

Title: Quantum Field Theory for Cosmology - Achim Kempf - Lecture 12

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Abstract:

Recall strategy:

SR, 1st Q
Hamiltonian
formalism

step 1
Legendre transform
(equivalence) →

SR, 1st Q.
Lagrangian
formalism

step 2
allow
curvature ↓

GR, 1st Q
Hamiltonian
formalism

step 3
Legendre transform
(equivalence) ←

GR, 1st Q
Lagrangian
formalism

we are here

step 4 ↓

GR, 2nd Q
Hamiltonian
formalism

Dyson Schwinger eqns are same
(equivalence) ←

GR, 2nd Q
Lagrangian formalism
(Path integral of QFT)

$$S_{\text{KG}} = \frac{1}{2} \int_{\mathbb{R}^4} \left(g^{\mu\nu}(x) \phi_{,\mu}(x) \phi_{,\nu}(x) - m^2 \phi^2(x) \right) \sqrt{|g|} d^4x$$

↖ we assume that the coordinate system is such, for simplicity.

It yields, via $\frac{\delta S_{\text{KG}}}{\delta \phi} = 0$ the Klein Gordon eqn:

$$\frac{1}{\sqrt{|g(x)|}} \frac{\partial}{\partial x^\nu} \left(g^{\mu\nu}(x) \sqrt{|g(x)|} \phi_{,\mu}(x) \right) + m^2 \phi(x) = 0 \quad (\text{KG})$$

* We read off the Lagrangian:

$$L_{\text{KG}}(t) = \frac{1}{2} \int_{\mathbb{R}^3} \left(g^{\mu\nu}(x,t) \phi_{,\mu}(x,t) \phi_{,\nu}(x,t) - m^2 \phi^2(x,t) \right) \sqrt{|g|} d^3x$$

Step 3: Legendre transform back to the Hamiltonian form

* The transform:

$$H(\phi, \pi, t) \xleftarrow[\text{Legendre transform}]{\pi(x,t) := \frac{\delta L}{\delta \phi_{,\mu}(x,t)}} L(\phi, \phi_{,\mu}, t)$$

* Thus, the canonically conjugate field $\pi(x,t)$ reads:

$$\pi(x,t) = \frac{\delta L}{\delta \phi_{,\mu}(x,t)} = \sqrt{|g(x,t)|} g^{\mu\nu}(x,t) \phi_{,\nu}(x,t)$$

$$L_{\text{ve}}(t) = \frac{1}{2} \int_{\mathbb{R}^3} \left(g^{\mu\nu}(x,t) \phi_{,\mu}(x,t) \phi_{,\nu}(x,t) - m^2 \phi^2(x,t) \right) \sqrt{|g|} d^3x$$

Step 3: Legendre transform back to the Hamiltonian form

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* Explicitly:

* The transform:

$$H(\phi, \pi, t) \xleftarrow[\text{Legendre transform}]{\pi(x,t) := \frac{\delta L}{\delta \dot{\phi}_r(x,t)}} L(\phi, \dot{\phi}_r, t)$$

* Thus, the canonically conjugate field π

$$\pi(x,t) = \frac{\delta L}{\delta \dot{\phi}_r(x,t)} = \sqrt{|g(x,t)|} g^{0r}(x,t) \dot{\phi}_r(x,t)$$

* Explicitly:

$$\pi(x,t) = \sqrt{|g|} g^{00} \dot{\phi}_0 + \sum_{i=1}^3 \sqrt{|g|} g^{0i} \dot{\phi}_i$$

* Explicitly:

$$\pi(x,t) = \sqrt{|g|} g^{00} \phi_{,0} + \sum_{i=1}^3 \sqrt{|g|} g^{0i} \phi_{,i}$$

* Thus, we can also express $\phi_{,0}(x,t)$ in terms of $\phi_{,i}(x,t)$ and $\pi(x,t)$ (as will be necessary after the Legendre transform):

$$\phi_{,0}(x,t) = \frac{\pi(x,t)}{g^{00} \sqrt{|g|}} - \sum_{i=1}^3 \frac{g^{0i}}{g^{00}} \phi_{,i}(x,t)$$

(of course, g^{00} depends on x,t too)

* The Hamiltonian:

$$H(\phi, \pi) = \int \left(\pi(x,t) \dot{\phi}(x,t) - \mathcal{L}(\phi, \pi, x, t) \right) dx$$

* Thus, we can also express $\phi_{,a}(x,t)$ in terms of $\phi(x,t)$ and $\pi(x,t)$ (as will be necessary after the Legendre transform):

$$\phi_{,a}(x,t) = \frac{\pi(x,t)}{g^{00} \sqrt{|g|}} - \sum_{i=1}^3 \frac{g^{0i}}{g^{00}} \phi_{,i}(x,t) \quad (V)$$

↖
(of course, $g^{\mu\nu}$ depends on x, t too)

* The Hamiltonian:

$$H(\phi, \pi) = \int_{\mathbb{R}^3} \pi(x,t) \phi_{,0}(x,t) d^3x - \frac{1}{2} \int_{\mathbb{R}^3} \left(g^{\mu\nu} \phi_{,\mu} \phi_{,\nu} - m^2 \phi^2 \right) \sqrt{|g|} d^3x$$

↑
 Why don't we need a factor of $\sqrt{|g|}$ for covariance here? Because π has it built in!

$= L(\phi, \phi_{,a}(\phi, \pi), t)$

$$H(\phi, \pi) = \int_{\mathbb{R}^3} \pi(x, t) \phi_{,0}(x, t) d^3x - \frac{1}{2} \int_{\mathbb{R}^3} \left(g^{\mu\nu} \phi_{,\mu} \phi_{,\nu} - m^2 \phi^2 \right) \sqrt{|g|} d^3x$$

Why don't we need a factor of $\sqrt{|g|}$ for covariance here? Because π has it built in!

$$= L(\phi, \phi_{,\nu}(\phi, \pi), t)$$

* In H , one needs to express all occurring $\phi_{,\nu}$ in terms of the new variables ϕ and π , by using (V), to obtain $H(\phi, \pi, t)$.

→ Exercise: Calculate $H(\phi, \pi, t)$ and simplify the expression as far as possible.

* The equations of motion:

We know from the general properties of the Legendre transform that the equations of motion now take the form:

$$\frac{d}{dt} \phi(x,t) = \frac{\delta H(\phi, \pi, t)}{\delta \pi(x,t)}, \quad \frac{d}{dt} \pi(x,t) = - \frac{\delta H(\phi, \pi, t)}{\delta \phi(x,t)}$$

* Exercise: Verify that these eqns are equivalent to (KG).

We are now ready to 2nd quantize:

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* Exercise: Verify that these eqns are equivalent to (KG).

We are now ready to 2nd quantize:

SR, 1st Q

step 1
Legendre transform

SR, 1st Q

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Solving the quantized theory is to solve:

1.) Commutation relations:

$$[\hat{\phi}(x,t), \hat{\pi}(x',t)] = i \hbar \delta^3(x-x')$$

$$[\hat{\phi}(x,t), \hat{\phi}(x',t)] = 0$$

$$[\hat{\pi}(x,t), \hat{\pi}(x',t)] = 0$$

} (CCRs)

2.) Hermiticity:

$$\hat{\phi}^\dagger(x,t) = \hat{\phi}(x,t), \quad \hat{\pi}^\dagger(x,t) = \hat{\pi}(x,t) \quad (HC)$$

3.) Equations of motion:

In the Heisenberg picture, they are formally unchanged:

$$\frac{d}{dt} \hat{f}(t, \pi) = \frac{i}{\hbar} [\hat{f}, \hat{H}] \quad \text{for } \hat{f} = \hat{\phi}, \hat{\pi}, \text{ etc}$$

Namely:

$$\left(\frac{1}{\sqrt{|g|}} \frac{\partial}{\partial x^\mu} g^{\mu\nu} \sqrt{|g|} \frac{\partial}{\partial x^\nu} + m^2 \right) \hat{\phi}(x, t) = 0 \quad (\text{EOM})$$

and:

In the Heisenberg picture, they are formally unchanged:

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Namely:

$$\left(\frac{1}{\sqrt{|g|}} \frac{\partial}{\partial x^\mu} g^{\mu\nu} \sqrt{|g|} \frac{\partial}{\partial x^\nu} + m^2 \right) \hat{\phi}(x, t) = 0 \quad (\text{E.O.M.1})$$

and:

$$\hat{\pi}(x, t) = \sqrt{|g|} g^{\mu\nu} \frac{\partial}{\partial x^\nu} \hat{\phi}(x, t) \quad (\text{E.O.M.2})$$

How to solve the CCR, HC and EoM equations?

Recall: the solution we obtained on Minkowski space:

$$\hat{\phi}(x,t) = \int_{\mathbb{R}^3} \frac{1}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_k}} \left(e^{-i\omega_k t + ikx} a_k + e^{i\omega_k t - ikx} a_k^\dagger \right)$$

Number-valued solutions to the K.G. equation

The a_k, a_k^\dagger take care of the CCRs

and $\hat{\pi}(x,t) = \dot{\hat{\phi}}(x,t)$

Strategy:

- * ensure hermiticity, HC, by construction
- * separate the CCR and EoM problems

Ansatz:

$$\hat{\phi}(x,t) := \sum_k u_k(x,t) a_k + u_k^\dagger(x,t) a_k^\dagger$$

Recall: the solution we obtained on Minkowski space:

$$\hat{\phi}(x,t) = \int_{\mathbb{R}^3} \frac{1}{(2\pi)^3} \frac{1}{\sqrt{2\omega_k}} \left(e^{-i\omega_k t + ikx} + e^{i\omega_k t - ikx} a_k^+ \right) dk$$

Number-valued solutions to the K.G. equation

care of the LLCS

and $\hat{\pi}(x,t) = \dot{\hat{\phi}}(x,t)$

Strategy:

- * ensure hermiticity, HC, by ...
- * separate the CCR and E...

Ansatz:

$$\hat{\phi}(x,t) := \sum_k u_k b_k + u_k^* a_k^+$$

$$\hat{\pi}(x,t) := \sqrt{|g|} g^{0\nu} \frac{\partial}{\partial x^\nu} \hat{\phi}(x,t)$$

(not be "momentum"!) |

$$\hat{\phi}(x,t) = \int_{\mathbb{R}^3} \frac{1}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_k}} \left(e^{-i\omega_k t + ikx} a_k + e^{i\omega_k t - ikx} a_k^\dagger \right) d^3k$$

Number-valued solutions to the K.G. equation

and $\hat{\pi}(x,t) = \dot{\hat{\phi}}(x,t)$

Strategy:

- * ensure hermiticity, HC, by construction
- * separate the CCR and EoM problems:

Ansatz:

$$\hat{\phi}(x,t) := \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^\dagger$$

$$\hat{\pi}(x,t) := \sqrt{|g|} g^{0\nu} \frac{\partial}{\partial x^\nu} \hat{\phi}(x,t)$$

(k must be
some direction)

* The 1st equation of motion: ✓

$$\left(\frac{1}{\sqrt{|g|}} \frac{\partial}{\partial x^\mu} g^{\mu\nu} \sqrt{|g|} \frac{\partial}{\partial x^\nu} + m^2 \right) \hat{\phi}(x,t) = 0 \quad (\text{EOM})$$

This eqn holds because in our ansatz,

$$\hat{\phi}(x,t) := \sum_{\mathbf{k}} u_{\mathbf{k}}(x,t) a_{\mathbf{k}} + u_{\mathbf{k}}^*(x,t) a_{\mathbf{k}}^*$$

the $a_{\mathbf{k}}$ are constant operators while the

functions $u_{\mathbf{k}}(x,t)$ are assumed to solve (EOM)

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Strategy:

- * ensure normality, P.C., by construction
- * separate the CCR and EoM problems:

Ansatz:

$$\hat{\phi}(x,t) = \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^\dagger$$

$$\hat{\pi}(x,t) := \sqrt{|g|} g^{0\nu} \frac{\partial}{\partial x^\nu} \hat{\phi}(x,t)$$

(k need not be a "momentum"!)

□ Here, we use the easy-to-compute operators that obey

$$[a_k, a_{k'}^\dagger] = \delta_{kk'}$$

□ And, we use some number solutions $u_k(x,t)$

Now check CCR:

$$[\hat{\phi}(x,t), \hat{\pi}(x,t)]$$

$$= \left[\sum_i u_i(x,t) a_i + u_i^*(x,t) a_i^\dagger, \sqrt{|g|} g^{00} \sum_{k'} \left(\frac{\partial}{\partial x^k} u_k(x,t) a_{k'} + \frac{\partial}{\partial x^k} u_k^*(x,t) a_{k'}^\dagger \right) \right]$$

$$= \sqrt{|g|} g^{00}(x,t) \sum_{k,k'} \left(u_k(x,t) \frac{\partial}{\partial x^k} u_{k'}^*(x,t) - u_k^*(x,t) \frac{\partial}{\partial x^k} u_{k'}(x,t) \right) \delta_{k,k'}$$

$$= \sqrt{|g|} g^{00} \sum \left(u_k(x,t) \frac{\partial}{\partial x^k} u_k^*(x,t) - u_k^*(x,t) \frac{\partial}{\partial x^k} u_k(x,t) \right) \stackrel{!}{=} i \delta^3(\vec{x} - \vec{x}')$$

Now check CCR:

$$[\hat{\phi}(x,t), \hat{\pi}(x,t)]$$

$$= \left[\sum_{\nu} u_{\nu}(x,t) a_{\nu} + u_{\nu}^*(x,t) a_{\nu}^{\dagger}, \sqrt{|g|} g^{\mu\nu} \sum_{\kappa} \left(\frac{\partial}{\partial x^{\mu}} u_{\kappa}(x,t) \right) a_{\kappa} + \left(\frac{\partial}{\partial x^{\mu}} u_{\kappa}^*(x,t) \right) a_{\kappa}^{\dagger} \right]$$

$$= \sqrt{|g|} g^{\mu\nu}(x,t) \sum_{\nu, \kappa} \left(u_{\nu}(x,t) \frac{\partial}{\partial x^{\mu}} u_{\kappa}^*(x,t) - u_{\kappa}^*(x,t) \frac{\partial}{\partial x^{\mu}} u_{\nu}(x,t) \right)$$

$$= \sqrt{|g|} g^{\mu\nu} \sum_{\nu, \kappa} \left(u_{\nu}(x,t) \frac{\partial}{\partial x^{\mu}} u_{\kappa}^*(x,t) - u_{\kappa}^*(x,t) \frac{\partial}{\partial x^{\mu}} u_{\nu}(x,t) \right) \stackrel{!}{=} i \delta^{\mu\nu}$$

$$[\hat{\phi}(x,t), \hat{\pi}(x,t)]$$

$$= \left[\sum_i u_i(x,t) a_i + u_i^*(x,t) a_i^\dagger, \sqrt{|g|} g^{uv} \sum_{k'} \left(\frac{\partial}{\partial x^v} u_k(x,t) \right) a_{k'} + \left(\frac{\partial}{\partial x^v} u_k^*(x,t) \right) a_{k'}^\dagger \right]$$

$$= \sqrt{|g|} g^{uv}(x,t) \sum_{k,k'} \left(u_k(x,t) \frac{\partial}{\partial x^v} u_{k'}^*(x,t) - u_{k'}^*(x,t) \frac{\partial}{\partial x^v} u_k(x,t) \right) \delta_{k,k'}$$

$$= \sqrt{|g|} g^{uv} \sum_k \left(u_k(x,t) \frac{\partial}{\partial x^v} u_k^*(x,t) - u_k^*(x,t) \frac{\partial}{\partial x^v} u_k(x,t) \right) = i \delta^3(\vec{x} - \vec{x}')$$

Conclusion so far:

Our ansatz

$$\hat{\psi}(x,t) := \sum u_n(x,t) a_n + u_n^*(x,t) a_n^+$$

solves the QFT, i.e., HC, EoM and CCR

if we can find a set of number-valued solutions

$$\{u_n(x,t)\}$$

of the Klein Gordon equation that always:

When do such $\{u_\alpha(u, t)\}$ exist? I.e., when does the ansatz succeed?

Proposition: \square Assume spacetime is "globally hyperbolic",
i.e., that it possesses a foliation by Cauchy surfaces,
i.e., that it is topologically of the form:

$$\mathbb{R} \times \mathcal{M}$$

\uparrow any 3-dim differentiable manifold

\square In this case, spacetime possesses no closed timelike curves (no travel into the past), i.e., initial conditions set on the Cauchy surfaces determine the solution everywhere

\square Then, such a set of functions $\{u_\alpha\}$ can be shown to exist

Proof:

□ Consider the vector space, V , of all real-valued solutions of the Klein Gordon equations.

□ We define a bi-linear form $(,)$ on V . For all $f, h \in V$:

$$(f, h) := \int_{\Sigma} d\Sigma_{\nu} \sqrt{g} g^{\mu\nu} (f \partial_{\nu} h - h \partial_{\nu} f)$$

↳ any spacelike hypersurface
i.e. set of points of equal time.

□ Proposition: (f, h) is independent of choice of Σ .

Proof: Later (uses Stokes' theorem and K.G. equation)

Theorem (Darboux):

For any nondegenerate symplectic form (\cdot, \cdot) , there exists a basis $\{v_m\}$ such that, in this basis, (\cdot, \cdot) takes the matrix form:

$$\begin{pmatrix} 0 & 1 & & & & \\ -1 & 0 & & & & \\ & & 0 & 1 & & \\ & & -1 & 0 & & \\ 0 & & & & 0 & 1 \\ & & & & -1 & 0 \end{pmatrix}$$

i.e., such that $(v_{2m}, v_{2m+1}) = 1, (v_{2m+1}, v_{2m}) = -1$ and all other pairings vanish.

□ Thus, if we expand $a, b \in V$ as: $\overset{V}{\cup} \overset{\mathbb{R}}{\cup} \overset{V}{\cup} \quad a = a_m v_m, \quad b = b_m v_m$

Then: $(a, b) = \sum_{m=0}^{\infty} a_{2m} b_{2m+1} - a_{2m+1} b_{2m}$

Theorem (Darboux):

For any nondegenerate symplectic form $(,)$, there exists a basis $\{v_n\}$ such that, in this basis, $(,)$ takes the matrix form:

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□ Thus, if we expand $a, b \in V$ as: $\overset{V}{\cup} \overset{\mathbb{R}}{\cup} \overset{V}{\cup} \quad a = a_n v_n, \quad b = b_m v_m \quad \overset{\mathbb{R}}{\cup}$

$$\text{Then: } (a, b) = \sum_{n=0}^{\infty} a_{2n} b_{2n+1} - a_{2n+1} b_{2n}$$

$$\begin{pmatrix} 0 & 1 & 0 & \dots \\ -1 & 0 & 0 & \dots \\ 0 & 0 & 0 & \dots \\ \vdots & \vdots & \vdots & \ddots \end{pmatrix}$$

i.e., such that $(v_{2n}, v_{2n+1}) = 1$, $(v_{2n+1}, v_{2n}) = -1$
and all other pairings vanish.

□ Thus, if we expand $a, b \in V$ as: $a = \sum_{n=0}^{\infty} a_n v_n$, $b = \sum_{n=0}^{\infty} b_n v_n$

$$\text{Then: } (a, b) = \sum_{n=0}^{\infty} a_{2n} b_{2n+1} - a_{2n+1} b_{2n}$$

□ Now assume we picked such a basis $\{v_n\}$ in V .

□ Recall: $V =$ space of real-valued solutions of K.G. eqn.

□ What is a natural product \langle, \rangle on \bar{V} ?

□ On \bar{V} we define:

$$\langle f, h \rangle = i \int_{\Sigma} d\Sigma_{\nu} \sqrt{|g|} g^{\mu\nu} (j^{\mu} \partial_{\nu} h - (\partial_{\nu} j^{\mu}) h)$$

□ Then, $(,)$ yields:

$$\langle u_{\mu}, u_{\nu} \rangle = -\delta_{\mu\nu}, \quad \langle u_{\mu}^{\dagger}, u_{\nu}^{\dagger} \rangle = +\delta_{\mu\nu}, \quad \langle u_{\mu}, u_{\nu}^{\dagger} \rangle = 0 \quad (\text{I})$$

Exercise: verify this.

□ What is a natural product \langle, \rangle on \bar{V} ?

□ On \bar{V} we define:

$$\langle f, h \rangle = i \int_{\Sigma} d\Sigma_{\nu} \sqrt{g} g^{\mu\nu} (f^{\dagger} \partial_{\nu} h - (\partial_{\nu} f^{\dagger}) h)$$

□ Then, $(,)$ yields:

$$\langle u_m, u_m \rangle = -\delta_{m,m}, \quad \langle u_m^{\dagger}, u_m^{\dagger} \rangle = +\delta_{m,m}, \quad \langle u_m, u_m^{\dagger} \rangle = 0 \quad (\text{I})$$

Exercise: verify this.

$$\mathbb{1} = \sum_{\alpha} |u_{\alpha}\rangle\langle u_{\alpha}| + \sum_{\alpha} |u_{\alpha}^{\perp}\rangle\langle u_{\alpha}^{\perp}|$$

Remark: One can also turn \mathbb{V} into a Hilbert space, namely the Krein space. Let P and P^{\perp} be the projectors on the spaces spanned by the u_{α} and the u_{α}^{\perp} respectively. \mathbb{V} is a positive definite inner product, and the Krein space $(\mathbb{V}, \langle\langle \cdot, \cdot \rangle\rangle)$ is a Hilbert space.

Proof:

Included, $\mathbb{1}|u_{\alpha}\rangle = |u_{\alpha}\rangle \sum_{\beta} \langle u_{\beta}, u_{\alpha}\rangle = |u_{\alpha}\rangle$

$$\mathbb{1}|u_{\alpha}^{\perp}\rangle = |u_{\alpha}^{\perp}\rangle \sum_{\beta} \langle u_{\beta}^{\perp}, u_{\alpha}^{\perp}\rangle = |u_{\alpha}^{\perp}\rangle$$

so that for any $|f\rangle \in \mathbb{V}$ we have:

$$= \sum_{\alpha} |u_{\alpha}\rangle\langle u_{\alpha}|f\rangle + \sum_{\alpha} |u_{\alpha}^{\perp}\rangle\langle u_{\alpha}^{\perp}|f\rangle = |f\rangle$$

$$(f, h) := \int_{\Sigma} d\Sigma_{\nu} \sqrt{|g|} g^{\nu\sigma} (f \partial_{\sigma} h - h \partial_{\sigma} f)$$

← any spacelike hypersurface
i.e. set of points of equal time.

□ Proposition: (f, h) is independent of choice of Σ .

Proof: Later (uses Stokes' theorem and K.G. equation)

□ (f, h) is a symplectic form, i.e.: $(f, h) = - (h, f)$.

↑
easy to see

□ What can we do with $(,)$? No diagonalization?

Theorem (Darboux):

Proof:

Indeed, $\mathbb{1}|u_n\rangle = |u_n\rangle$ using $\langle u_n, u_n \rangle = -1$

$\mathbb{1}|u_n^*\rangle = |u_n^*\rangle$ using $\langle u_n^*, u_n^* \rangle = 1$.

so that for any $|f\rangle \in \bar{V}$ we have:

$$-\sum_n |u_n\rangle \langle u_n|f\rangle + |u_n^*\rangle \langle u_n^*|f\rangle = |f\rangle \quad (P)$$

Writing this out, we will now show that it yields (W), i. e.:

$$\sqrt{|g|} g^{00} \sum_k \left(u_k(x,t) \frac{\partial}{\partial x'^0} u_k^*(x',t) - u_k^*(x,t) \frac{\partial}{\partial x'^0} u_k(x',t) \right) = i \delta^3(x-x')$$

Proof:Indeed, $\mathbb{1}|u_n\rangle = |u_n\rangle$ using $\langle u_n, u_n \rangle = -1$ $\mathbb{1}|u_n^*\rangle = |u_n^*\rangle$ using $\langle u_n^*, u_n^* \rangle = 1$.so that for any $|f\rangle \in \bar{V}$ we have:

$$-\sum_n |u_n\rangle \langle u_n | f \rangle + |u_n^*\rangle \langle u_n^* | f \rangle = |f\rangle \quad (P)$$

Writing this out, we will now show that it yields (W), i.e.:

$$\mathbb{1} \int g^{\mu\nu} \sum_n \left(u_{\mu}(x) \frac{\partial}{\partial x^{\nu}} u_{\nu}^*(x) - u_{\nu}^*(x) \frac{\partial}{\partial x^{\nu}} u_{\mu}(x) \right) = i \delta^3(x-x')$$

reads:

$$\sum_n u_n(x,t) i \int_{\Sigma} d^3x' \sqrt{g} g^{00} (u_n^* \partial_{x'} f - (\partial_{x'} u_n^*) f)$$

$$- \sum_n u_n^*(x,t) i \int_{\Sigma} d^3x' \sqrt{g} g^{00} (u_n \partial_{x'} f - (\partial_{x'} u_n) f) = f(x,t)$$

Now, interchanging Σ and \int_{Σ} yields $\forall f \in \bar{V}$:

$$i \int_{\Sigma} d^3x \sqrt{|g(x)|} g^{00} \sum_n (u_n(x,t) u_n^*(x',t) \partial_{x'} - u_n(x,t) \partial_{x'} u_n^*(x',t))$$

□ Now choose an arbitrary function $g(x)$.

□ Then there exists a solution $f(x', t)$ of the Klein Gordon equation obeying:

$$1) \quad f(x', t_0) = g(x')$$

$$2) \quad g''(x') \partial_{x^0} f(x', t_0) = 0$$

(Because the 2nd order K.G. equation on a globally hyperbolic spacetime has a well-defined Cauchy problem)

□ Therefore, (*) yields, for all choices of $g(x)$:

$$i \int_{\Sigma} d^3x' \sqrt{|g_{3D}|} g''(x') \sum_{\alpha} \left(-u_{\alpha}(x', t) \partial_{x^{\alpha}} u^{\alpha}(x', t) + u_{\alpha}^{\alpha}(x', t) \partial_{x^{\alpha}} u_{\alpha}(x', t) \right) g(x') = g(x) \quad \forall g(x)$$

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Conclusion so far:

1.) Commutation relations:

$$[\hat{\phi}(x,t), \hat{\pi}(x',t)] = i\hbar \delta^3(x-x')$$

$$[\hat{\phi}(x,t), \hat{\phi}(x',t)] = 0, \quad [\hat{\pi}(x,t), \hat{\pi}(x',t)] = 0$$

2.) Hermiticity: $\hat{\phi}^\dagger(x,t) = \hat{\phi}(x,t), \quad \hat{\pi}^\dagger(x,t) = \hat{\pi}(x,t)$

3.) Equations of motion:

$$\left(\frac{1}{\sqrt{|g|}} \partial_{x^\mu} g^{\mu\nu} \sqrt{|g|} \partial_{x^\nu} + m^2 \right) \hat{\phi}(x,t) = 0 \quad (KG)$$

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on any globally hyperbolic spacetime:

$$\hat{\phi}(x,t) := \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^+$$

$$\hat{\pi}(x,t) := \sqrt{|g|} g^{0\nu} \frac{\partial}{\partial x^\nu} \hat{\phi}(x,t)$$

where $[a_k, a_{k'}^+] = \delta(k-k')$ and where the $u_k(x,t)$ are number-valued solutions to (KG) which also obey (R1):

$$\sqrt{|g|} g^{0\nu} \sum_k \left(u_k(x,t) \frac{\partial}{\partial x^\nu} u_k^*(x',t) - u_k^*(x,t) \frac{\partial}{\partial x^\nu} u_k(x',t) \right) = i \delta^3(x-x')$$

Q: We showed that (W) ensures the CCRs at one time.

What guarantees conservation of the CCRs?

A: Stokes' theorem and unitarity

Q: Are the u, v unique?

A: No! Math: Bogolubov transformations

↑

Q: What's the physics?

A: Vacuum ambiguity.