

Title: Victoria Kaspi: The Cosmic Gift of Neutron Stars

Date: Feb 03, 2016 07:00 PM

URL: <http://pirsa.org/16020084>

Abstract: 

Neutron stars are a celestial gift to scientists. These incredibly dense collapsed stars act as very precise cosmic beacons that help shed light on some of the most challenging problems in modern physics.

In her Feb. 3 talk at Perimeter Institute, astrophysicist Victoria Kaspi will explore these strange objects, explain how astronomers are using them to study issues ranging from the origins of the universe to the very nature of matter, and even let the audience hear the cosmic symphony they create.

# Arecibo Radio Telescope in Puerto Rico



**Diameter:  
305 m**



# Pulsar "Sounds"



*PSR B0329+54 P=0.7s*



*PSR B0833-45 P=0.089s*



*PSR B0531+21 P=0.033s*

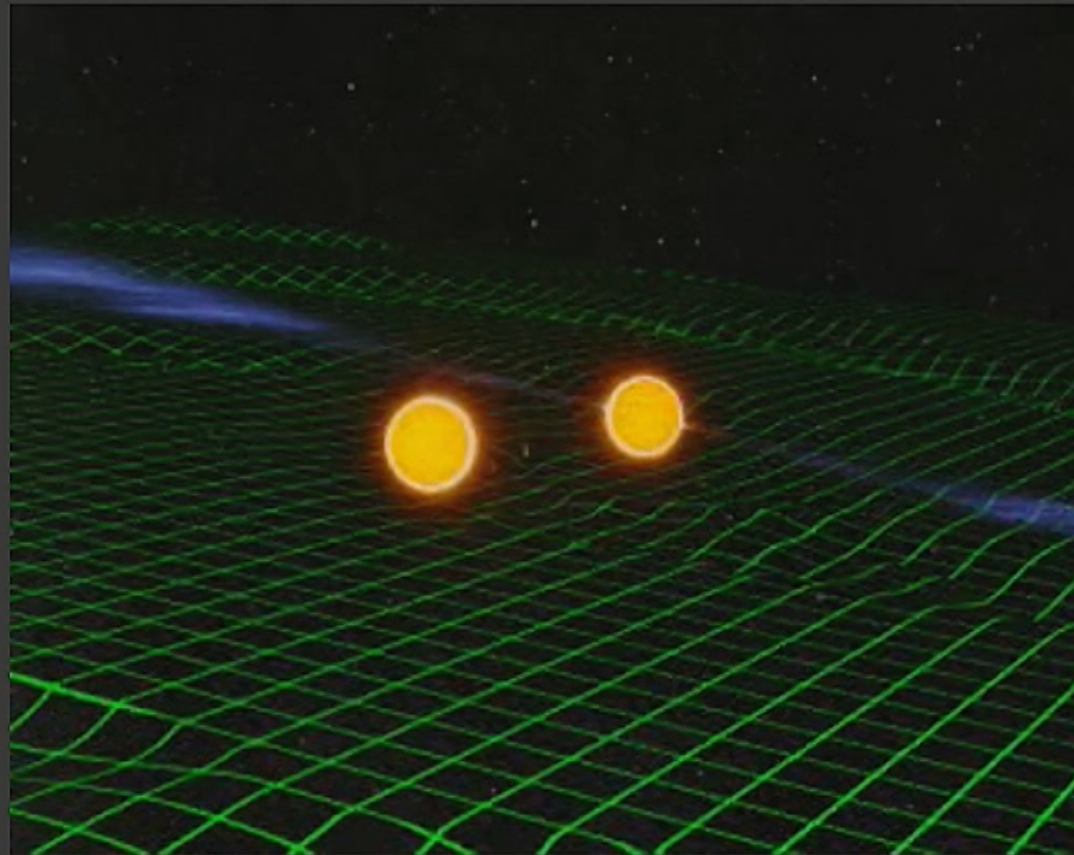
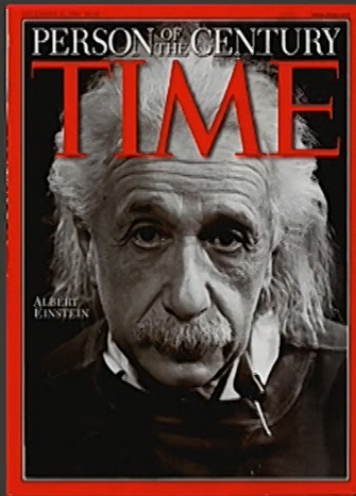


*PSR B1937+21 P=0.0015s*

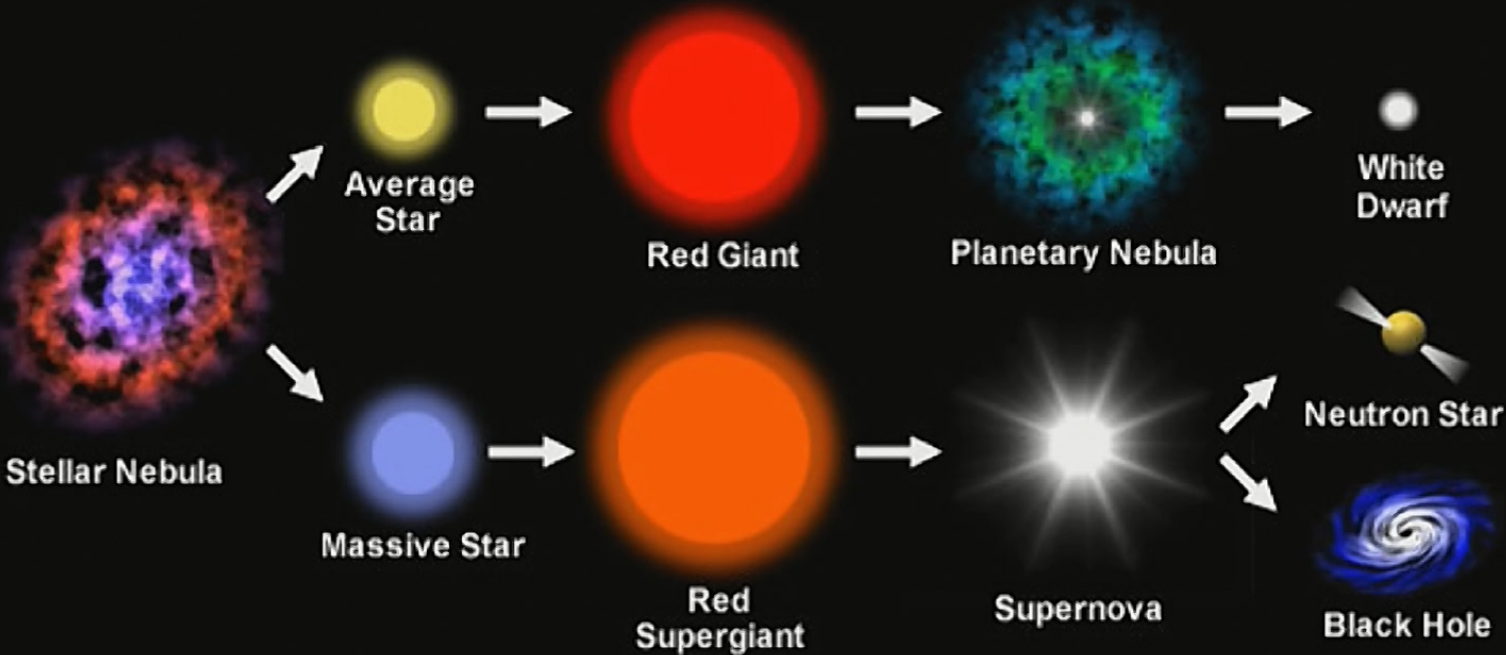


# Binary Pulsar and Gravitational Waves

*Pulsar  
orbit  
should  
decay!*

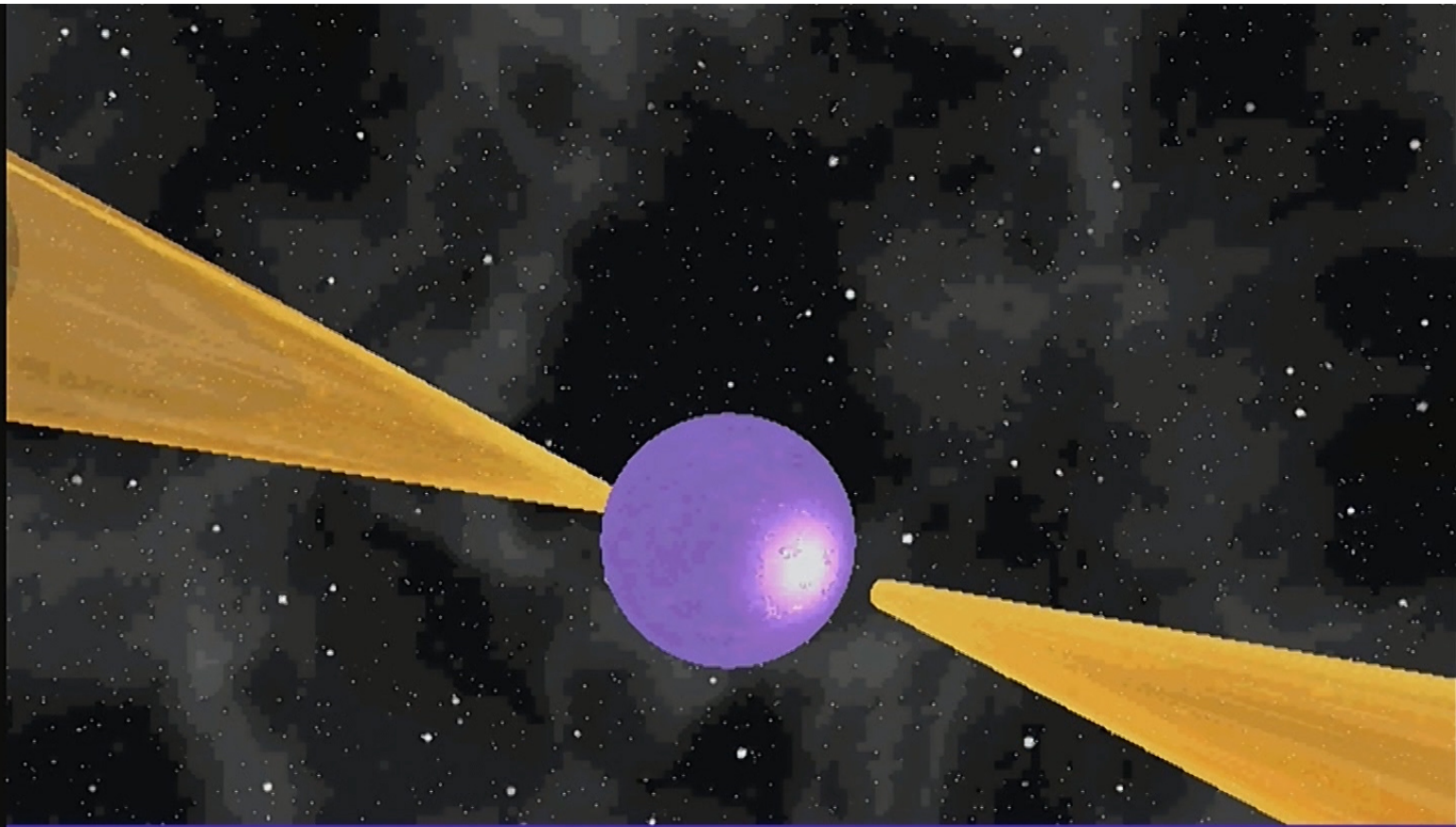


# Life Cycle of a Star

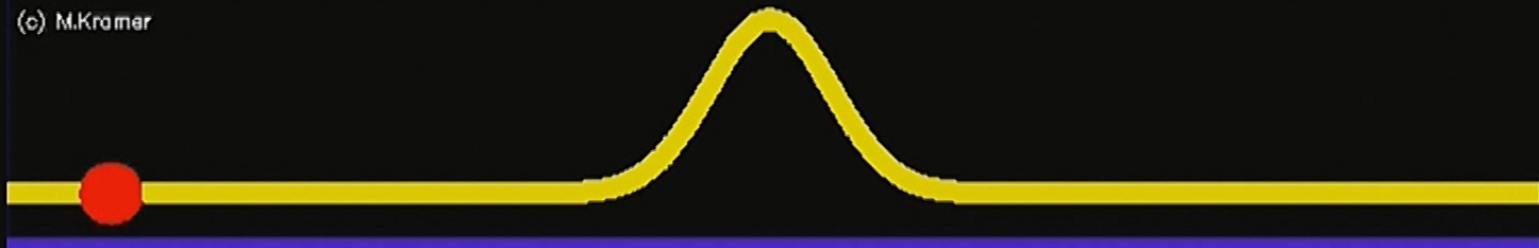


**Crab  
Nebula:  
Remnant of  
Supernova  
in 1054 AD**





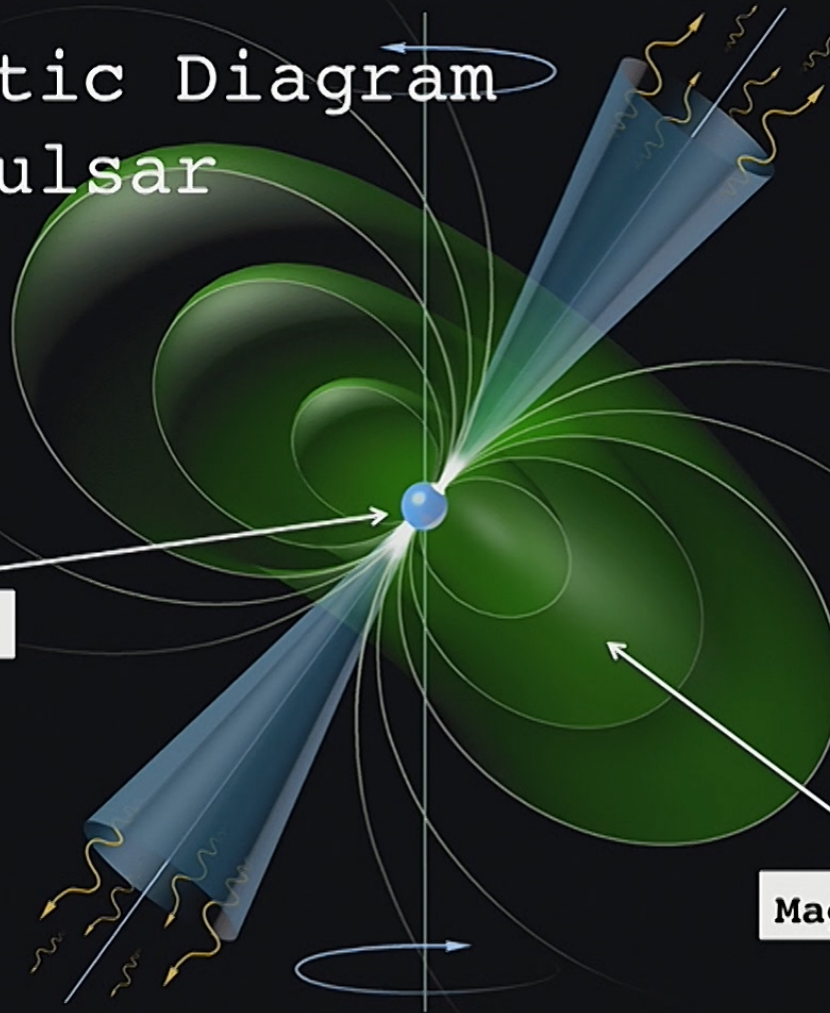
(c) M.Kramer



# Schematic Diagram of a Pulsar

Neutron Star

Magnetosphere





# Neutron Star Rotation



- Conservation of Angular Momentum
- Progenitor star's spin gets amplified in collapse



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**Maximum  
625 rps**



What is the speed at the surface of a millisecond pulsar?



$$v = \frac{2\pi R}{P} = \frac{2\pi 10000}{0.0014} = 4.5 \times 10^7 \text{ m/s} = 0.15c$$

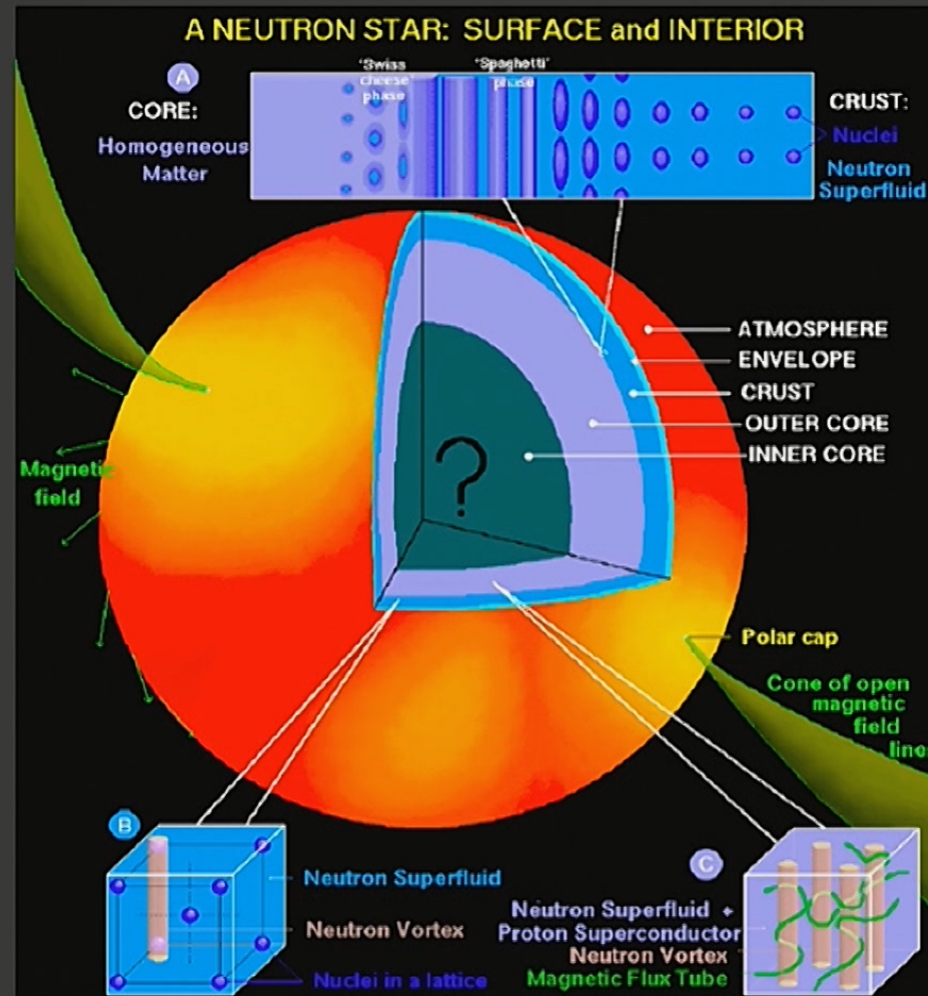
↑  
velocity

↑  
very fast!

HOW FAST CAN A PULSAR SPIN?



# What's Inside a Neutron Star?



# Arecibo Radio Telescope in Puerto Rico



iameter:  
5 m



# Fastest Pulsar Yet!

ASTRONOMY

## A Neutron Star in F-sharp

Jonathan E. Grindlay

Millisecond pulsars are extreme examples of what can happen when stars evolve into neutron stars in compact binary systems. These rotating objects are spun up by accretion of matter from their binary companions, producing luminous x-ray emission, and later become detectable as pulsars with periods of a few milliseconds (1). As a result, these "fast pulsars" may offer some of the best probes to study matter and space in the relativistic regime of strong gravity. On page 1901, Hessels *et al.* (2) report the discovery of pulsar PSR J1748-2446ad in the dense globular cluster Terzan 5 (Ter5-ad). This object, detected with the Green Bank radio telescope, holds the new record for the fastest spinning neutron star (or, indeed, any object of

stellar mass or larger). Its spin period is 1.396 ms, even shorter than that of the first millisecond pulsar discovered in 1967 (3). With a rotation frequency of 71 Hz, it reaches a new high note for the celestial spheres—between F-sharp and G-sharp, whereas B1937+21 (at 642 Hz) can be placed between D-sharp and E.

Since their discovery in 1967, millisecond pulsars have been the gateway to the study of matter and space at the extremes found in compact stars (4). Such stars are nature's laboratories for the study of the limits of matter and are only a few kilometers larger in radius than an object that would collapse to a black hole. With  $\sim 1.4$  solar masses ( $M_{\odot}$ ) packed into a radius of  $\sim 10$  km, neutron stars are the ultimate laboratories for the astrophysics and physics of matter. Neutron stars can exhibit magnetic fields  $10^{15}$  times that of Earth, as revealed

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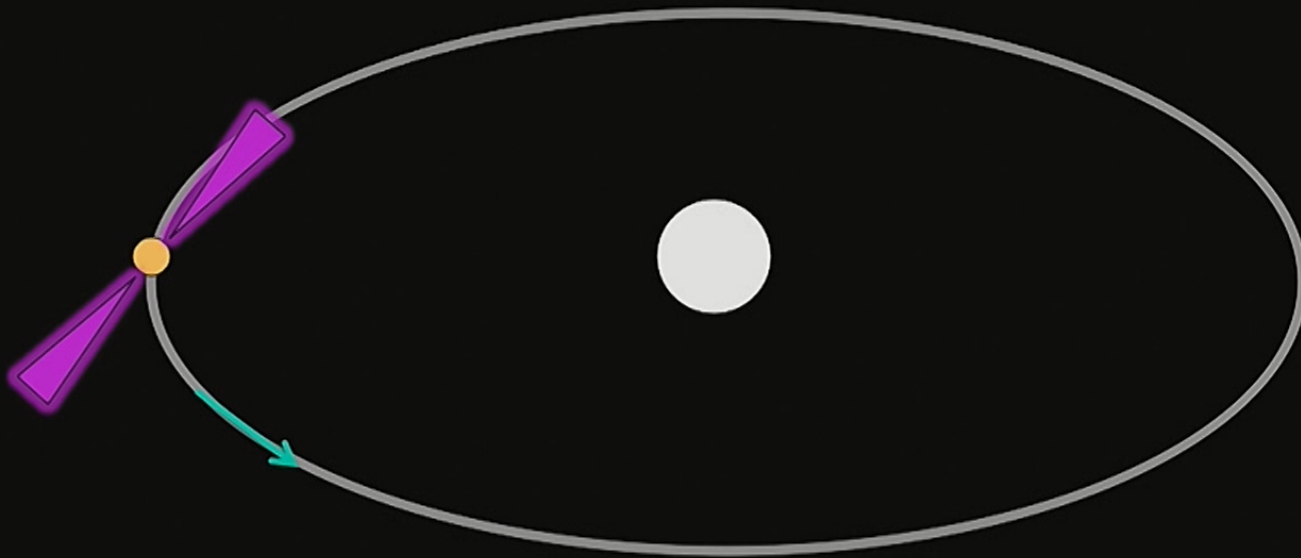
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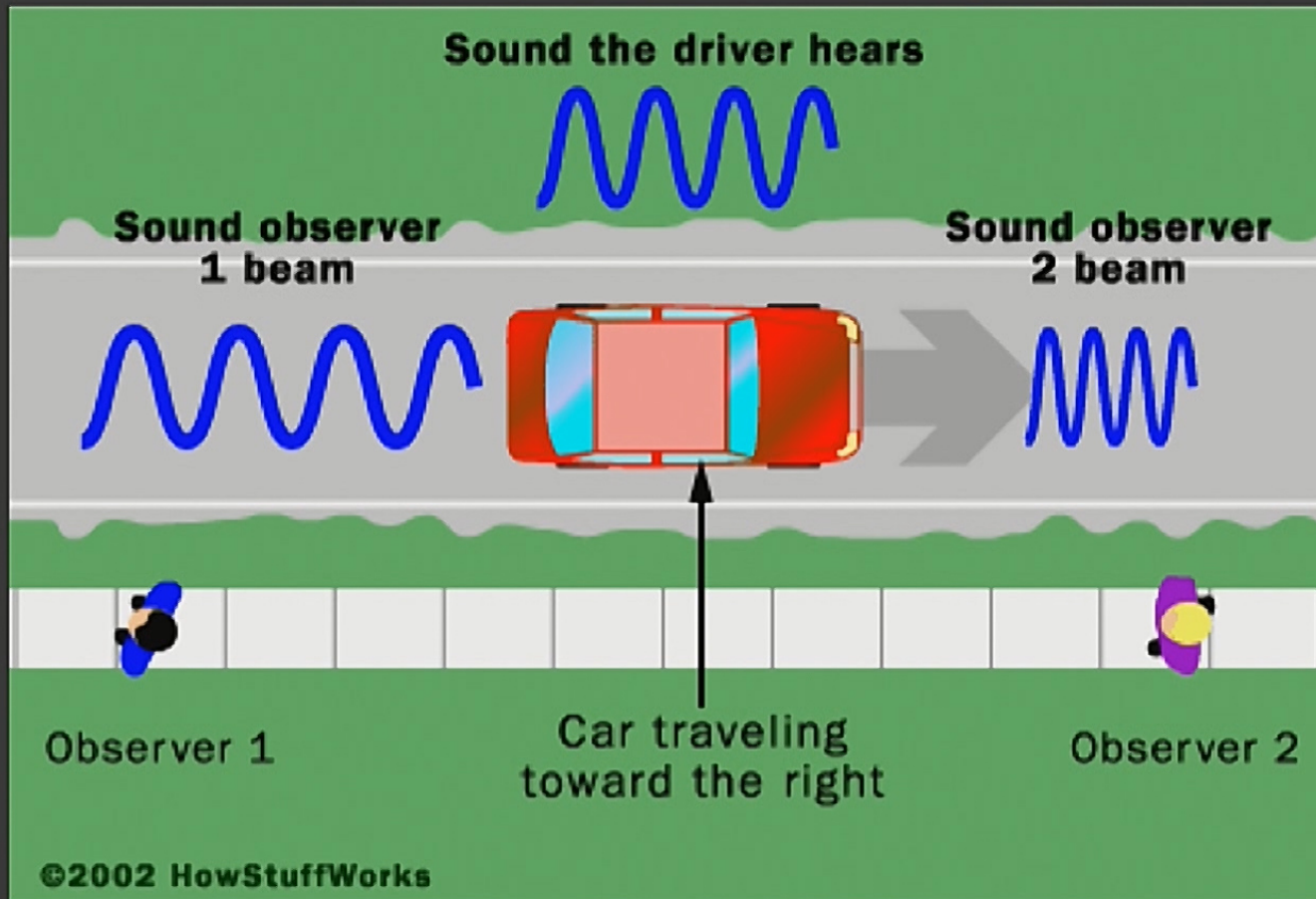


# Binary Pulsars



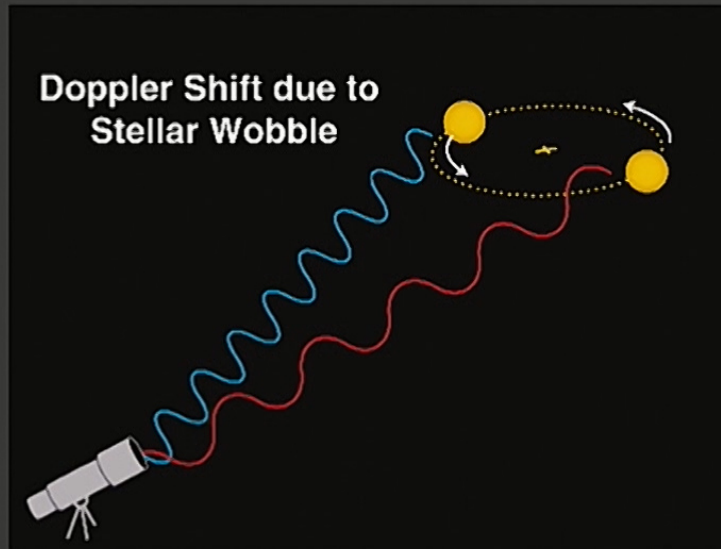
The pulsar orbits another star

# Doppler Shift



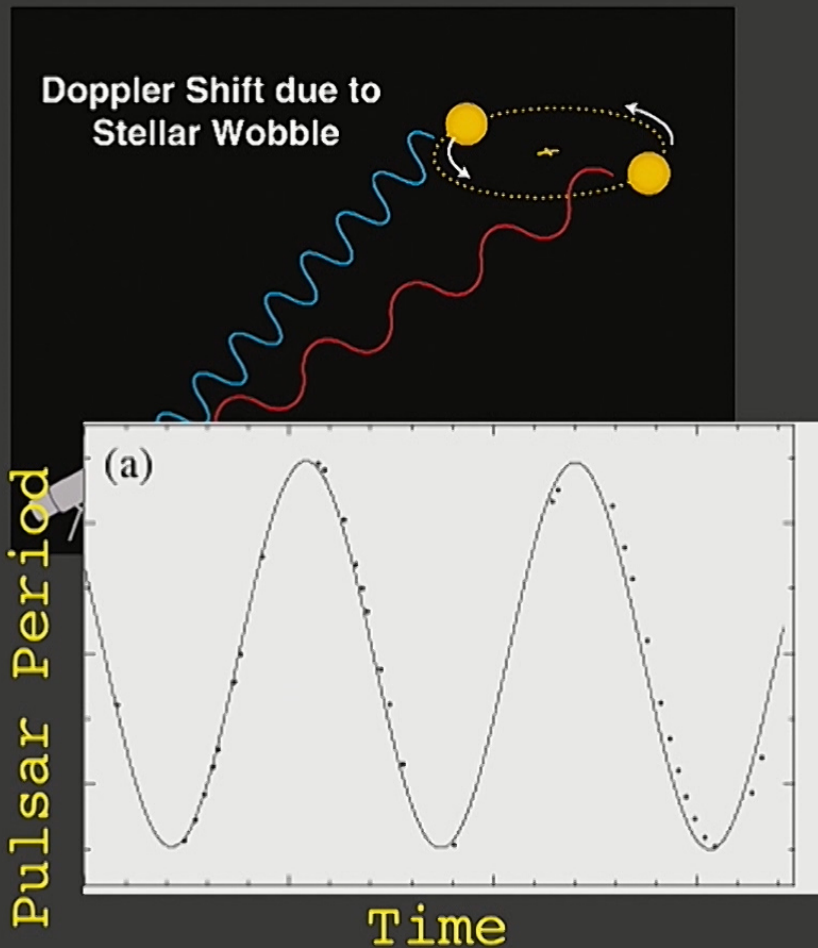


# Doppler Shift in Binary Stars



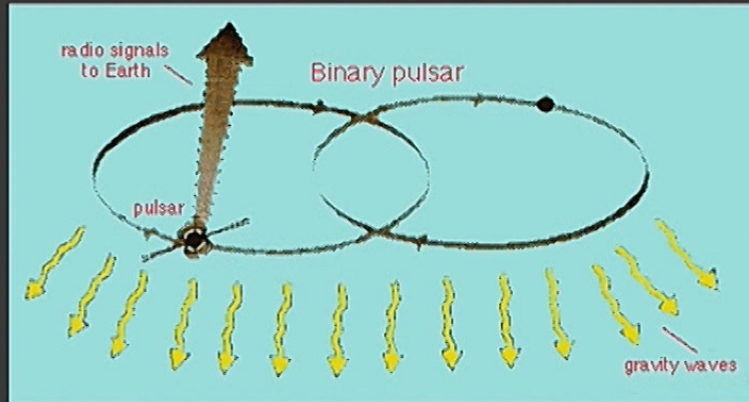
- Pulse period in binary system changes!
- Can study binary motion with *exquisite* precision

# Doppler Shift in Binary Stars



- Pulse period in binary system changes!
- Can study binary motion with *exquisite* precision

# 1974: Discovery of first Binary Pulsar



- PSR B1913+16 in orbit around a second neutron star that is not a pulsar



Russell Hulse Joe Taylor

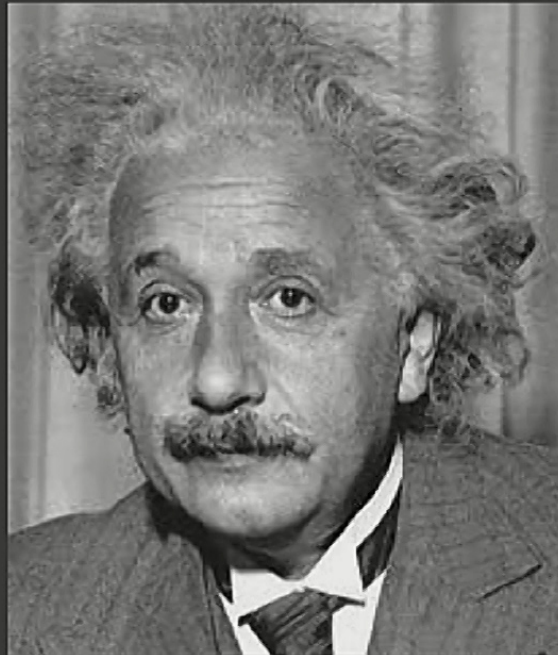
**PSR B1913+16: Orbital period 8hrs!**  
**HIGHLY RELATIVISTIC**

## Classical Dynamics



$$\mathbf{F} = m\mathbf{a}$$

## Relativistic Dynamics



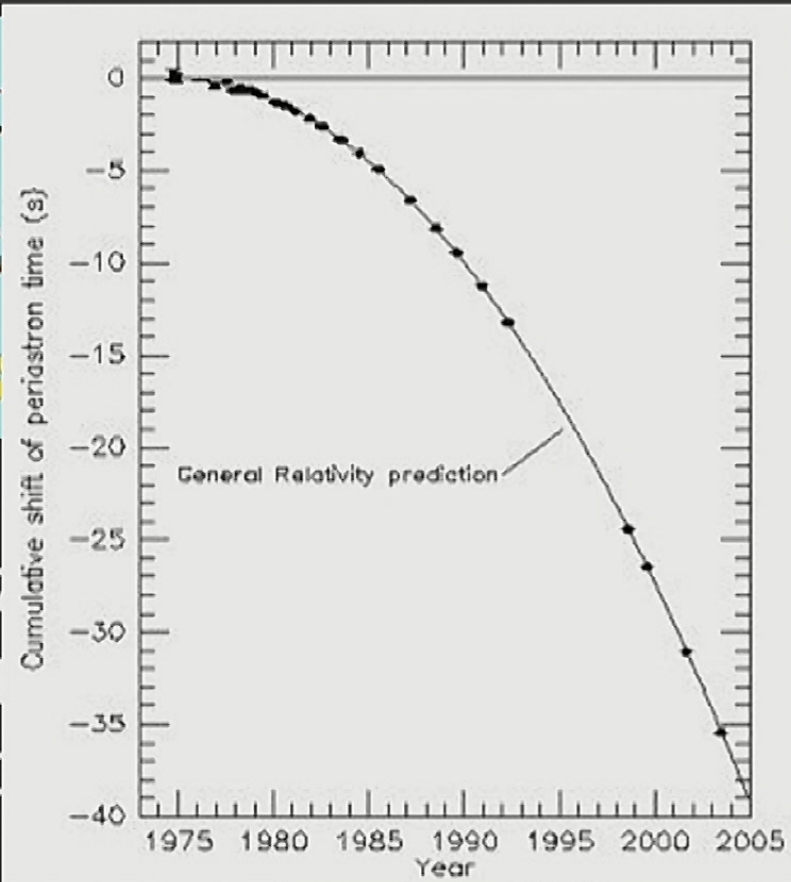
*Einstein's Equation:*

$$R^{\mu\nu} - (1/2)(g^{\mu\nu}R) + \Lambda g^{\mu\nu} = 8\pi G T^{\mu\nu}$$

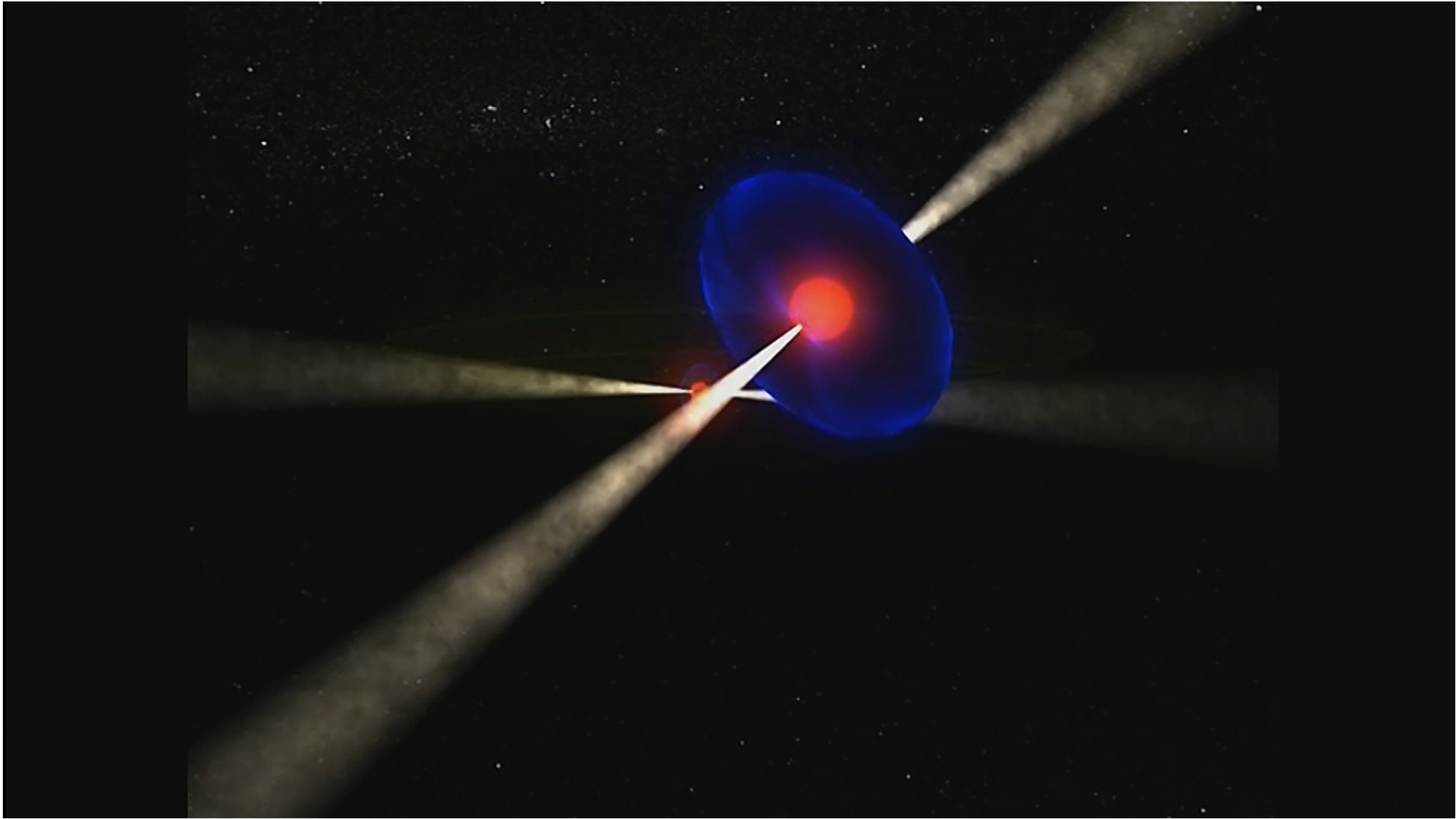
$$ds^2 = [1 - (\Lambda/3)r^2 - r_g/r] dt^2 - [1 - (\Lambda/3)r^2 - r_g/r]^{-1} dr^2 - r^2(d\theta^2 + \sin^2\theta d\phi^2)$$

Schwarzschild metric with a cosmological constant

# Binary Pulsars and Einstein

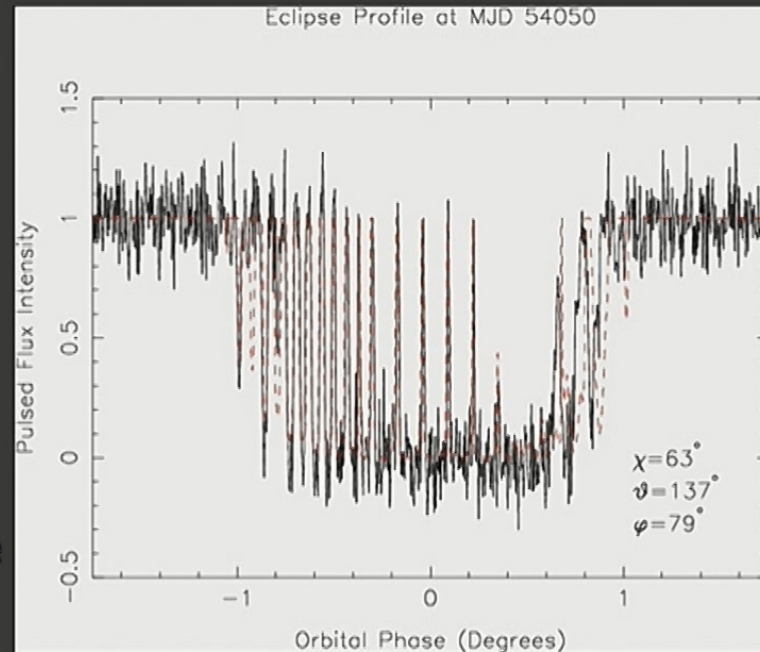


EIN  
RI



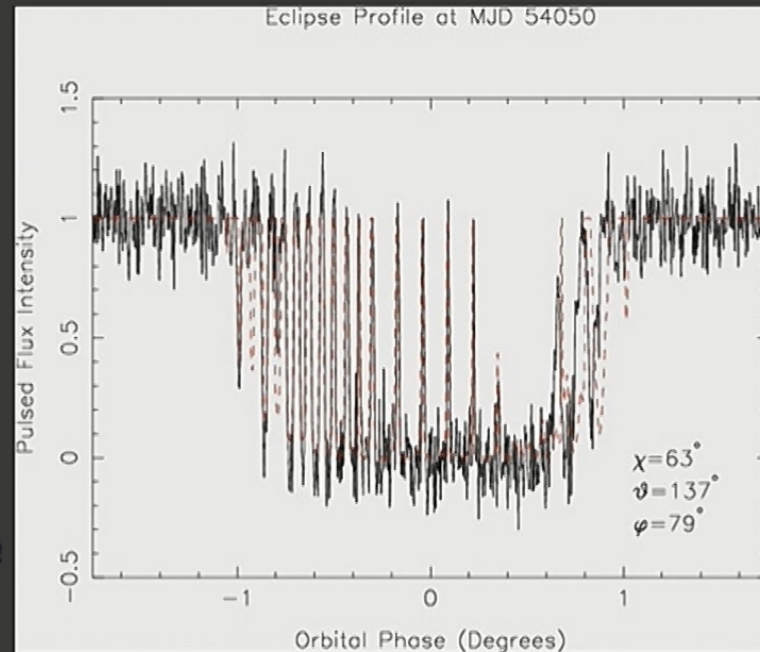
# Double Pulsar: GR Laboratory

- Unparalleled lab for GR
- Use eclipses of pulsar A to study B's orientation in space
- GR:B should change if GR correct



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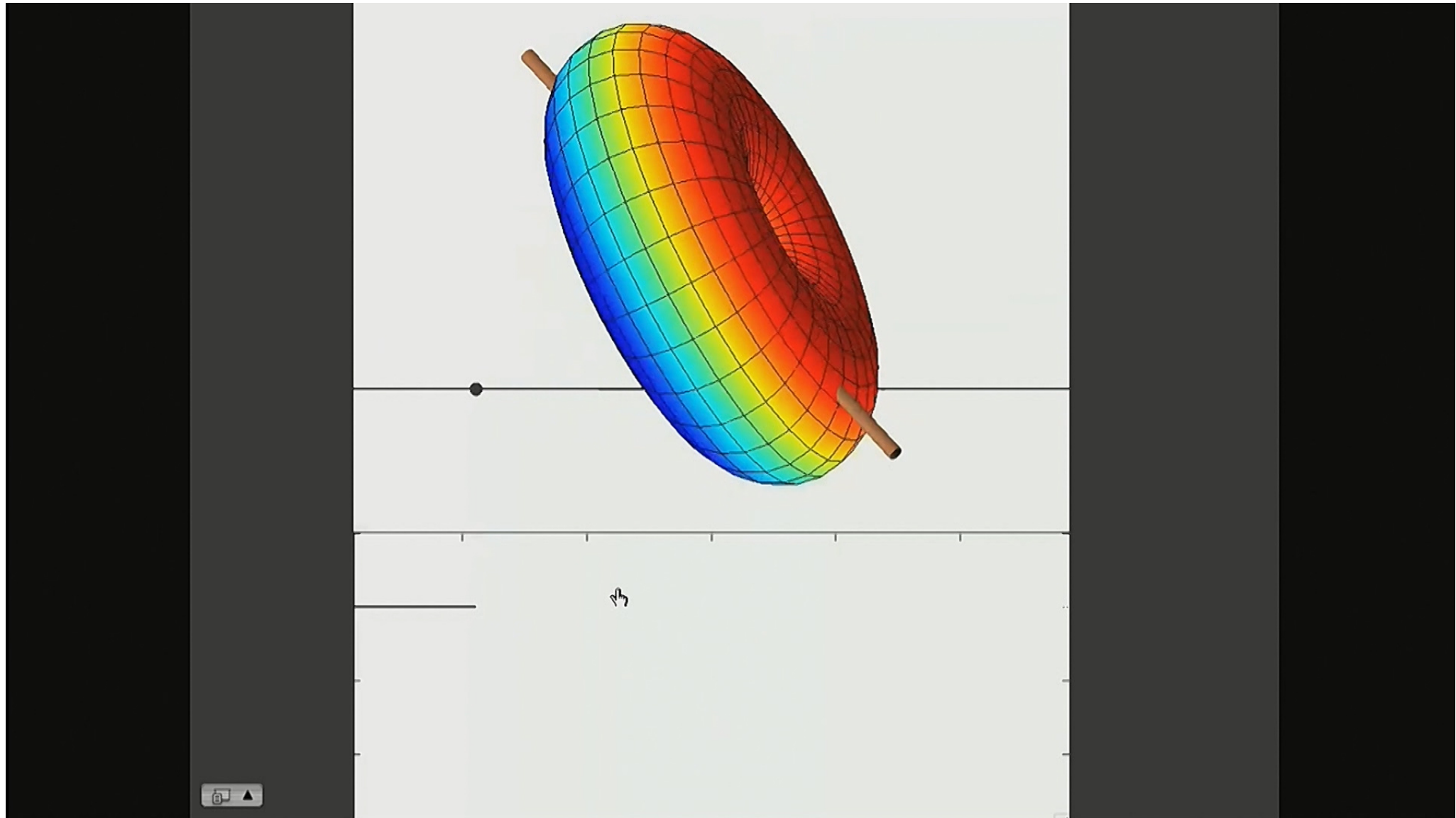
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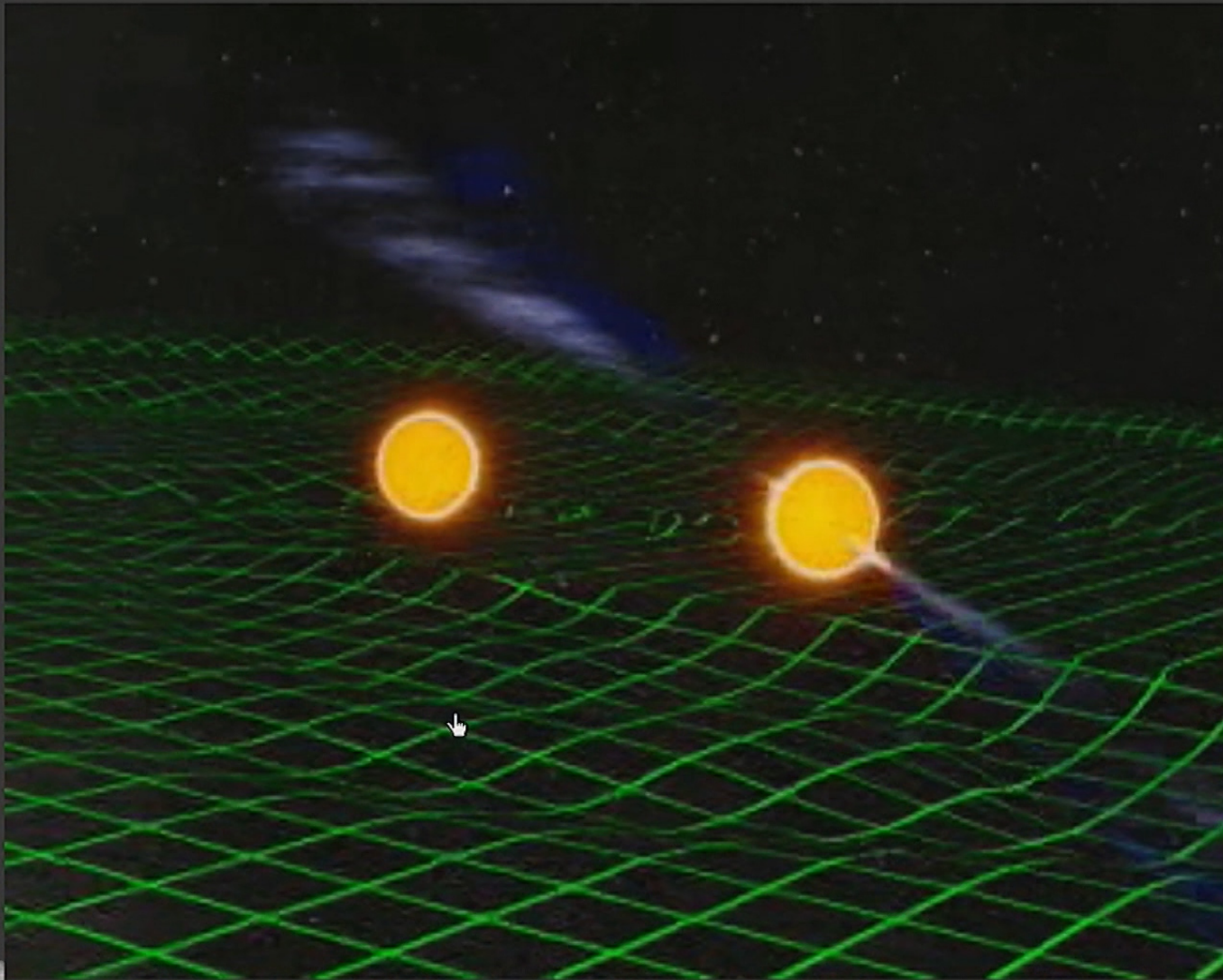


PhD thesis of  
Rene Breton,  
McGill U.









# Neutron Stars: Gifts that Keeps Giving

- Pulsars are sources of gravitational waves...and
- Pulsars can be detectors of gravitational waves!
  - Pulsars' signals affected by other sources of gravitational waves
  - Example: merging supermassive black holes



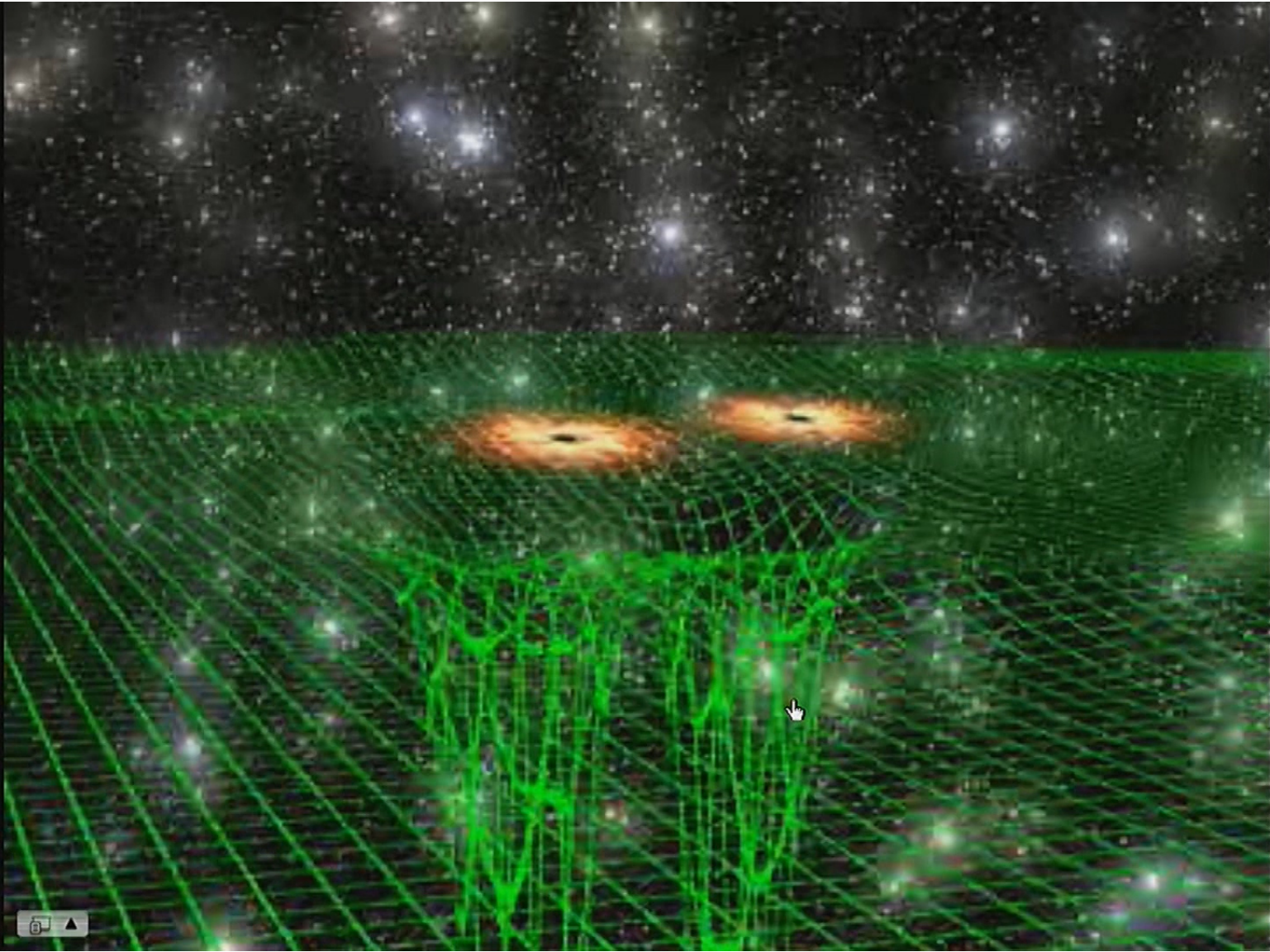
# Merging Galaxies



Optical image of  
merging galaxies  
"The Mice"



Computer Simulation  
of merging galaxies



# Cosmic Gifts from Pulsars

- Pulsar music
- Perfect clocks
- Binary pulsars
- Tests of Einstein's theories
  - Indirect detection of gravitational waves
- Double pulsar
- Soon to come: the next gift?
  - Direct detection of gravitational waves from supermassive black holes?



# McGill Pulsar Group

