Title: Cosmology â€“ Observations: Gravitational lensing

Date: Jul 15, 2015 04:00 PM

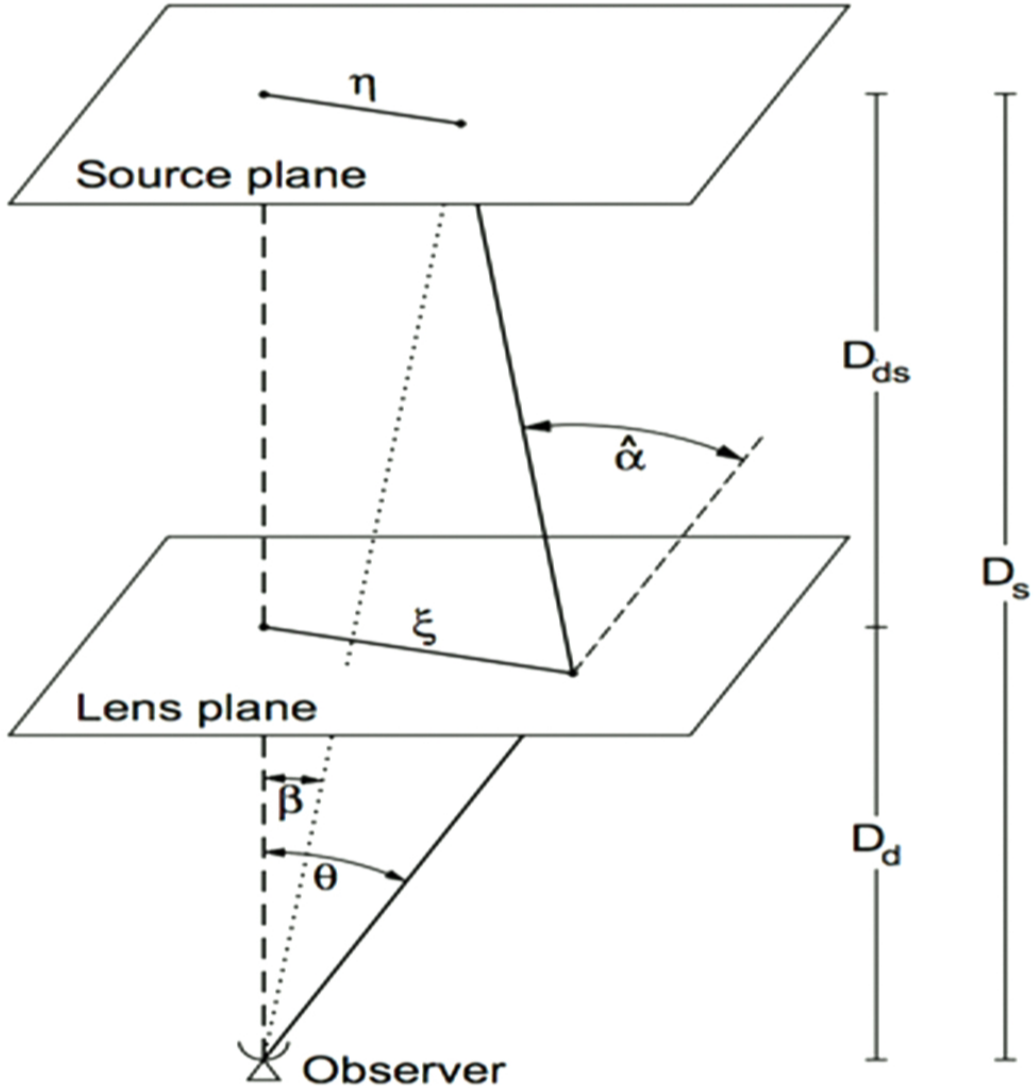
URL: http://pirsa.org/15070051

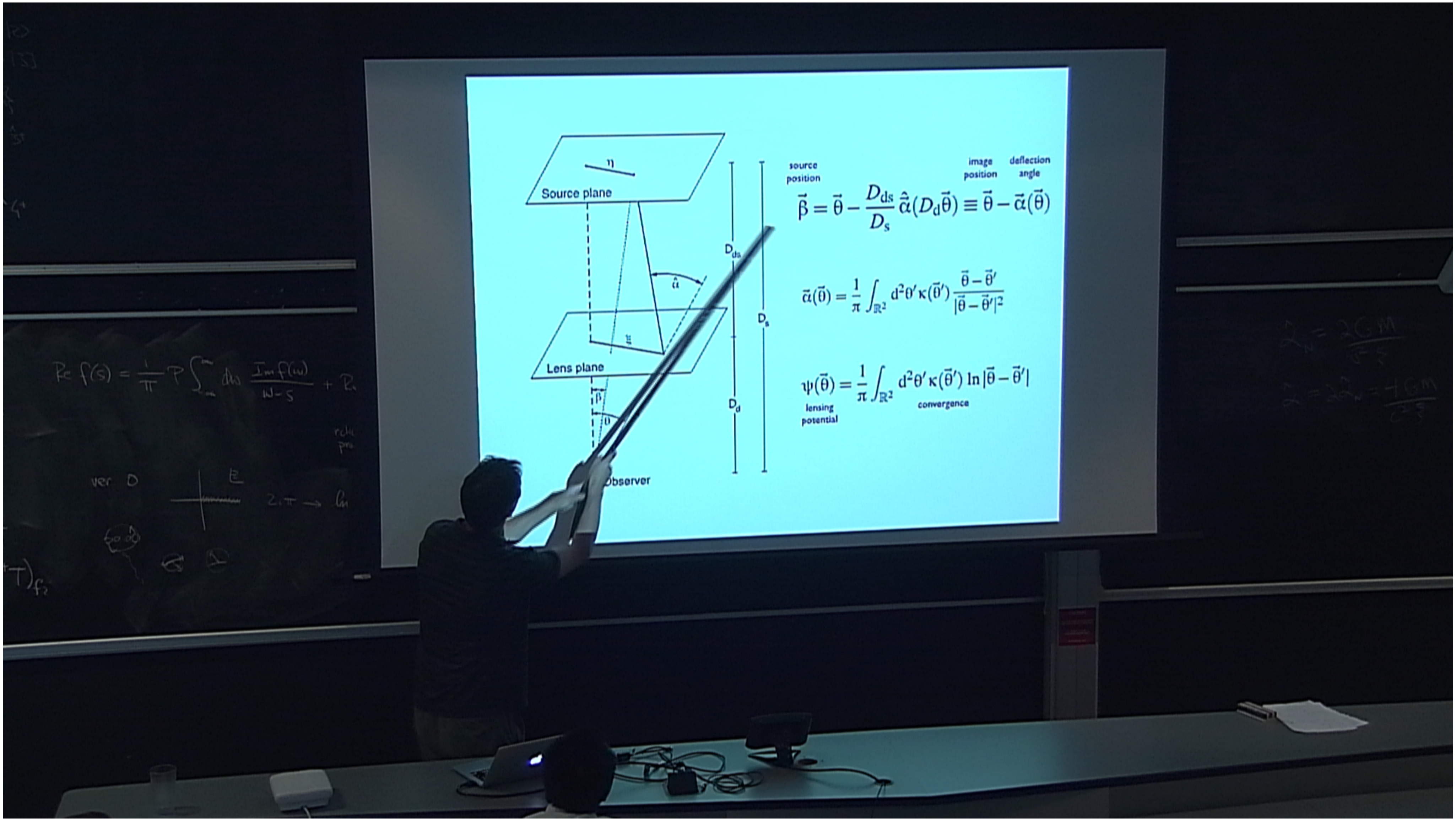
Abstract:

Gravitational lensing

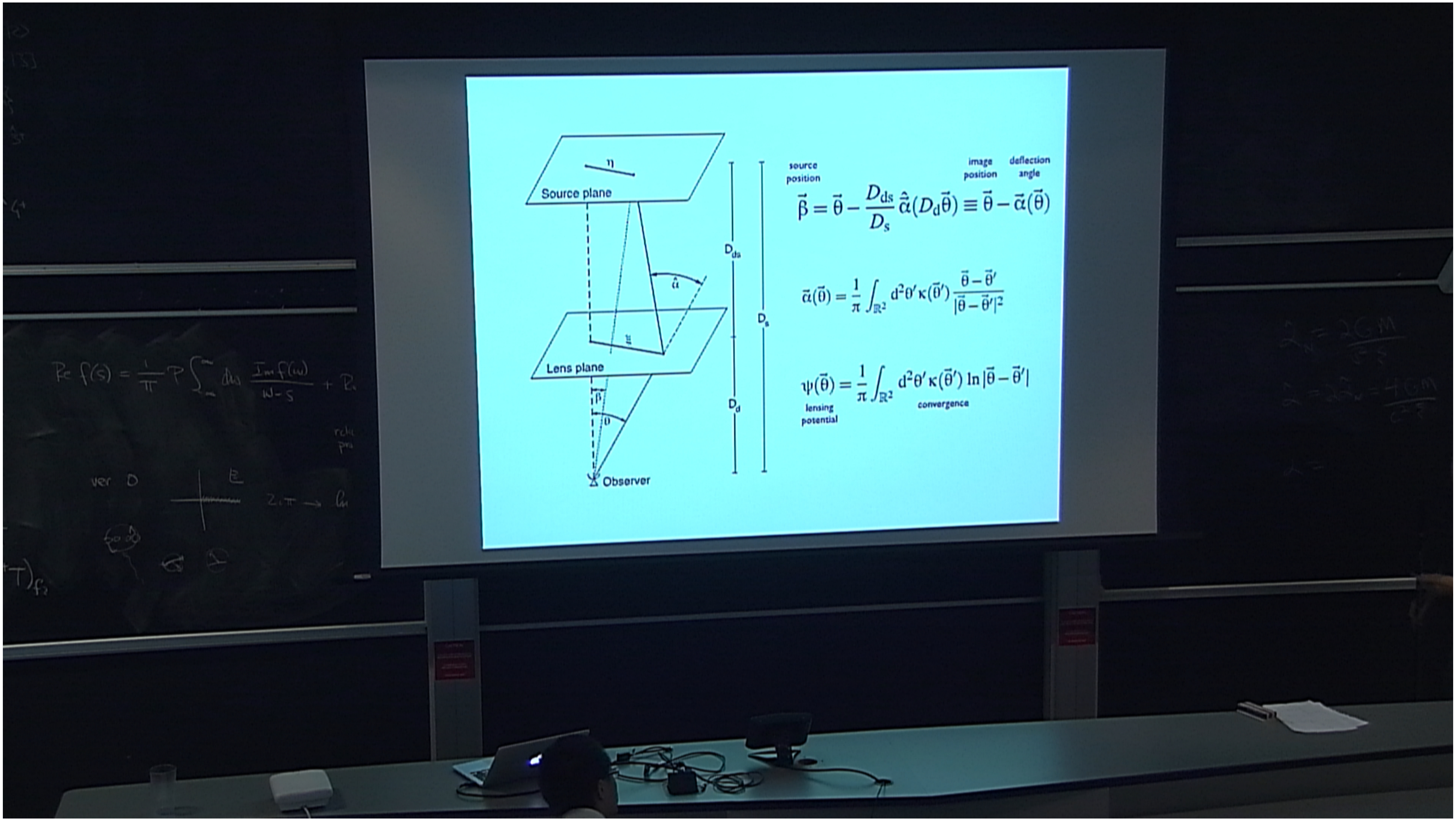
$$\hat{\alpha} = \frac{4GM}{c^2 \xi}$$

Bartelmann & Schneider
http://arxiv.org/abs/astro-ph/9912508





source position
image position
deflection angle
$\vec{\beta} = \vec{\theta} - \frac{D_{ds}}{D_s}\hat{\vec{\alpha}}(D_d\vec{\theta}) \equiv \vec{\theta} - \vec{\alpha}(\vec{\theta})$
$\vec{\alpha}(\vec{\theta}) = \frac{1}{\pi}\int_{\mathbb{R}^2} d^2\theta' \kappa(\vec{\theta}') \frac{\vec{\theta}-\vec{\theta}'}{|\vec{\theta}-\vec{\theta}'|^2}$
$\psi(\vec{\theta}) = \frac{1}{\pi}\int_{\mathbb{R}^2} d^2\theta' \kappa(\vec{\theta}') \ln|\vec{\theta}-\vec{\theta}'|$
lensing potential
convergence
Source plane
Lens plane
Observer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Lens plane
Observer
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D_s
D_d
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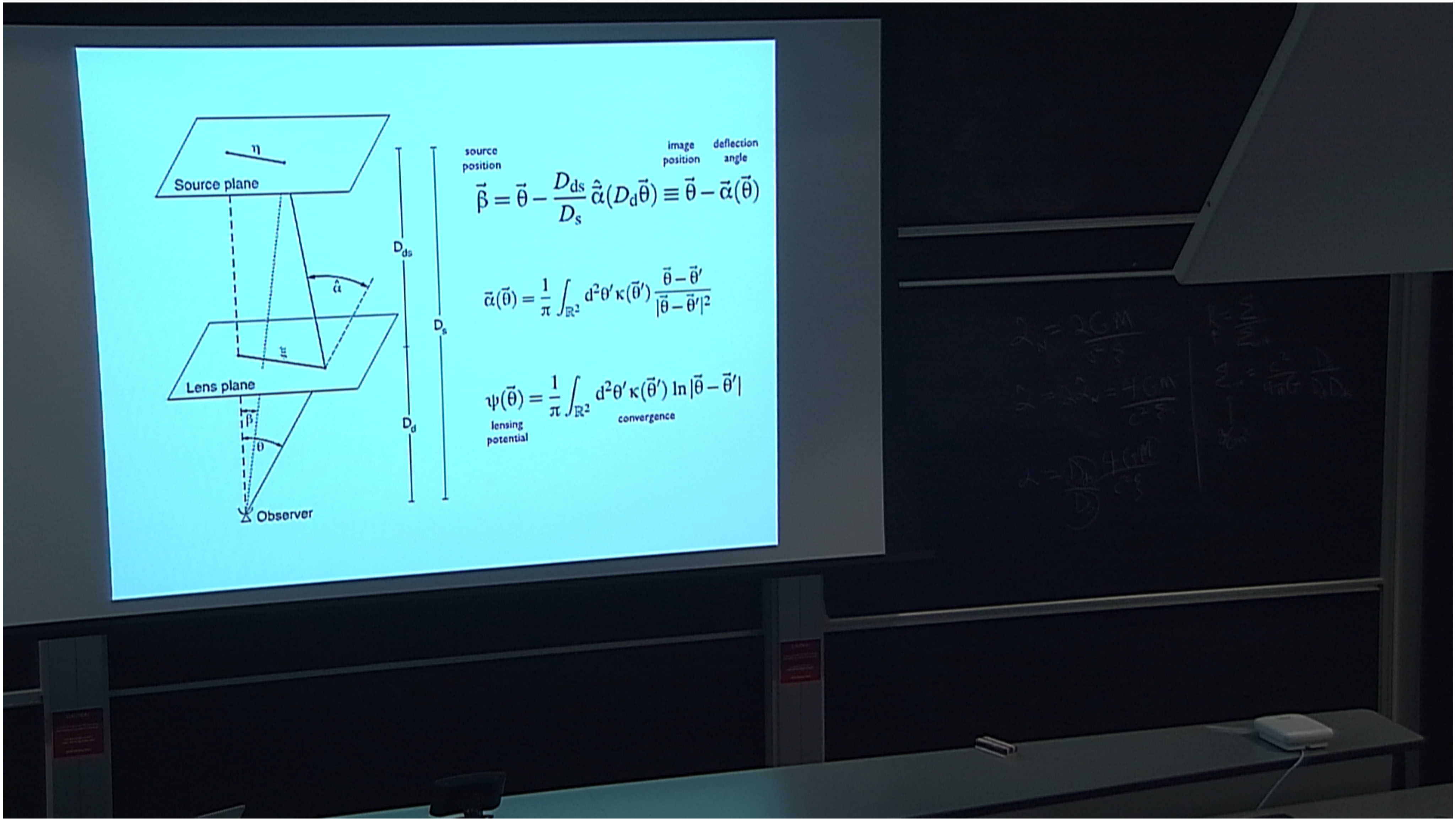
η
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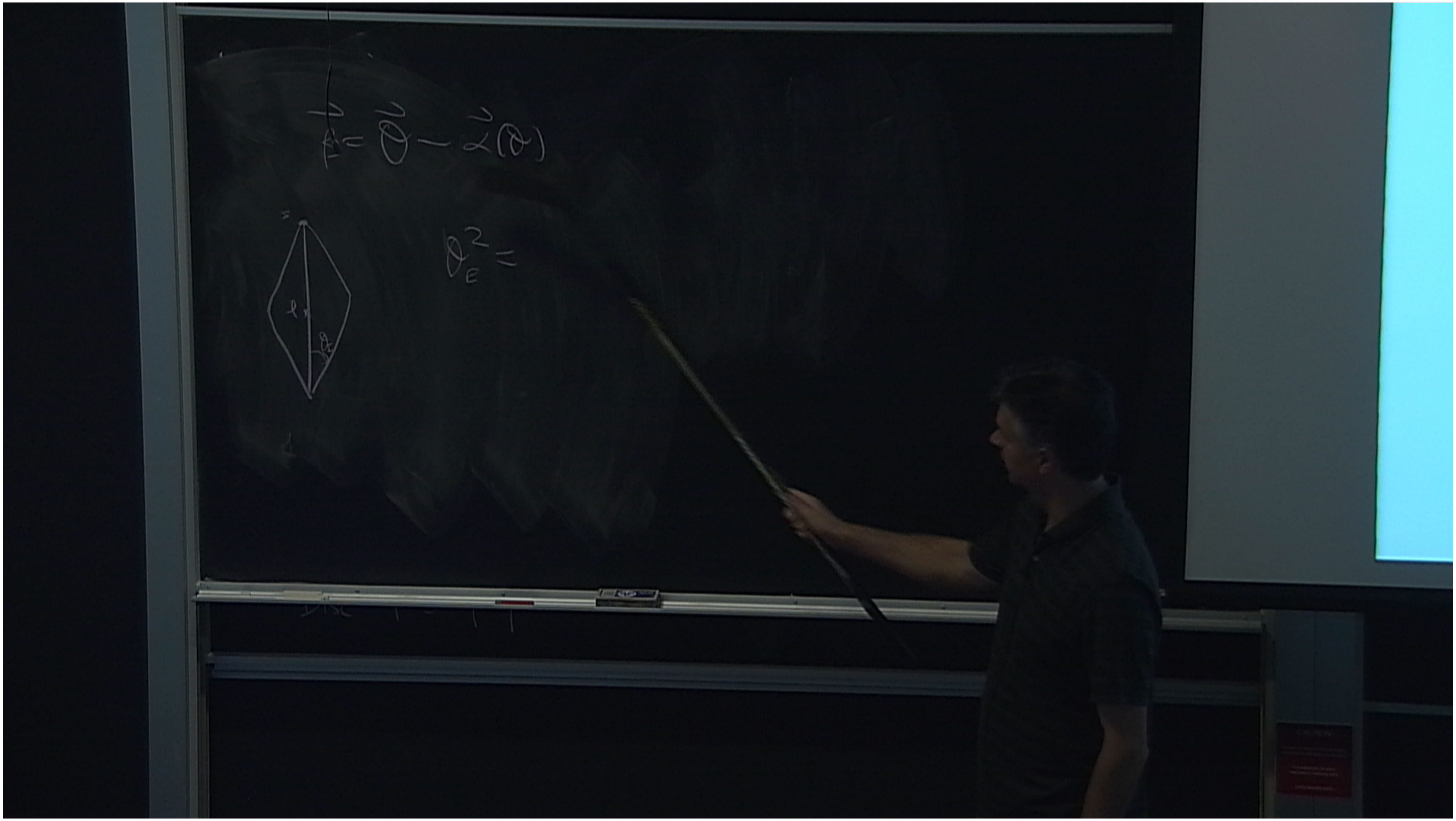
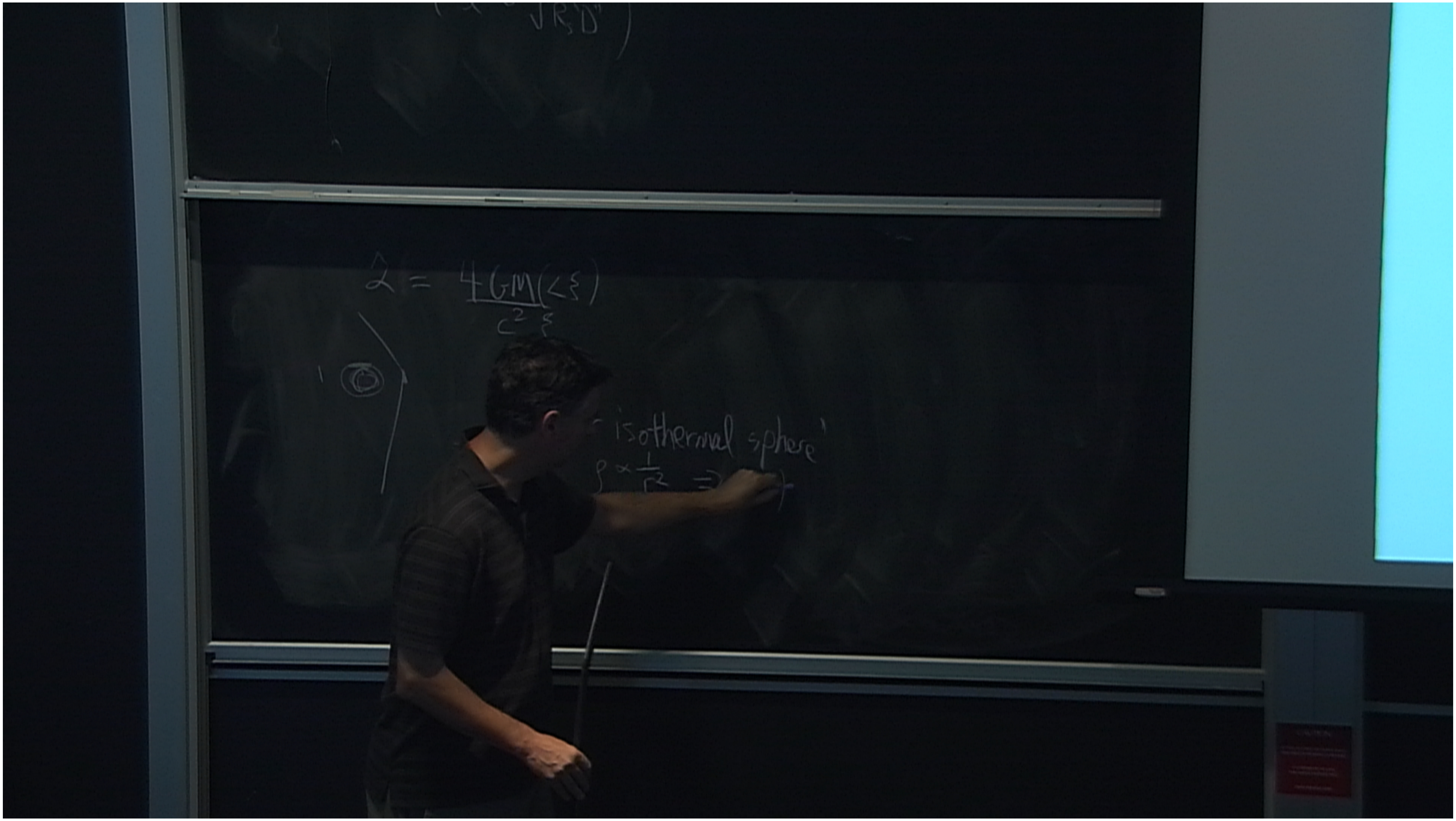
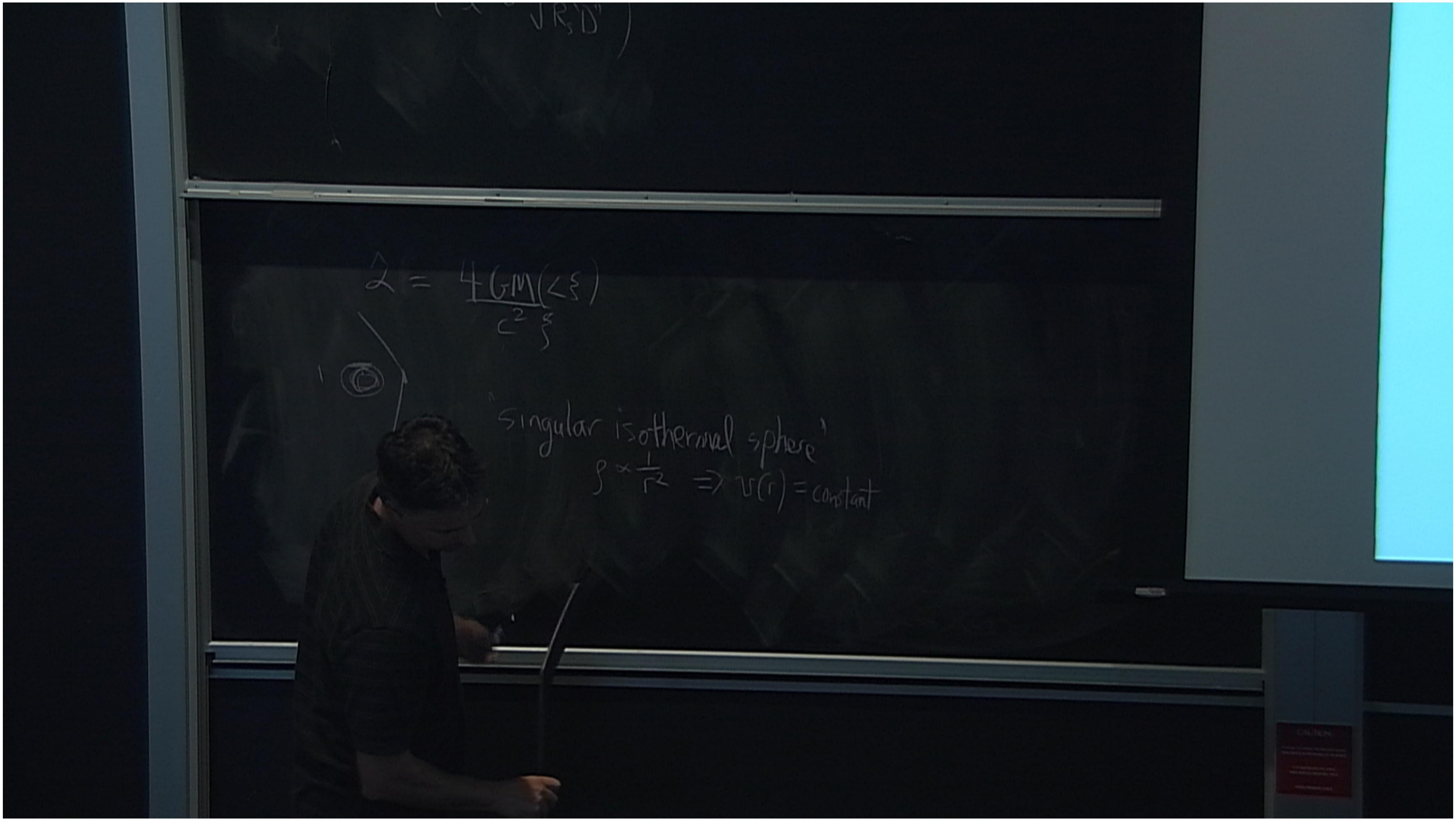
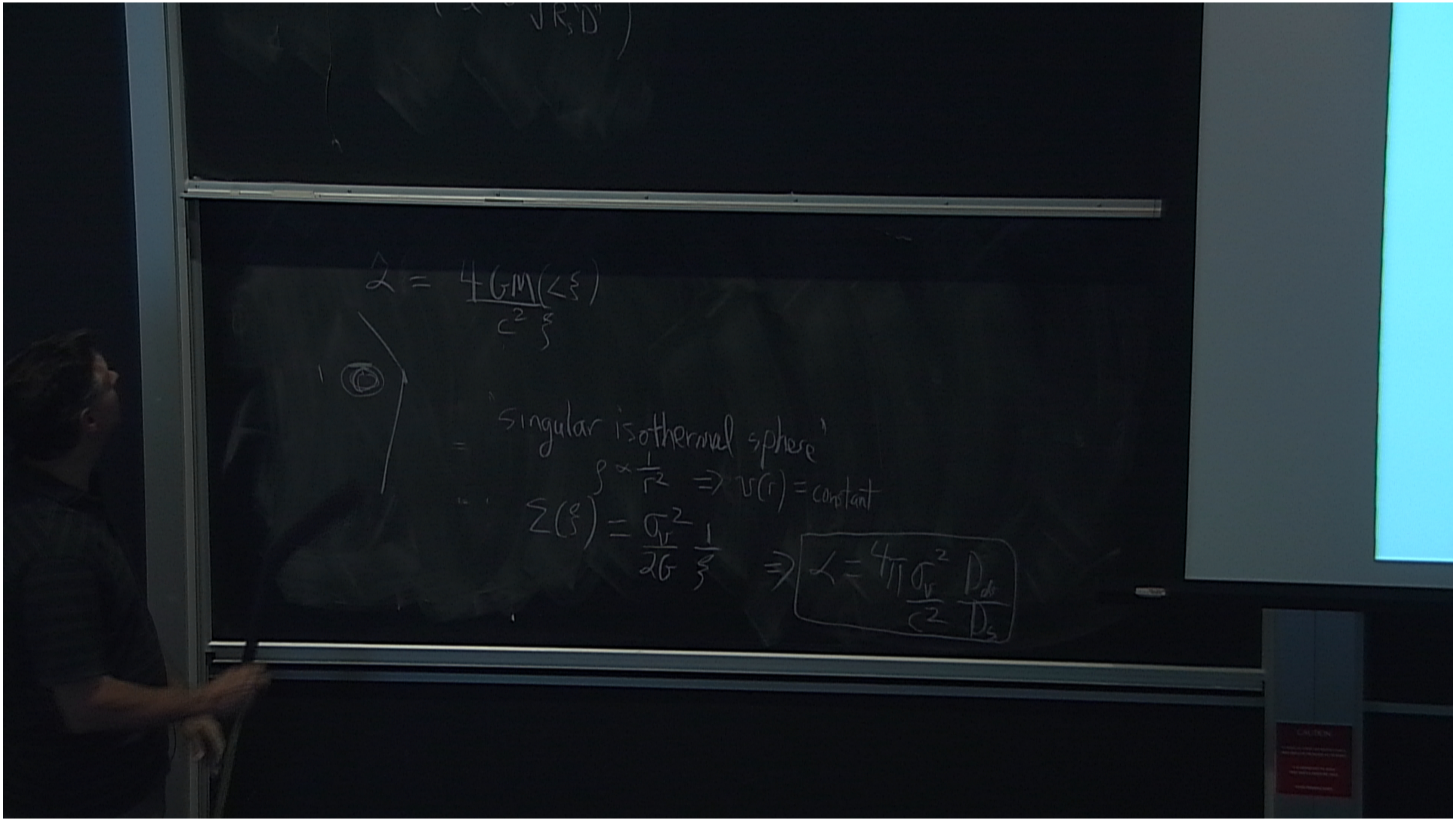

β = θ − α(θ)
θE2 =


isothermal sphere


α = 4GM(<ξ) / c² ξ
"singular isothermal sphere"
ρ ∝ 1/r² ⇒ v(r) = constant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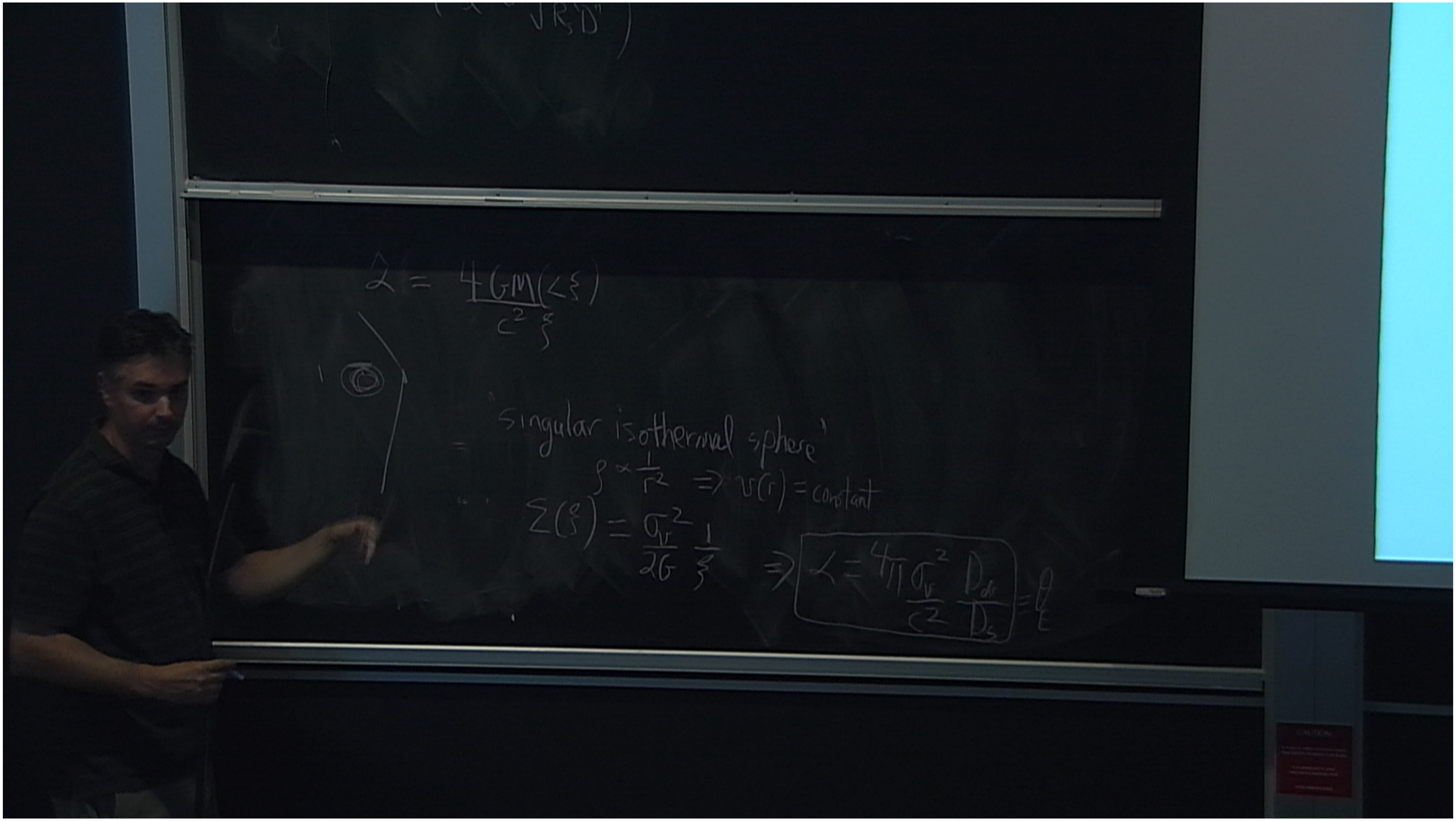
⇒ α = 4π σ_v²/c² D_ds/D_s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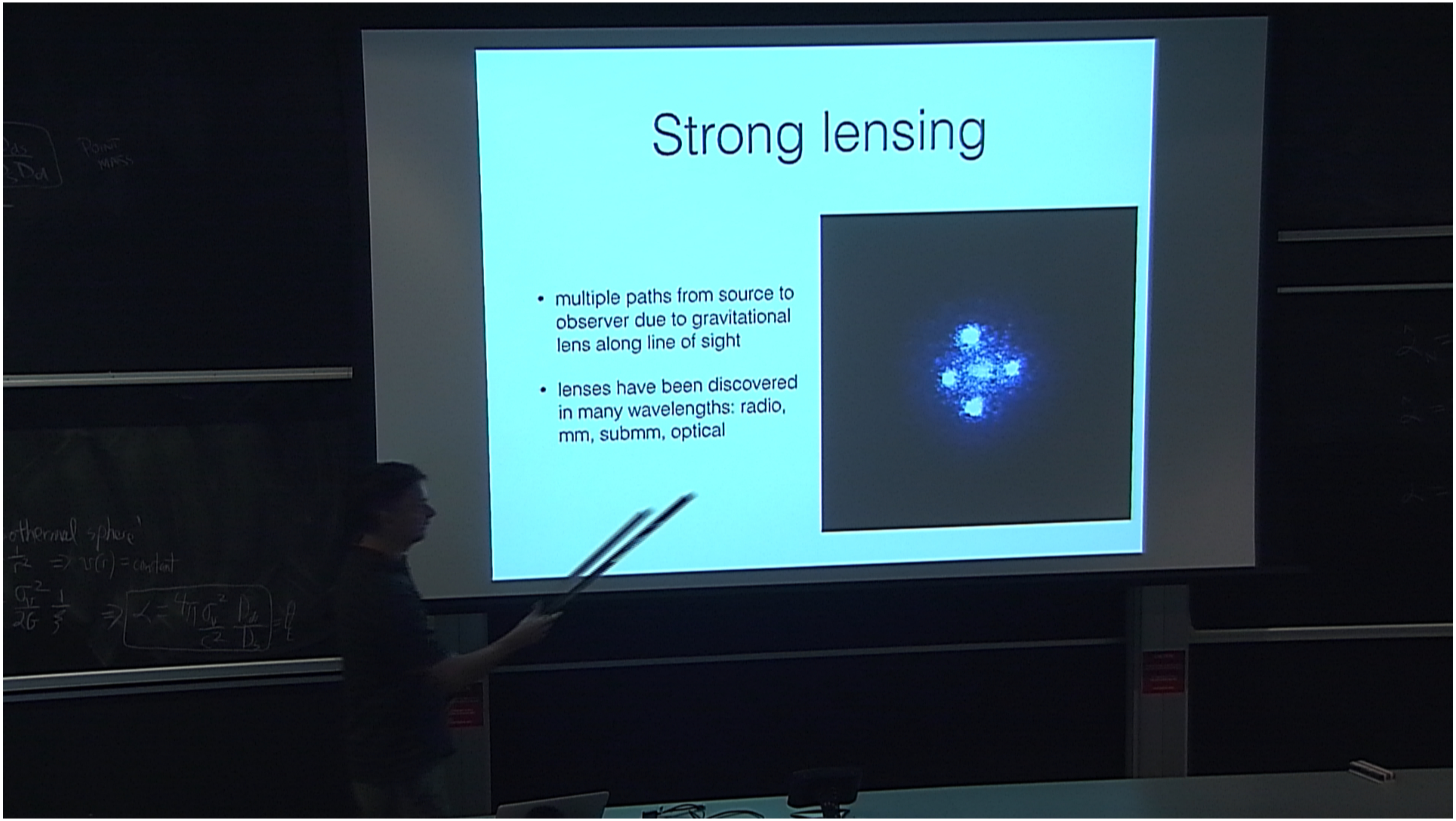
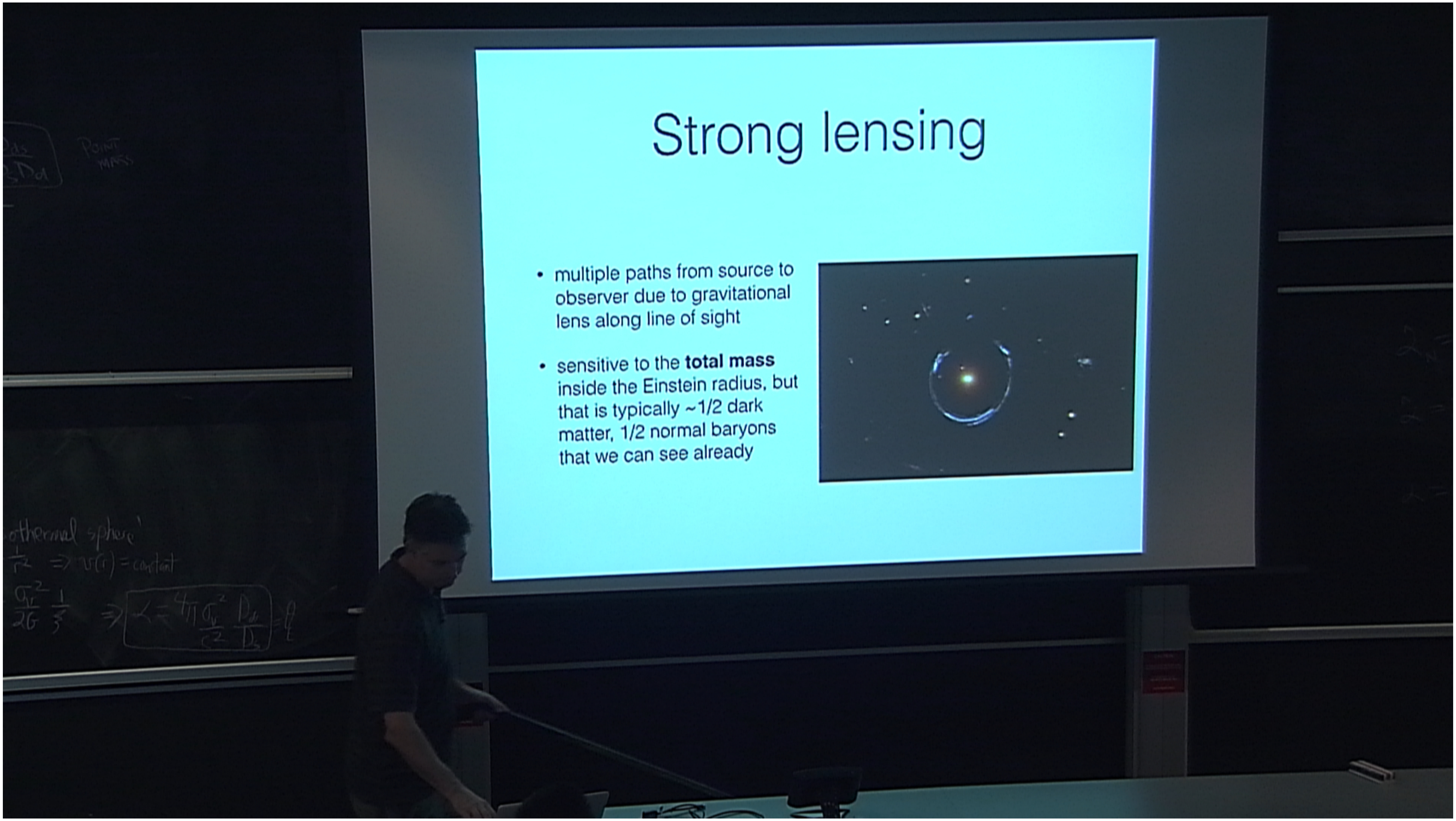
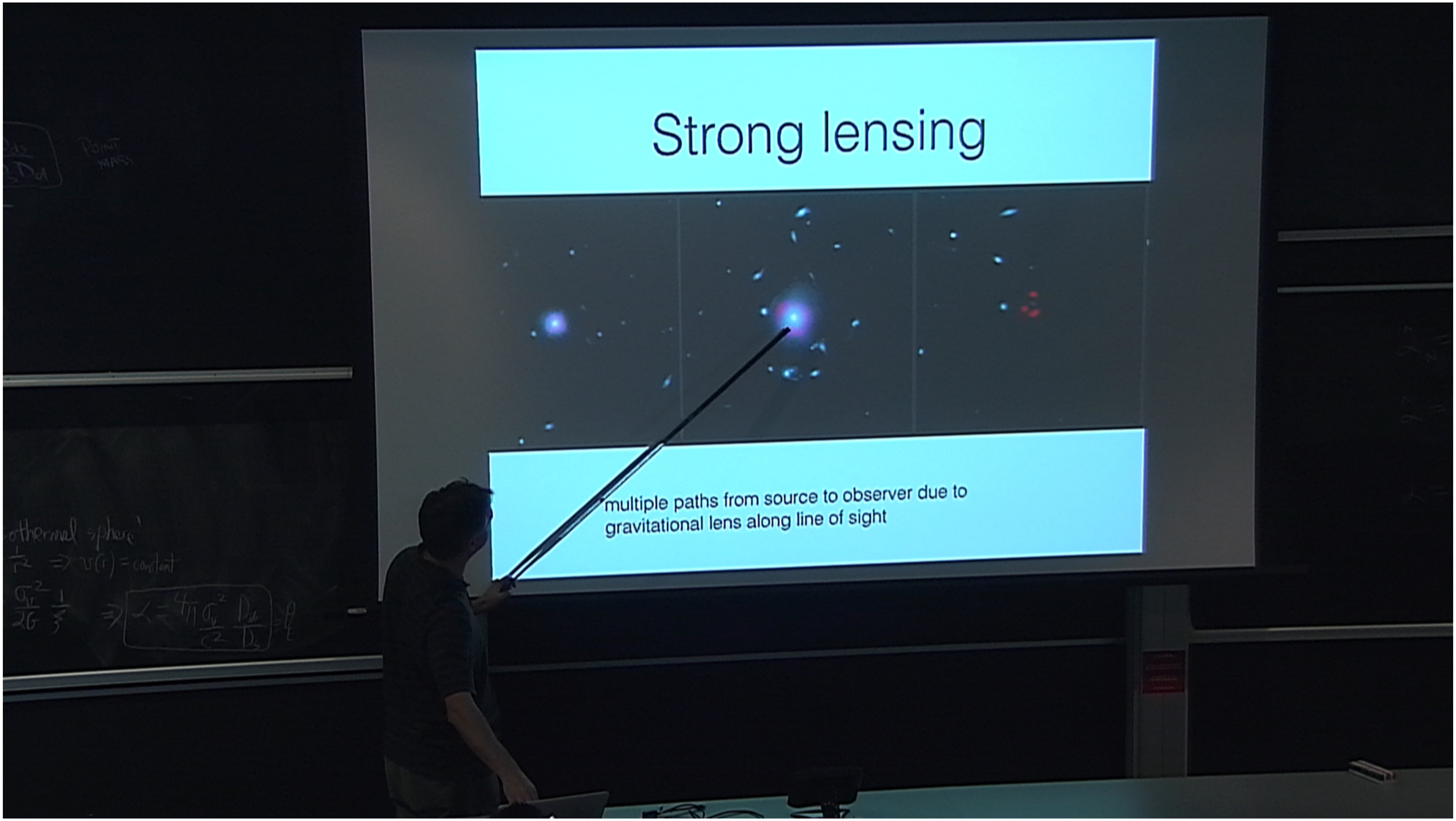

Strong lensing
• multiple paths from source to observer due to gravitational lens along line of sight
• lenses have been discovered in many wavelengths: radio, mm, submm, optical


Strong lensing
multiple paths from source to observer due to gravitational lens along line of sight
sensitive to the **total mass** inside the Einstein radius, but that is typically ~1/2 dark matter, 1/2 normal baryons that we can see already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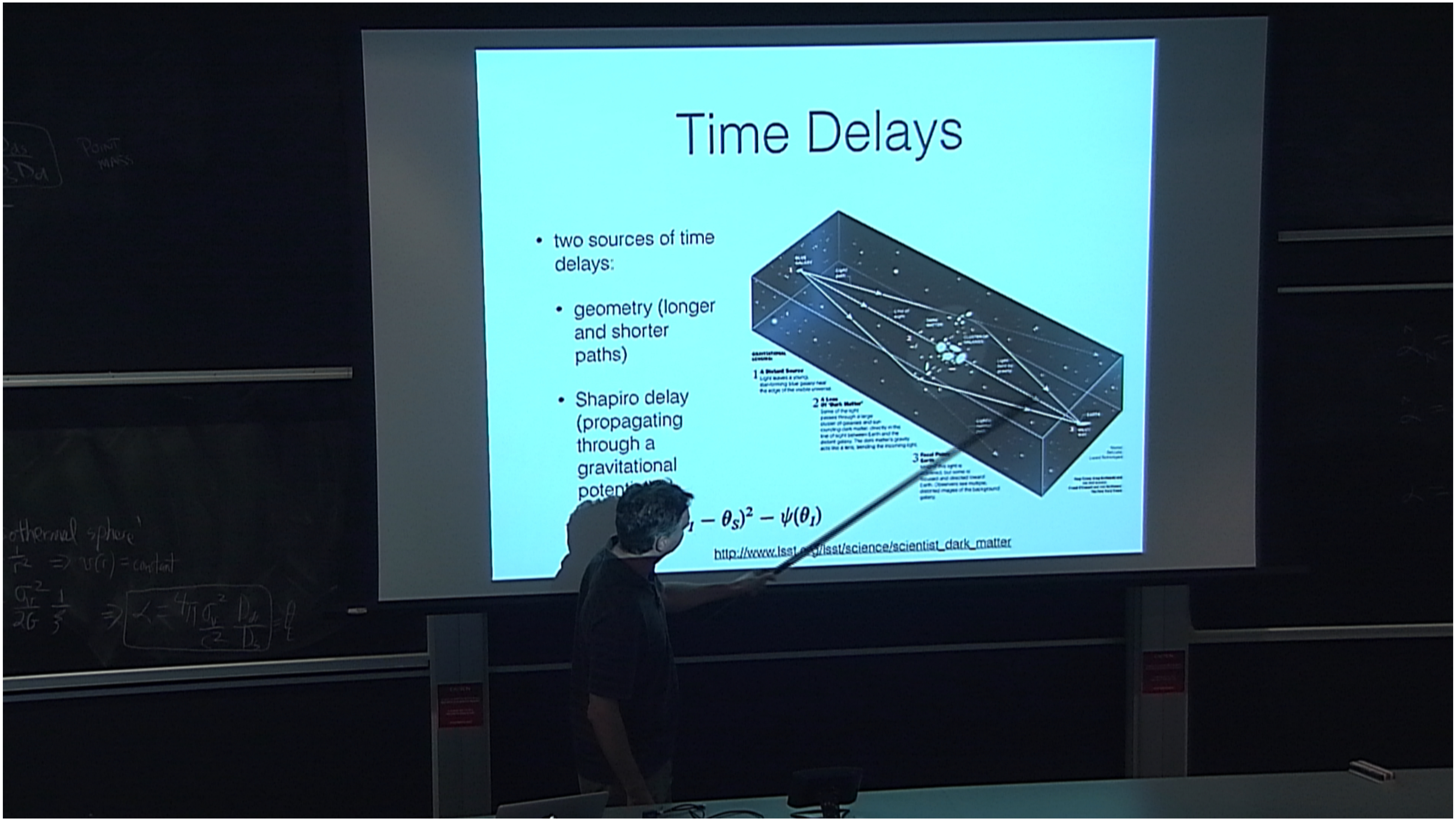
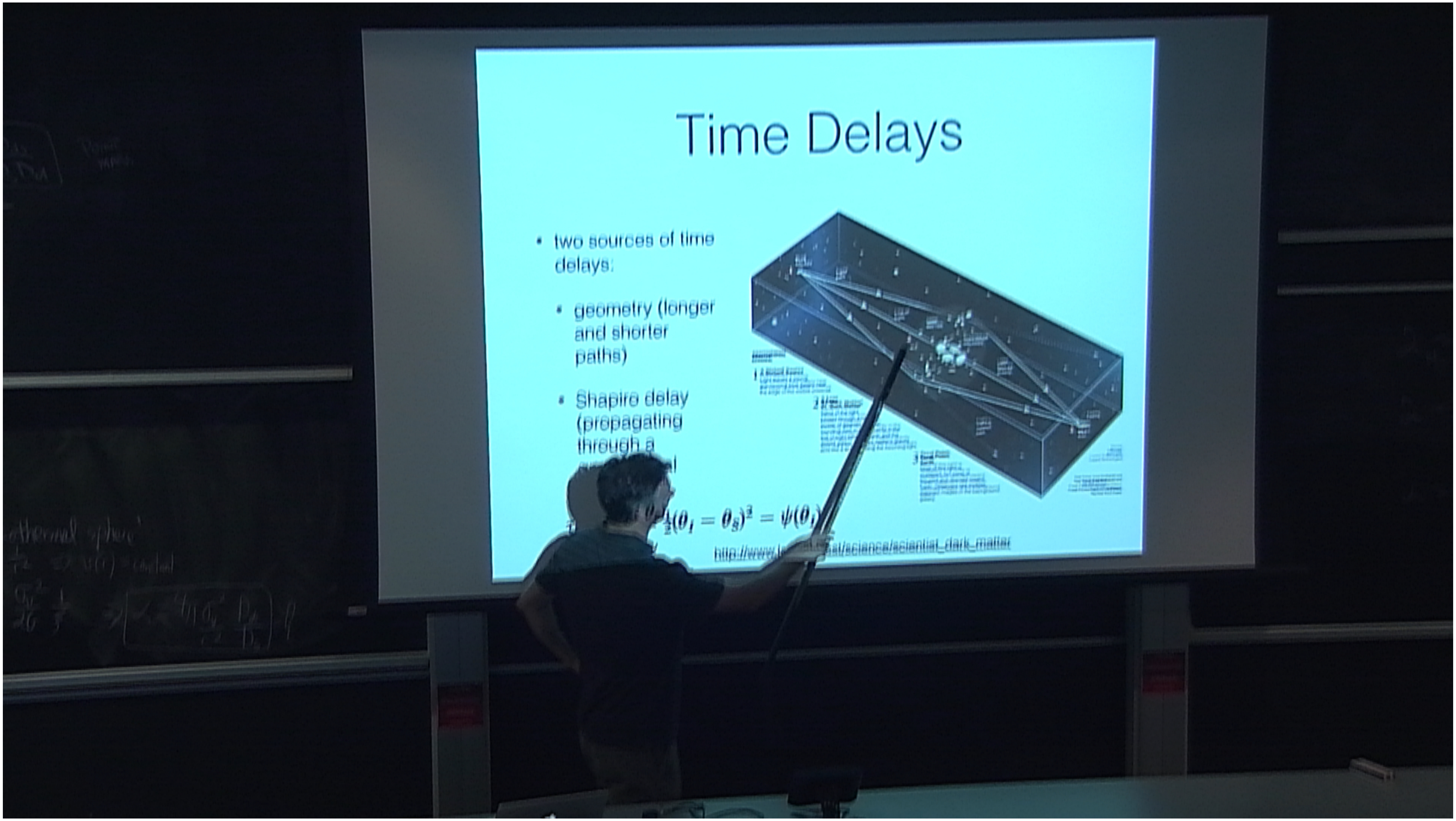

Time Delays
• two sources of time delays:
• geometry (longer and shorter paths)
• Shapiro delay (propagating through a gravitational poten
http://www.lsst.org/lsst/science/scientist_dark_matter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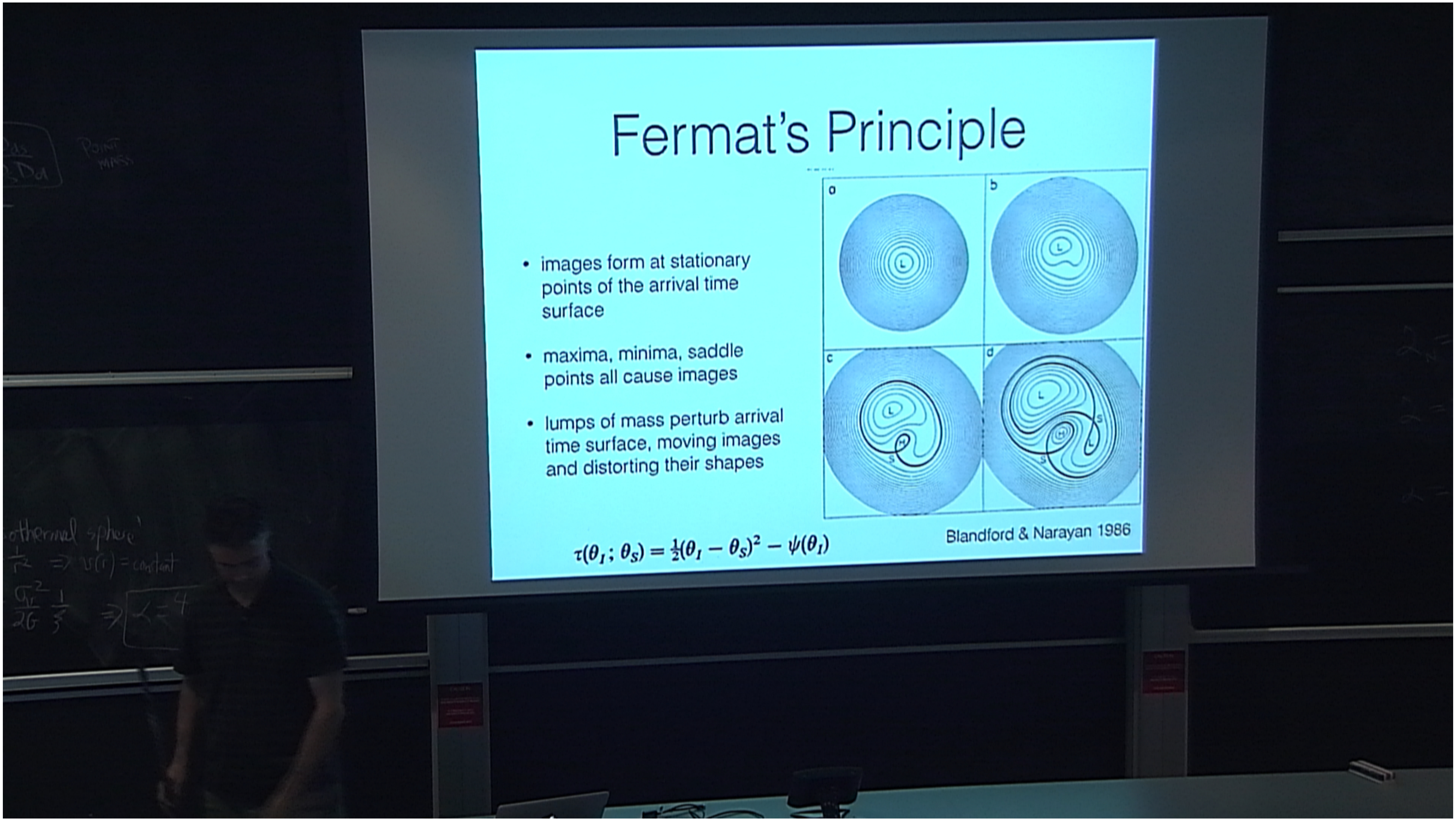
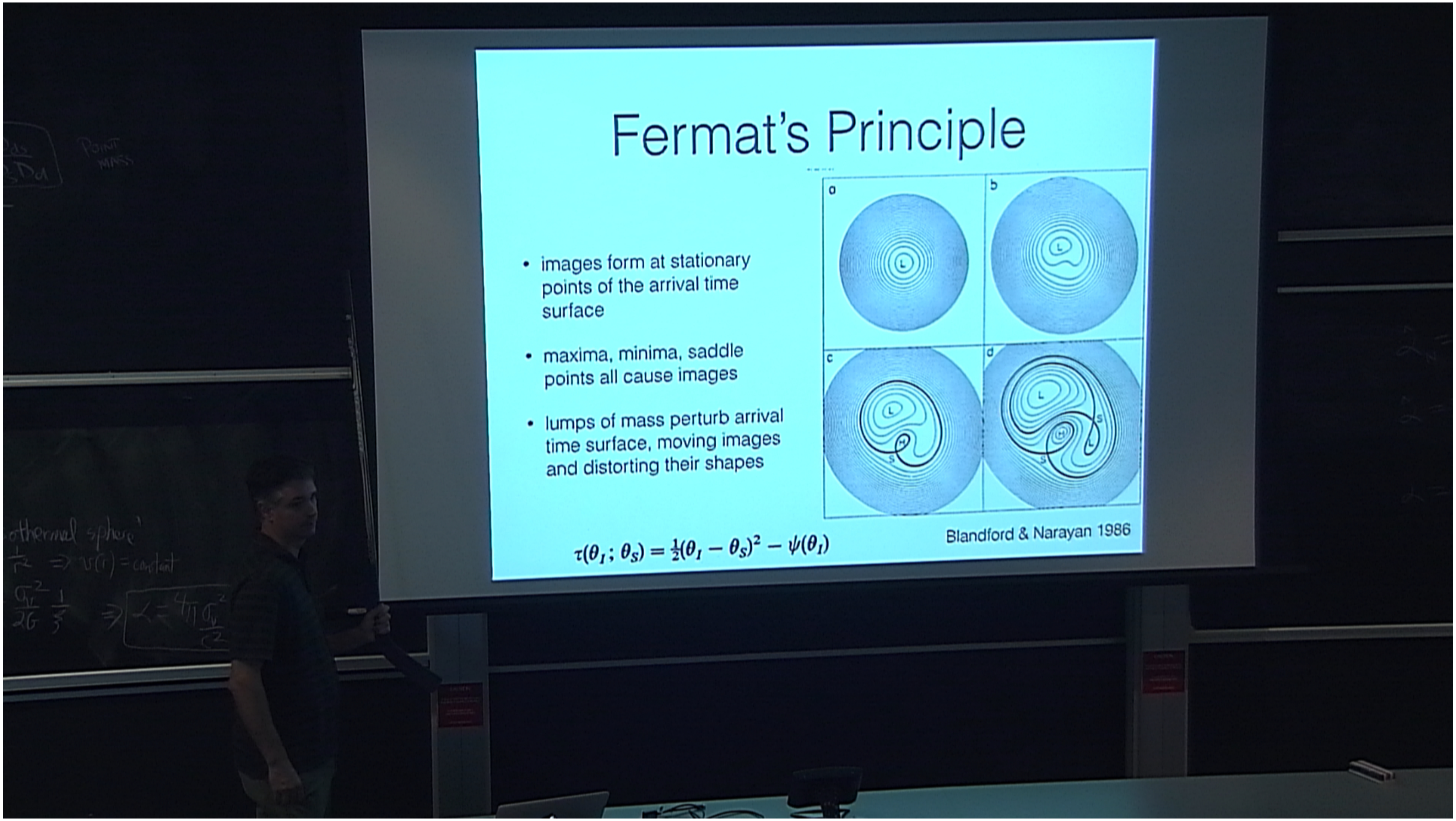

Fermat's Principle
• images form at stationary points of the arrival time surface
• maxima, minima, saddle points all cause images
• lumps of mass perturb arrival time surface, moving images and distorting their shapes
a
b
c
d
Blandford & Narayan 1986
$\tau(\theta_I;\theta_S) = \frac{1}{2}(\theta_I - \theta_S)^2 - \psi(\theta_I)$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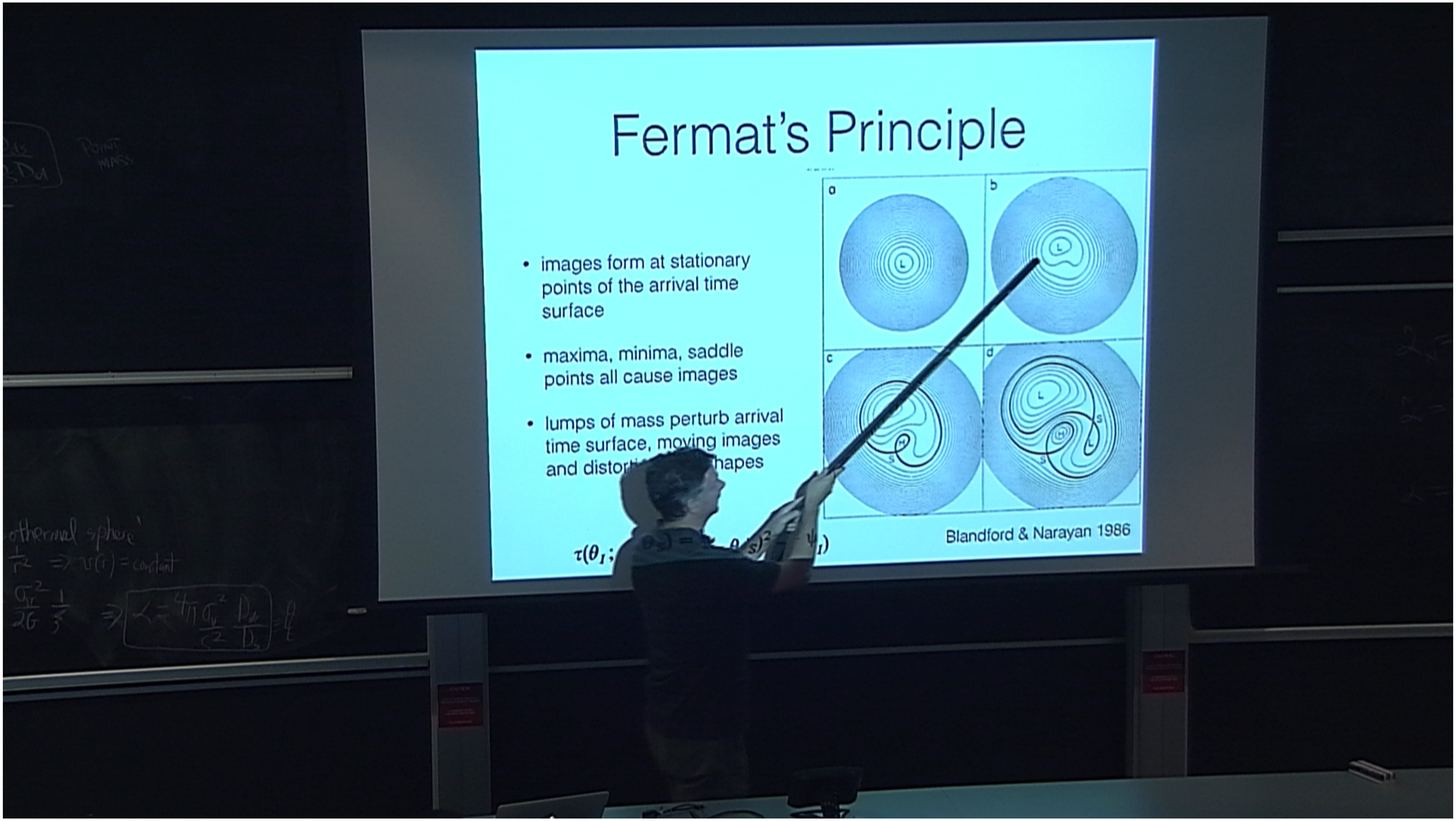
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lumps of mass perturb arrival time surface, moving images
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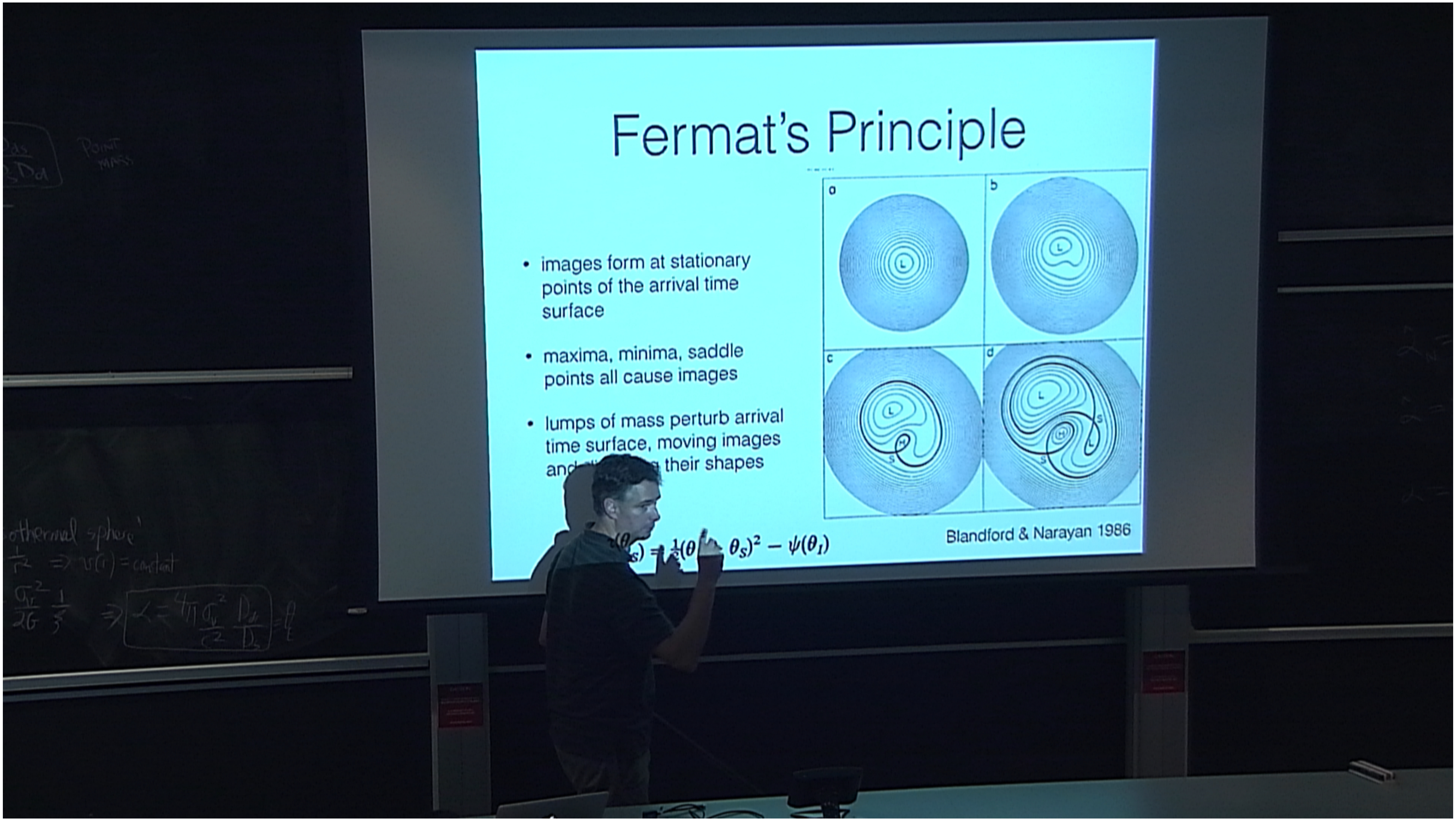
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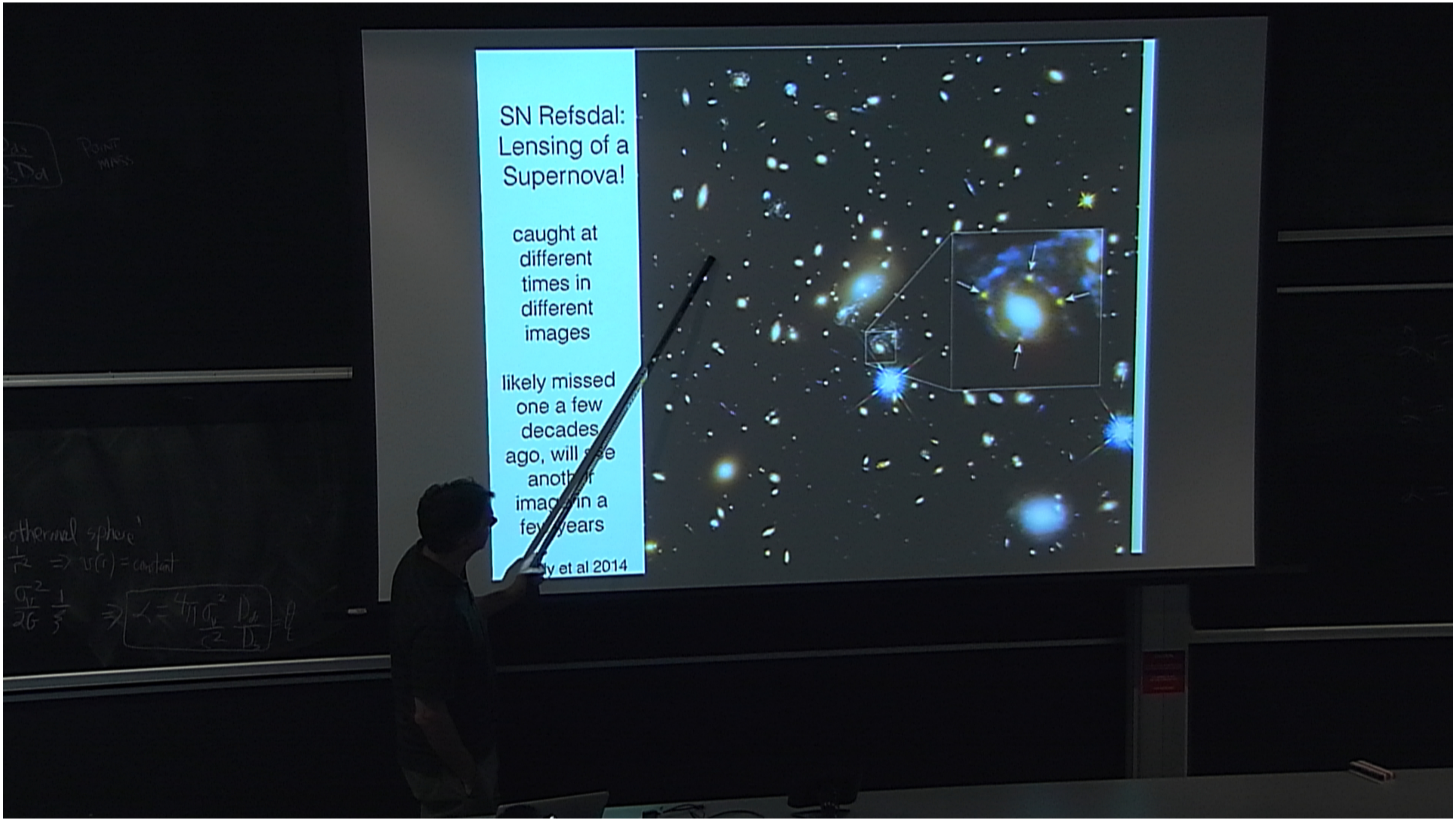
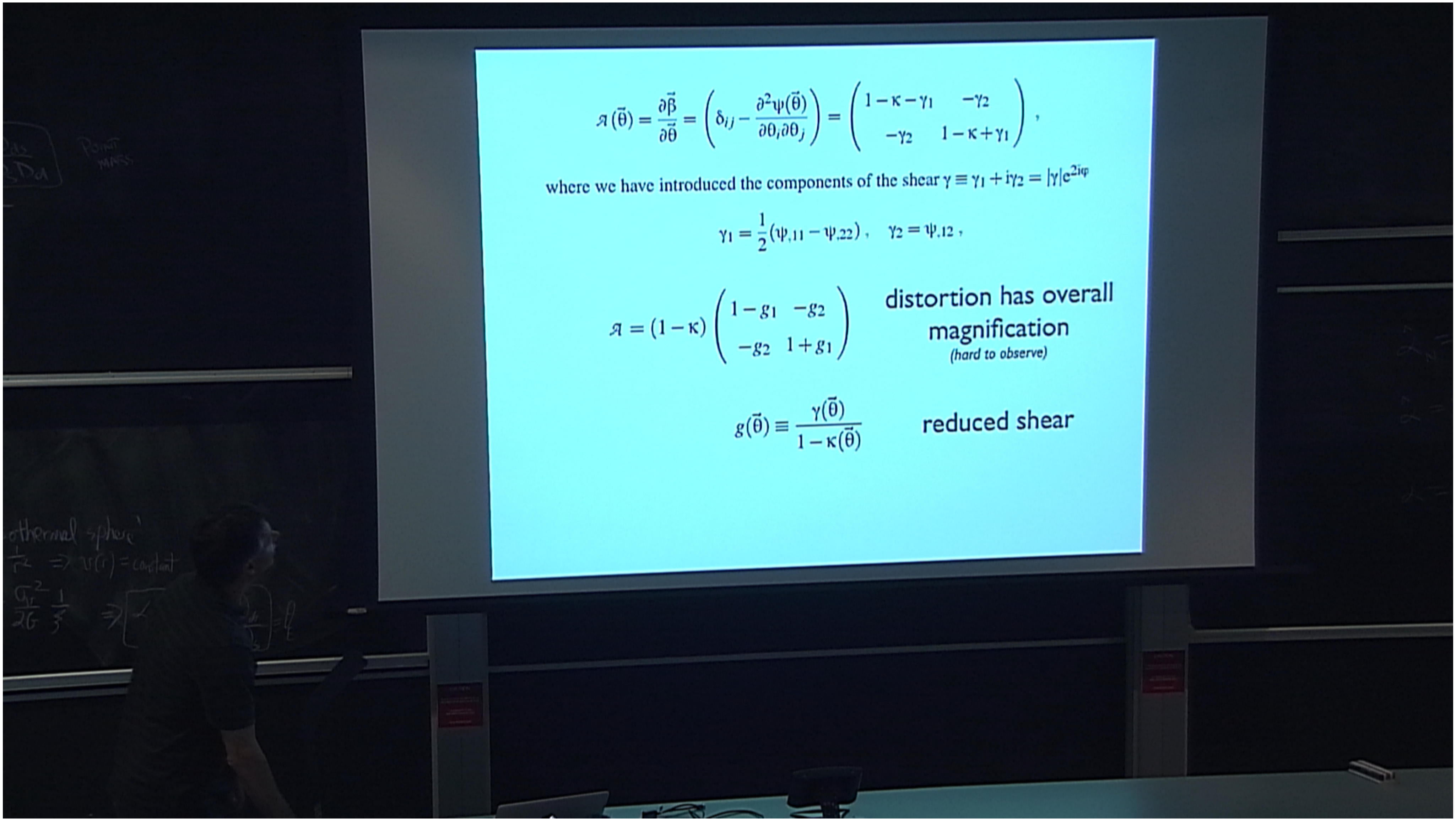
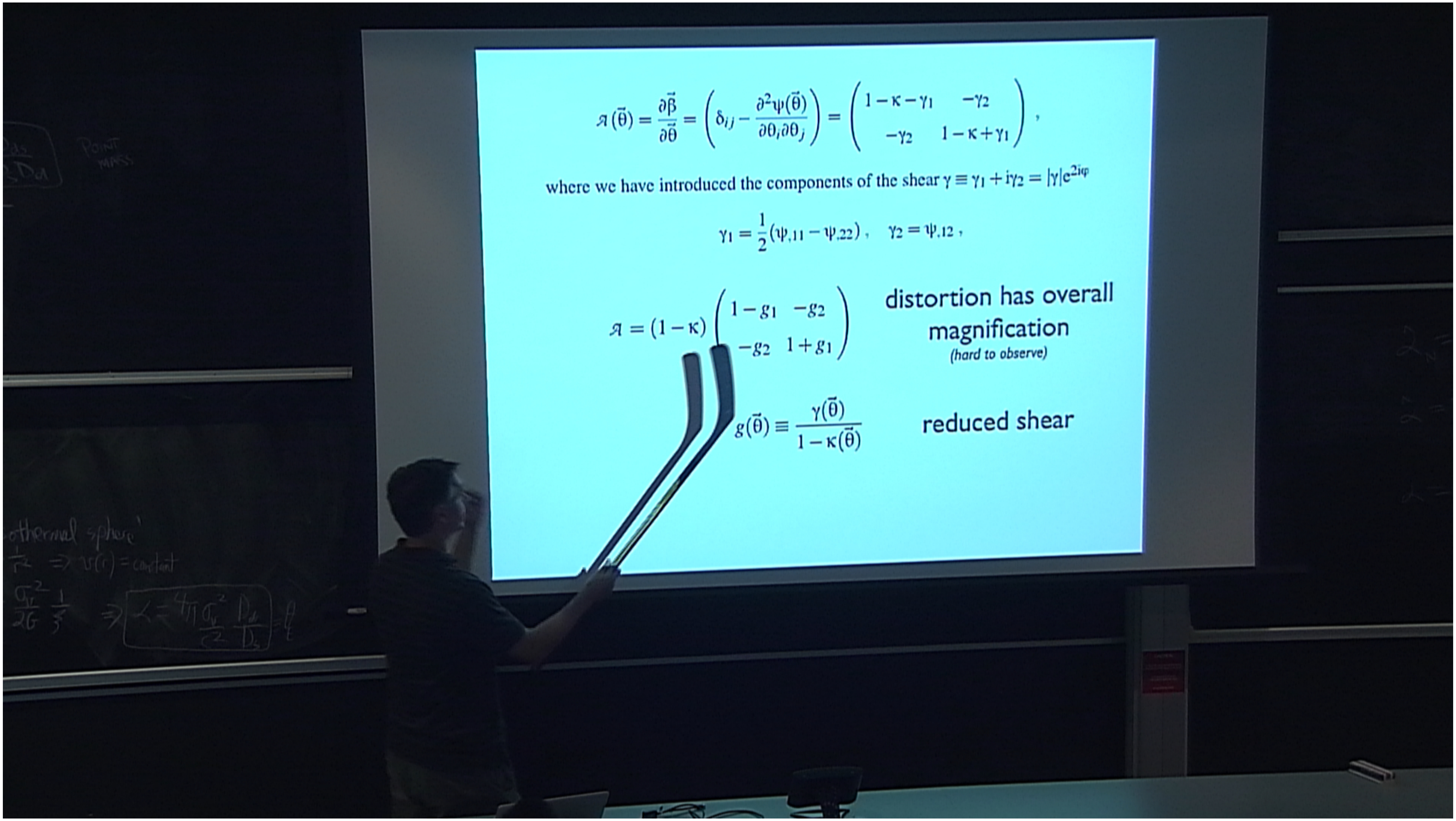


SN Refsdal:
Lensing of a
Supernova!
caught at different times in different images
likely missed one a few decades ago, will
et al 2014



$\mathcal{A}(\vec{\theta}) = \frac{\partial\vec{\beta}}{\partial\vec{\theta}} = \left(\delta_{ij} - \frac{\partial^2\psi(\vec{\theta})}{\partial\theta_i\partial\theta_j}\right) = \begin{pmatrix} 1-\kappa-\gamma_1 & -\gamma_2 \\ -\gamma_2 & 1-\kappa+\gamma_1 \end{pmatrix}$,
where we have introduced the components of the shear $\gamma \equiv \gamma_1 + i\gamma_2 = |\gamma|e^{2i\varphi}$
$\gamma_1 = \frac{1}{2}(\psi_{,11} - \psi_{,22})$, $\gamma_2 = \psi_{,12}$,
$\mathcal{A} = (1-\kappa)\begin{pmatrix} 1-g_1 & -g_2 \\ -g_2 & 1+g_1 \end{pmatrix}$
distortion has overall magnification
(hard to observe)
$g(\vec{\theta}) \equiv \frac{\gamma(\vec{\theta})}{1-\kappa(\vec{\theta})}$
reduced shear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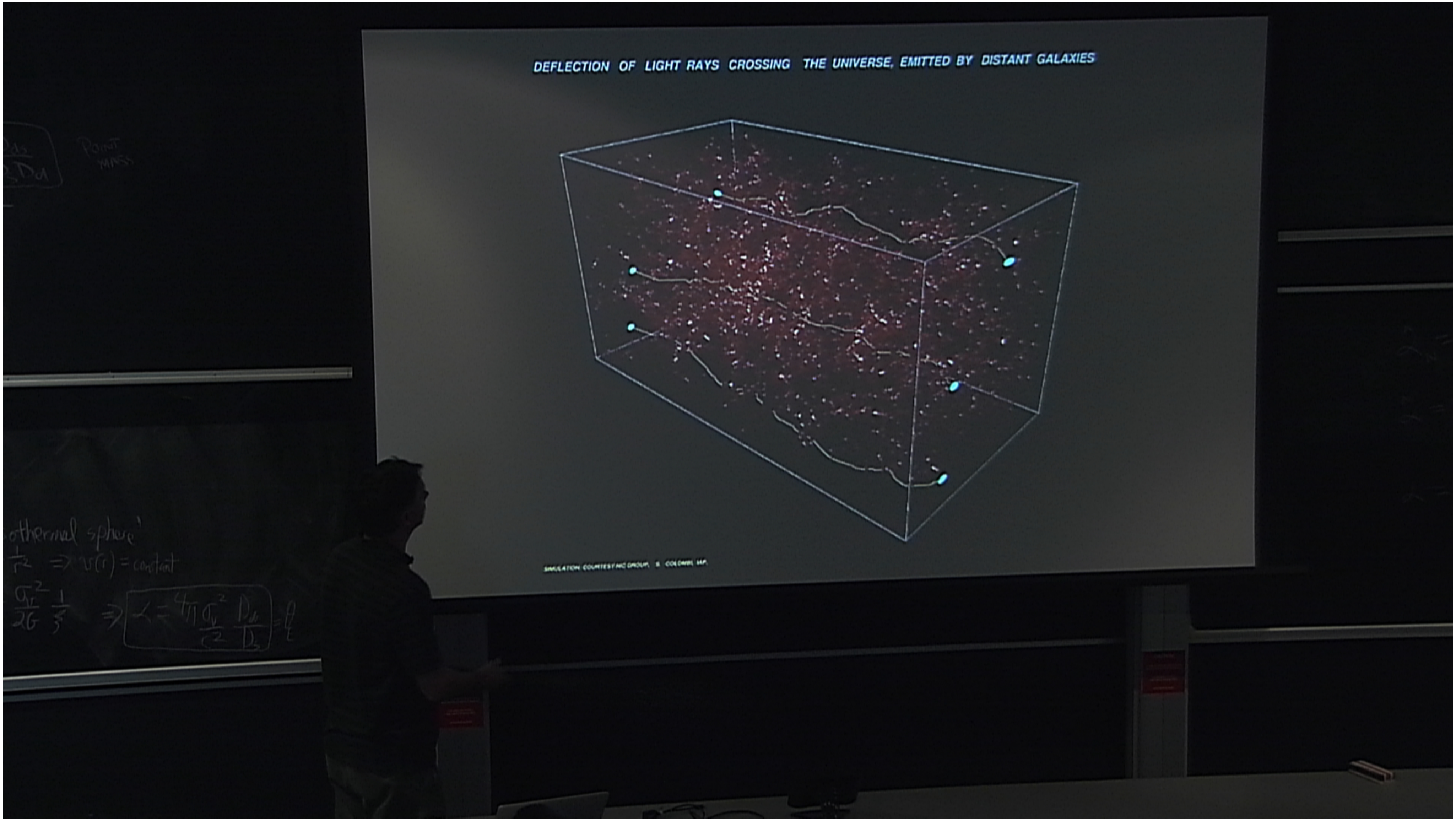
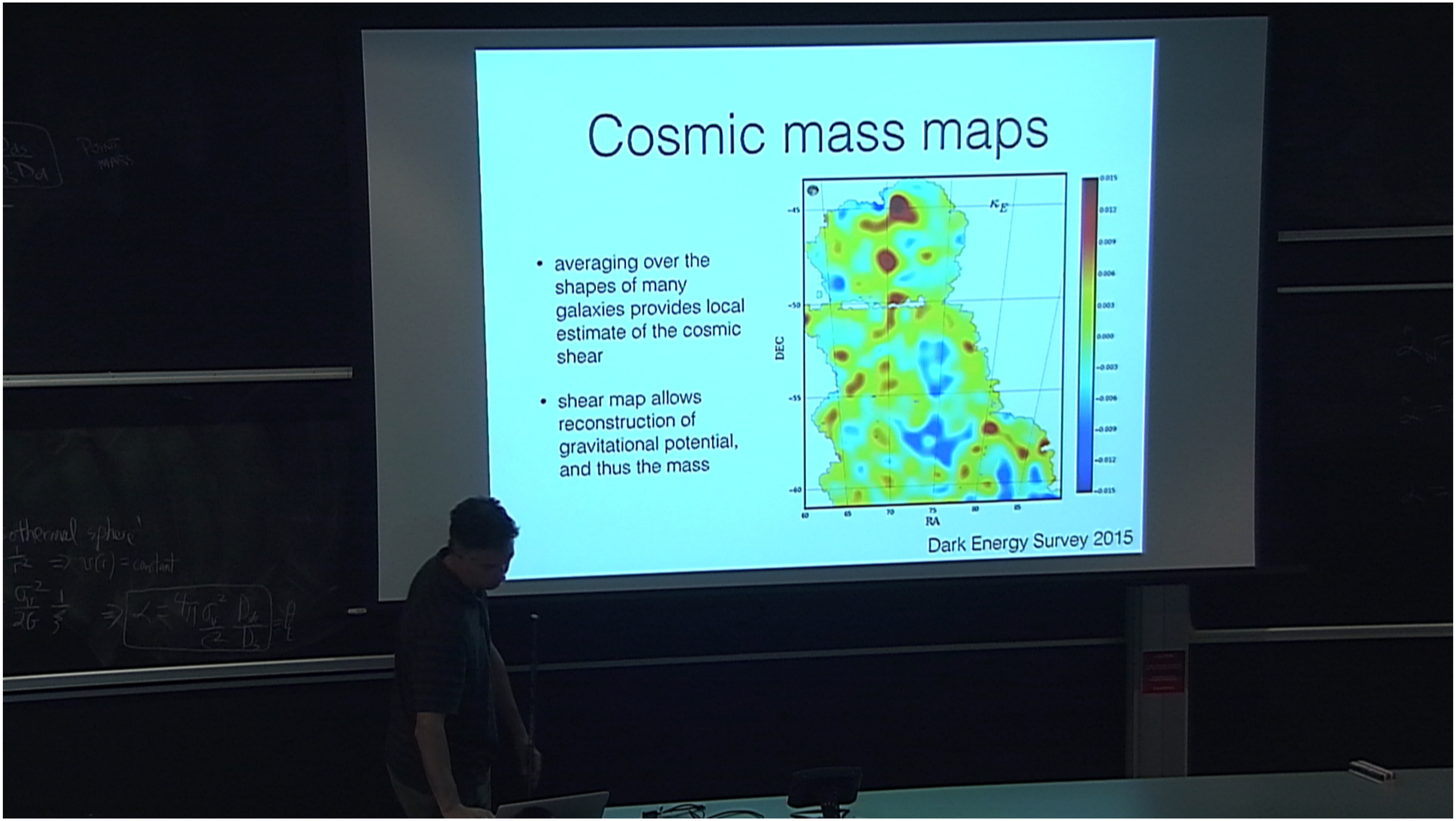
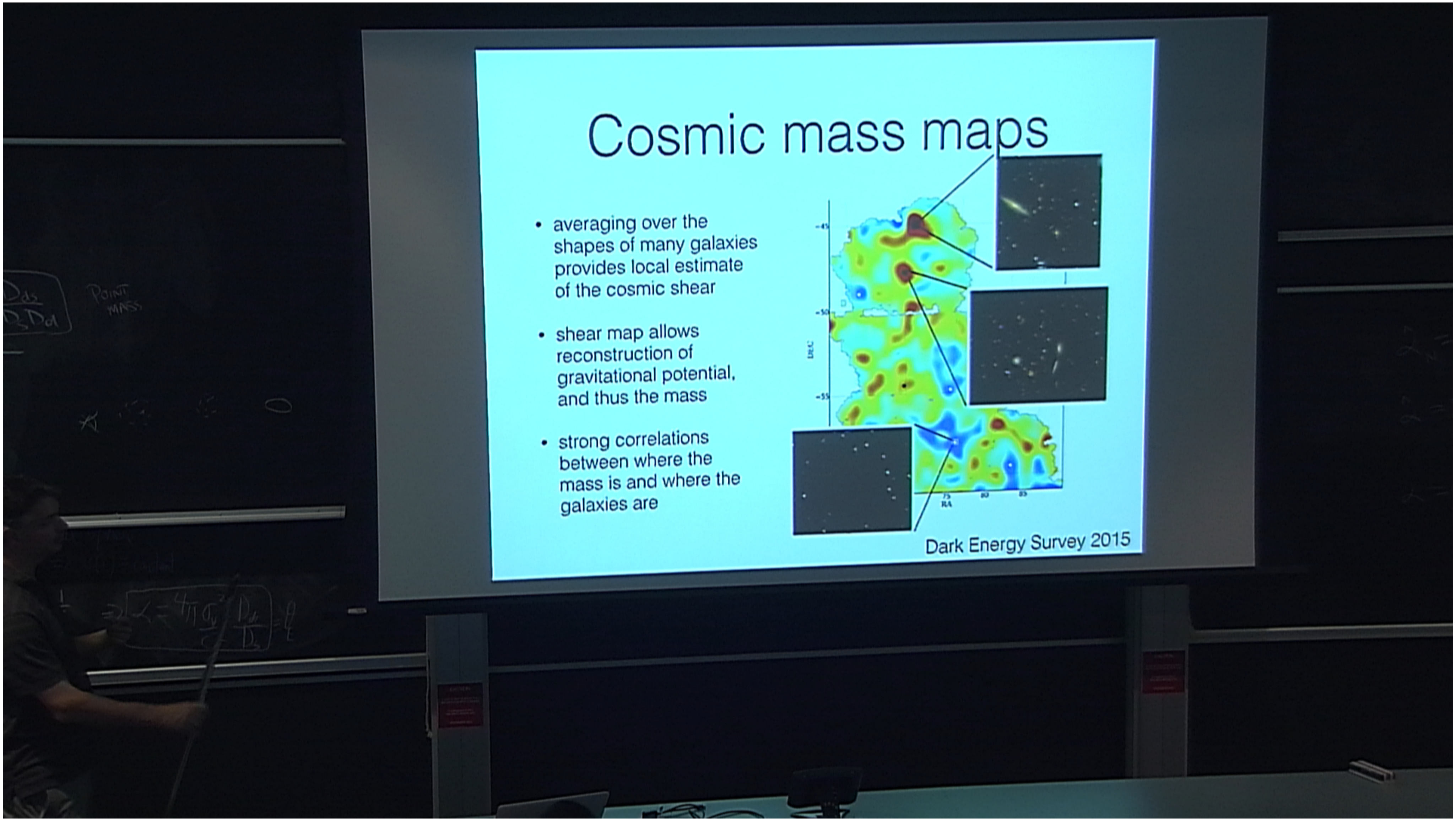
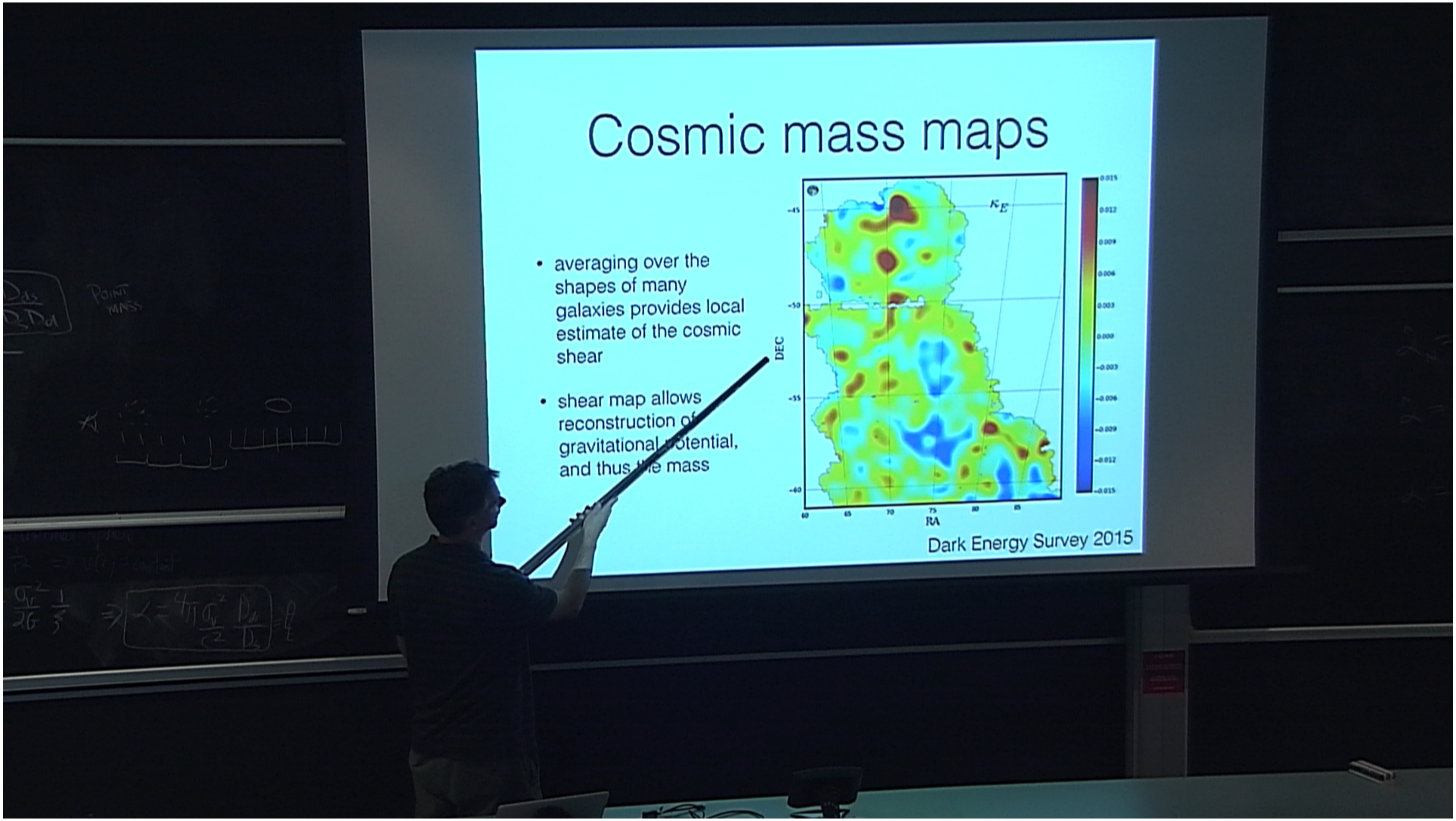

DEFLECTION OF LIGHT RAYS CROSSING THE UNIVERSE, EMITTED BY DISTANT GALAXIES


Galaxies are not round
individual galaxies have complex morphologies
solution: average over many galaxies
© Anglo-Australian Observatory


Cosmic mass maps
• averaging over the shapes of many galaxies provides local estimate of the cosmic shear
• shear map allows reconstruction of gravitational potential, and thus the mass
DEC
RA
κ_E
Dark Energy Survey 2015


Cosmic mass maps
• averaging over the shapes of many galaxies provides local estimate of the cosmic shear
• shear map allows reconstruction of gravitational potential, and thus the mass
• strong correlations between where the mass is and where the galaxies are
Dark Energy Survey 2015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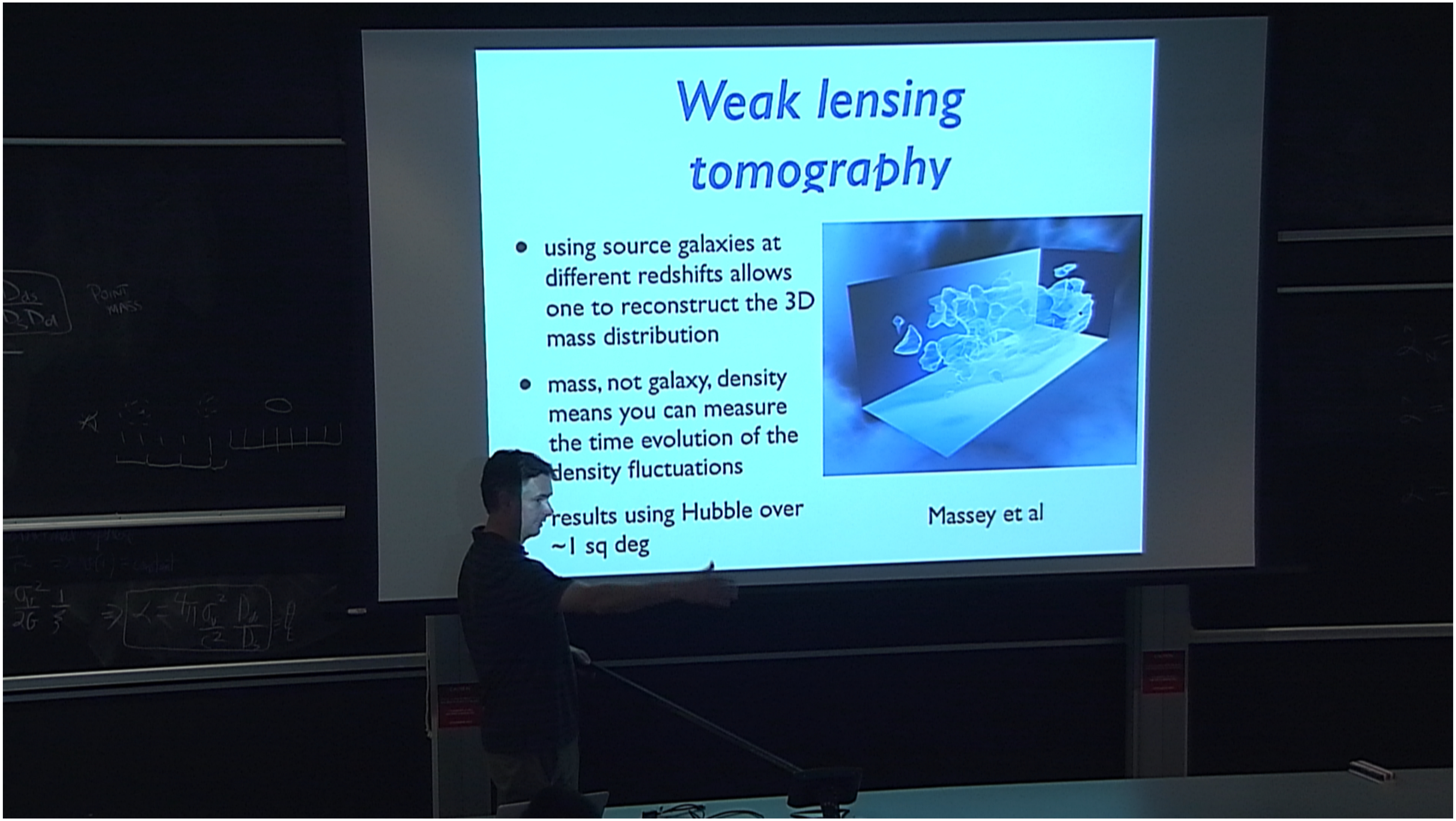
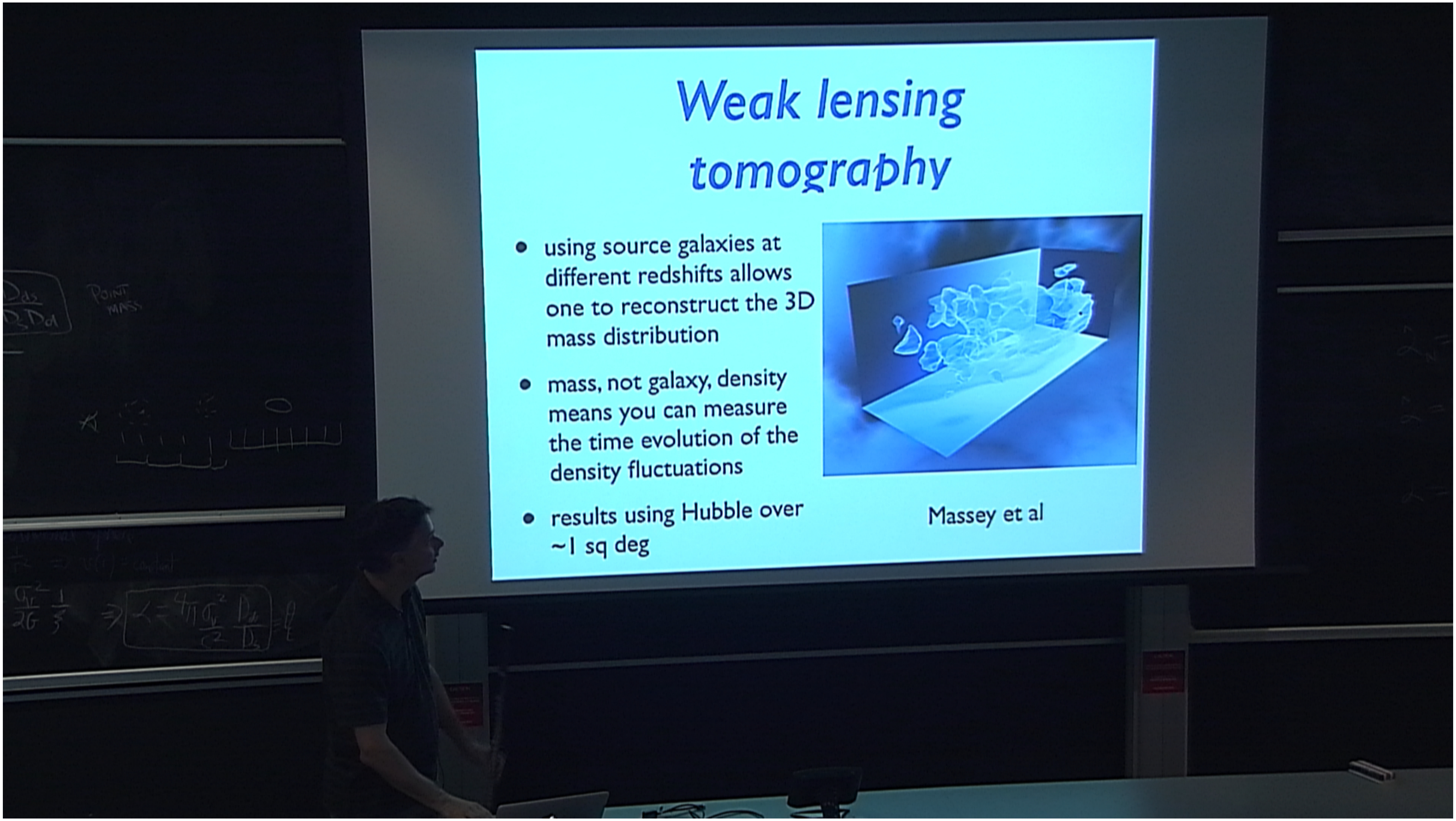
DEC
RA
Dark Energy Survey 2015


DEFLECTION OF LIGHT RAYS CROSSING THE UNIVERSE, EMITTED BY DISTANT GALAXIES
POINT MASS


Weak lensing tomography
using source galaxies at different redshifts allows one to reconstruct the 3D mass distribution
mass, not galaxy, density means you can measure the time evolution of the density fluctuations
results using Hubble over ~1 sq deg
Massey et al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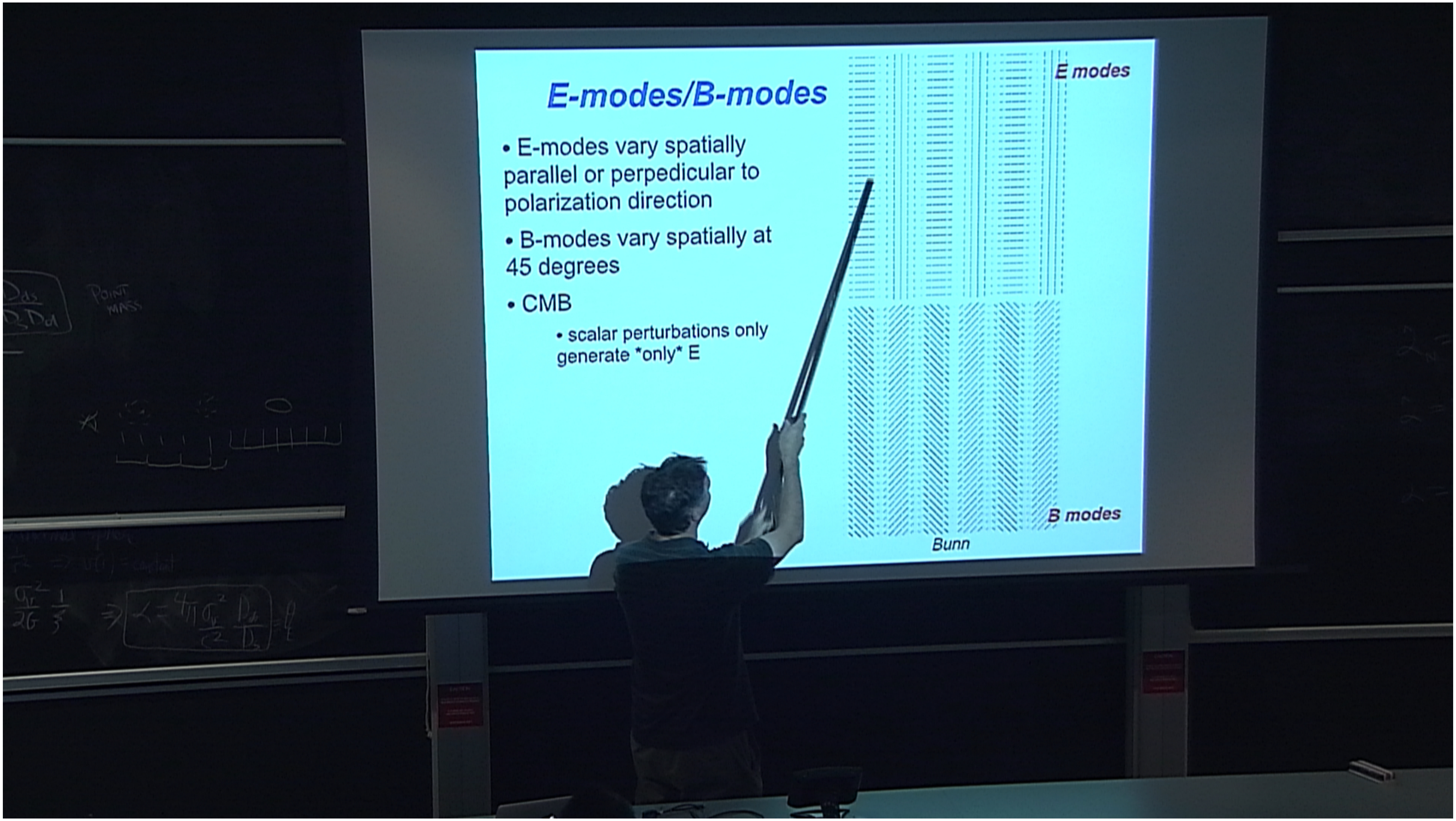
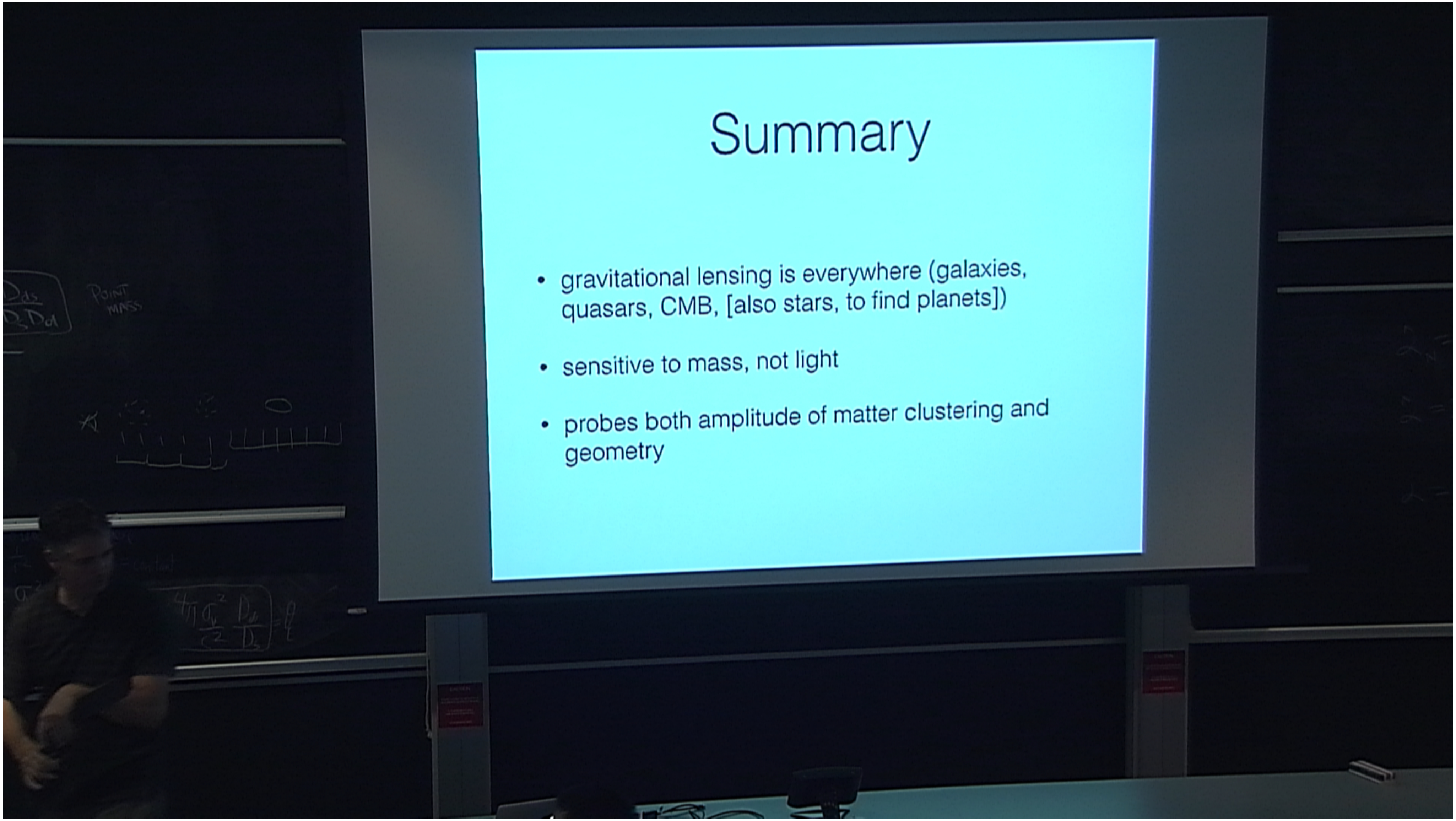

E-modes/B-modes
• E-modes vary spatially parallel or perpedicular to polarization direction
• B-modes vary spatially at 45 degrees
• CMB
• scalar perturbations only generate *only* E
E modes
B modes
Bunn


Summary
• gravitational lensing is everywhere (galaxies, quasars, CMB, [also stars, to find planets])
• sensitive to mass, not light
• probes both amplitude of matter clustering and geometry