

Title: PSI 2015/2016 Quantum Gravity - Lecture 2

Date: Feb 23, 2016 10:15 AM

URL: <http://pirsa.org/16020073>

Abstract:

II. Action principle - 1st order action

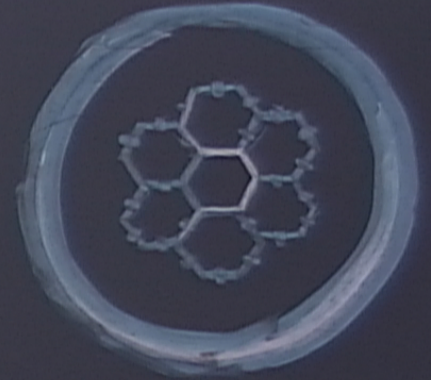
A. Review of the 4D Einstein action and equations of motion

B. 3D case

C. Triads

D. Spin connection

E. First order action for 3D gravity
= Palatini action



II. A.

Einstein-Hilbert action. $S = \frac{1}{\kappa} \int \sqrt{g} R d^4x$

$$\kappa = \frac{16\pi G}{c^3} = 1$$

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Einstein-Hilbert action. $S = \frac{1}{\kappa} \int \sqrt{g} R d^4x = S[\underline{g_{\mu\nu}}]$.

$$\kappa = \frac{16\pi G}{c^3} = 1$$

\Rightarrow Equations of motion

$$K d^4x = S[\underline{g_{\mu\nu}}].$$

laws of motion

$$| \quad g_{\mu\alpha} g^{\alpha\nu} = \delta_{\mu}^{\nu}$$

$$\delta(g_{\mu\alpha} g^{\alpha\nu}) = 0 \Rightarrow \delta g_{\mu\nu} = -g_{\mu\alpha} \delta g^{\alpha\beta} g_{\beta\nu}$$

ction. $S = \frac{1}{2} \int \sqrt{g} R d^4x = S[\underline{g_{\mu\nu}}].$

$$= \frac{16\pi G}{c^3} = 1$$

⇒ Equations of motion
 variations of the action
 w.r.t metric & inverse metric.

$$g_{\mu\alpha} g^{\alpha\nu} = \delta_{\mu}^{\nu}$$

$$\delta(g_{\mu\alpha} g^{\alpha\nu}) = 0 \Rightarrow \delta$$

$$\delta S = \delta S_1 + \delta S_2 + \delta S_3.$$

$$R = R_{\mu\nu} g^{\mu\nu}$$

$$\delta S_1 = \int (\delta R_{\mu\nu}) g^{\mu\nu} \sqrt{g} d^4x$$

see
tutorial

$$\delta S_1 = \int \nabla$$

$$\delta S = \delta S_1 + \delta S_2 + \delta S_3.$$

$$\delta S_1 = \int (\delta R_{\mu\nu}) g^{\mu\nu} \sqrt{g} d^4x$$

see
tutorial
←

$$\delta S_1 = \int (\nabla^\mu N_\mu) \sqrt{g} d^4x.$$

$$\delta S_2 = \int R_{\mu\nu} \delta(g^{\mu\nu}) \sqrt{g} d^4x.$$

$$S = \delta S_1 + \delta S_2 + \delta S_3$$

$$\delta S_1 = \int (\delta R_{\mu\nu}) g^{\mu\nu} \sqrt{g} d^4x$$

see
tutorial

$$\delta S_1 = \int (\nabla^\mu N_\mu) \sqrt{g} d^4x$$

$$\delta S_2 = \int R_{\mu\nu} \delta(g^{\mu\nu}) \sqrt{g} d^4x$$

determinant

express g in terms of the metric

$$\delta S_3 =$$

$$\delta \sqrt{g} = \frac{1}{2\sqrt{g}} \delta g = -\frac{1}{2} \frac{1}{\sqrt{g}} g_{\mu\nu} \delta g^{\mu\nu}$$

... of the metric
 w.r.t metric & inverse metric $d \ln \sqrt{g} = -\frac{1}{2} g^{\mu\nu} \delta g_{\mu\nu}$

$$\delta S = \delta S_1 + \delta S_2 + \delta S_3$$

$$R = R_{\mu\nu} g^{\mu\nu}$$

$$\delta S_1 = \int (\delta R_{\mu\nu}) g^{\mu\nu} \sqrt{g} d^4x$$

$$\delta S_1 = \int (\nabla^\mu \delta x^\mu) \sqrt{g} d^4x$$

see tutorial

$$\delta S_2 = \int R_{\mu\nu} \delta(g^{\mu\nu}) \sqrt{g} d^4x$$

determinant

express g in terms of the metric

$$\delta S_3 = \int -\frac{1}{2} R g_{\mu\nu} \sqrt{g} \delta g^{\mu\nu}$$

$$\delta \sqrt{g} = \frac{1}{2\sqrt{g}} \delta g = -\frac{1}{2} \frac{1}{\sqrt{g}} g^{\mu\nu} \delta g_{\mu\nu}$$

$$\delta S = \int d^4x$$



$$S_S = \int d^4x \underbrace{\left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right)}_{G_{\mu\nu}} \sqrt{g} \delta g^{\mu\nu}$$

Vacuum Einstein

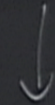
$$SS = \int d^4x \left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right) \sqrt{g} \delta g^{\mu\nu}$$

$G_{\mu\nu}$

Vacuum Einstein equation

$$G_{\mu\nu} = 0$$

trace of $G_{\mu\nu} = 0$



$$\Rightarrow R_{\mu\nu} = 0$$

$$G_{\mu\nu} = \left(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right) \sqrt{|g|} \delta g^{\mu\nu}$$

$G_{\mu\nu}$

Einstein equation

$$G_{\mu\nu} = 0$$

$$\Rightarrow R_{\mu\nu} = 0$$

trace of $G_{\mu\nu} = 0$

($R = 0$ when $n > 2$)

B. 3D case

• $S = \int \sqrt{g} R \, d^3x$.

• equations of motion $G_{\mu\nu} = 0 \Rightarrow R_{\mu\nu} = 0$.

- $S = \int \sqrt{g} R \, d^3x$.

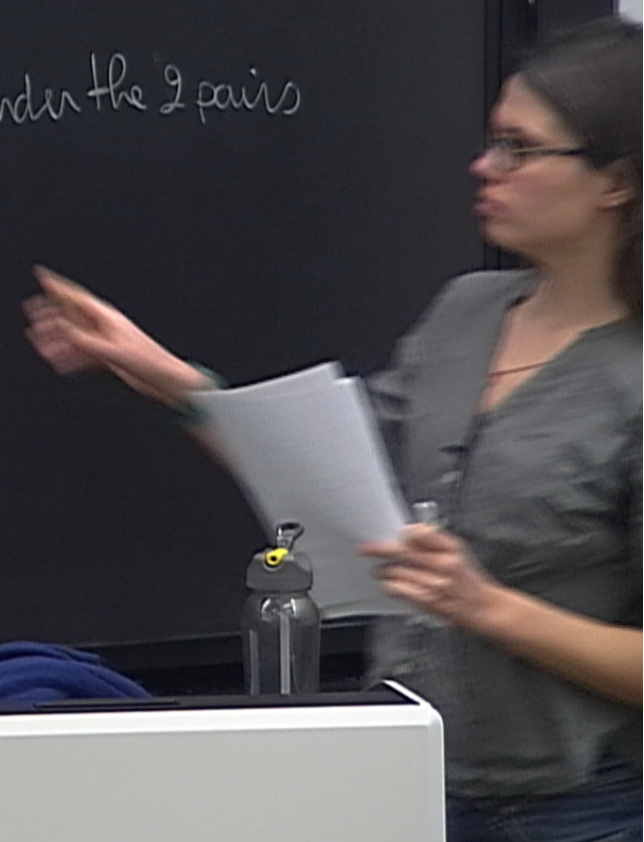
- equations of motion $G_{\mu\nu} = 0 \Rightarrow R_{\mu\nu} = 0$.

- $R_{\mu\nu} = 0 \Rightarrow R_{\mu\nu\rho\sigma} = 0$.

$R_{\mu\nu\sigma}$

- antisymmetric in the 1st 2 indices: 3 pairs & one index running from 1 to 3.
- antisymmetric in the last 2 indices \rightarrow //
- symmetric under the exchange under the 2 pairs

\downarrow \downarrow
 1 to 3 1 to 3



$R_{\nu\mu\alpha\beta}$

\downarrow \downarrow
 $\nu\beta$ $\mu\alpha$

\downarrow
 $\nu\beta \quad \mu\alpha$

\Downarrow
6 dof

antisymmetric in the 1st 2 indices : 3 pairs & one index running from 1-3

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$R_{\mu\nu\alpha\beta}$

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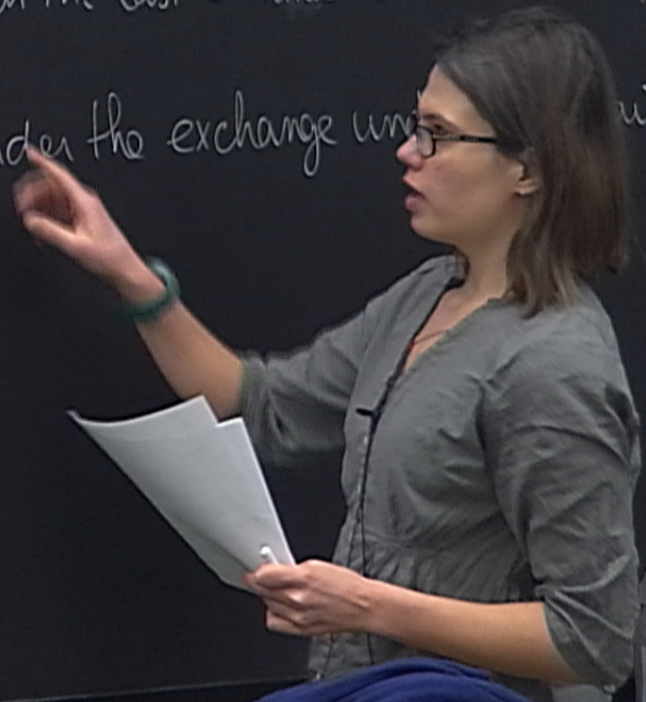
\downarrow
 6 dof

$R_{\mu\nu}$

$R_{\mu\nu\alpha\beta}$

- antisymmetric in the 1st 2 indices: 3 pairs & one index running from 1-3
- antisymmetric in the last 2 indices → //
- symmetric under the exchange of the first two indices

\downarrow \downarrow
 1 to 3 1 to 3
 \downarrow
 \equiv
 6 dof
 $R_{\mu\nu} = 6 \text{ dof}$



$R_{\mu\nu\sigma\tau}$

- antisymmetric in the 1st 2 indices: 3 pairs & one index running from 1-3
- antisymmetric in the last 2 indices → //
- symmetric under the exchange under the 2 pairs

$(R_{[\mu\nu\sigma\tau]} = 0)$

$\downarrow \quad \downarrow$
 1 to 3 1 to 3
 \downarrow
 6 dof
 $R_{\mu\nu} = 6 \text{ dof}$

⇒ 3D gravity: no local degree of freedom

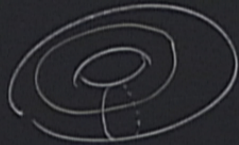
Global dof depend on the topology of the manifold.

↳ fundamental group.

\Rightarrow 3D gravity: no local degree of freedom

Global dof depend on the topology of the manifold.

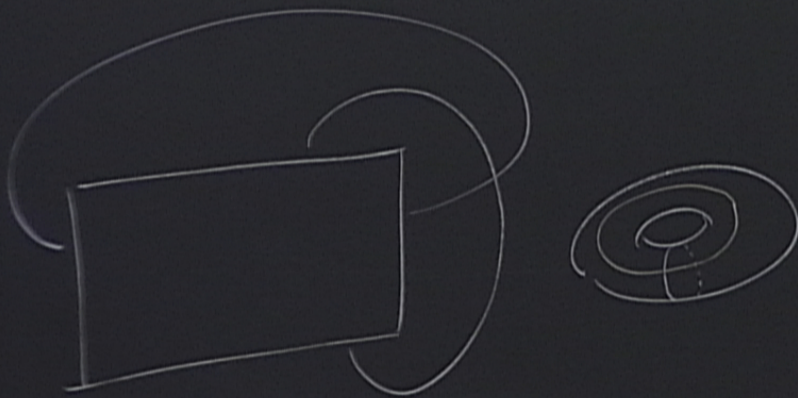
\hookrightarrow fundamental group.



\Rightarrow 3D gravity: no local degree of freedom

Global dof depend on the topology of the manifold.

\hookrightarrow fundamental group



$$M = \mathbb{R} \times \Sigma$$

S^2 , g = nb of holes
= genus of Σ .



II. C Triads n-beins.

orthonormal frame at each point of the manifold.
field of 3 vectors: e_i where $i=1-3$

II. C Triads

n-beins.

orthonormal frame at each point of the manifold.

field of 3 vectors: e_i^{μ} where $i=1..3$.

orthogonality
 \Rightarrow

$$g_{\mu\nu} e_i^{\mu} e_j^{\nu} = e_{i\alpha} e_j^{\alpha} = \delta_{ij}$$

\uparrow
co-triad

II. C Triads n-beins.

orthonormal frame at each point of the manifold

field of 3 vectors: e_i^{μ} where $i=1, 2, 3$ (if Minkowski signature)

orthogonality $g_{\mu\nu} e_i^{\mu} e_j^{\nu} = e_i^{\mu} e_j^{\nu} = \delta_{ij}$ ← Euclidean signature

↑
C-triad

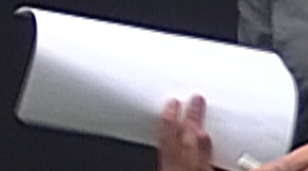
ω^i, ω_i

the manifold.

$\omega = e^i$ where $i=1..3$ (Minkowski signature)

$\omega_j = e_i \omega_j = \delta_{ij}$ ← Euclidean signature.

↑
co-triad



at each point of the manifold.

field of 3 vectors: $\boxed{e_i^\mu}$ where $i=1..3$ (Minkowski signature)

orthogonality

$$g_{\mu\nu} e_i^\mu e_j^\nu = \boxed{e_{i\nu}} e_j^\nu = \delta_{ij} \leftarrow \text{Euclidean signature}$$

↑
co-triad

frame at each point of the manifold.

field of 3 vectors: $\boxed{e_i^\mu}$ where $i=1..3$ $(\eta_{ij} \leftarrow \text{if Minkowski signature})$

orthogonality \Rightarrow

$$g_{\mu\nu} e_i^\mu e_j^\nu = \boxed{e_{ij}^\nu} e_j^\nu = \delta_{ij} \leftarrow \text{Euclidean signature}$$

co-triad

co-triad $e_i^\nu = g_{\mu\nu} e_j^\mu \delta^{ij}$

orthonormal frame at each point of the manifold.

field of 3 vectors: $\boxed{e_i^\mu}$ where $i=1..3$. $(\eta_{ij} \leftarrow \text{if Minkowski})$

orthogonality $\Rightarrow g_{\mu\nu} e_i^\mu e_j^\nu = \boxed{e_{ij}^\nu} e_j^\nu = \delta_{ij} \leftarrow \text{Euclidean}$

\uparrow
co-triad

Contracting by e_k^ν $\Rightarrow g_{\mu\nu} = e_\mu^i e_\nu^d \delta_{ij}$

co-triad

$$g_{\mu\nu} = e^i_{\mu} e^j_{\nu} \delta_{ij}$$

$$g_{\mu\nu} = e^i_{\mu} e^j_{\nu} \delta_{ij}$$

↓
6

↓
g.

dof

$$g_{\mu\nu} = e^i{}_\nu e^j{}_\mu \delta_{ij}$$

↓
6

↓
g.

dof

additional symmetry (SO(3) gauge symmetry)

$$e^i{}_\nu(\alpha) = R^i{}_j e^j{}_\nu(\alpha)$$

$$g_{\mu\nu} = e^i_{\nu} e^j_{\mu} \delta_{ij}$$

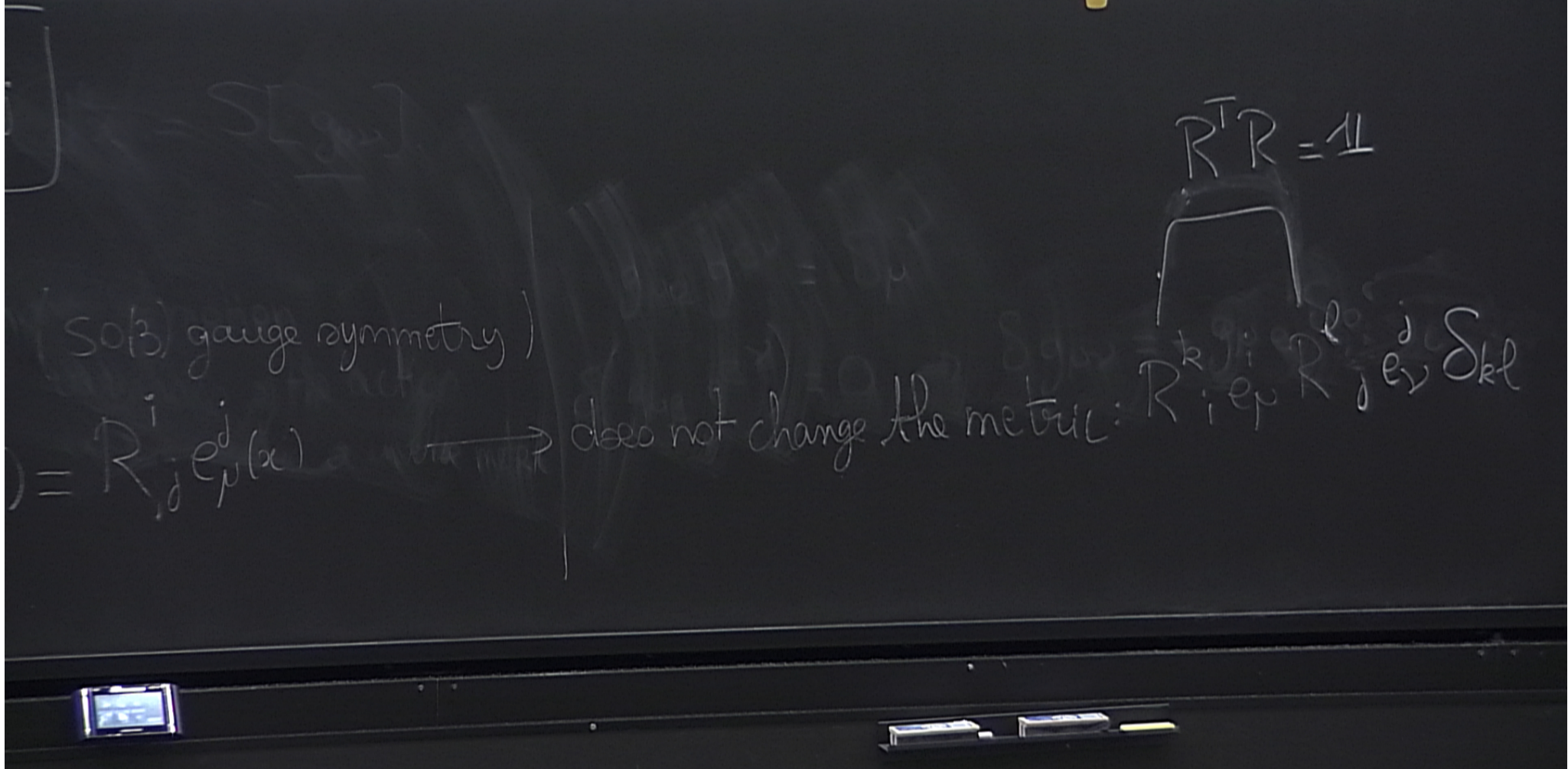
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6

g

additional symmetry (SO(3) gauge symmetry)

$$e^{i'}_{\mu}(\alpha) = R^i_{j'} e^j_{\mu}(\alpha) \quad \text{does not change}$$



$SO(3)$

(SO(3) gauge symmetry)

$$e^j_\mu(x) \rightarrow R^i_j e^i_\mu(x)$$

does not change the metric:

$$R^T R = 1$$
$$S^{kl} = R^k_j R^l_i e^j_\mu e^i_\nu S^{kl}$$

$SL(2, \mathbb{R})$

(SO(3) gauge symmetry)

$$R^i_j e^j_\mu(x)$$

does not change the metric:

$$R^T R = \mathbb{1}$$
$$R^k_j R^l_i e^j_\nu e^i_\mu \delta_{kl} = g_{\mu\nu}$$

objects with internal indices: i, j ← transformation rule w/ internal rotation
with space-time: μ, ν ← spacetime tensor.

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→ covariant derivative ∇_μ ($\Gamma_{\mu\nu}^\alpha$: L-C connection)

with internal indices: i, j \leftarrow transformation rule w/ internal rotation

with space-time: μ, ν \leftarrow space time tensor.

\rightarrow standard covariant derivative ∇_μ ($\Gamma_{\mu\nu}^\sigma$: L-C connection)

Δ internal indices: covariant derivative / internal indices: $\omega_{\mu k}^i$: spin-connection

with internal indices: i, j \leftarrow transformation rule w/ internal rotation

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standard covariant derivative ∇_μ ($\Gamma_{\mu\nu}^\sigma$: L-C connection)

Δ internal indices: covariant derivative / internal indices: w_{jk}^i : spin-connection
 $\hookrightarrow D_\mu$.

with internal indices: i, j \leftarrow transformation rule w/ internal rotation

with space-time: μ, ν \leftarrow space time tensor.

\rightarrow standard covariant derivative ∇_μ ($\Gamma_{\mu\nu}^\sigma$: L-C connection)

Δ internal indices: covariant derivative / internal indices: ω spin-connection

$$\underline{D_\mu \Phi^a = \partial_\mu \Phi^a + \omega_{\mu b}^a \Phi^b}$$

$$D_\mu N_\nu^j = \partial_\mu N_\nu^j - \Gamma_{\mu\nu}^e N_e^j + \omega_{\mu k}^j N_\nu^k$$

internal indices: covariant derivative / internal indices: $\omega_{\mu k}^i$: spin-connection

$$\underline{D_{\mu} \Phi^{\dot{a}} = \partial_{\mu} \Phi^{\dot{a}} + \omega_{\mu k}^{\dot{a}} \Phi^{\dot{k}}}$$

$\rightarrow \nabla_{\mu} \Phi^{\dot{a}} = \partial_{\mu} \Phi^{\dot{a}}$

$$D_{\mu} N_{\dot{a} \dot{b}}^{\dot{c}} = \partial_{\mu} N_{\dot{a} \dot{b}}^{\dot{c}} - \Gamma_{\mu \nu}^{\dot{c}} N_{\dot{a} \dot{b}}^{\dot{e}} + \omega_{\mu k}^{\dot{c}} N_{\dot{a} \dot{b}}^{\dot{k}}$$

$$D_\mu N_{\nu}^d = \partial_\mu N_{\nu}^d - \Gamma_{\mu\nu}^e N_e^d + \omega_{\mu k}^d N_{\nu}^k$$

Compatibility condition $D_\mu e_{\nu}^d = 0$

$$D_\mu N_{\nu}^d = \partial_\mu N_{\nu}^d - \Gamma_{\mu\nu}^e N_e^d + \omega_{\mu k}^d N_{\nu}^k$$

Compatibility condition $\boxed{D_\mu e_{\nu}^d = 0}$ → makes sure that the derivative commutes with the contraction.

$$\partial_\mu e_{\nu}^d - \Gamma_{\mu\nu}^e e_e^d + \omega_{\mu k}^d e_{\nu}^k = 0$$

contraction
by e_k^{ν}

$$\boxed{\omega_{\mu k}^d = -e_k^{\nu} \nabla_\mu e_{\nu}^d}$$

$$D_\mu N_{\nu}^d = \partial_\mu N_{\nu}^d - \Gamma_{\mu\nu}^e N_e^d + \omega_{\mu k}^d N_{\nu}^k$$

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contraction
by e_k^{ν}

$$\omega_{\mu k}^d = -e_k^{\nu} \nabla_{\mu} e_{\nu}^d \quad (\Gamma)$$

$$\partial_\mu \Omega_{\nu}^d = \partial_\mu \Omega_{\nu}^d - \Gamma_{\mu\nu}^e \Omega_e^d + \omega_{\mu\nu}^d \Omega_{\nu}^k$$

Compatibility condition $\boxed{D_\mu e_{\nu}^d = 0}$ \rightarrow makes sure that the derivative commutes with the contraction

$$\partial_\mu e_{\nu}^d - \Gamma_{\mu\nu}^e e_e^d + \omega_{\mu\nu}^d e_{\nu}^k = 0$$

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