Title: The universe as a quantum gravity condensate.

Date: Nov 03, 2016 02:30 PM

URL: http://pirsa.org/16110048

Abstract: I describe how, within the group field theory (GFT) formalism for quantum gravity, we can:

1) provide a candidate description of the quantum building blocks of spacetime, bringing together ideas and mathematical structures from other quantum gravity formalisms;

2) apply powerful tools from quantum field theory, like the (perturbative and non-perturbative) renormalization group, to establish the quantum consistency of given GFT models and to study their continuum limit and phase structure;

3) extract, from the full theory, an effective cosmological dynamics for the universe described as a quantum condensate of GFT building blocks; in the simplest approximation, this dynamics reduces to the Friedmann equations at large scales but replaces the classical big bang singularity with a quantum bounce.

Pirsa: 16110048 Page 1/154





The Universe as a Quantum Gravity Condensate

Daniele Oriti

Albert Einstein Institute

Perimeter Institute for Theoretical Physics 3/10/2016





Pirsa: 16110048 Page 2/154

Plan of the talk

- disappearance and emergence of Space and Time in Quantum Gravity
- GFTs: what are they?
 - general formalism
 - relation with other QG approaches
- continuum limit in GFT and GFT renormalization
- effective continuum physics
 - cosmology as Quantum Gravity hydrodynamics
 - GFT condensate cosmology

Pirsa: 16110048 Page 3/154

Intro: disappearance and emergence of Space and Time in Quantum Gravity

Pirsa: 16110048 Page 4/154

Hints for the Disappearance of Space and Time

challenges to "localization" in semi-classical GR

minimal length scenarios non-commutative spacetimes

spacetime singularities in GR

need just quantum corrections to classical GR?

breakdown of continuum itself?

black hole thermodynamics

solution of information loss paradox require non-locality?

if spacetime itself has entropy, it has microstructure if entropy is finite, this implies discreteness

Einstein's equations as equation of state (Jacobson et al)

GR dynamics is effective equation of state for any microscopic dofs collectively described by a spacetime, a metric and some matter fields

insights from analog gravity models in condensed matter physics

effective curved metric and matter fields from non-geometric atomic theory

Pirsa: 16110048 Page 5/154

Hints for the Disappearance of Space and Time

challenges to "localization" in semi-classical GR

minimal length scenarios non-commutative spacetimes

spacetime singularities in GR

need just quantum corrections to classical GR?

breakdown of continuum itself?

black hole thermodynamics

solution of information loss paradox require non-locality?

if spacetime itself has entropy, it has microstructure if entropy is finite, this implies discreteness

Einstein's equations as equation of state (Jacobson et al)

GR dynamics is effective equation of state for any microscopic dofs collectively described by a spacetime, a metric and some matter fields

insights from analog gravity models in condensed matter physics

effective curved metric and matter fields from non-geometric atomic theory

Space and Time disappear in QG



Space and Time emerge from (discrete?) non-spatiotemporal entities

Pirsa: 16110048 Page 6/154



Pirsa: 16110048 Page 7/154

Beyond the spacetime continuum?

already imagining and constructing a consistent pre-geometric, pre-continuum picture of spacetime, is highly non-trivial

Einstein (1936): "the introduction of a space-time continuum may be considered as contrary to nature in view of the molecular structure of everything which happens on a small scale. [...] perhaps the success of the Heisenberg method points to a purely algebraic method of description of nature, that is to the elimination of continuous functions from physics. Then, however, we must also give up, by principle, the space-time continuum. It is not unimaginable that human ingenuity will some day find methods which will make it possible to proceed along such a path. At the present time, however, such a program looks like an attempt to breathe in empty space."

Pirsa: 16110048 Page 8/154

Spacetime emergence: phase transition + coarse graining?

if spacetime is made of discrete, pre-geometric building blocks, why does it look geometric and continuous?

Pirsa: 16110048 Page 9/154

Spacetime emergence: phase transition + coarse graining?

if spacetime is made of discrete, pre-geometric building blocks, why does it look geometric and continuous?

guiding hypotheses

space, time and geometry are the result of the collective behaviour of the microscopic building blocks ("QG atoms")

the universe and its smooth, macroscopic geometry are the result of a phase transition (geometrogenesis) of QG system, from a non-geometric, non-spatio-temporal phase, to a geometric one

the emergent, continuum dynamics of geometry (and GR) should be looked for in the coarse grained description of the fundamental dynamics, in "geometric phase"

in particular, cosmology is QG hydrodynamics

Pirsa: 16110048 Page 10/154

Spacetime emergence: phase transition + coarse graining?

if spacetime is made of discrete, pre-geometric building blocks, why does it look geometric and continuous?

guiding hypotheses

space, time and geometry are the result of the collective behaviour of the microscopic building blocks ("QG atoms")

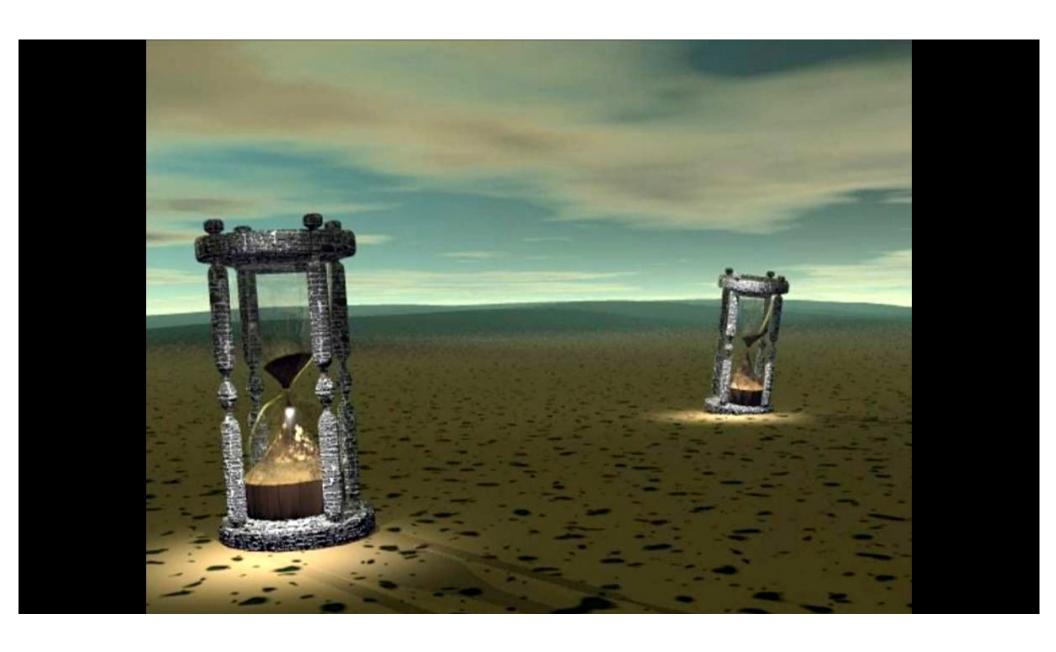
the universe and its smooth, macroscopic geometry are the result of a phase transition (geometrogenesis) of QG system, from a non-geometric, non-spatio-temporal phase, to a geometric one

the emergent, continuum dynamics of geometry (and GR) should be looked for in the coarse grained description of the fundamental dynamics, in "geometric phase"

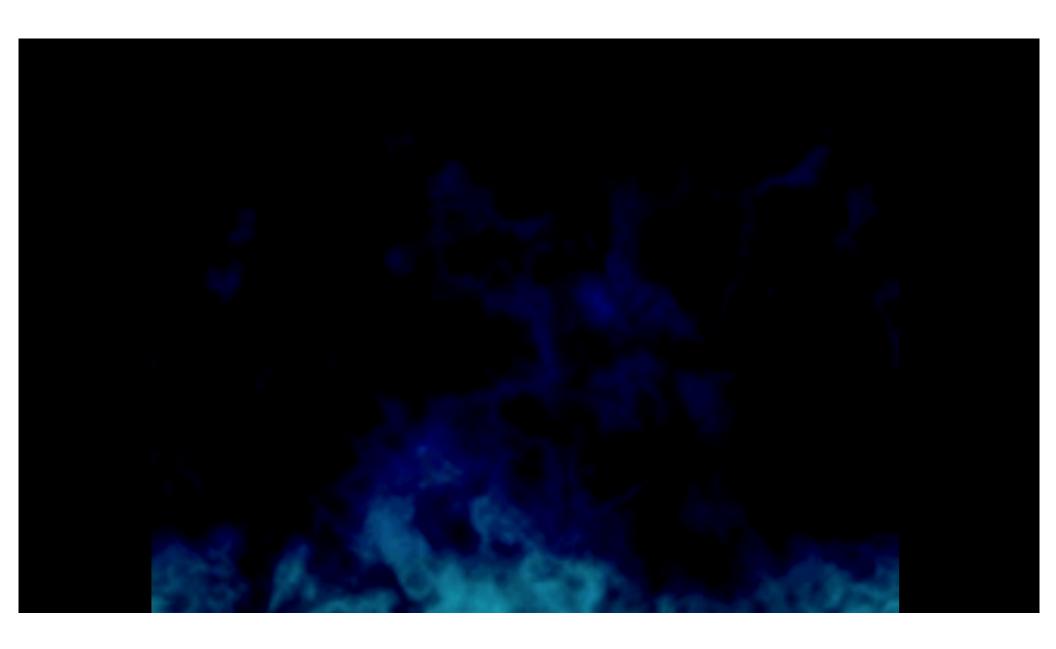
in particular, cosmology is QG hydrodynamics

analogy: spacetime is like a condensed matter system, arising from a dynamical "condensation" of QG building blocks (~ "atoms of space")

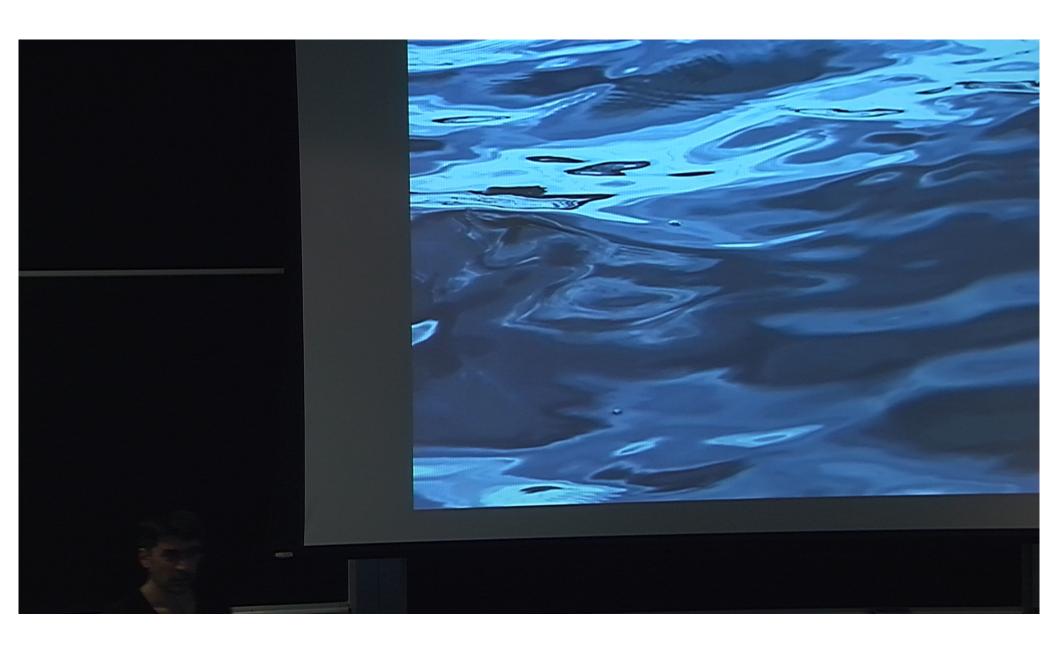
Pirsa: 16110048 Page 11/154



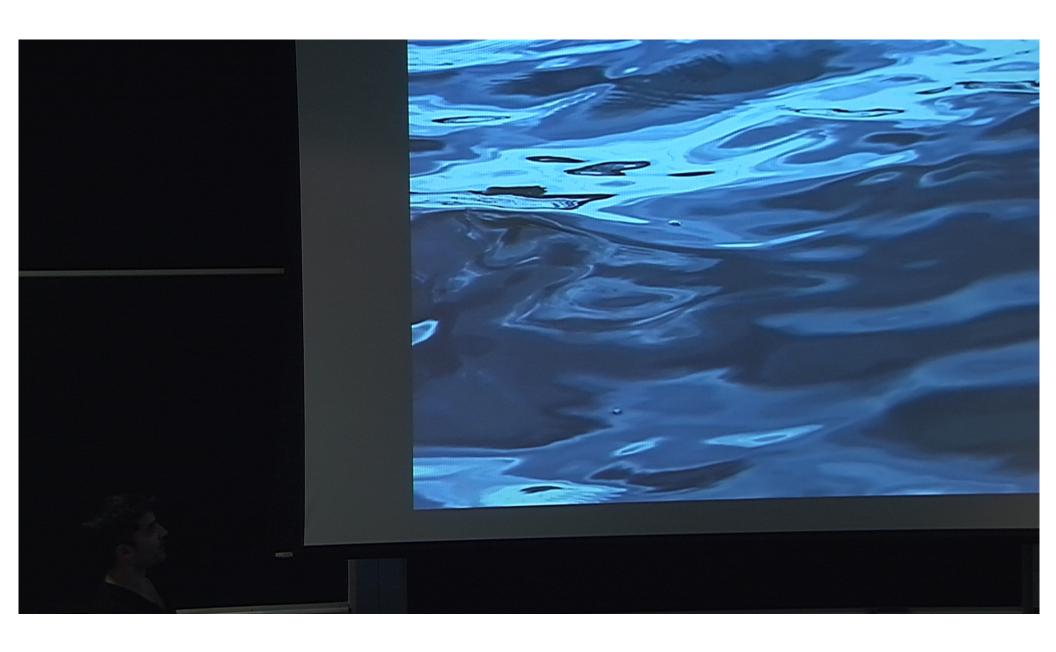
Pirsa: 16110048 Page 12/154



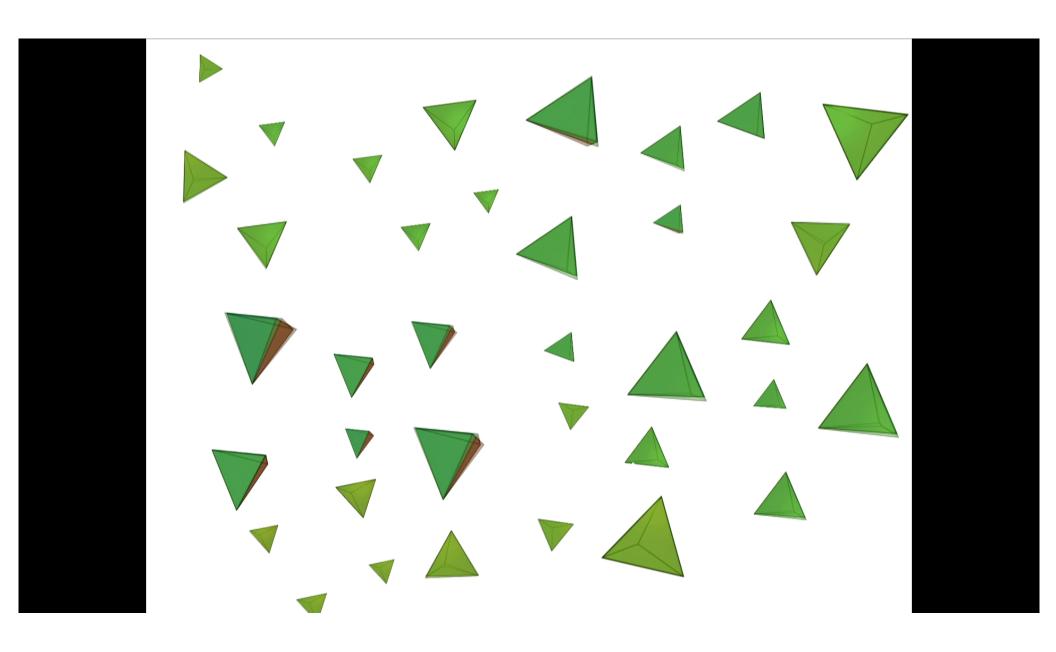
Pirsa: 16110048 Page 13/154



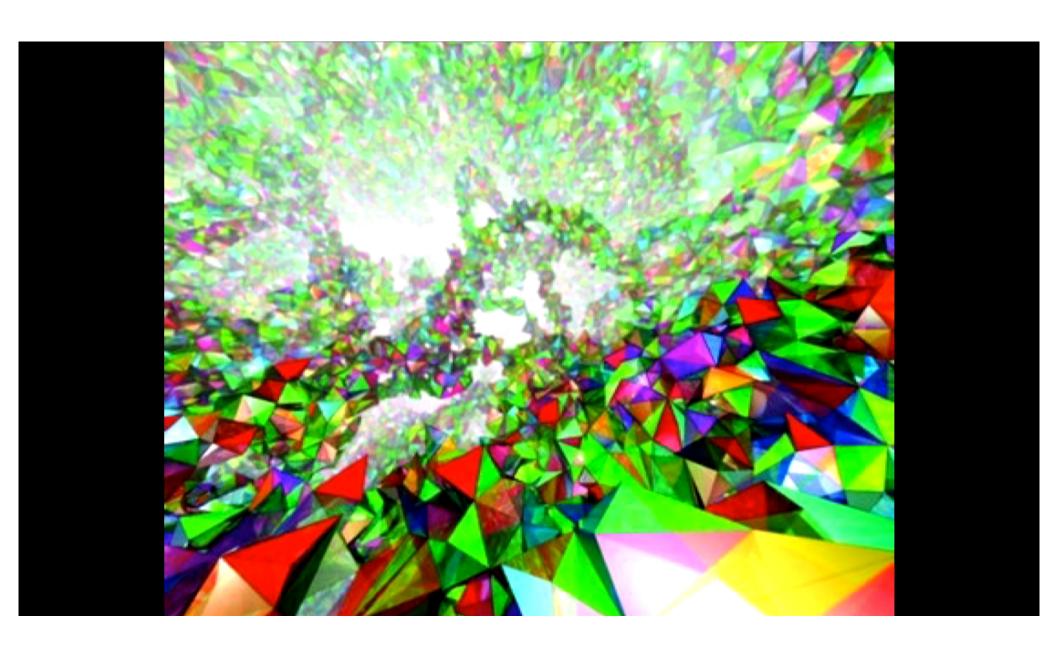
Pirsa: 16110048 Page 14/154



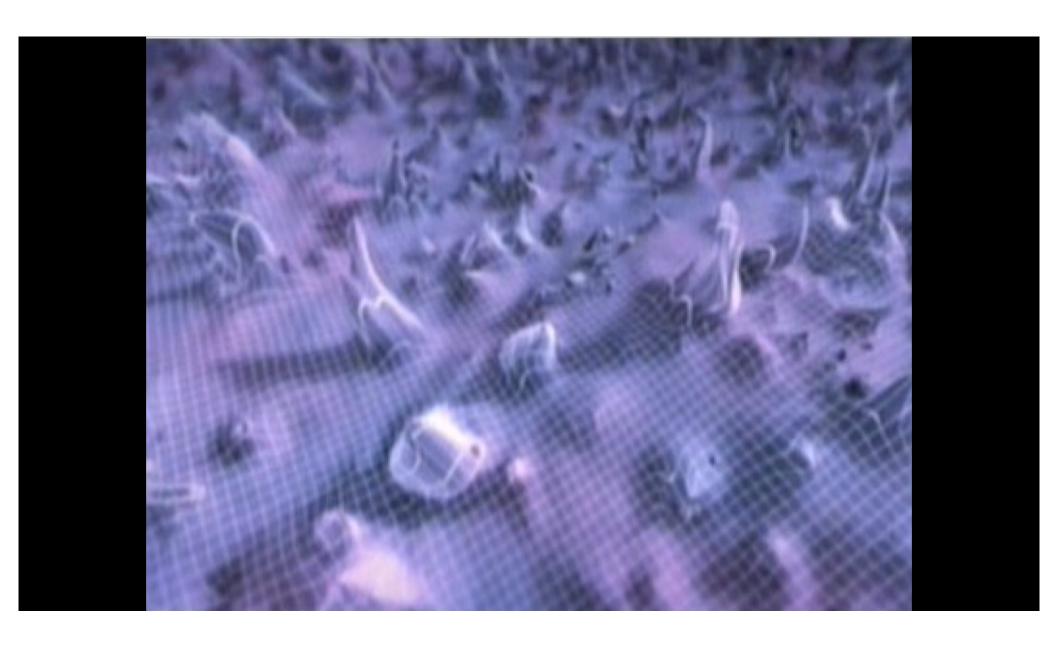
Pirsa: 16110048 Page 15/154



Pirsa: 16110048 Page 16/154



Pirsa: 16110048 Page 17/154



Pirsa: 16110048 Page 18/154

Emergence for space and time in QG

There is no Space and no Time in Quantum Gravity, both have to emerge in some approximation

The world is fundamentally Quantum and its building blocks do not have spatiotemporal features

Continuum, spatiotemporal physics to be looked for in collective behaviour of fundamental building blocks

Cosmology is sector of such collective physics, hydrodynamics of microscopic building blocks





Pirsa: 16110048 Page 19/154



Pirsa: 16110048 Page 20/154

Part 1: the GFT formalism

Pirsa: 16110048 Page 21/154

(Boulatov, Ooguri, De Pietri, Freidel, Krasnov, Rovelli, Perez, DO, Livine, Baratin,)

QFT -of- spacetime, not -on- spacetime

Pirsa: 16110048 Page 22/154

(Boulatov, Ooguri, De Pietri, Freidel, Krasnov, Rovelli, Perez, DO, Livine, Baratin,)

QFT -of- spacetime, not -on- spacetime

a QFT for the building blocks of (quantum) space

Quantum field theories over group manifold G (or corresponding Lie algebra) $\varphi:G^{ imes d} o \mathbb{C}$

relevant classical phase space for "GFT quanta":

$$(\mathcal{T}^*G)^{\times d} \simeq (\mathfrak{g} \times G)^{\times d}$$

can reduce to subspaces in specific models depending on conditions on the field

d is dimension of "spacetime-to-be"; for gravity models, G = local gauge group of gravity (e.g. Lorentz group)

example: d=4 $\varphi(g_1,g_2,g_3,g_4) \leftrightarrow \varphi(B_1,B_2,B_3,B_4) \to \mathbb{C}$

(Boulatov, Ooguri, De Pietri, Freidel, Krasnov, Rovelli, Perez, DO, Livine, Baratin,)

QFT -of- spacetime, not -on- spacetime

a QFT for the building blocks of (quantum) space

Quantum field theories over group manifold G (or corresponding Lie algebra) $\varphi:G^{ imes d} o \mathbb{C}$

relevant classical phase space for "GFT quanta":

$$(\mathcal{T}^*G)^{\times d} \simeq (\mathfrak{g} \times G)^{\times d}$$

can reduce to subspaces in specific models depending on conditions on the field

d is dimension of "spacetime-to-be"; for gravity models, G = local gauge group of gravity (e.g. Lorentz group)

example: d=4 $\varphi(g_1,g_2,g_3,g_4) \leftrightarrow \varphi(B_1,B_2,B_3,B_4) \to \mathbb{C}$

very general framework; interest rests on specific models/use (most interesting QG models are for Lorentz group in 4d)

QFT of spacetime, not defined on spacetime

a QFT for the building blocks of (quantum) space

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$

boson statistics is -assumption-(can construct, e.g., fermionic models)

$$\left[\hat{\varphi}(\vec{g})\,,\,\hat{\varphi}^{\dagger}(\vec{g}')\right]\,=\,\mathbb{I}_{G}(\vec{g},\vec{g}')\qquad \left[\hat{\varphi}(\vec{g})\,,\,\hat{\varphi}(\vec{g}')\right]\,=\,\left[\hat{\varphi}^{\dagger}(\vec{g})\,,\,\hat{\varphi}^{\dagger}(\vec{g}')\right]\,=\,0$$

additional conditions (e.g. symmetries) can be imposed on fields restrictions on Hilbert space

QFT of spacetime, not defined on spacetime

a QFT for the building blocks of (quantum) space

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$

boson statistics is -assumption-(can construct, e.g., fermionic models)

$$\left[\hat{\varphi}(\vec{g})\,,\,\hat{\varphi}^{\dagger}(\vec{g}')\right] \,=\, \mathbb{I}_{G}(\vec{g},\vec{g}') \qquad \left[\hat{\varphi}(\vec{g})\,,\,\hat{\varphi}(\vec{g}')\right] = \left[\hat{\varphi}^{\dagger}(\vec{g})\,,\,\hat{\varphi}^{\dagger}(\vec{g}')\right] \,=\, 0$$

additional conditions (e.g. symmetries) can be imposed on fields restrictions on Hilbert space

in GFT models, this is fundamental Hilbert space of dofs of universe

spacetime, geometry and matter fields should emerge from these quantum data

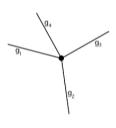
a QFT for the building blocks of (quantum) space

Fock vacuum: "no-space" ("emptiest") state | 0 >

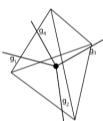
(d=4)

single field "quantum": spin network vertex or tetrahedron ("building block of space")

$$\varphi(g_1, g_2, g_3, g_4) \leftrightarrow \varphi(B_1, B_2, B_3, B_4) \to \mathbb{C}$$







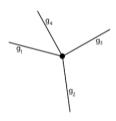
a QFT for the building blocks of (quantum) space

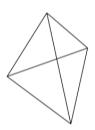
Fock vacuum: "no-space" ("emptiest") state | 0 >

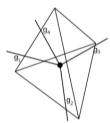
(d=4)

single field "quantum": spin network vertex or tetrahedron ("building block of space")

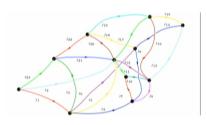
$$\varphi(g_1, g_2, g_3, g_4) \leftrightarrow \varphi(B_1, B_2, B_3, B_4) \to \mathbb{C}$$

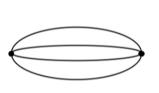




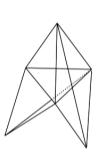


generic quantum state: arbitrary collection of spin network vertices (including glued ones) or tetrahedra (including glued ones)









a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi, \overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia}, \bar{g}_{iD}) + c.c.$$

Pirsa: 16110048 Page 29/154

a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality" in pairing of field arguments

Pirsa: 16110048 Page 30/154

a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality" in pairing of field arguments

specific combinatorics depends on model

simplest example (case d=4): simplicial setting

a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality" in pairing of field arguments

specific combinatorics depends on model

simplest example (case d=4): simplicial setting

combinatorics of field arguments in interaction: gluing of 5 tetrahedra across common triangles, to form 4-simplex ("building block of spacetime")

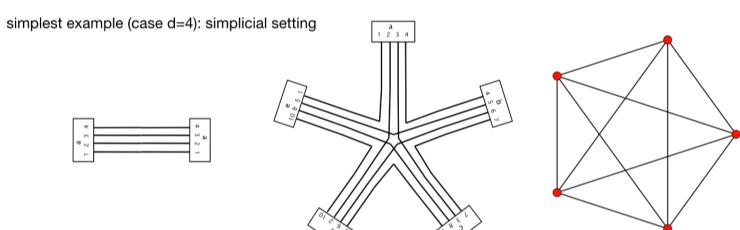
Pirsa: 16110048 Page 32/154

a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality" in pairing of field arguments

specific combinatorics depends on model



Pirsa: 16110048 Page 33/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Pirsa: 16110048 Page 34/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Pirsa: 16110048 Page 35/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

(generalisation of matrix models for 2d gravity/string worldsheet)

Pirsa: 16110048 Page 36/154

a QFT for the building blocks of (quantum) space

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi, \overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia}, \bar{g}_{iD}) + c.c.$$

Pirsa: 16110048 Page 37/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Pirsa: 16110048 Page 38/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Pirsa: 16110048 Page 39/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

(generalisation of matrix models for 2d gravity/string worldsheet)

Feynman amplitudes (model-dependent):

equivalently:

spin foam models (in group irreps)

Reisenberger, Rovelli, '00

 lattice gauge theories (with group/Lie algebra variables)

A. Baratin, DO, '11

Pirsa: 16110048 Page 40/154

a QFT for the building blocks of (quantum) space

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

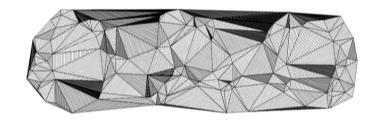
(simplicial case: simplicial complexes obtained by gluing d-simplices)

(generalisation of matrix models for 2d gravity/string worldsheet)

Feynman amplitudes (model-dependent):

equivalently:

- spin foam models (in group irreps)
 Reisenberger, Rovelli, '00
- lattice gauge theories
 (with group/Lie algebra variables)
 A. Baratin, DO, '11



Pirsa: 16110048 Page 41/154



Pirsa: 16110048 Page 42/154

(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

same combinatorics (of states/observables and histories/Feynman diagrams), no group-theoretic data

Pirsa: 16110048 Page 43/154

(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

same combinatorics (of states/observables and histories/Feynman diagrams), no group-theoretic data

example: d=3

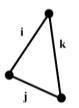
dropping group/algebra data (or restricting to finite group)

$$\varphi(g_1, g_2, g_3): G^{\times 3} \to \mathbb{C}$$

$$T_{ijk}: \mathbb{Z}_N^{\times 3} \to \mathbb{C}$$

 $T_{ijk}: X^{\times 3} \to \mathbb{C}$ $X = 1, 2, ..., N$

$$X = 1, 2, ..., N$$



(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

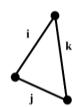
same combinatorics (of states/observables and histories/Feynman diagrams), no group-theoretic data

example: d=3

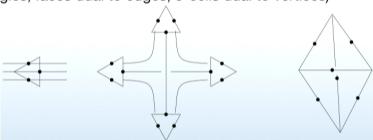
dropping group/algebra data (or restricting to finite group)

$$X = 1, 2, ..., N$$

$$S(T) = \frac{1}{2} \sum_{i,j,k} T_{ijk} T_{kji} - \frac{\lambda}{4!\sqrt{N^3}} \sum_{ijklmn} T_{ijk} T_{klm} T_{mjn} T_{nli}$$



Feynman diagrams are stranded graphs dual to 3d simplicial complexes (nodes dual to tetrahedra, lines dual to triangles, faces dual to edges, 3-cells dual to vertices)



(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

Quantum dynamics (purely combinatorial - sum over random triangulations):

purely combinatorial - sum over random triangulations):
$$Z = \int \mathcal{D}T \, e^{-S(T,\lambda)} \, = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} \, Z_{\Gamma} = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} \, N^{F_{\Gamma} \, - \, \frac{3}{2} V_{\Gamma}}$$

can be recast in terms of Regge action for gravity discretised on equilateral triangulation

Pirsa: 16110048 Page 46/154

(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

Quantum dynamics (purely combinatorial - sum over random triangulations):

purely combinatorial - sum over random triangulations):
$$Z = \int \mathcal{D}T \, e^{-S(T,\lambda)} \, = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} \, Z_{\Gamma} = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} N^{F_{\Gamma} - \frac{3}{2}V_{\Gamma}}$$

can be recast in terms of Regge action for gravity discretised on equilateral triangulation

Random tensors — -> random geometries

Pirsa: 16110048 Page 47/154

(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

Quantum dynamics (purely combinatorial - sum over random triangulations):

purely combinatorial - sum over random triangulations):
$$Z = \int \mathcal{D}T \, e^{-S(T,\lambda)} \, = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} \, Z_{\Gamma} = \sum_{\Gamma} \frac{\lambda^{V_{\Gamma}}}{sym(\Gamma)} N^{F_{\Gamma} \, - \, \frac{3}{2} V_{\Gamma}}$$

can be recast in terms of Regge action for gravity discretised on equilateral triangulation

Random tensors — -> random geometries

most (combinatorial) results of tensor models also apply to GFTs

- use of colors (colored tensors) to encode topology
- large-N expansion
- double scaling
- universality of random tensors

.

(Ambjorn, Durhuus, Sasakura, ..., Gurau, Rivasseau, Bonzom, Ryan,)

Quantum dynamics (purely combinatorial - sum over random triangulations):

$$Z=\int \mathcal{D}T\,e^{-S(T,\lambda)}\,=\,\sum_{\Gamma}rac{\lambda^{V_{\Gamma}}}{sym(\Gamma)}\,Z_{\Gamma}=\sum_{\Gamma}rac{\lambda^{V_{\Gamma}}}{sym(\Gamma)}\,N^{F_{\Gamma}\,-\,rac{3}{2}V_{\Gamma}}$$

can be recast in terms of Regge action for gravity discretised on equilateral triangulation

Random tensors — -> random geometries

most (combinatorial) results of tensor models also apply to GFTs

- use of colors (colored tensors) to encode topology
- large-N expansion
- double scaling
 - universality of random tensors

.....

GFTs = tensor models + group data

richer models, richer dynamics

model building guided by simplicial geometry:

GFT quanta are discrete geometric structures with group-theoretic variables

Pirsa: 16110048 Page 50/154

model building guided by simplicial geometry:

GFT quanta are discrete geometric structures with group-theoretic variables

 $\varphi(q_1,q_2,q_3,q_4) \leftrightarrow \varphi(B_1,B_2,B_3,B_4) \rightarrow \mathbb{C}$ example: 4d quantum gravity

d = 4; G = local gauge group of gravity = <math>SO(3,1)(in riemannian signature, SO(4))

phase space before geometricity constraints:

$$[\mathcal{T}^*Spin(4)]^{\times 4} \simeq [\mathcal{T}^*SU(2) \times \mathcal{T}^*SU(2)]^{\times 4}$$

classical tetrahedron in 4d:

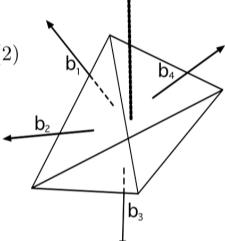
$$A_i\,n_i^I\,=\,b_i^I\in\mathbb{R}^4 \qquad b_i\cdot N\,=\,0 \qquad \sum_i\,b_i\,=\,0 \quad igg] \quad b_i\,\simeq\,\mathfrak{su}(2)$$



unique intrinsic geometry (up to rotations)

$$B_i^{IJ} \simeq N^I \wedge b_i^J$$

$$\left(B_i^{IJ} \in \wedge^2 \mathbb{R}^4 \simeq \mathfrak{so}(4) \,,\, N^I \in \mathbb{R}^4\right) \qquad N_I\left(*B_i^{IJ}\right) \,=\, 0 \qquad \sum_i B_i^{IJ} \,=\, 0$$



model building guided by simplicial geometry:

GFT quanta are discrete geometric structures with group-theoretic variables

 $\varphi(q_1,q_2,q_3,q_4) \leftrightarrow \varphi(B_1,B_2,B_3,B_4) \rightarrow \mathbb{C}$ example: 4d quantum gravity

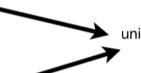
d = 4; G = local gauge group of gravity = <math>SO(3,1)(in riemannian signature, SO(4))

phase space before geometricity constraints:

$$[\mathcal{T}^*Spin(4)]^{\times 4} \simeq [\mathcal{T}^*SU(2) \times \mathcal{T}^*SU(2)]^{\times 4}$$

classical tetrahedron in 4d:

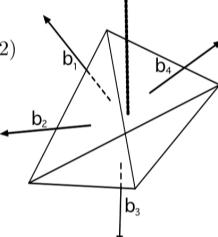
$$A_i\,n_i^I\,=\,b_i^I\in\mathbb{R}^4 \qquad b_i\cdot N\,=\,0 \qquad \sum_i\,b_i\,=\,0 \quad \Bigg] \quad b_i\,\simeq\,\mathfrak{su}(2)$$



unique intrinsic geometry (up to rotations)

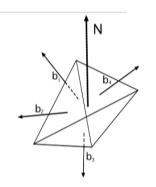
$$B_i^{IJ} \simeq N^I \wedge b_i^J$$

$$\left(B_i^{IJ} \in \wedge^2 \mathbb{R}^4 \simeq \mathfrak{so}(4) \,,\, N^I \in \mathbb{R}^4
ight) \qquad N_I \left(*B_i^{IJ}
ight) \,=\, 0 \qquad \sum_i B_i^{IJ} \,=\, 0
ight)$$



(in riemannian signature, SO(4))

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$



additional geometricity conditions

d = 4; G = SO(3,1)

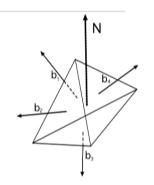
restrictions on Hilbert space

(can be imposed at dynamical level)

Pirsa: 16110048 Page 53/154

(in riemannian signature, SO(4))

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$



additional geometricity conditions

d = 4; G = SO(3,1)

restrictions on Hilbert space

(can be imposed at dynamical level)

Pirsa: 16110048 Page 54/154

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$

d = 4; G = SO(3,1) (in riemannian signature, SO(4))

additional geometricity conditions

restrictions on Hilbert space

(can be imposed at dynamical level)



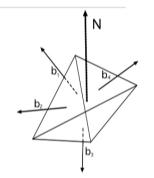
GFT quanta are discrete geometric simplices

this is candidate fundamental Hilbert space of dofs of universe

spacetime, geometry and matter fields should emerge from these quantum data

Pirsa: 16110048 Page 55/154

$$\mathcal{F}(\mathcal{H}_v) = \bigoplus_{V=0}^{\infty} sym \left\{ \left(\mathcal{H}_v^{(1)} \otimes \mathcal{H}_v^{(2)} \otimes \cdots \otimes \mathcal{H}_v^{(V)} \right) \right\}$$
$$\mathcal{H}_v = L^2 \left(G^d; d\mu_{Haar} \right)$$



d = 4; G = SO(3,1)

additional geometricity conditions

restrictions on Hilbert space

(can be imposed at dynamical level)

GFT quanta are discrete geometric simplices

this is candidate fundamental Hilbert space of dofs of universe

spacetime, geometry and matter fields should emerge from these quantum data

(in riemannian signature, SO(4))

Pirsa: 16110048 Page 56/154

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality"
in pairing of field arguments

+ appropriate "geometricity conditions"

Pirsa: 16110048 Page 57/154

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality" in pairing of field arguments

+ appropriate "geometricity conditions"

Pirsa: 16110048 Page 58/154

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality"
in pairing of field arguments

+ appropriate "geometricity conditions"

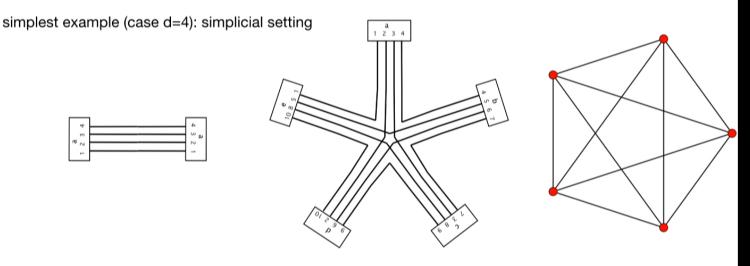
Pirsa: 16110048 Page 59/154

classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\bar{g}_{iD}) \mathcal{V}(g_{ia},\bar{g}_{iD}) + c.c.$$
"combinatorial non-locality"
in pairing of field arguments

specific combinatorics depends on model

+ appropriate "geometricity conditions"



Pirsa: 16110048 Page 60/154

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Feynman amplitudes (model-dependent):

equivalently:

• spin foam model (sum-over-histories of spin

networks ~ covariant LQG)

Reisenberger, Rovelli, '00

 lattice gravity path integral (with group+Lie algebra variables)

A. Baratin, DO, '11

Pirsa: 16110048 Page 61/154

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma}$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Feynman amplitudes (model-dependent):

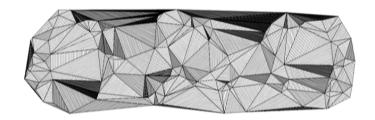
equivalently:

 spin foam model (sum-over-histories of spin networks ~ covariant LQG)

Reisenberger, Rovelli, '00

 lattice gravity path integral (with group+Lie algebra variables)

A. Baratin, DO, '11



Pirsa: 16110048 Page 62/154

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma} \, \bullet$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Feynman amplitudes (model-dependent):

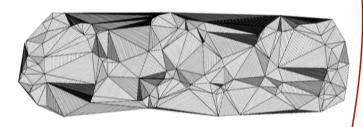
equivalently:

 spin foam model (sum-over-histories of spin networks ~ covariant LQG)

Reisenberger, Rovelli, '00

 lattice gravity path integral (with group+Lie algebra variables)

A. Baratin, DO, '11



GFT as lattice quantum gravity:

dynamical triangulations + quantum Regge calculus

Pirsa: 16110048 Page 63/154

Feynman perturbative expansion around trivial vacuum

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i \, S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \, \mathcal{A}_{\Gamma} \, \bullet$$

Feynman diagrams (obtained by convoluting propagators with interaction kernels) =

= stranded diagrams dual to cellular complexes of arbitrary topology

(simplicial case: simplicial complexes obtained by gluing d-simplices)

Feynman amplitudes (model-dependent):

equivalently:

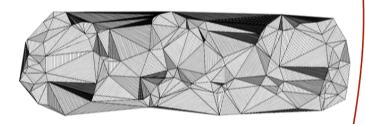
 spin foam model (sum-over-histories of spin networks ~ covariant LQG)

Reisenberger, Rovelli, '00

 lattice gravity path integral (with group+Lie algebra variables)

A. Baratin, DO, '11

discrete semiclassical limit --> Regge calculus



GFT as lattice quantum gravity:

dynamical triangulations + quantum Regge calculus

Pirsa: 16110048 Page 64/154

appropriate conditions on GFT fields or GFT dynamics (and choice of data) turn GFT Feynman amplitudes into lattice gauge theories/discrete gravity path integrals/spin foam models

e.g. gauge invariance of GFT fields under diagonal action of group G

example: d=3
$$\varphi_{\ell}: SO(3)^3/SO(3) \to \mathbb{R}$$
 + simplicial interaction $\forall h \in SO(3), \quad \varphi_{\ell}(hg_1, hg_2, hg_3) = \varphi_{\ell}(g_1, g_2, g_3)$ with only delta functions

valid for GFT definition of BF theory in any dimension

can be computed in different (equivalent) representations (group, spin, Lie algebra)

Pirsa: 16110048 Page 65/154

appropriate conditions on GFT fields or GFT dynamics (and choice of data) turn GFT Feynman amplitudes into lattice gauge theories/discrete gravity path integrals/spin foam models

e.g. gauge invariance of GFT fields under diagonal action of group G

example: d=3
$$\varphi_{\ell}: SO(3)^3/SO(3) \to \mathbb{R}$$
 + simplicial interaction $\forall h \in SO(3), \quad \varphi_{\ell}(hg_1,hg_2,hg_3) = \varphi_{\ell}(g_1,g_2,g_3)$ with only delta functions

valid for GFT definition of BF theory in any dimension

can be computed in different (equivalent) representations (group, spin, Lie algebra)

$$\mathcal{A}_{\Gamma} = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(H_{f}(h_{l})\right) = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(\overrightarrow{\prod}_{l \in \partial f} h_{l}\right) = \underbrace{\int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(\overrightarrow{\prod}_{l \in \partial f} h_{l}\right)}_{\text{3d gravity/BF theory}} = \sum_{\{j_{e}\}} \prod_{e} d_{j_{e}} \prod_{\tau} \left\{ \begin{array}{c} j_{1}^{\tau} & j_{2}^{\tau} & j_{3}^{\tau} \\ j_{4}^{\tau} & j_{5}^{\tau} & j_{6}^{\tau} \end{array} \right\} = \int \prod_{l} [\mathrm{d}h_{l}] \prod_{e} [\mathrm{d}^{3}x_{e}] \, e^{i\sum_{e} \mathrm{Tr}\,x_{e}H_{e}}$$

Pirsa: 16110048 Page 66/154

appropriate conditions on GFT fields or GFT dynamics (and choice of data) turn GFT Feynman amplitudes into lattice gauge theories/discrete gravity path integrals/spin foam models

e.g. gauge invariance of GFT fields under diagonal action of group G

example: d=3
$$\varphi_{\ell}: SO(3)^3/SO(3) \to \mathbb{R}$$
 + simplicial interaction $\forall h \in SO(3), \quad \varphi_{\ell}(hg_1, hg_2, hg_3) = \varphi_{\ell}(g_1, g_2, g_3)$ with only delta functions

valid for GFT definition of BF theory in any dimension

can be computed in different (equivalent) representations (group, spin, Lie algebra)

$$\mathcal{A}_{\Gamma} = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(H_{f}(h_{l})\right) = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(\overrightarrow{\prod}_{l \in \partial f} h_{l}\right) = \underbrace{\int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(\overrightarrow{\prod}_{l \in \partial f} h_{l}\right)}_{\text{3d gravity/BF theory}}$$

$$= \sum_{\{j_{e}\}} \prod_{e} d_{j_{e}} \prod_{\tau} \left\{ \begin{array}{c} j_{1}^{\tau} & j_{2}^{\tau} & j_{3}^{\tau} \\ j_{4}^{\tau} & j_{5}^{\tau} & j_{6}^{\tau} \end{array} \right\} = \int \prod_{l} [\mathrm{d}h_{l}] \prod_{e} [\mathrm{d}^{3}x_{e}] \, e^{i\sum_{e} \mathrm{Tr}\,x_{e}H_{e}}$$

spin foam formulation of 3d gravity/BF theory

Pirsa: 16110048 Page 67/154

appropriate conditions on GFT fields or GFT dynamics (and choice of data) turn GFT Feynman amplitudes into lattice gauge theories/discrete gravity path integrals/spin foam models

e.g. gauge invariance of GFT fields under diagonal action of group G

example: d=3
$$\varphi_{\ell}: SO(3)^3/SO(3) \to \mathbb{R}$$
 + simplicial interaction $\forall h \in SO(3), \quad \varphi_{\ell}(hg_1, hg_2, hg_3) = \varphi_{\ell}(g_1, g_2, g_3)$ with only delta functions

valid for GFT definition of BF theory in any dimension

can be computed in different (equivalent) representations (group, spin, Lie algebra)

$$\mathcal{A}_{\Gamma} = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(H_{f}(h_{l})\right) = \int \prod_{l} \mathrm{d}h_{l} \prod_{f} \delta\left(\overrightarrow{\prod}_{l \in \partial f} h_{l}\right) = \text{lattice gauge theory formulation of } 3\mathrm{d} \ \mathrm{gravity/BF} \ \mathrm{theory}$$

$$= \sum_{\{j_{e}\}} \prod_{e} d_{j_{e}} \prod_{\tau} \left\{ \begin{array}{c} j_{1}^{\mathsf{T}} & j_{2}^{\mathsf{T}} & j_{3}^{\mathsf{T}} \\ j_{4}^{\mathsf{T}} & j_{5}^{\mathsf{T}} & j_{6}^{\mathsf{T}} \end{array} \right\} = \int \prod_{l} [\mathrm{d}h_{l}] \prod_{e} [\mathrm{d}^{3}x_{e}] \, e^{i\sum_{e} \mathrm{Tr}\,x_{e}H_{e}}$$

$$\mathrm{discrete} \ \mathrm{1st} \ \mathrm{order} \ \mathrm{path} \ \mathrm{integral} \ \mathrm{for} \ \mathrm{3d} \ \mathrm{gravity/BF} \ \mathrm{theory}$$

$$\mathrm{on} \ \mathrm{simplicial} \ \mathrm{complex} \ \mathrm{dual} \ \mathrm{to} \ \mathrm{GFT} \ \mathrm{Feynman} \ \mathrm{diagram}$$

Pirsa: 16110048 Page 68/154

second quantized version of Loop Quantum Gravity but dynamics not derived from canonical quantization of GR

DO, 1310.7786 [gr-qc]

DO, J. Ryan, J. Thurigen, '14

(LQG spin network states ~ many-particles states, "particle" ~ spin network vertex)

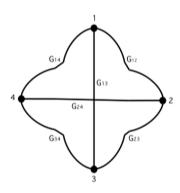
Pirsa: 16110048 Page 69/154

second quantized version of Loop Quantum Gravity but dynamics not derived from canonical quantization of GR

DO, 1310.7786 [gr-qc]

DO, J. Ryan, J. Thurigen, '14

(LQG spin network states ~ many-particles states, "particle" ~ spin network vertex)



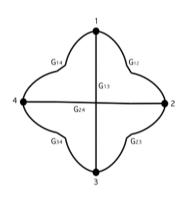
Pirsa: 16110048 Page 70/154

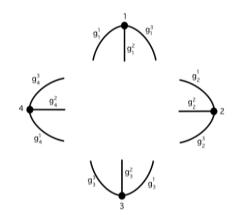
second quantized version of Loop Quantum Gravity but dynamics not derived from canonical quantization of GR

DO, 1310.7786 [gr-qc]

DO, J. Ryan, J. Thurigen, '14

(LQG spin network states ~ many-particles states, "particle" ~ spin network vertex)





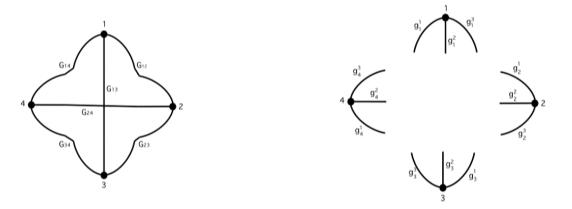
Pirsa: 16110048 Page 71/154

second quantized version of Loop Quantum Gravity but dynamics not derived from canonical quantization of GR

DO, 1310.7786 [gr-qc]

DO, J. Ryan, J. Thurigen, '14

(LQG spin network states ~ many-particles states, "particle" ~ spin network vertex)



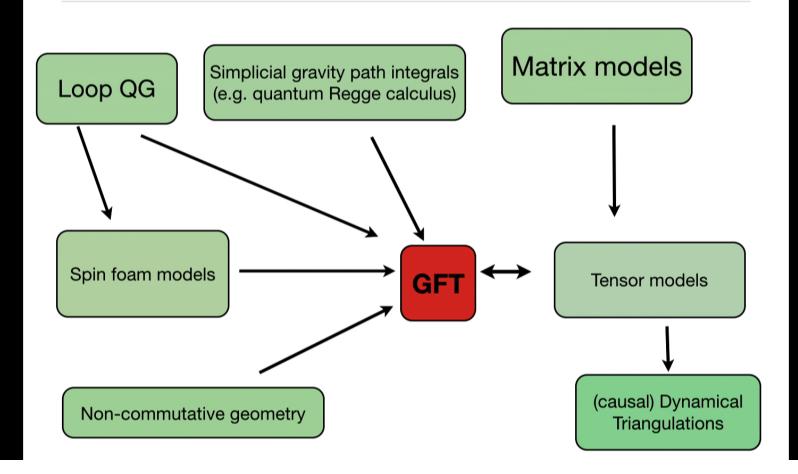
GFT Hilbert space = Fock space of open spin network vertices - contains any LQG state (all spin networks)

any LQG observable has a 2nd quantised, GFT counterpart

choice of LQG dynamics (Hamiltonian constraint operator) translates into choice of GFT action

Pirsa: 16110048 Page 72/154

Group Field Theory: crossroad of approaches



Pirsa: 16110048 Page 73/154



Pirsa: 16110048 Page 74/154

Part 2: the continuum limit of GFTs

GFT renormalization

Pirsa: 16110048 Page 75/154

Part 2: the continuum limit of GFTs

GFT renormalization

Pirsa: 16110048 Page 76/154

The problem of the continuum limit in QG

new (non-geometric, non-spatio-temporal) physical degrees of freedom ("building blocks") for space-time

new direction to explore: number of fundamental degrees of freedom

(quantum) continuum, geometric space-time should be recovered in the regime of large number N of non-spatio-temporal d.o.f.s

Pirsa: 16110048 Page 77/154

The problem of the continuum limit in QG

new (non-geometric, non-spatio-temporal) physical degrees of freedom ("building blocks") for space-time

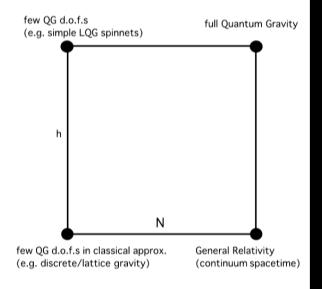
new direction to explore: number of fundamental degrees of freedom

(quantum) continuum, geometric space-time should be recovered in the regime of large number N of non-spatio-temporal d.o.f.s

continuum approximation very different (conceptually, technically) from classical approximation

N-direction
(collective behaviour of many interacting degrees of freedom):
continuum approximation

h-direction: classical approximation



Pirsa: 16110048 Page 78/154

The problem of the continuum limit in QG

new (non-geometric, non-spatio-temporal) physical degrees of freedom ("building blocks") for space-time

new direction to explore: number of fundamental degrees of freedom

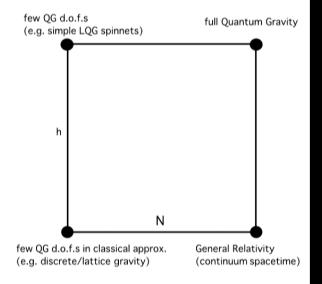
(quantum) continuum, geometric space-time should be recovered in the regime of large number N of non-spatio-temporal d.o.f.s

continuum approximation very different (conceptually, technically) from classical approximation

N-direction (collective behaviour of many interacting degrees of freedom): continuum approximation

h-direction: classical approximation

"well-understood" in spin foam models and discrete gravity



Pirsa: 16110048 Page 79/154

Problem of the continuum in QG: role of RG

Renormalization Group is crucial tool

for taking into account the physics of more and more d.o.f.s

- for our QG models, do not expect to have a unique continuum limit

 collective behaviour of (interacting) fundamental d.o.f.s should lead to different macroscopic phases,

 separated by phase transitions
- for a non-spatio-temporal QG system (e.g. LQG in GFT formulation), which of the macroscopic phases is described by a smooth geometry with matter fields?
- need to understand effective dynamics at different "GFT scales":
 RG flow of effective actions & phase structure & phase transitions
 _{Koslowski, '07; DO, '07}

Pirsa: 16110048 Page 80/154

Problem of the continuum in QG: role of RG

Renormalization Group is crucial tool

for taking into account the physics of more and more d.o.f.s

- for our QG models, do not expect to have a unique continuum limit

 collective behaviour of (interacting) fundamental d.o.f.s should lead to different macroscopic phases,

 separated by phase transitions
- for a non-spatio-temporal QG system (e.g. LQG in GFT formulation),
 which of the macroscopic phases is described by a smooth geometry with matter fields?
- need to understand effective dynamics at different "GFT scales":

 RG flow of effective actions & phase structure & phase transitions

 Koslowski, '07; DO, '07

many results in related formalisms:

- renormalization in SF models (~ lattice gauge theories) Dittrich, Bahr, Steinhaus, Martin-Benito,
- different (kinematical) phases in LQG Ashtekar-Lewandowski, Koslowski-Sahlmann, Dittrich-Geiller)
- phase diagrams (numerically) in (causal) dynamical triangulations Ambjorn, Loll, Jurkiewicz,

Pirsa: 16110048 Page 81/154

GFT renormalisation - general scheme

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i S_{\lambda}(\varphi, \overline{\varphi})} = \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \mathcal{A}_{\Gamma}$$
$$S(\varphi, \overline{\varphi}) = \frac{1}{2} \int [dg_{i}] \overline{\varphi(g_{i})} \mathcal{K}(g_{i}) \varphi(g_{i}) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\overline{g}_{iD}) \mathcal{V}(g_{ia}, \overline{g}_{iD}) + c.c.$$

general strategy:

treat GFTs as ordinary QFTs defined on Lie group manifold use group structures (Killing form, topology, etc) to define notion of scale and to set up mode integration subtleties of quantum gravity context at the level of interpretation

scales:

defined by propagator: e.g. spectrum of Laplacian on G = indexed by group representations

Pirsa: 16110048 Page 82/154

GFT renormalisation - general scheme

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i S_{\lambda}(\varphi, \overline{\varphi})} = \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \mathcal{A}_{\Gamma}$$
$$S(\varphi, \overline{\varphi}) = \frac{1}{2} \int [dg_{i}] \overline{\varphi(g_{i})} \mathcal{K}(g_{i}) \varphi(g_{i}) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \varphi(\overline{g}_{iD}) \mathcal{V}(g_{ia}, \overline{g}_{iD}) + c.c.$$

general strategy:

treat GFTs as ordinary QFTs defined on Lie group manifold use group structures (Killing form, topology, etc) to define notion of scale and to set up mode integration subtleties of quantum gravity context at the level of interpretation

scales:

defined by propagator: e.g. spectrum of Laplacian on G = indexed by group representations

- need to have control over "theory space" (e.g. via symmetries)

 A. Kegeles, DO, 15, 16
- main difficulty:
 controlling the combinatorics of GFT Feynman diagrams and interactions to control RG flow and divergences
 need to adapt/redefine many QFT notions: connectedness, subgraph contraction, Wick ordering,

Pirsa: 16110048 Page 83/154

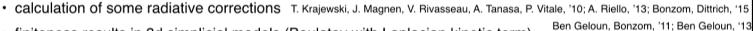


Pirsa: 16110048 Page 84/154

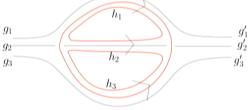
GFT perturbative renormalisation

step by step, towards renormalizable 4d gravity models:

- scale indexed by group representations
- interplay between algebraic data and combinatorics of diagrams



• finiteness results in 3d simplicial models (Boulatov with Laplacian kinetic term)



• renormalizable TGFT models (3d, 4d, and higher) - Laplacian + tensorial interactions

Ben Geloun, Rivasseau, '11 Carrozza, DO, Rivasseau, '12. '13

$$S(arphi,\overline{arphi}) = \sum_{b \in \mathcal{B}} t_b I_b(arphi,\overline{arphi})$$

Carrozza, DO, Rivasseau, '12. '13 -> with gauge invariance



--> SO(4) or SO(3,1) with simplicity constraints: first results on BC-like 4d models

———> generic (and robust?) asymptotic freedom

Ben Geloun. '12; Carrozza, '14

Lahoche, DO, '15; Carrozza, Lahoche, DO, '16

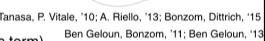
GFT perturbative renormalisation

step by step, towards renormalizable 4d gravity models:

- scale indexed by group representations
- interplay between algebraic data and combinatorics of diagrams



finiteness results in 3d simplicial models (Boulatov with Laplacian kinetic term)



· renormalizable TGFT models (3d, 4d, and higher) - Laplacian + tensorial interactions

Ben Geloun, Rivasseau, '11 Carrozza, DO, Rivasseau, '12, '13

$$S(arphi,\overline{arphi}) = \sum_{b \in \mathcal{B}} t_b I_b(arphi,\overline{arphi})$$

-> with gauge invariance

-> non-abelian (SU(2))

--> SO(4) or SO(3,1) with simplicity constraints: first results on BC-like 4d models

Lahoche, DO, '15; Carrozza, Lahoche, DO, '16 ---> generic (and robust?) asymptotic freedom Ben Geloun, '12; Carrozza, '14

many important lessons

(e.g. learnt to deal with combinatorics and topology of spin foam complex)

main open issues:

characterise better theory space (kinetic term, combinatorics of interactions, ...)

 g_2

 deal with non-group structures (due to Immirzi parameter) understand in full the geometric interpretation of UV/IR and of RG flow

Pirsa: 16110048 Page 86/154

GFT non-perturbative renormalisation

the GFT proposal:

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \mathcal{A}_{\Gamma}$$

controlling the continuum limit ~ evaluating GFT path integral (in some non-perturbative approximation)

Benedetti, Ben Geloun, DO, Martini, Lahoche, Carrozza, Ousmane-Samary, Duarte,

Freidel, Louapre, Noui, Magnen, Smerlak, Gurau, Rivasseau, Tanasa, Dartois, Delpouve,

Pirsa: 16110048

GFT non-perturbative renormalisation

$$\mathcal{Z} = \int \mathcal{D}\varphi \mathcal{D}\overline{\varphi} \ e^{i S_{\lambda}(\varphi, \overline{\varphi})} \quad = \quad \sum_{\Gamma} \frac{\lambda^{N_{\Gamma}}}{sym(\Gamma)} \mathcal{A}_{\Gamma}$$

controlling the continuum limit ~ evaluating GFT path integral (in some non-perturbative approximation)

two directions:

· GFT non-perturbative renormalization and "IR" fixed points (e.g. FRG analysis - e.g. a la Wetterich

Benedetti, Ben Geloun, DO, Martini, Lahoche, Carrozza, Ousmane-Samary, Duarte,

GFT constructive analysis

Freidel, Louapre, Noui, Magnen, Smerlak, Gurau, Rivasseau, Tanasa, Dartois, Delpouve,

non-perturbative resummation of perturbative (SF) series

variety of techniques:

- intermediate field method
- · loop-vertex expansion
- · Borel summability

FRG analysis of GFT models

D. Benedetti, J. Ben Geloun, DO, '14

regularised path integral: $\mathcal{Z}_k\left[J,\overline{J}\right] = e^{W_{k'}\left[J,\overline{J}\right]} = \int d\phi d\overline{\phi} \; e^{-S\left[\phi,\overline{\phi}\right] - \Delta S_{k'}\left[\phi,\overline{\phi}\right] + \mathrm{Tr}\left(J\cdot\overline{\phi}\right) + \mathrm{Tr}\left(\overline{J}\cdot\phi\right)}$

regulator cutting off IR modes (UV well-defined with appropriate choice of IR regulator)

$$\Delta S_{k}[\phi, \overline{\phi}] = \text{Tr}(\overline{\phi} \cdot R_{k} \cdot \phi) = \sum_{\mathbf{P}, \mathbf{P}'} \overline{\phi}_{\mathbf{P}} R_{k}(\mathbf{P}; \mathbf{P}') \phi_{\mathbf{P}'}$$

$$R_{k}(\mathbf{p}, \mathbf{p}') = \theta(k^{2} - \Sigma_{s} \rho_{s}^{2}) Z_{k}(k^{2} - \Sigma_{s} \rho_{s}^{2}) \delta(\mathbf{p} - \mathbf{p}')$$

 $\text{effective action:} \qquad \Gamma_{\textstyle k}[\varphi,\overline{\varphi}] = \sup_{J,\overline{J}} \biggl\{ \mathrm{Tr}(J\cdot\overline{\varphi}) + \mathrm{Tr}(\overline{J}\cdot\varphi) - W_{\textstyle k}[J,\overline{J}] - \Delta S_{\textstyle k}[\varphi,\overline{\varphi}] \biggr\}$

Wetterich equation:

$$\left\{ \partial_t \mathsf{\Gamma}_k = ext{Tr} [\partial_t R_k \cdot (\mathsf{\Gamma}_k^{(2)} + R_k)^{-1}]
ight.
ight\} t = \log k$$

boundary conditions: $\Gamma_{k=0}[\varphi,\overline{\varphi}] = \Gamma[\varphi,\overline{\varphi}], \qquad \Gamma_{k=\Lambda}[\varphi,\overline{\varphi}] = S[\varphi,\overline{\varphi}] \qquad \varphi = \langle \phi \rangle$

computing the effective action solving the Wetterich equation amounts to solving the GFT path integral

need truncation of effective action up to some order of interactions

FRG analysis of GFT models

D. Benedetti, J. Ben Geloun, DO, '14

regularised path integral: $\mathcal{Z}_{k}\left[J,\overline{J}\right] = e^{W_{k'}\left[J,\overline{J}\right]} = \int d\phi d\overline{\phi} \; e^{-S\left[\phi,\overline{\phi}\right] - \Delta S_{k'}\left[\phi,\overline{\phi}\right] + \mathrm{Tr}\left(J\cdot\overline{\phi}\right) + \mathrm{Tr}\left(\overline{J}\cdot\phi\right)}$

regulator cutting off IR modes (UV well-defined with appropriate choice of IR regulator)

$$\Delta S_{k}[\phi, \overline{\phi}] = \text{Tr}(\overline{\phi} \cdot R_{k} \cdot \phi) = \sum_{\mathbf{P}, \mathbf{P}'} \overline{\phi}_{\mathbf{P}} R_{k}(\mathbf{P}; \mathbf{P}') \phi_{\mathbf{P}'}$$

$$R_{k}(\mathbf{p}, \mathbf{p}') = \theta(k^{2} - \Sigma_{s} \rho_{s}^{2}) Z_{k}(k^{2} - \Sigma_{s} \rho_{s}^{2}) \delta(\mathbf{p} - \mathbf{p}')$$

 $\text{effective action:} \qquad \Gamma_{\underline{k}}[\varphi,\overline{\varphi}] = \sup_{J,\overline{J}} \biggl\{ \mathrm{Tr}(J \cdot \overline{\varphi}) + \mathrm{Tr}(\overline{J} \cdot \varphi) - W_{\underline{k}}[J,\overline{J}] - \Delta S_{\underline{k}}[\varphi,\overline{\varphi}] \biggr\}$

Wetterich equation:

$$\left\{ \partial_t \mathsf{\Gamma}_k = ext{Tr} [\partial_t R_k \cdot (\mathsf{\Gamma}_k^{(2)} + R_k)^{-1}]
ight.
ight\} t = \log k$$

boundary conditions: $\Gamma_{k=0}[\varphi,\overline{\varphi}] = \Gamma[\varphi,\overline{\varphi}], \qquad \Gamma_{k=\Lambda}[\varphi,\overline{\varphi}] = S[\varphi,\overline{\varphi}] \qquad \varphi = \langle \phi \rangle$

computing the effective action solving the Wetterich equation amounts to solving the GFT path integral

need truncation of effective action up to some order of interactions

GFT non-perturbative renormalisation

recent results:

FRG for (tensorial) GFT models

Pirsa: 16110048 Page 91/154

GFT non-perturbative renormalisation

recent results:

FRG for (tensorial) GFT models

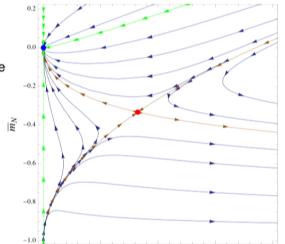
(similar to matrix model but distinctively field-theoretic)

Benedetti, Lahoche, '15; Duarte, DO, '16

Eichhorn, Koslowski, '14

- Polchinski formulation based on SD equations
- Krajewski, Toriumi, '14
- · general set-up for Wetterich formulation based on effective action
 - analysis of TGFT on compact U(1)^d
 - · RG flow and phase diagram established
 - · analysis of TGFT on non-compact R^d
 - · RG flow and phase diagram established
 - analysis of TGFT on non-compact R^d with gauge invariance
 - · RG flow and phase diagram established
 - analysis of TGFT on SU(2)^3 Carrozza, Lahoche, '16

generically (so far):
two FPs (Gaussian-UV, Wilson-Fisher-IR)
asymptotic freedom
one symmetric phase
one broken or condensate phase



 $\overline{\lambda}_N$

Benedetti, Ben Geloun, DO, '14; Ben Geloun, Martini, DO, '15, '16,

Pirsa: 16110048 Page 92/154



Pirsa: 16110048 Page 93/154

Part IV: effective cosmology from GFTs Cosmology as QG hydrodynamics

Pirsa: 16110048 Page 94/154

Quantum spacetime: the difficult path from microstructure to cosmology

the issue:

identify relevant phase for effective continuum geometry extract effective continuum dynamics and relate it to GR

Quantum Gravity problem:

identify microscopic d.o.f. of quantum spacetime and their fundamental dynamics



derive effective (QG-inspired) models for fundamental (quantum) cosmology: explain features of early Universe, obtain testable QG predictions

various models: loop quantum cosmology,

also work by:

C. Rovelli, F. Vidotto (perturbative GFT (spin foam) context); E. Alesci, F. Cianfrani (canonical LQG context);

Pirsa: 16110048 Page 95/154

What is cosmology, then?

two views on quantum gravity:

- 1. quantum gravity = quantum theory of gravitational field ~ quantum General Relativity
- 2. quantum gravity = microscopic theory of pre-geometric quantum degrees of freedom ("quantum (field) theory of atoms of space")



gravitational field result of collective dynamics spacetime and geometry are emergent entities

in case 2.

cosmology is necessarily to be understood as result of coarse graining of microscopic dofs up to global observables only, homogeneous sector of (quantum) GR

(quantum) cosmological degrees of freedom governed by statistical distribution not quantum theory of homogeneous geometries

cosmological dynamics to be looked for in the hydrodynamic approximation of full quantum gravity (most macroscopic, coarse grained, global description of the microscopic pre-geometric system)

Pirsa: 16110048 Page 96/154

What is cosmology, then?

two views on quantum gravity:

- 1. quantum gravity = quantum theory of gravitational field ~ quantum General Relativity
- 2. quantum gravity = microscopic theory of pre-geometric quantum degrees of freedom ("quantum (field) theory of atoms of space")



gravitational field result of collective dynamics spacetime and geometry are emergent entities

in case 2.

cosmology is necessarily to be understood as result of coarse graining of microscopic dofs up to global observables only, homogeneous sector of (quantum) GR

(quantum) cosmological degrees of freedom governed by statistical distribution not quantum theory of homogeneous geometries

cosmological dynamics to be looked for in the hydrodynamic approximation of full quantum gravity (most macroscopic, coarse grained, global description of the microscopic pre-geometric system)

Pirsa: 16110048 Page 97/154

Cosmology as hydrodynamics of (quantum) spacetime

re-thinking the "Cosmological Principle":

"every point is equivalent to any other" ~ homogeneity of space

really means: a certain approximation is assumed valid:

universe is in state where inhomogeneities can be neglected, in relation to dynamics of homogeneous modes

~ universe is in state where effects on largest wavelengths of shorter wavelengths is negligible

~ can neglect wavelengths (much) shorter than scale factor

very similar in spirit to hydrodynamic approximation:

dynamics of microscopic degrees of freedom can be neglected + effects of small wavelengths can be neglected

degrees of freedom of local region can describe whole of system (in a coarse grained, statistical sense)

i.e. whole universe (dynamics) well-approximated by local patch (dynamics)

Pirsa: 16110048 Page 98/154

Cosmology as hydrodynamics of (quantum) spacetime

re-thinking the "Cosmological Principle":

"every point is equivalent to any other" ~ homogeneity of space

really means: a certain approximation is assumed valid:

universe is in state where inhomogeneities can be neglected, in relation to dynamics of homogeneous modes

~ universe is in state where effects on largest wavelengths of shorter wavelengths is negligible

~ can neglect wavelengths (much) shorter than scale factor

very similar in spirit to hydrodynamic approximation:

dynamics of microscopic degrees of freedom can be neglected + effects of small wavelengths can be neglected

degrees of freedom of local region can describe whole of system (in a coarse grained, statistical sense)

i.e. whole universe (dynamics) well-approximated by local patch (dynamics)

end result:

basic variable is "fluid density" with arguments the geometric data of minisuperspace

cosmology is (non-linear) dynamics for such density and for geometric (global) observables computed from it

Pirsa: 16110048 Page 99/154

key strategy:

coarse graining of QG configurations



coarse graining of QG (quantum) dynamics

Pirsa: 16110048 Page 100/154

key strategy:

coarse graining of QG configurations



coarse graining of QG (quantum) dynamics

very difficult in general
(see comparatively simpler problem of coarse graining classical GR)
(see also analogous problem in condensed matter theory)

Pirsa: 16110048 Page 101/154

key strategy:

coarse graining of QG configurations



coarse graining of QG (quantum) dynamics

very difficult in general
(see comparatively simpler problem of coarse graining classical GR)
(see also analogous problem in condensed matter theory)

Pirsa: 16110048 Page 102/154

key strategy:

coarse graining of QG configurations



coarse graining of QG (quantum) dynamics

very difficult in general
(see comparatively simpler problem of coarse graining classical GR)
(see also analogous problem in condensed matter theory)

one special case:

quantum condensates (BEC)

effective hydrodynamics directly read out of microscopic quantum dynamics (in simplest approximation)

Pirsa: 16110048 Page 103/154

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14; S. Gielen, '14; S. Gielen, '15; DO, L. Sindoni, E. Wilson-Ewing, '16; M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16; M. De Cesare, A. Pithis, M. Sakellariadou, '16

Pirsa: 16110048 Page 104/154

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14; S. Gielen, '14; S. Gielen, '15; DO, L. Sindoni, E. Wilson-Ewing, '16; M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16; M. De Cesare, A. Pithis, M. Sakellariadou, '16

problem 1:

identify quantum states in fundamental theory with continuum spacetime interpretation

Pirsa: 16110048 Page 105/154

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14; S. Gielen, '14; S. Gielen, '15; DO, L. Sindoni, E. Wilson-Ewing, '16; M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16; M. De Cesare, A. Pithis, M. Sakellariadou, '16

problem 1:

identify quantum states in fundamental theory with continuum spacetime interpretation

Quantum GFT condensates are continuum homogeneous (quantum) spaces

Pirsa: 16110048 Page 106/154

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14; S. Gielen, '14; S. Gielen, '15; DO, L. Sindoni, E. Wilson-Ewing, '16; M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16; M. De Cesare, A. Pithis, M. Sakellariadou, '16

problem 1:

identify quantum states in fundamental theory with continuum spacetime interpretation

Quantum GFT condensates are continuum homogeneous (quantum) spaces

e.g. (simplest):
$$|\sigma\rangle := \exp\left(\hat{\sigma}\right)|0
angle$$

GFT field coherent state
$$\hat{\sigma} := \int d^4 g \; \sigma(g_I) \hat{arphi}^\dagger(g_I) \;\;\;\; \sigma(g_I k) \; = \; \sigma(g_I)$$

superposition of infinitely many SN dofs

 $\sigma\left(\mathcal{D}\right)$ described by single collective wave function (depending on homogeneous anisotropic geometric data)

 $\{ \mbox{geometries of tetrahedron} \} \simeq \\ \{ \mbox{continuum spatial geometries at a point} \} \simeq \\ \mbox{minisuperspace of homogeneous geometries}$

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14; S. Gielen, '14; S. Gielen, '15; DO, L. Sindoni, E. Wilson-Ewing, '16; M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16; M. De Cesare, A. Pithis, M. Sakellariadou, '16

problem 1:

identify quantum states in fundamental theory with continuum spacetime interpretation

Quantum GFT condensates are continuum homogeneous (quantum) spaces

e.g. (simplest):
$$|\sigma\rangle := \exp\left(\hat{\sigma}\right)|0
angle$$

GFT field coherent state
$$\hat{\sigma} := \int d^4 g \; \sigma(g_I) \hat{arphi}^\dagger(g_I) \;\;\;\; \sigma(g_I k) \; = \; \sigma(g_I)$$

superposition of infinitely many SN dofs

 $\sigma\left(\mathcal{D}\right)\qquad\mathcal{D}$ described by single collective wave function (depending on homogeneous anisotropic geometric data)

 $\{\text{geometries of tetrahedron}\} \simeq$ $\{\text{continuum spatial geometries at a point}\} \simeq$ minisuperspace of homogeneous geometries

Homogeneous geometries & GFT condensates

Pirsa: 16110048 Page 109/154

Homogeneous geometries & GFT condensates

• lift homogeneity criterion to quantum level (and include conjugate information):

all GFT quanta have the same (gauge invariant) "wave function", i.e. are in the same quantum state

$$\Psi(B_{i(1)},....,B_{i(N)}) = \frac{1}{N!} \prod_{m=1}^{N} \Phi(B_{i}(m))$$

- in GFT: such states can be expressed in 2nd quantized language and one can consider superpositions of states of arbitrary N
- sending N to infinity means improving arbitrarily the accuracy of the sampling



quantum GFT condensates are continuum homogeneous (quantum) spaces

Pirsa: 16110048 Page 110/154

follow closely procedure used in real BECs

single-particle GFT condensate:

$$|\sigma\rangle := \exp\left(\hat{\sigma}\right)|0\rangle$$

$$|\sigma\rangle := \exp{(\hat{\sigma})} |0\rangle$$
 $\hat{\sigma} := \int d^4 g \, \sigma(g_I) \hat{\varphi}^{\dagger}(g_I)$ $\sigma(g_I k) = \sigma(g_I)$

$$\sigma(g_I k) = \sigma(g_I)$$

S. Gielen, DO, L. Sindoni,

PRL, arXiv:1303.3576 [gr-qc];

JHEP, arXiv:1311.1238 [gr-qc]

superposition of infinitely many SN dofs

Pirsa: 16110048 Page 111/154

follow closely procedure used in real BECs

S. Gielen, DO, L. Sindoni,

PRL, arXiv:1303.3576 [gr-qc];

JHEP, arXiv:1311.1238 [gr-qc]

superposition of infinitely many SN dofs

single-particle GFT condensate:

$$|\sigma\rangle := \exp(\hat{\sigma})|0\rangle$$

$$|\sigma\rangle := \exp{(\hat{\sigma})} |0\rangle$$
 $\hat{\sigma} := \int d^4g \ \sigma(g_I) \hat{\varphi}^\dagger(g_I)$ $\sigma(g_I k) = \sigma(g_I)$

$$\sigma(g_I k) = \sigma(g_I)$$

from truncation of SD equations for GFT model applied to (coherent) GFT condensate state. gives equation for "wave function":

$$\int [dg_i'] \,\tilde{\mathcal{K}}(g_i, g_i') \sigma(g_i') + \lambda \frac{\delta \tilde{\mathcal{V}}}{\delta \varphi(g_i)}|_{\varphi \equiv \sigma} = 0$$

basically (up to some approximations), the "classical GFT eqns"

similar equations to M. Bojowald et al., arXiv:1210.8138 [gr-qc]

follow closely procedure used in real BECs

S. Gielen, DO, L. Sindoni,

PRL, arXiv:1303.3576 [gr-qc];

JHEP, arXiv:1311.1238 [gr-qc]

superposition of infinitely many SN dofs

single-particle GFT condensate:

$$|\sigma\rangle := \exp{(\hat{\sigma})} |0\rangle$$
 $\hat{\sigma} := \int d^4g \ \sigma(g_I) \hat{\varphi}^\dagger(g_I)$ $\sigma(g_I k) = \sigma(g_I)$

$$\sigma(g_I k) \, = \, \sigma(g_I)$$

from truncation of SD equations for GFT model applied to (coherent) GFT condensate state. gives equation for "wave function":

$$\int [dg_i'] \,\tilde{\mathcal{K}}(g_i, g_i') \sigma(g_i') + \lambda \frac{\delta \tilde{\mathcal{V}}}{\delta \varphi(g_i)}|_{\varphi \equiv \sigma} = 0$$

no perturbative (spin foam) expansion infinite superposition of SF amplitudes

basically (up to some approximations), the "classical GFT egns"

similar equations to M. Bojowald et al., arXiv:1210.8138 [gr-qc]

follow closely procedure used in real BECs

S. Gielen, DO, L. Sindoni,

PRL, arXiv:1303.3576 [gr-qc];

JHEP, arXiv:1311.1238 [gr-qc]

superposition of infinitely

many SN dofs

single-particle GFT condensate:

$$|\sigma\rangle := \exp(\hat{\sigma})|0\rangle$$

$$|\sigma\rangle := \exp{(\hat{\sigma})} |0\rangle$$
 $\hat{\sigma} := \int d^4g \ \sigma(g_I) \hat{\varphi}^{\dagger}(g_I)$ $\sigma(g_I k) = \sigma(g_I)$

$$\sigma(g_I k) = \sigma(g_I)$$

from truncation of SD equations for GFT model applied to (coherent) GFT condensate state. gives equation for "wave function":

$$\int [dg_i'] \,\tilde{\mathcal{K}}(g_i, g_i') \sigma(g_i') + \lambda \frac{\delta \tilde{\mathcal{V}}}{\delta \varphi(g_i)}|_{\varphi \equiv \sigma} = 0$$

no perturbative (spin foam) expansion infinite superposition of SF amplitudes

basically (up to some approximations), the "classical GFT egns"

similar equations to M. Bojowald et al., arXiv:1210.8138 [gr-qc]

non-linear and non-local extension of quantum cosmology-like equation for "collective wave function"

QG (GFT) analogue of Gross-Pitaevskii hydrodynamic equation in BECs

Cosmology from GFT condensates

summary of recent results:

· general scheme, geometric interpretation and effective dynamics (GP and dipole condensates)

S. Gielen, DO, L. Sindoni, '13

generalised condensate states (also for spherical black holes)

DO, D. Pranzetti, J. Ryan, L. Sindoni, '15; DO, D. Pranzetti, L. Sindoni, '15

 lattice refinement and GFT cosmological observables S. Gielen, DO, '14

 relation with LQC S. Gielen, '14, '15, '16; G. Calcagni, '14

effective cosmological dynamics from EPRL model

DO, L. Sindoni, E. Wilson-Ewing, '16;

M. De Cesare, M. Sakellariadou, '16; S. Gielen, '16

- · isotropic reduction, scalar field coupling, relational observables
- generalised Friedmann equations
- · generic big bounce resolution of classical singularity
- reduction to LQC dynamics
- · effect of GFT interaction in emergent cosmological dynamics
 - long-lasting acceleration after bounce (no inflation)

M. De Cesare, A. Pithis, M. Sakellariadou, '16

• non-normalisable condensate states (hints of GFT phase transition?)

A. Pithis, M. Sakellariadou, P. Tomov, '16

· first steps with cosmological perturbations S. Gielen, '14, '15

Pirsa: 16110048 Page 115/154

DO, Sindoni, Wilson-Ewing, '16

 starting from (generalised) EPRL model for 4d Lorentzian QG (simplicial interactions, G=SU(2), dynamics encodes embedding into SL(2,C) ~ simplicity constraints)

Engle, Pereira, Rovelli, Livine, '07; Freidel, Krasnov, '07

Pirsa: 16110048 Page 116/154

DO, Sindoni, Wilson-Ewing, '16

 starting from (generalised) EPRL model for 4d Lorentzian QG (simplicial interactions, G=SU(2), dynamics encodes embedding into SL(2,C) ~ simplicity constraints)

Engle, Pereira, Rovelli, Livine, '07; Freidel, Krasnov, '07

• coupling of free massless scalar field (+ truncation at lowest order ~ slowly varying field)

$$\hat{\varphi}(g_{v}) \rightarrow \hat{\varphi}(g_{v}, \phi)$$
 $K_{2}(g_{v_{1}}, g_{v_{2}}, \phi_{1}, \phi_{2}) = K_{2}(g_{v_{1}}, g_{v_{2}}, (\phi_{1} - \phi_{2})^{2})$

$$\mathcal{V}_5(g_{m{v_a}},\phi_{m{a}})=\mathcal{V}_5(g_{m{v_a}})\prod\delta(\phi_{m{a}}-\phi_1)$$

Pirsa: 16110048 Page 117/154

DO, Sindoni, Wilson-Ewing, '16

 starting from (generalised) EPRL model for 4d Lorentzian QG (simplicial interactions, G=SU(2), dynamics encodes embedding into SL(2,C) ~ simplicity constraints)

Engle, Pereira, Rovelli, Livine, '07; Freidel, Krasnov, '07

• coupling of free massless scalar field (+ truncation at lowest order ~ slowly varying field)

$$\hat{\varphi}(g_{v}) \rightarrow \hat{\varphi}(g_{v}, \phi)$$
 $K_{2}(g_{v_{1}}, g_{v_{2}}, \phi_{1}, \phi_{2}) = K_{2}(g_{v_{1}}, g_{v_{2}}, (\phi_{1} - \phi_{2})^{2})$

$$\mathcal{V}_{5}(g_{oldsymbol{v_a}},\phi_{oldsymbol{a}})=\mathcal{V}_{5}(g_{oldsymbol{v_a}})\prod\delta(\phi_{oldsymbol{a}}-\phi_{1})$$

• reduction to isotropic condensate configurations (depending on single spin variable j):

$$|\sigma
angle \sim \exp\left(\int \mathrm{d} g_{f v} \mathrm{d} \phi \ \sigma(g_{f v},\phi) \hat{\phi}^\dagger(g_{f v},\phi)
ight) |{f 0}
angle \qquad \sigma(g_{f v},\phi)
ightarrow \sigma_j(\phi)$$

Pirsa: 16110048 Page 118/154

DO, Sindoni, Wilson-Ewing, '16

 starting from (generalised) EPRL model for 4d Lorentzian QG (simplicial interactions, G=SU(2), dynamics encodes embedding into SL(2,C) ~ simplicity constraints)

Engle, Pereira, Rovelli, Livine, '07; Freidel, Krasnov, '07

coupling of free massless scalar field (+ truncation at lowest order ~ slowly varying field)

$$\hat{\varphi}(g_{v}) \rightarrow \hat{\varphi}(g_{v}, \phi)$$
 $K_{2}(g_{v_{1}}, g_{v_{2}}, \phi_{1}, \phi_{2}) = K_{2}(g_{v_{1}}, g_{v_{2}}, (\phi_{1} - \phi_{2})^{2})$

$$\mathcal{V}_5(g_{v_a},\phi_a) = \mathcal{V}_5(g_{v_a}) \prod \delta(\phi_a - \phi_1)$$

reduction to isotropic condensate configurations (depending on single spin variable j):

$$|\sigma
angle \sim \exp\left(\int \mathrm{d} g_{f v} \mathrm{d} \phi \ \sigma(g_{f v},\phi) \hat{\phi}^\dagger(g_{f v},\phi)
ight) |{f 0}
angle \qquad \sigma(g_{f v},\phi)
ightarrow \sigma_j(\phi)$$

effective condensate hydrodynamics (non-linear quantum cosmology):

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

functions A, B, w define the details of the EPRL model

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_j(\phi) = \rho_j(\phi)e^{i\theta_j(\phi)}$$

$$m_j^2 = B_j/A_j$$

$$\qquad \left[\rho_j'' - \frac{Q_j^2}{\rho_j^3} - m_j^2 \rho_j \approx 0 \right]$$

 $E_j=A_j|\partial_\phi\sigma_j(\phi)|^2-B_j|\sigma_j(\phi)|^2+rac{2}{5}{
m Re}\left(w_j\sigma_j(\phi)^5
ight)$

$$Q_{j} = -\frac{i}{2} \left[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \right]$$

$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

· key relational observables (expectation values in condensate state) with scalar field as clock:

universe volume (at fixed "time")

$$V(\phi) = \sum_{j} V_j \bar{\sigma}_j(\phi) \sigma_j(\phi) = \sum_{j} V_j \rho_j(\phi)^2 \qquad V_j \sim j^{3/2} \ell_{\rm Pl}^3$$

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_j(\phi) = \rho_j(\phi)e^{i\theta_j(\phi)}$$

onsistent with simple approximation of vacuum state $E_j=A_j|\partial_\phi\sigma_j(\phi)|^2-B_j|\sigma_j(\phi)|^2+rac{2}{5}{
m Re}\left(w_j\sigma_j(\phi)^5
ight)$

$$Q_{j} = -\frac{i}{2} \left[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \right]$$

$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

· key relational observables (expectation values in condensate state) with scalar field as clock:

universe volume (at fixed "time")

$$V(\phi) = \sum_{j} V_j \bar{\sigma}_j(\phi) \sigma_j(\phi) = \sum_{j} V_j \rho_j(\phi)^2 \qquad V_j \sim j^{3/2} \ell_{\rm Pl}^3$$

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_j(\phi) = \rho_j(\phi)e^{i\theta_j(\phi)}$$

 $E_{j} = A_{j} |\partial_{\phi} \sigma_{j}(\phi)|^{2} - B_{j} |\sigma_{j}(\phi)|^{2} + \frac{2}{5} \operatorname{Re} \left(w_{j} \sigma_{j}(\phi)^{5} \right)$

$$Q_{j} = -\frac{i}{2} \Big[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \Big]$$

$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

· key relational observables (expectation values in condensate state) with scalar field as clock:

universe volume (at fixed "time")

$$V(\phi) = \sum_{j} V_{j} \bar{\sigma}_{j}(\phi) \sigma_{j}(\phi) = \sum_{j} V_{j} \rho_{j}(\phi)^{2} \qquad V_{j} \sim j^{3/2} \ell_{\text{Pl}}^{3}$$

momentum of scalar field (at fixed "time") $\pi_{\phi} = \langle \sigma | \hat{\pi}_{\phi}(\phi) | \sigma \rangle = \hbar \sum_{i} Q_{i}$

constant of motion ~ continuity equation

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_i(\phi) = \rho_i(\phi)e^{i\theta_j(\phi)}$$

$$m_j^2 = B_j/A_j \qquad \qquad \boxed{ \rho_j'' - \frac{Q_j^2}{\rho_j^3} - m_j^2 \rho_j \approx 0 }$$

 $E_j = A_j |\partial_{\phi} \sigma_j(\phi)|^2 - B_j |\sigma_j(\phi)|^2 + \frac{2}{5} \operatorname{Re} \left(w_j \sigma_j(\phi)^5 \right)$

$$Q_{j} = -\frac{i}{2} \left[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \right]$$

$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

· key relational observables (expectation values in condensate state) with scalar field as clock:

universe volume (at fixed "time")

$$V(\phi) = \sum_{j} V_{j} \bar{\sigma}_{j}(\phi) \sigma_{j}(\phi) = \sum_{j} V_{j} \rho_{j}(\phi)^{2} \qquad V_{j} \sim j^{3/2} \ell_{\text{Pl}}^{3}$$

momentum of scalar field (at fixed "time") $\pi_{\phi} =$

$$\pi_{\phi} = \langle \sigma | \hat{\pi}_{\phi}(\phi) | \sigma \rangle = \hbar \sum_{j} Q_{j}$$

constant of motion ~ continuity equation

energy density of scalar field (at fixed "time")

$$\rho = \frac{\pi_{\phi}^2}{2V^2} = \frac{\hbar^2 (\sum_j Q_j)^2}{2(\sum_j V_j \rho_j^2)^2}$$

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_i(\phi) = \rho_i(\phi)e^{i\theta_j(\phi)}$$

$$m_j^2 = B_j/A_j \qquad \qquad \boxed{\rho_j'' - \frac{Q_j^2}{\rho_j^3} - m_j^2 \rho_j \approx 0}$$

$$E_{j} = A_{j} |\partial_{\phi} \sigma_{j}(\phi)|^{2} - B_{j} |\sigma_{j}(\phi)|^{2} + \frac{2}{5} \operatorname{Re} \left(w_{j} \sigma_{j}(\phi)^{5} \right)$$
$$Q_{j} = -\frac{i}{2} \left[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \right]$$

$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

· key relational observables (expectation values in condensate state) with scalar field as clock:

universe volume (at fixed "time")

$$V(\phi) = \sum_{j} V_j \bar{\sigma}_j(\phi) \sigma_j(\phi) = \sum_{j} V_j \rho_j(\phi)^2 \qquad V_j \sim j^{3/2} \ell_{\text{Pl}}^3$$

momentum of scalar field (at fixed "time") $\pi_{\phi} =$

$$\pi_{\phi} = \langle \sigma | \hat{\pi}_{\phi}(\phi) | \sigma \rangle = \hbar \sum_{j} Q_{j}$$

constant of motion ~ continuity equation

energy density of scalar field (at fixed "time")

$$\rho = \frac{\pi_{\phi}^2}{2V^2} = \frac{\hbar^2 (\sum_j Q_j)^2}{2(\sum_j V_j \rho_j^2)^2}$$

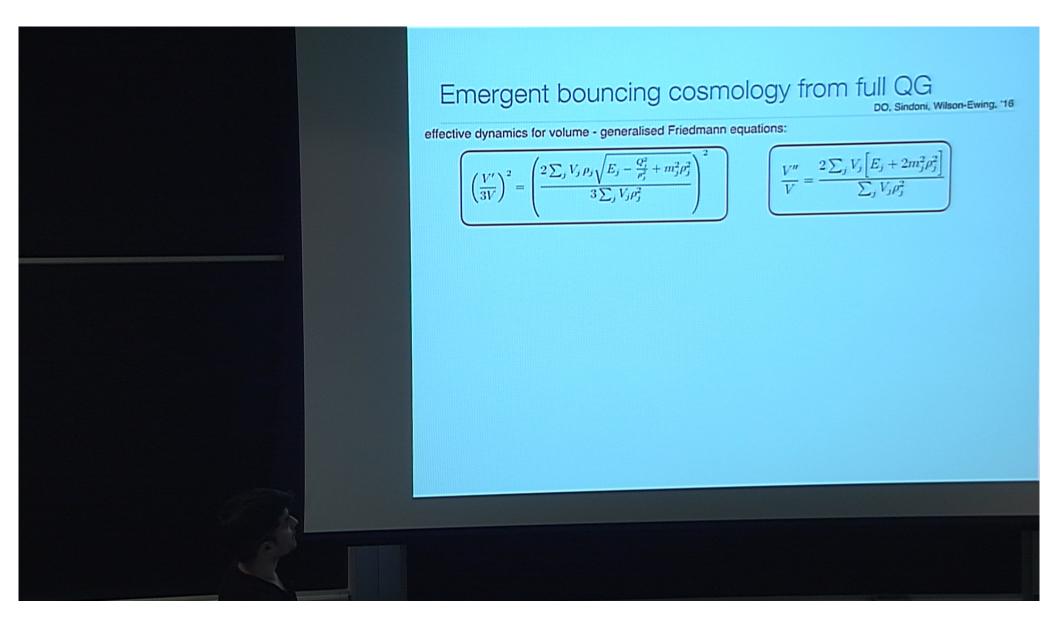
DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$

Pirsa: 16110048 Page 125/154



Pirsa: 16110048 Page 126/154

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$

Pirsa: 16110048 Page 127/154

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_j \left[E_j + 2m_j^2 \rho_j^2 \right]}{\sum_{j} V_j \rho_j^2}$$

 $\exists j / \rho_j(\phi) \neq 0 \ \forall \phi \longrightarrow V = \sum_j V_j \rho_j^2$

$$V = \sum_{j} V_{j} \hat{\rho}_{j}^{2}$$

remains positive at all times

generic quantum bounce!

Pirsa: 16110048 Page 128/154

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_j \left[E_j + 2m_j^2 \rho_j^2 \right]}{\sum_{j} V_j \rho_j^2}$$

Pirsa: 16110048 Page 129/154

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$

 $\exists j / \rho_j(\phi) \neq 0 \ \forall \phi \longrightarrow V = \sum_j V_j \rho_j^2$

$$V = \sum_{j} V_{j} \hat{\rho}_{j}^{2}$$

remains positive at all times

generic quantum bounce!

Pirsa: 16110048 Page 130/154

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j} - rac{Q_{j}^{2}}{
ho_{j}^{2}} + m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$



$$V = \sum_j V_j \hat{
ho}_j^2$$
 remains positive at all times

generic quantum bounce!

+ primordial accelleration De Cesare, Sakellariadou, '16

• classical approx. $ho_j^2 \gg |E_j|/m_j^2 ext{ and }
ho_j^4 \gg Q_j^2/m_j^2$



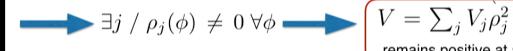
$$\frac{V''}{V} = \frac{4\sum_{j} V_j m_j^2 \rho_j^2}{\sum_{j} V_j \rho_j^2}$$

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j} - rac{Q_{j}^{2}}{
ho_{j}^{2}} + m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$



$$V = \sum_{j} V_{j} \rho_{j}^{2}$$

remains positive at all times

generic quantum bounce!

+ primordial accelleration De Cesare, Sakellariadou, '16

• classical approx. $ho_j^2 \gg |E_j|/m_j^2 ext{ and }
ho_j^4 \gg Q_j^2/m_j^2$



$$\frac{V''}{V} = \frac{4\sum_{j} V_j m_j^2 \rho_j^2}{\sum_{j} V_j \rho_j^2}$$

DO, Sindoni, Wilson-Ewing, '16

$$A_j \partial_{\phi}^2 \sigma_j(\phi) - B_j \sigma_j(\phi) + w_j \sigma_j(\phi)^4 = 0$$

interaction terms sub-dominant (dilute-gas approx., consistent with simple approximation of vacuum state)

• two (approximately) conserved quantities (per mode):

$$\sigma_j(\phi) = \rho_j(\phi)e^{i\theta_j(\phi)}$$

$$m_j^2 = B_j/A_j$$

$$m_j^2 = B_j/A_j \qquad \qquad \boxed{\rho_j'' - \frac{Q_j^2}{\rho_j^3} - m_j^2 \rho_j \approx 0}$$

$$E_{j} = A_{j} |\partial_{\phi} \sigma_{j}(\phi)|^{2} - B_{j} |\sigma_{j}(\phi)|^{2} + \frac{2}{5} \operatorname{Re} \left(w_{j} \sigma_{j}(\phi)^{5} \right)$$
$$Q_{j} = -\frac{i}{2} \left[\bar{\sigma}_{j}(\phi) \partial_{\phi} \sigma_{j}(\phi) - \sigma_{j}(\phi) \partial_{\phi} \bar{\sigma}_{j}(\phi) \right]$$

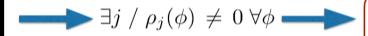
$$E_j \approx (\rho_j')^2 + \rho_j^2 (\theta_j')^2 - m_j^2 \rho^2$$
 $Q_j \approx \rho_j^2 \theta_j'$

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(rac{V'}{3V}
ight)^2 = \left(rac{2\sum_{j}V_{j}\,
ho_{j}\sqrt{E_{j}-rac{Q_{j}^{2}}{
ho_{j}^{2}}+m_{j}^{2}
ho_{j}^{2}}}{3\sum_{j}V_{j}
ho_{j}^{2}}
ight)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$



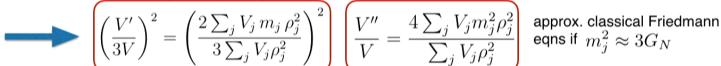
$$V = \sum_{j} V_{j} \rho_{j}^{2}$$

remains positive at all times

generic quantum bounce!

+ primordial accelleration De Cesare, Sakellariadou, '16

• classical approx. $ho_j^2 \gg |E_j|/m_j^2 ext{ and }
ho_j^4 \gg Q_j^2/m_j^2$



$$\left(\frac{V''}{V} = \frac{4\sum_{j} V_j m_j^2 \rho_j^2}{\sum_{j} V_j \rho_j^2}\right)$$

· simple condensate:

$$\sigma_j(\phi) = 0$$
, for all $j \neq j_o$

$$\left(\frac{V'}{3V}\right)^2 = \frac{4\pi G}{3} \left(1 - \frac{\rho}{\rho_c}\right) + \frac{V_{j_o} E_{j_o}}{9V}$$
$$\rho_c = 6\pi G \hbar^2 / V_{j_o}^2 \sim (6\pi / j_o^3) \rho_{\text{Pl}}$$

LQC-like modified dynamics!

DO, Sindoni, Wilson-Ewing, '16

effective dynamics for volume - generalised Friedmann equations:

$$\left(\frac{V'}{3V}\right)^2 = \left(\frac{2\sum_{j} V_j \, \rho_j \sqrt{E_j - \frac{Q_j^2}{\rho_j^2} + m_j^2 \rho_j^2}}{3\sum_{j} V_j \rho_j^2}\right)^2$$

$$\frac{V''}{V} = \frac{2\sum_{j} V_{j} \left[E_{j} + 2m_{j}^{2} \rho_{j}^{2} \right]}{\sum_{j} V_{j} \rho_{j}^{2}}$$



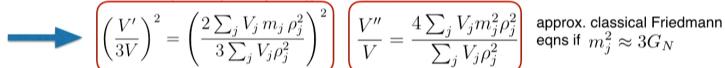
$$V = \sum_{j} V_{j} \rho_{j}^{2}$$

remains positive at all times

generic quantum bounce!

+ primordial accelleration De Cesare, Sakellariadou, '16

• classical approx. $ho_j^2 \gg |E_j|/m_j^2 ext{ and }
ho_j^4 \gg Q_j^2/m_j^2$



$$\frac{V''}{V} = \frac{4\sum_{j} V_j m_j^2 \rho_j^2}{\sum_{j} V_j \rho_j^2}$$

· simple condensate:

$$\sigma_j(\phi) = 0$$
, for all $j \neq j_o$



$$\left(\frac{V'}{3V}\right)^2 = \frac{4\pi G}{3} \left(1 - \frac{\rho}{\rho_c}\right) + \frac{V_{j_o} E_{j_o}}{9V} \qquad \begin{array}{c} \text{LQC-like} \\ \text{modified} \\ \rho_c = 6\pi G \hbar^2/V_{j_o}^2 \sim (6\pi/j_o^3)\rho_{\text{Pl}} \end{array} \qquad \begin{array}{c} \text{dynamics!} \\ \end{array}$$

can show that

- 1) generic solutions approximate such simple condensates at late times
- Gielen, '16
- De Cesare, Pithis, Sakellariadou, '16
- 2) GFT interactions can make primordial acceleration last enough e-folds to avoid need for inflation



Pirsa: 16110048 Page 136/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, very similar to LQC

Pirsa: 16110048 Page 137/154

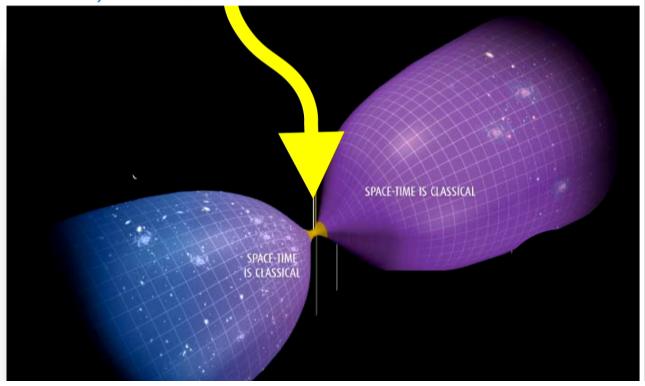
Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, very similar to LQC

(Big Bounce from the full theory!)



Pirsa: 16110048 Page 138/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC

Pirsa: 16110048 Page 139/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC (Big Bounce from the full theory!)

Pirsa: 16110048 Page 140/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC

(Big Bounce from the full theory!)

.... provided the GFT hydrodynamics approximation (and other assumptions) does not break down in that regime

Pirsa: 16110048 Page 141/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC

(Big Bounce from the full theory!)

.... provided the GFT hydrodynamics approximation (and other assumptions) does not break down in that regime scenario of Big Bang as cosmological phase transition (geometrogenesis) suggests it should break down geometrogenesis scenario similar to "emergent universe" scenario with degenerate scale factor before transition

Pirsa: 16110048 Page 142/154

Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC

(Big Bounce from the full theory!)

.... provided the GFT hydrodynamics approximation (and other assumptions) does not break down in that regime scenario of Big Bang as cosmological phase transition (geometrogenesis) suggests it should break down geometrogenesis scenario similar to "emergent universe" scenario with degenerate scale factor before transition if it does break, one has to go back to the full GFT theory, and improve the construction (ansatz for vacuum, approximation of SD equations,)

and then try again

Pirsa: 16110048 Page 143/154

```
Big Bounce?

DO, L. Sindoni, E. Wilson-Ewing, '16

M. De Cesare, A. Pithis, M. Sakellariadou, '16

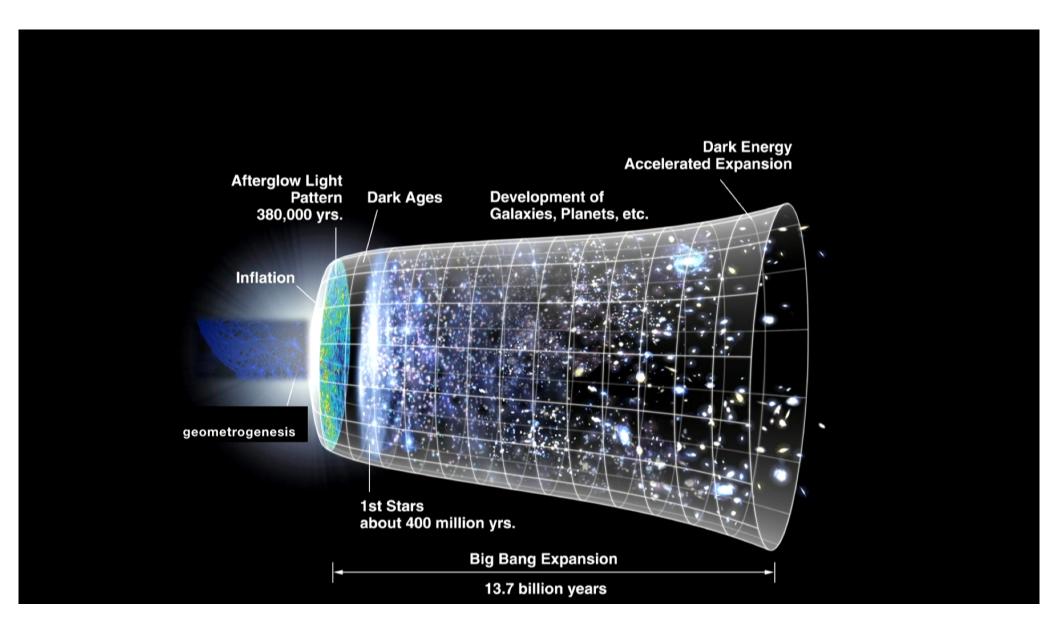
given effective cosmological equations for GFT condensates, it can be derived via the same type of calculations done in LQC ....

(Big Bounce from the full theory!)

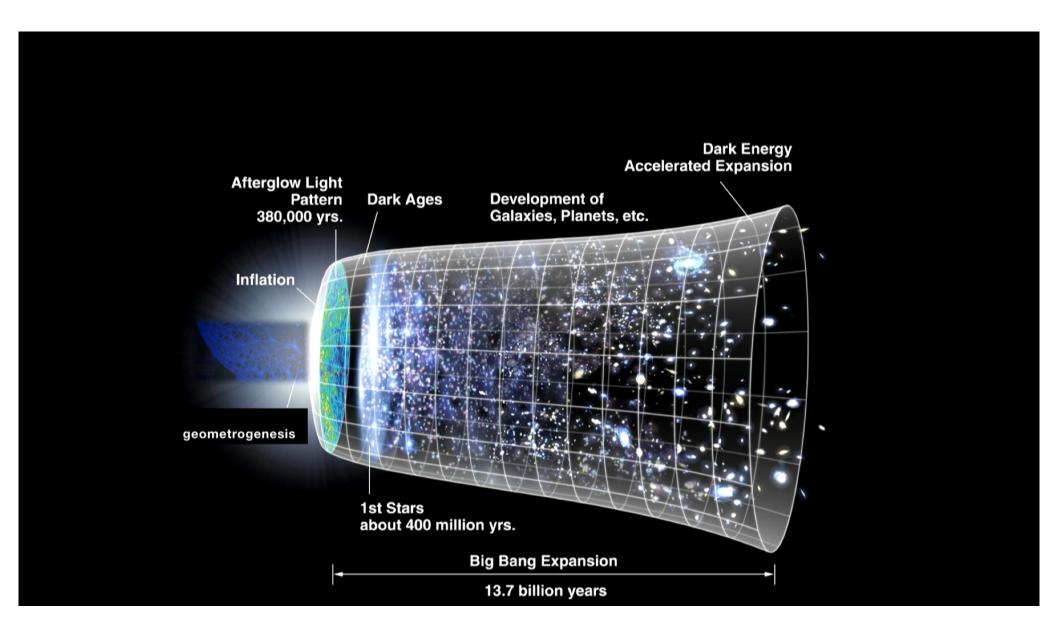
.... provided the GFT hydrodynamics approximation (and other assumptions) does not break down in that regime scenario of Big Bang as cosmological phase transition (geometrogenesis) suggests it should break down geometrogenesis scenario similar to "emergent universe" scenario with degenerate scale factor before transition if it does break, one has to go back to the full GFT theory, and improve the construction (ansatz for vacuum, approximation of SD equations, ....)

and then try again
```

Pirsa: 16110048 Page 144/154



Pirsa: 16110048 Page 145/154



Pirsa: 16110048 Page 146/154

goal: effective dynamics of cosmological perturbations from first principles, i.e. from full QG formalism

needed for computation of CMB spectrum and for tests of fate of Lorentz invariance

Pirsa: 16110048 Page 147/154

goal: effective dynamics of cosmological perturbations from first principles, i.e. from full QG formalism

needed for computation of CMB spectrum and for tests of fate of Lorentz invariance

several possible strategies:

"cheap" - dressed metric approach:

1) define modified FRW metric from expectation values for cosmological variables derived from GFT 2) just use it inside standard effective QFT for fields

"intermediate" - "separate universe approach":

several homogeneous patches, each satisfying modified Friedmann eon compute effective dynamics of volume differences (scalar perturbations, valid at long wavelength)

Pirsa: 16110048 Page 148/154

goal: effective dynamics of cosmological perturbations from first principles, i.e. from full QG formalism

needed for computation of CMB spectrum and for tests of fate of Lorentz invariance

several possible strategies:

"cheap" - dressed metric approach:

1) define modified FRW metric from expectation values for cosmological variables derived from GFT 2) just use it inside standard effective QFT for fields

"intermediate" - "separate universe approach":

several homogeneous patches, each satisfying modified Friedmann eon compute effective dynamics of volume differences (scalar perturbations, valid at long wavelength)

"ambitious":

- 1) derive effective dynamics for GFT fluctuations above condensate from full theory
- 2) recast it in standard spacetime-based QFT form using information from background GFT condensate (difficulty is: the formalism naturally gives it in diffeo-invariant variables, spacetime-free form) need to add material reference frame to define local regions

Pirsa: 16110048 Page 149/154

goal: effective dynamics of cosmological perturbations from first principles, i.e. from full QG formalism

needed for computation of CMB spectrum and for tests of fate of Lorentz invariance

several possible strategies:

"cheap" - dressed metric approach:

1) define modified FRW metric from expectation values for cosmological variables derived from GFT 2) just use it inside standard effective QFT for fields

"intermediate" - "separate universe approach":

several homogeneous patches, each satisfying modified Friedmann eon compute effective dynamics of volume differences (scalar perturbations, valid at long wavelength)

"ambitious":

- 1) derive effective dynamics for GFT fluctuations above condensate from full theory
- 2) recast it in standard spacetime-based QFT form using information from background GFT condensate (difficulty is: the formalism naturally gives it in diffeo-invariant variables, spacetime-free form) need to add material reference frame to define local regions

Pirsa: 16110048 Page 150/154

goal: effective dynamics of cosmological perturbations from first principles, i.e. from full QG formalism

needed for computation of CMB spectrum and for tests of fate of Lorentz invariance

several possible strategies:

"cheap" - dressed metric approach:

define modified FRW metric from expectation values for cosmological variables derived from GFT
 just use it inside standard effective QFT for fields

"intermediate" - "separate universe approach":

several homogeneous patches, each satisfying modified Friedmann eon compute effective dynamics of volume differences (scalar perturbations, valid at long wavelength)

"ambitious":

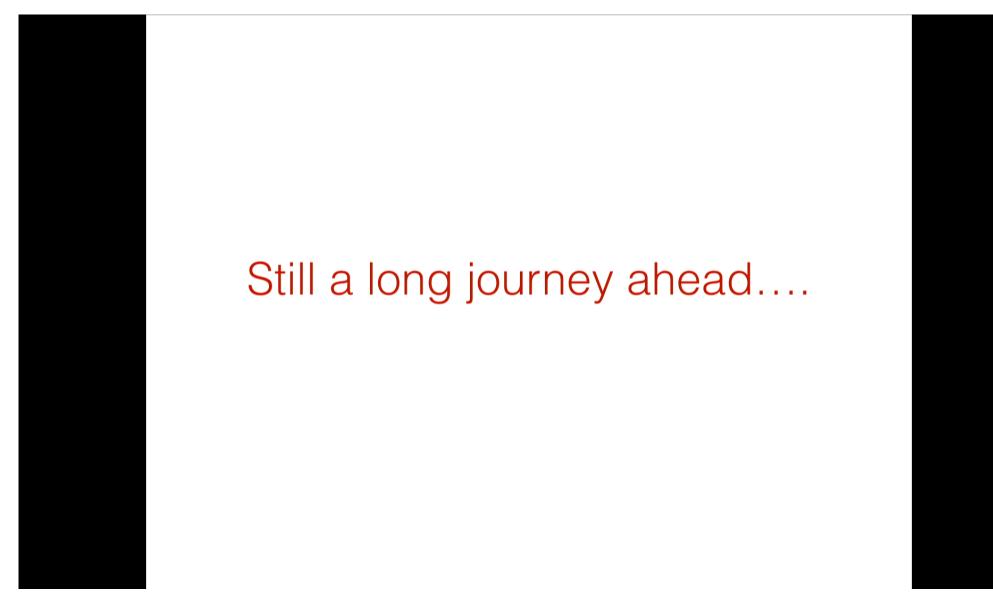
- derive effective dynamics for GFT fluctuations above condensate from full theory
- recast it in standard spacetime-based QFT form using information from background GFT condensate (difficulty is: the formalism naturally gives it in diffeo-invariant variables, spacetime-free form) need to add material reference frame to define local regions

expect deformation of standard QFT:

- from holonomization of the connection and non-commutativity of triad variables
- derivation of effective dynamics of perturbations around mean field in topological GFT:
 non-commutative scalar field theory on non-commutative flat space

W. Fairbairn, E. Livine, '07; F. Girelli, E. Livine, DO, '09

Pirsa: 16110048 Page 151/154



Pirsa: 16110048 Page 152/154



Pirsa: 16110048



Pirsa: 16110048 Page 154/154