

Title: PSI 16/17 Explorations in Cosmology (Kendrick Smith) - Lecture 6

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URL: <http://pirsa.org/17040058>

Abstract:

$$\langle \phi_i(\vec{r}) \phi_j(\vec{r}') \rangle = S_{ij}(|\vec{r} - \vec{r}'|)$$

$$\langle \phi_i(k) \phi_j(k') \rangle = P_{ij}(k) (2\pi)^3 \delta^3(k - k')$$

$\vec{r}$

$$\phi = \phi(\vec{x})$$

$$\phi_i(\vec{x}) = (\phi_1(\vec{x}), \phi_2(\vec{x}))$$

TRANSLATION + ROTATION INVARIANCE

$$S_{ij}(s) = \begin{pmatrix} S_{11}(s) & S_{12}(s) \\ S_{21}(s) & S_{22}(s) \end{pmatrix}$$

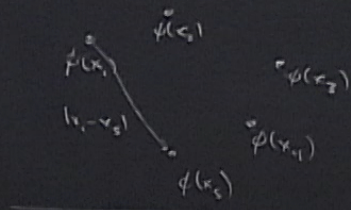
$$S_{21} = S_{12}$$

$P_{ij}(k)$ = SYMMETRIC REAL MATRIX

IN VARIANCE

$$S_{21} = S_{12}$$

$\psi(\vec{r})$ GWF $P(k)$



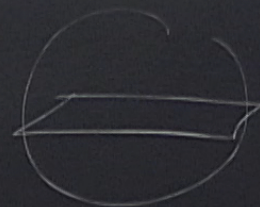
$$C_{11} = \text{VAR}(\psi(x_1)) = \langle \psi(x_1)^2 \rangle = \langle S(0) \rangle$$

$$C_{ij} = \begin{pmatrix} \langle x_i, x_j \rangle \\ \langle x_i, x_j \rangle^2 \end{pmatrix}$$

$$C_{ij} = \langle \psi(x_i) \psi(x_j) \rangle$$

$(\psi(x_1), \dots, \psi(x_n)) = \text{MULTIVARIATE GAUSSIAN R.V.}$

LINEAR COMBINATIONS OF GAUSSIANS ARE GAUSSIAN



$$S(\vec{s}) = \int \frac{d^3k}{(2\pi)^3} P(k) e^{i\vec{k} \cdot \vec{s}}$$

$$= \langle \psi(\vec{r}) \psi(\vec{r} + \vec{s}) \rangle$$

$$\phi = \phi(\vec{r}) \quad \phi_1(\vec{r}) = (\phi_1(x), \phi_1(y))$$

$P(\vec{r}) = \text{SYMMETRIC REAL FUNCTION}$

$(\phi(x_1), \dots, \phi(x_n)) = \text{MULTIVARIATE GAUSSIAN P.D.F.}$
 LINEAR COMBINATIONS OF GAUSSIANS ARE GAUSSIAN

$$C_{11} = \text{VAR}(\phi(x_1)) = \langle \phi(x_1)^2 \rangle = \langle \phi(x_1) \phi(x_1) \rangle = \langle \phi(x_1)^2 \rangle$$

$$S(\vec{q}) = \int \frac{d^3r}{(2\pi)^3} P(\vec{r}) e^{i\vec{q}\cdot\vec{r}} = \langle \phi(\vec{r}) \phi(\vec{r} + \vec{s}) \rangle$$

RANDOM FIELDS ON A 2D SPHERE

2D PLANE

2D SPHERE

$C_0, C_2, C_4, \dots$

E.G. CMB TEMPERATURE $T(\hat{n})$

CORRELATION FN $\langle T(\hat{n}) T(\hat{n}') \rangle = S(|\hat{n} - \hat{n}'|)$

$\langle T(\hat{n}) T(\hat{n}') \rangle = f(\Theta_{nn'})$

WHERE $\Theta_{nn'} = \cos^{-1}(\hat{n} \cdot \hat{n}')$

FOURIER TRANSFORM

$$T(\vec{r}) = \int \frac{d^2\ell}{(2\pi)^2} \tilde{T}(\vec{\ell}) e^{i\vec{\ell}\cdot\vec{r}}$$

$$T(\hat{n}) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} a_{\ell m} Y_{\ell m}(\hat{n})$$

"SPHERICAL HARMONIC TRANSFORM"
 $T(\hat{n}) \longleftrightarrow a_{\ell m}$

REALITY CONDITION

$$T(\vec{\ell})^* = T(-\vec{\ell})$$

$$a_{\ell m}^* = (-1)^m a_{\ell, -m}$$

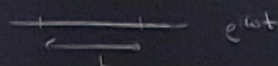
$$[Y_{\ell m}^* = (-1)^m Y_{\ell, -m}]$$

POWER SPECTRUM

$$\langle T(\vec{\ell}) T(\vec{\ell}') \rangle = C_\ell / (2\pi)^2 \delta^2(\vec{\ell} - \vec{\ell}')$$

$$\langle a_{\ell m}^* a_{\ell' m'} \rangle = C_\ell \delta_{\ell\ell'} \delta_{mm'}$$

DEFINITION OF C_ℓ



$$e^{i\omega t}$$

$$e^{in(\pi/2) +}$$

$$\omega = n(\pi/2)$$

$$a_{00} = 2.726 \text{ K}$$

$$a_{1m} \sim 3 \text{ mK}$$

$$a_{2m} \sim 1.72$$

$$(\Omega_{\text{rad}} h^2)$$

(DIPOLE MODES: MOTION OF EARTH)

(COSMOLOGICAL FLUCTUATIONS)

2D PLANE

2D SPHERE

$c_2, c_3, c_4, \dots$

$T(\vec{n})$

CORRELATION FN $\langle T(\vec{x}) T(\vec{y}) \rangle = S(|\vec{x} - \vec{y}|)$

FOURIER TRANSFORM

$$T(\vec{x}) = \int \frac{d^2\ell}{(2\pi)^2} \tilde{T}(\ell) e^{i\vec{\ell} \cdot \vec{x}}$$

REALITY CONDITION

$$T(\vec{\ell})^* = T(-\vec{\ell})$$

POWER SPECTRUM

$$\langle T(\vec{\ell})^* T(\vec{\ell}') \rangle = C(\ell) (2\pi)^2 \delta^2(\ell - \ell')$$

REAL-SPACE VARIANCE

$$\begin{aligned} \langle T(\vec{n})^2 \rangle &= \int \frac{d^2\ell}{(2\pi)^2} C(\ell) \\ &= \int d\ell \frac{C(\ell)}{2\pi} \end{aligned}$$

$\langle T(\vec{n}) T(\vec{n}') \rangle = \rho(\Theta_{nn'})$

$$T(\vec{n}) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} a_{\ell m} Y_{\ell m}(\vec{n})$$

$$a_{\ell m}^* = (-1)^m a_{\ell, -m}$$

$$\langle a_{\ell m}^* a_{\ell' m'} \rangle = C_{\ell} \delta_{\ell\ell'} \delta_{mm'}$$

$$\langle T(\vec{n})^2 \rangle = \sum_{\ell} \frac{2\ell+1}{4\pi} C_{\ell}$$

WHERE $\Theta_{nn'} = \cos^{-1}(\vec{n} \cdot \vec{n}')$

"SPHERICAL HARMONIC TRANSFORM"

$$T(\vec{n}) \longleftrightarrow a_{\ell m}$$

$$[Y_{\ell m}^* = (-1)^m Y_{\ell, -m}]$$

DEFINITION OF C_{ℓ}

"FISHER MATRIX"

STATISTICS

- BIASED COIN W/ UNKNOWN PROBABILITY Θ OF HEADS. SUPPOSE THAT IN 100 FLIPS, 64 HEADS OCCUR.

- WHAT IS 95% CONFIDENCE REGION ON Θ ?

$$\Theta = 0.2$$

- IF THE
WHAT'S T

P

$p(\Theta)$

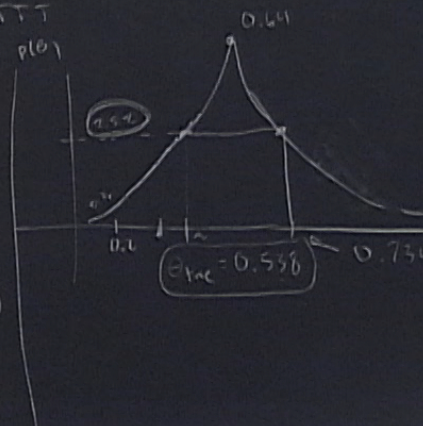
"FREQUENTIST STATISTICS"

- IF THE TRUE θ WERE $\theta_{true} = 0.2$,
WHAT'S THE PROBABILITY P OF ≥ 64 HEADS?

$$P = \sum_{k \geq 64} \binom{100}{k} (0.2)^k (0.8)^{100-k}$$

HHH - H TTTT

$$p(\theta_{true}) = \begin{cases} \sum_{k \geq 64} \binom{100}{k} \theta_{true}^k (1 - \theta_{true})^{100-k} & (\theta_{true} \leq 0.64) \\ \sum_{k \leq 64} \binom{100}{k} \theta_{true}^k (1 - \theta_{true})^{100-k} & (\theta_{true} \geq 0.64) \end{cases}$$



95% CONFIDENCE REGION

$$0.538 \leq \theta_{true} \leq 0.734$$