 20, 2017 10:15 AM

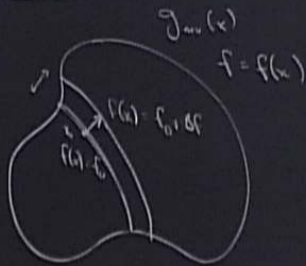
URL: <http://pirsa.org/17040061>

Abstract:



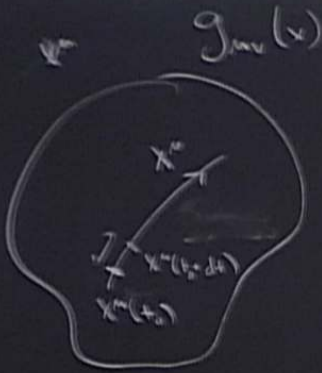
$$\text{DISTANCE} = \left[g_{\mu\nu} \frac{dx^\mu}{dt} \frac{dx^\nu}{dt} \right]^{1/2} (\Delta t)$$

$$g_{\mu\nu} = \delta_{\mu\nu} \quad \|\dot{x}\| (\Delta t)$$



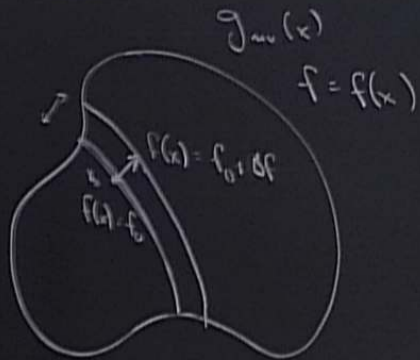
$$\text{DISTANCE} = \left[g^{\mu\nu} \partial_\mu f \partial_\nu f \right]^{-1/2} (\Delta f)$$

$$g_{\mu\nu} = \delta_{\mu\nu} \quad \|\nabla f\|^{-1} (\Delta f)$$



$$\text{DISTANCE} = \left[g_{\mu\nu} \frac{dx^\mu}{dt} \frac{dx^\nu}{dt} \right]^{1/2} (\Delta t)$$

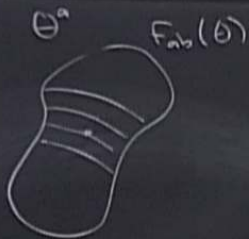
$$\sim \|\dot{x}\| (\Delta t) \quad \text{IF } g_{\mu\nu} = \delta_{\mu\nu}$$



$$\text{DISTANCE} = \left[g^{\mu\nu} \partial_\mu f \partial_\nu f \right]^{-1/2} (\Delta f)$$

$$\sim \|\nabla f\|^{-1} (\Delta f) \quad \text{IF } g_{\mu\nu} = \delta_{\mu\nu}$$





LET $f = f(\theta)$

$F_{ab} = \delta_{ab}$
 $f = \theta_1$

$\sigma(f) = [(F^{-1})^{ab}]^{1/2}$

$F_{ab}^{prior} = \begin{pmatrix} \sigma_f^{-2} & & & \\ & 0 & & \\ & & 0 & \\ & & & c \end{pmatrix}$

"FULLY MARGINALIZED"
 STATISTICAL
 ERROR

$\sigma(f) = \left[\left(\frac{\partial f}{\partial \theta_a} \right) \left(\frac{\partial f}{\partial \theta_b} \right) (F^{-1})^{ab} \right]^{1/2}$

GAUSSIAN PRIOR
 ON f WITH ERROR
 σ_f

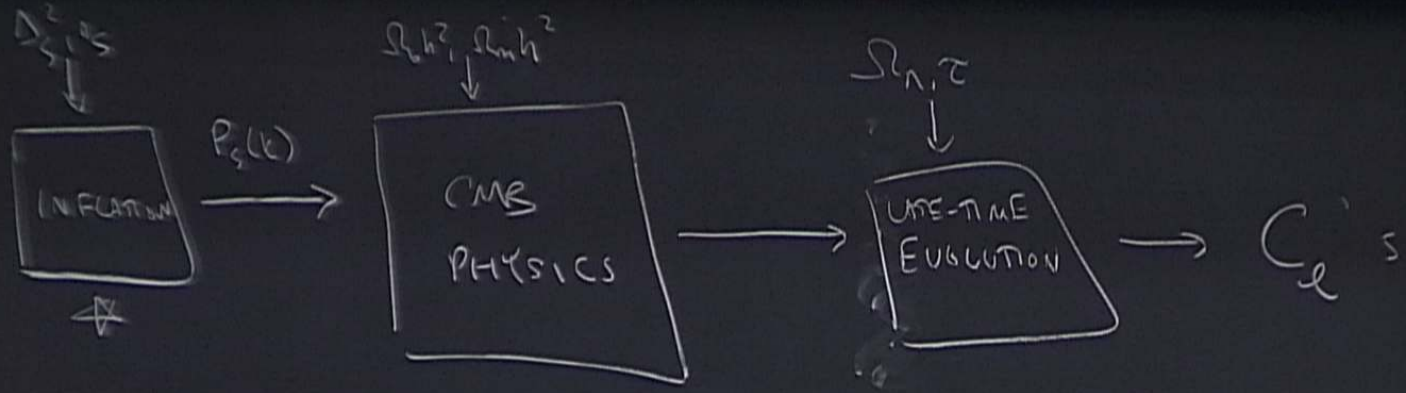
$F_{ab}^{prior} = \sigma_f^{-2} \frac{\partial f}{\partial \theta_a} \frac{\partial f}{\partial \theta_b}$

$F = F_1^{prior} + F_2^T \dots$

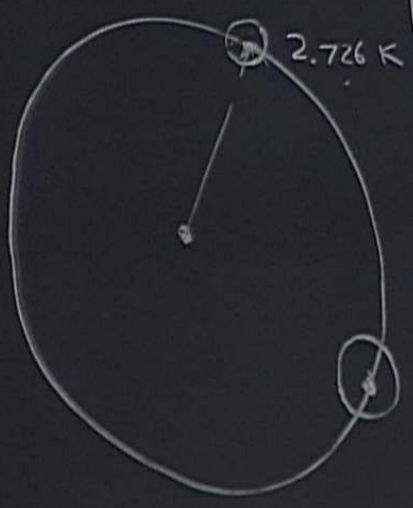
$\Theta = \{M_1, M_2, a_1, a_2, \dots\}$



$\int d^n \Theta \sqrt{F}$



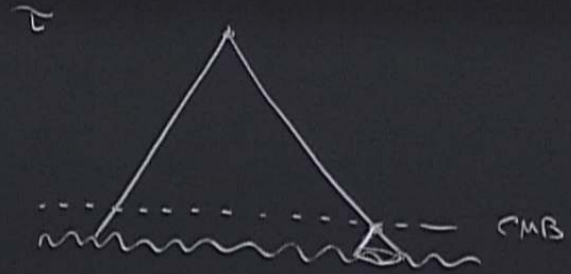
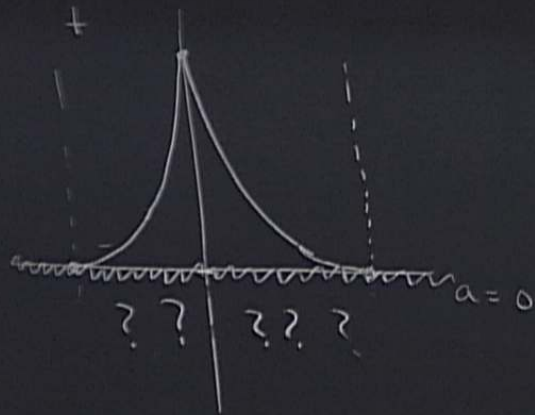
INFLATION: SOLVES "HORIZON PROBLEM"



"WEAK HORIZON PROBLEM":
 CMB IS NEARLY ISOTHERMAL
 EVEN THOUGH CAUSAL HORIZON IS $\sim 1^\circ$
 "STRONG HORIZON PROBLEM":
 LONG-RANGE CORRELATIONS
 IN FLUCTUATIONS

UNIVERSE HAS A FINITE CONFORMAL AGE

→ C_e^3



GENERAL
UNIVERSE IS $\sim t^2$
CMB

$$F = \begin{pmatrix} F_1 \\ F_2 \\ \vdots \\ F_n \end{pmatrix}$$

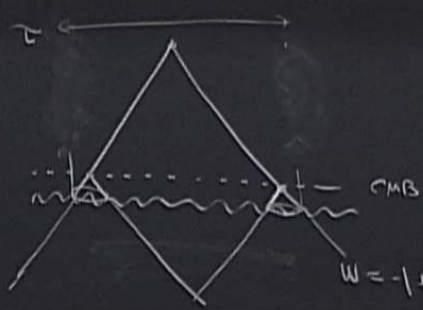
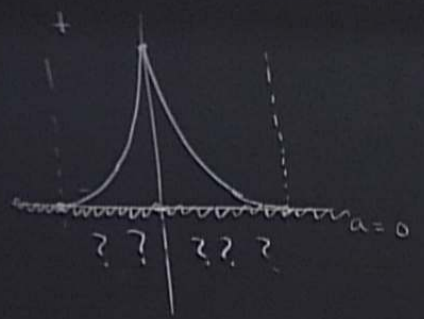
$$F = F_1 + F_2 + \dots$$

$$f(\theta) = p(\theta|d)$$

$$H(q, p)$$

UNIVERSE HAS A FINITE COMING AGE

NEARLY FLAT POTENTIAL

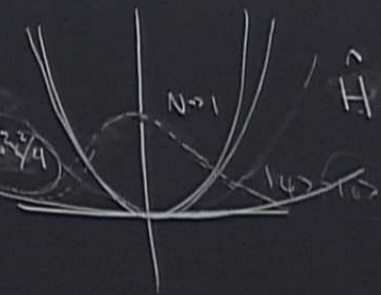


$$S = \int d^4x \sqrt{g} \left(\frac{M_{pl}^2}{2} R - \frac{1}{2} (\nabla^\mu \psi)(\nabla_\mu \psi) - V(\psi) \right)$$

$$D_s^2 \left(\frac{k}{k_0} \right)^{n_s-1}$$

$$w = -1 + \epsilon$$

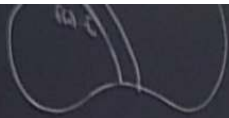
$$\psi(x) = e^{-\frac{2}{\sqrt{3}} \frac{x}{M_{pl}}}$$



$$\hat{H} = \frac{1}{2} \dot{\psi}^2 + \frac{1}{2} w(t)^2 \frac{\psi^2}{x^2}$$

$$\dot{\psi} \ll w^2$$

$$\dot{\psi} \gg w^2$$



$$\sim \|\nabla F\|^{-1} (\Delta f) \quad \text{if } g_{\mu\nu} = \delta_{\mu\nu}$$

$$\sigma(F) = (F^{-1})^{1/2}$$

$$F_{ab}^{pr} = \begin{pmatrix} g^{\mu\nu} & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

$$F = F_1 + F_2 + \dots$$

$$\Gamma_{\mu\nu}^{\rho} = \frac{1}{2} g^{\rho\lambda} (\partial_{\mu} g_{\lambda\nu} + \partial_{\nu} g_{\lambda\mu} - \partial_{\lambda} g_{\mu\nu})$$

$$R_{\mu\nu}^{\rho} = \partial_{\nu} \Gamma_{\mu\lambda}^{\rho} - \partial_{\lambda} \Gamma_{\mu\nu}^{\rho} + \Gamma_{\mu\lambda}^{\sigma} \Gamma_{\nu\sigma}^{\rho} - \Gamma_{\mu\nu}^{\sigma} \Gamma_{\lambda\sigma}^{\rho}$$

$$R_{\mu\nu} = R_{\mu\nu}^{\lambda}{}_{\lambda}$$

$$= \partial_{\lambda} \Gamma_{\mu\nu}^{\lambda} - \partial_{\nu} \Gamma_{\mu\lambda}^{\lambda} + \Gamma_{\mu\nu}^{\lambda} \Gamma_{\lambda\rho}^{\rho} - \Gamma_{\mu\lambda}^{\rho} \Gamma_{\nu\rho}^{\lambda}$$

$$R = g^{\mu\nu} R_{\mu\nu}$$

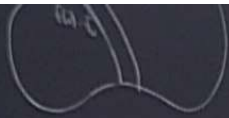
$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R$$

$$S = \underbrace{\left(\frac{M_{pl}^2}{2} \int d^4x \sqrt{g} R \right)}_{S_{EH}} + S_m$$

$$[M_{pl} = (8\pi G)^{-1/2}]$$

$$T_{\mu\nu} = -\frac{2}{\sqrt{g}} \frac{\delta S_m}{\delta g^{\mu\nu}}$$

$$G_{\mu\nu} = M_{pl}^{-2} T_{\mu\nu}$$



$$\sim \|\nabla F\|^{-1} (\Delta F) \quad \text{if } g_{\mu\nu} = \delta_{\mu\nu}$$

$$\sigma(F) = (F^{-1})^{1/2}$$

$$F_{ab}^{pr} = \begin{pmatrix} g^{\mu\nu} & & & \\ & 1 & & \\ & & 1 & \\ & & & 0 & \dots & 0 \end{pmatrix}$$

$$F = F_1^{pr} + F_2^T \dots$$

$$\Gamma_{\mu\nu}^{\rho} = \frac{1}{2} g^{\rho\lambda} (\partial_{\mu} g_{\lambda\nu} + \partial_{\nu} g_{\lambda\mu} - \partial_{\lambda} g_{\mu\nu})$$

$$R_{\mu\nu}^{\rho} = \partial_{\nu} \Gamma_{\mu\lambda}^{\rho} - \partial_{\lambda} \Gamma_{\mu\nu}^{\rho} + \Gamma_{\mu\lambda}^{\sigma} \Gamma_{\nu\sigma}^{\rho} - \Gamma_{\nu\lambda}^{\sigma} \Gamma_{\mu\sigma}^{\rho}$$

$$R_{\mu\nu} = R_{\nu\mu}$$

$$= \partial_{\lambda} \Gamma_{\mu\nu}^{\lambda} - \partial_{\nu} \Gamma_{\mu\lambda}^{\lambda} + \Gamma_{\mu\nu}^{\lambda} \Gamma_{\lambda\rho}^{\rho} - \Gamma_{\mu\lambda}^{\rho} \Gamma_{\nu\rho}^{\lambda}$$

$$R = g^{\mu\nu} R_{\mu\nu}$$

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R$$

$$S = \underbrace{\left(\frac{M_{pl}^2}{2} \int d^4x \sqrt{g} R \right)}_{S_{EH}} + S_m$$

$$[M_{pl} = (8\pi G)^{-1/2}]$$

$$T_{\mu\nu}^{DEF} = -\frac{2}{\sqrt{g}} \frac{\delta S_m}{\delta g^{\mu\nu}}$$

$$G_{\mu\nu} = M_{pl}^{-2} T_{\mu\nu}$$

$$\nabla_{\mu} f \stackrel{?}{=} \partial_{\mu} f \quad \text{TRUE}$$

$$\begin{aligned} \nabla_{\mu} \nabla_{\nu} f &= \nabla_{\mu} (\partial_{\nu} f) \\ &= \partial_{\mu} \partial_{\nu} f - \Gamma_{\mu\nu}^{\lambda} \partial_{\lambda} f \end{aligned}$$

$$\nabla_{\mu} \nabla^{\nu} f \stackrel{?}{=} g^{\nu\rho} \nabla_{\mu} \nabla_{\rho} f \quad \text{OR} \quad \nabla_{\mu} \nabla^{\nu} f = \nabla_{\mu} (g^{\nu\rho} \nabla_{\rho} f) \quad \nabla_{\mu} g_{\nu\rho} = 0$$

$$\nabla_{\mu} f \stackrel{?}{=} \partial_{\mu} f \quad \text{TRUE}$$

$$\begin{aligned} \nabla_{\mu} \nabla_{\nu} f &= \nabla_{\mu} (\partial_{\nu} f) \\ &= \partial_{\mu} \partial_{\nu} f - \Gamma_{\mu\nu}^{\lambda} \partial_{\lambda} f \end{aligned}$$

$$\nabla_{\mu} \nabla^{\nu} f \stackrel{?}{=} g^{\nu\rho} \nabla_{\mu} \nabla_{\rho} f \quad \text{OR} \quad \nabla_{\mu} \nabla^{\nu} f = \nabla_{\mu} (g^{\nu\rho} \nabla_{\rho} f)$$

"METRIC COMPATIBILITY"

↓

$$\nabla_{\mu} g_{\nu\rho} = 0$$

$$\nabla_M \nabla_\nu f = \nabla_\nu \nabla_M f \quad \text{TRUE} \quad \text{"TORSION FREE"}$$

$$\nabla_M \nabla_\nu X_\rho = \nabla_\nu \nabla_M X_\rho + R_{\mu\nu\rho}{}^\lambda X_\lambda$$

"DIVERGENCE FORMULA":

$$\nabla_M X^M = \frac{1}{\sqrt{-g}} \partial_M (\sqrt{-g} X^M)$$

$$\int d^4x \partial_m X^m \stackrel{?}{=} 0 \quad \checkmark$$

$$\int d^4x \sqrt{-g} \partial_m X^m \stackrel{?}{=} 0 \quad \times$$

$$\int d^4x (\nabla_m X^m) \stackrel{?}{=} 0 \quad \times$$

$$\int d^4x \sqrt{-g} (\nabla_m X^m) \stackrel{?}{=} 0 \quad \checkmark$$

$$F_{ab} = \begin{pmatrix} 0 & f & & \\ -f & 0 & & \\ & & 0 & 0 \\ & & 0 & 0 \end{pmatrix}$$

$$F = F_1 + F_2 + \dots$$

$$\int d^4x \sqrt{-g} (\nabla_m X^m) f = \int d^4x \sqrt{-g} \left[X^m (-\nabla_m f) + \cancel{\nabla_m (f X^m)} \right]$$

$$\int d^4x \sqrt{-g} (\nabla_m \phi) (\nabla^m \psi) = - \int \sqrt{-g} \phi (\nabla_m \nabla^m \psi)$$

$$\int d^4x \sqrt{-g} (\nabla_m X_\nu) (\nabla^\nu \psi^m) = \int \sqrt{-g} (\nabla^\nu X_\nu) (\nabla_m \psi^m)$$

$$\int d^4x \sqrt{-g} (\nabla_\mu X^\mu) f = \int d^4x \sqrt{-g} \left[X^\mu (-\nabla_\mu f) + \nabla_\mu (f X^\mu) \right]$$

$$\int d^4x \sqrt{-g} (\nabla_\mu \phi) (\nabla^\mu \psi) = - \int \sqrt{-g} \phi (\nabla_\mu \nabla^\mu \psi)$$

$$\begin{aligned} \int d^4x \sqrt{-g} (\nabla_\mu X_\nu) (\nabla^\nu Y^\mu) &= - \int d^4x \sqrt{-g} X_\nu (\nabla_\mu \nabla^\nu Y^\mu) \\ &= \int d^4x \sqrt{-g} X^\nu \left[\nabla_\nu \nabla_\mu Y^\mu - R_{\mu\nu\lambda}{}^\mu Y^\lambda \right] \\ &= \int d^4x \sqrt{-g} (\nabla_\nu X^\nu) (\nabla_\mu Y^\mu) + \int d^4x \sqrt{-g} R_{\nu\lambda} X^\nu Y^\lambda \end{aligned}$$