

Title: PSI 2015/2016 Cosmology - Lecture 4

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Abstract:

YESTERDAY: DYNAMICS

$$ds^2 = -dt^2 + a^2(t) g_{ij}^{(spac)} dx^i dx^j$$
$$T_{\mu\nu} = P g_{\mu\nu} + (\rho + P) u_\mu u_\nu$$

for ISO $\Rightarrow a(t), \rho(t), P(t)$

(FE) BECOME ODEs

$$H^2 = \frac{8\pi G}{3} \rho - \frac{K}{a^2}, \quad \dot{\rho} = -3H(\rho + P), \quad \rho = \rho(P)$$

$$t \rightarrow \eta, \quad a = A b^\alpha(\eta)$$

$$H = \frac{\dot{a}}{a} = \left(\frac{da}{d\eta} \frac{d\eta}{dt} \right) \frac{1}{a} = \frac{a'}{a^2} = \frac{A \alpha b^{\alpha-1} b'}{A^2 b^{2\alpha}} = \frac{\alpha b'}{A b^{1+\alpha}}$$

$$1 = \frac{3H^2}{8\pi G_N \rho} + \frac{3K}{8\pi G_N \rho a^2} = \frac{3\alpha^2 b'^2}{A^2 b^{2+2\alpha}} \frac{b^{3\alpha(1+\alpha)}}{8\pi G_N \left(\frac{A}{a_0} \right)^{3(1+\alpha)} \rho_0} + \frac{3K}{8\pi G_N \rho_0}$$

$$a^2 \quad A^2 b^{2\alpha} \quad \underline{A b^{1+\alpha}}$$

$$= \frac{3H^2}{8\pi G_N \rho} + \frac{3K}{8\pi G_N \rho a^2} = \frac{3\alpha^2 b'^2}{A^2 b^{2+2\alpha}} \frac{b^{3\alpha(1+n\omega)}}{8\pi G_N \left(\frac{A}{a_0}\right)^{3(1+n\omega)} \rho_0} + \frac{3K}{8\pi G_N \rho_0 \left(\frac{A}{a_0}\right)^{-3(1+n\omega)} A^2 b^{2\alpha}}$$

$$3\alpha(1+n\omega) - 2 - 2\alpha = 0 \Rightarrow \alpha = \frac{2}{1+3n\omega}$$

$$1 = \alpha^2 b'^2 + K b^2 \Rightarrow b = S_R (m/\alpha)$$

$$a|q| = \left[\frac{8\pi G_N}{3} \rho_0 a_0^{3(1+n\omega)} \right]^{\frac{\alpha}{2}} S_R^{\alpha} (m/\alpha)$$

$$\alpha = \frac{2}{1+3n\omega}$$

FOR $K=0$:

COMPONENT	w	DENSITY	$a(t)$	$a(t)$
RADIATION	$1/3$	$1/a^4$	$t^{-1/2}$	γ
MATTER	0	$1/a^3$	$t^{2/3}$	γ^2
Λ	-1	ρ^0	e^{Ht}	$-1/\gamma$

II) MATTER

a) BRIEF THERMAL HISTORY

$$a \leftrightarrow T \leftrightarrow z \leftrightarrow t$$

$$\frac{T}{T_0} = \frac{a_0}{a} = 1+z = \dots$$

• RAD DOM. ERA.

$$T [\text{MeV}] \approx \frac{0(1)}{\sqrt{t [\text{s}]}}$$

$10^{-12} (1 \text{ MeV})$ — $e^- e^+$ ANNIHILATION $\rightarrow \bar{\nu}_e / 10^{-9}$ ($T > T_\nu$)
 $1 (1 \text{ MeV})$ — NEUTRINO DECOUPLING $\rightarrow$ NEUTRON/PROTON RATIO FREEZE OUT
 2 — DN FREEZE OUT

$10^{-5} (150 \text{ GeV})$ — QCD PHASE TRANSITION
 $10^{-10} (100 \text{ GeV})$ — EW PHASE TRANSITION
 ACCESSIBLE BY LHC ... STANDARD MODEL
 $10^{-14} (10 \text{ TeV})$ — $\{$
 • BARYOGENESIS $\rightarrow$ BARYON ASYMMETRY $\eta = \frac{n_b}{n_{\text{gr}}} \sim 10^{-9}$
 • TOP DEFECTS, SUSY?
 • INFLATION
 $10^{-43} (10^{19} \text{ GeV})$ — PLANCK SCALE ... NON-PERTURBATIVE QG



10^{-10} (10.1 MeV) — BB-NUCLEOSYNTHESIS: H, He^{++}, Li^{++}
 6 (0.5 MeV) — $e^- e^+$ ANNIHILATION $\rightarrow e^-/10^9 n$ ($T_\gamma > T_\nu$)
 1 (1 MeV) — NEUTRINO DECOUPLING $\rightarrow$ NEUTRON/PROTON RATIO FREEZE OUT
 2 — DM FREEZE OUT

10^{-10} (100 GeV) — EW PHASE TRANSITION
 10^{-14} (10 TeV) — ACCESSIBLE BY LHC ... STANDARD MODEL
 10^{-43} (10^{13} GeV) — PLANCK SCALE ... NON-PERTURBATIVE QG

$(10^9 + 1) p \dots 10^9 \bar{p}$

$\eta = \frac{n_b}{n_{\gamma}} \sim 10^{-9}$

• BARYOGENESIS $\rightarrow$ BARYON ASYMMETRY
 • TOP DEFECTS, SUSY?
 • INFLATION

$$A^{-6^2} \quad \underline{A b^{1+2}}$$

$$1 = \frac{3H^2}{8\pi G \rho} + \frac{3K}{8\pi G \rho a^2} = \frac{3\alpha^2 b^2}{a^2 (1+2\alpha)} \frac{b^{3\alpha(1+n)}}{(A)^{3(1+n)}} + \frac{3K}{8\pi G \rho_0 \left(\frac{A}{a}\right)^{-3(1+n)} \frac{b^{3\alpha(1+n)}}{A^2} \frac{2\alpha}{A^2}} b^2$$

b) TDS IN EXPANDING UNIVERSE

- DISTRIBUTION FUNCTION
IN A PHASE SPACE.

NUMBER OF PARTICLES OF SPECIE 1

$$dN_1 = f_1(t, \vec{x}, \vec{p}) \frac{d^3x d^3p}{(2\pi\hbar)^3}$$

VOLUME OF SMALLEST
CELL A PARTICLE CAN BE
LOCALIZED.

PARTICLE DENSITY

$$n_1(t, \vec{x}) = \int \frac{d^3p}{(2\pi)^3} f_1(t, \vec{x}, \vec{p})$$

ENERGY DENSITY

$$\epsilon = \int \frac{d^3p}{(2\pi)^3} E(\vec{x}, \vec{p})$$

$\frac{1}{\Gamma} = \text{COLLISION TIME} \ll \text{EXPANSION TIME} \approx t_H = \frac{1}{H} = \frac{dt}{d \ln a}$

$\Rightarrow$ LOCAL TD EQ. ... MAXIMAL POSSIBLE ENTROPY

$$E = \sqrt{\vec{p}^2 + m^2}, \quad m: \dots \text{CHEM. POT.}$$

$$f_i = f_i(|\vec{p}|)$$

$$d^3p = d\Omega |\vec{p}|^2 d|\vec{p}| = d\Omega \sqrt{E^2 - m^2} E dE$$

$$n_i = \frac{1}{2\pi^2} \int_{m_i}^{\infty} \frac{\sqrt{E^2 - m_i^2} E}{e^{\frac{E - \mu_i}{T}} + 1} dE$$

$$\frac{1}{H} = \text{COLLISION TIME} \ll \text{EXPANSION TIME} \approx t_H = \frac{1}{H} = \frac{dt}{d \log a}$$

$\Rightarrow$ LOCAL TD EQ. ... MAXIMAL POSSIBLE ENTROPY

$$E = \sqrt{p^2 + m^2}, \quad \mu: \dots \text{CHE}$$

$$f_i = f_i(|\vec{p}|)$$

$$d^3p = d\Omega |\vec{p}|^2 d|\vec{p}| = d\Omega \sqrt{E^2 - m^2}$$

1) RELATIVISTIC LIMIT. $T \gg m_i, \mu_i$

$$n_i = \frac{1}{2\pi^2} \int_0^\infty \frac{E^2 dE}{e^{E/T} \pm 1} = \left| x = E/T \right|_{dE = T dx} \approx \frac{T^3}{2\pi^2} \int_0^\infty \frac{x^2 dx}{e^x \pm 1}$$

$$n_i = \begin{cases} \frac{\pi^2}{15} T^3 & \text{BOSONS} \\ \frac{3}{4} \frac{\pi^2}{15} T^3 & \text{FERMIONS} \end{cases}$$

DIM ANAL. $\hbar = c = k_B = 1$

$$\Rightarrow [E] = [T] = 1/L$$

$$\frac{1}{\Gamma} = \text{COLLISION TIME} \ll \text{EXPANSION TIME} \approx t_H = \frac{1}{H} = \frac{dt}{d \log a}$$

$\Rightarrow$ LOCAL TD EQ. ... MAXIMAL POSSIBLE ENTROPY

$$E = \sqrt{p^2 + m^2}, \quad \mu_i \dots \text{CHEM. POT.}$$

$$f_i = f_i(p_i)$$

$$d^3p = d\Omega |\vec{p}|^2 d|\vec{p}| = d\Omega \sqrt{E^2 - m^2} E dE$$

1) RELATIVISTIC LIMIT. $T \gg m_i, \mu_i$

$$n_i = \frac{1}{2\pi^2} \int_0^\infty \frac{E^2 dE}{e^{E/T} \pm 1} = \left| \begin{matrix} x = E/T \\ dE = T dx \end{matrix} \right| \approx \frac{T^3}{2\pi^2} \int_0^\infty \frac{x^2 dx}{e^x \pm 1}$$

$$n_i = \begin{cases} \frac{\pi^2}{15} T^3 & \text{BOSONS} \\ \frac{3}{4} \frac{\pi^2}{15} T^3 & \text{FERMIONS} \end{cases}$$

DIM ANAL. $\hbar = c = k_B = 1$

$$\Rightarrow [E] = [T] = 1/L$$

$$n_i = (\dots) T^3, \quad g_i = (\dots) T^4, \quad p_i = (\dots) T^4$$

- 10 - NO-TATTER EQ
- 3 MIN (0.1 MeV) - BB-NUCLEOSYNTHESIS: H, He^{++}, Li^{++}
- 6 (0.5 MeV) - $e^- e^+$ ANNIHILATION $\rightarrow 10^{-9} n$ ($T_p > T_\nu$)
- 1 (1 MeV) - NEUTRINO DECOUPLING $\rightarrow$ NEUTRON/PROTON RATIO FREEZE OUT
- 2 - DM FREEZE OUT

ii) NON-RELAT LIMIT $T \ll m_i$, $\mu_i < m_i$ (NEGLECT $\pm$)

$$n_i = \frac{1}{2\pi^2} \int_{m_i}^{\infty} \frac{\sqrt{E^2 - m_i^2} E dE}{e^{\frac{E - \mu_i}{T}}} = \left| \begin{array}{l} W = \frac{E - m_i}{T} \\ E = TW + m_i \end{array} \right| = \frac{1}{2\pi^2} e^{\frac{\mu_i - m_i}{T}} \int_0^{\infty} \sqrt{m_i^2 T^2 + 2m_i T W} dW$$