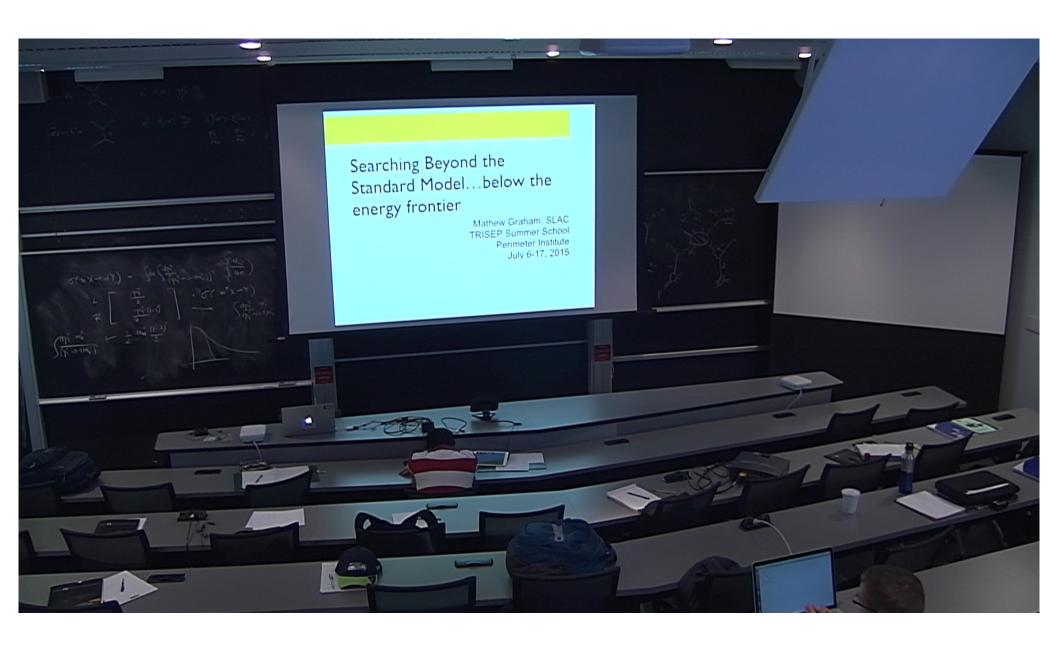
Date: Jul 08, 2015 10:15 AM

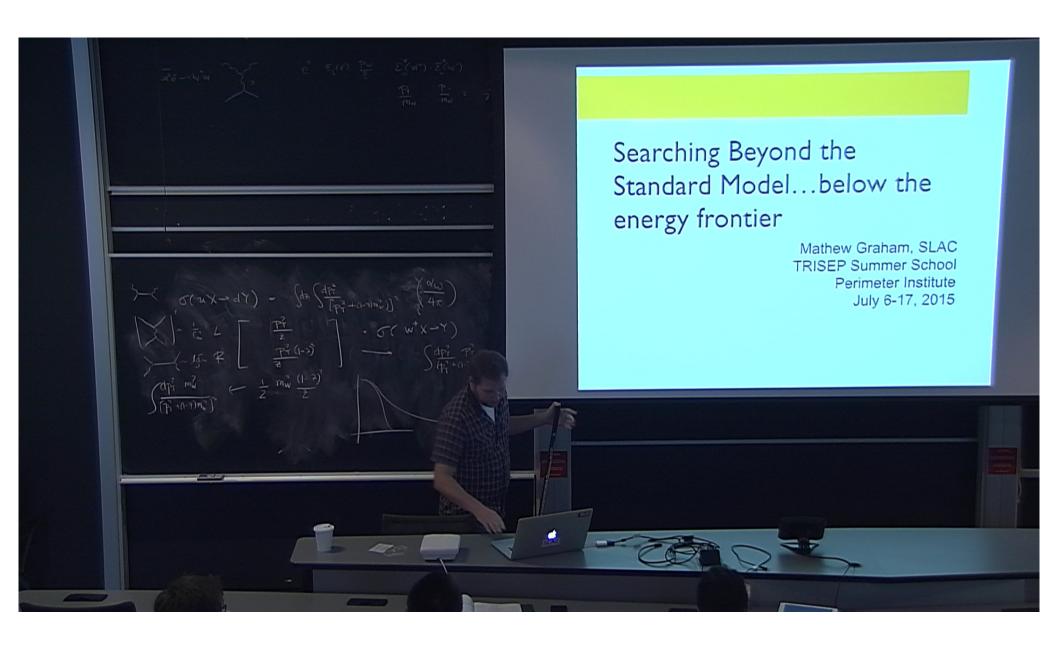
URL: http://pirsa.org/15070023

Abstract:

Pirsa: 15070023

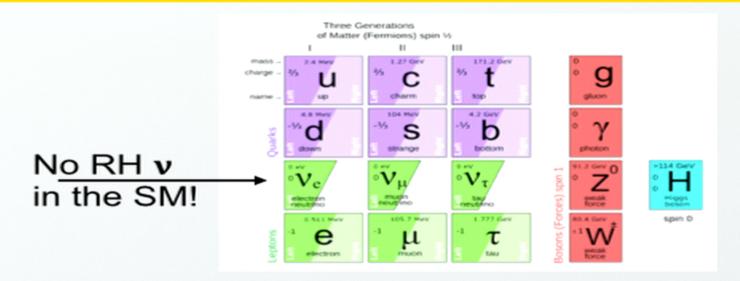


Pirsa: 15070023



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# Not quite done...



#### Big Problems!

- ...strong CP problem...
- ...dark matter...
- ...CP of the universe...
- ...origin & nature of neutrino mass...

minor issues...

... "naturalness" & hierarchy (problems)...

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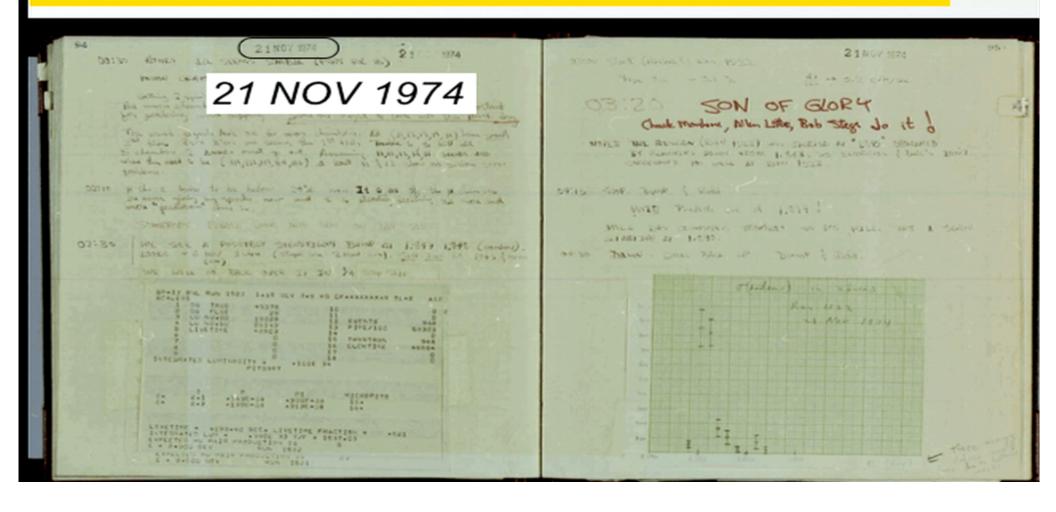
# How to look for new physics

- Go to higher and higher energies!
  - directly produce new particles at scales unobtainable before!
  - has worked in the past...

TRISEP 2015

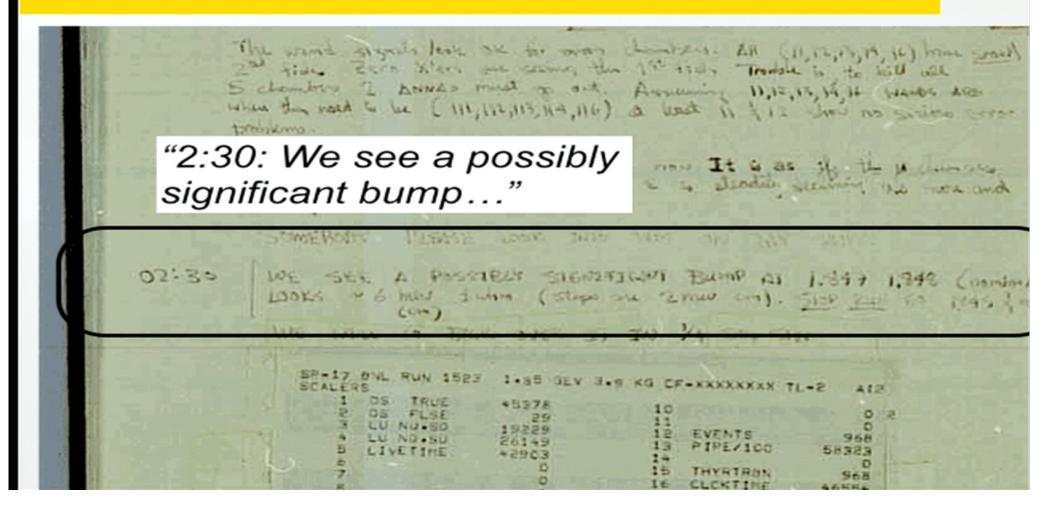
Pirsa: 15070023 Page 5/43

#### How do discover new physics @ low energies...

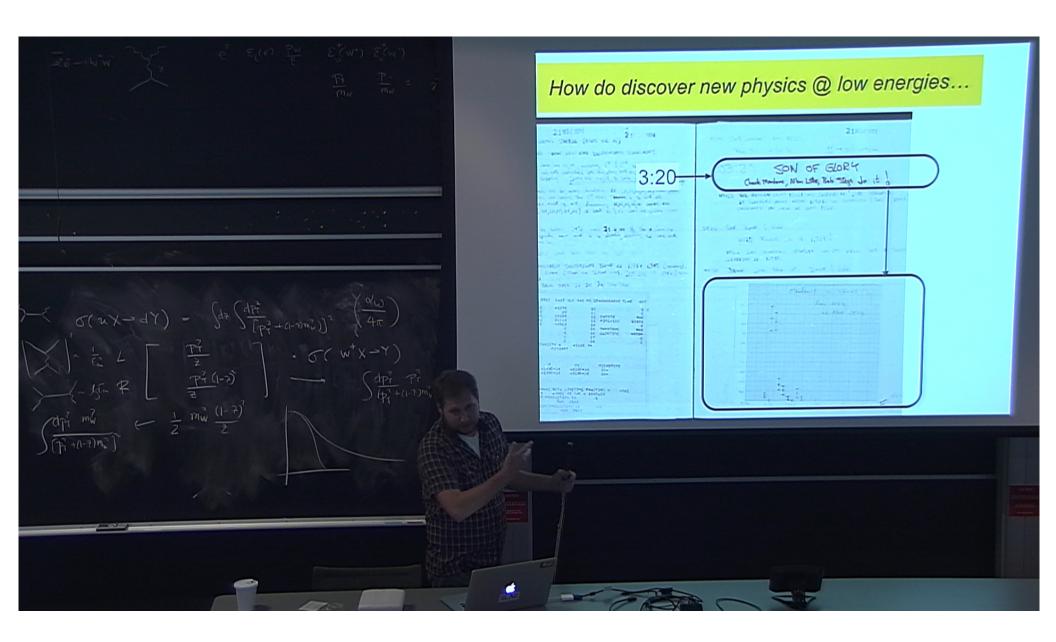


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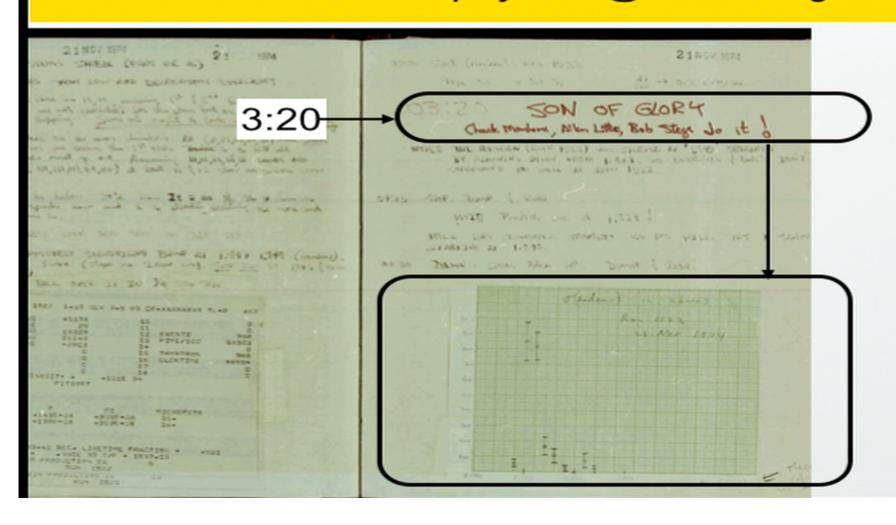


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#### How do discover new physics @ low energies...



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# How to look for new physics

- Go to higher and higher energies!
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- Go to higher and higher intensities!
  - rare production modes
  - rare decay modes
  - precise measurements of SM predictions
- Look somewhere "new"\*\*\*!
  - neutrinos!
  - particle-astrophysics

\*\*\*neither of these are new, but definitely room for growth

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# How to look for new physics

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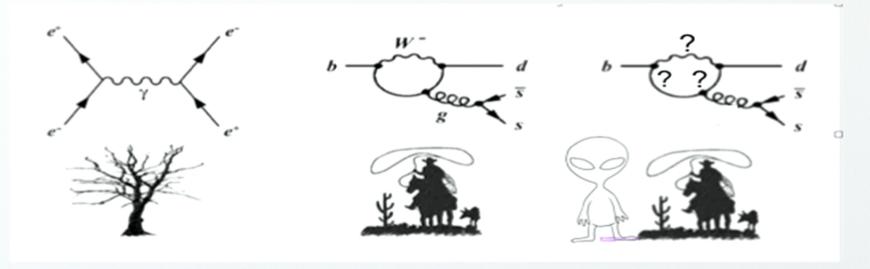
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# Searching For BSM @ "low-energies"

- Low Energy != Low Scale
  - Loops are our friend!



...high mass scales of NP particles in the loops is not suppressed like they are @ tree level (direct production)

Also! Typically don't have tree-level SM processes to compete against→lower background

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### Searching For BSM @ "low-energies"

- Low Energy != Low Scale
  - Loops are our friend!
  - Rare or forbidden decays
    - Lepton flavor violation searches
  - precision measurements of well predicted parameters
    - muon g-2
    - CKM unitarity triangle (& related measurements)
- Looking in new places Neutrinos!
  - we haven't fully explored neutrino sector (it's hard)
  - mixing matrix still has many open questions
    - is it unitary? Are there more "neutrinos"
    - what's the CP phase...is it 0? Some other interesting number?
  - What's the deal with the neutrino mass(es)?
    - ordering of masses? SM-like or inverted
    - where does it come from?
- Direct production, but with very intense sources
  - A ton of experiments were done in the old days that looked for crazy things...now we have much more intense beams, better detectors & electronics. Some crazy things may be there!
    - Dark sector searches

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# Searching For BSM @ "low-energies"

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Dark sector searches



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$$|\mathcal{A}_{Tot}|^2 = |\mathcal{A}_{SM}|^2 + |\mathcal{A}_{NP}|^2 + 2Re\left(\mathcal{A}_{SM}\mathcal{A}_{NP}^*\right)$$
If SM is large Tough...  $\mathcal{S} \sim \sqrt{\mathcal{L}} \times \frac{|\mathcal{A}_{NP}|^2}{|\mathcal{A}_{SM}|}$ 

$$S = Significance$$
 $\mathcal{L} = Integrated Luminosity$ 
TRISEP 2015

13

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13

$$|\mathcal{A}_{Tot}|^2 = |\mathcal{A}_{SM}|^2 + |\mathcal{A}_{NP}|^2 + 2Re\left(\mathcal{A}_{SM}\mathcal{A}_{NP}^*\right)$$
 If SM is (very) small 
$$\mathbf{Easy!} \ \ \mathcal{S} \sim \mathcal{L} \times |\mathcal{A}_{NP}|^2$$

$$S = Significance$$
 $L = Integrated Luminosity$ 
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$$|\mathcal{A}_{Tot}|^2 = |\mathcal{A}_{SM}|^2 + |\mathcal{A}_{NP}|^2 + 2Re\left(\mathcal{A}_{SM}\mathcal{A}_{NP}^*\right)$$

Get tricky, measure the interference. Sensitive to A<sub>NP</sub> (not ^2)...

If SM is ~ NP

$$S = Significance$$
  
 $L = Integrated Luminosity$   
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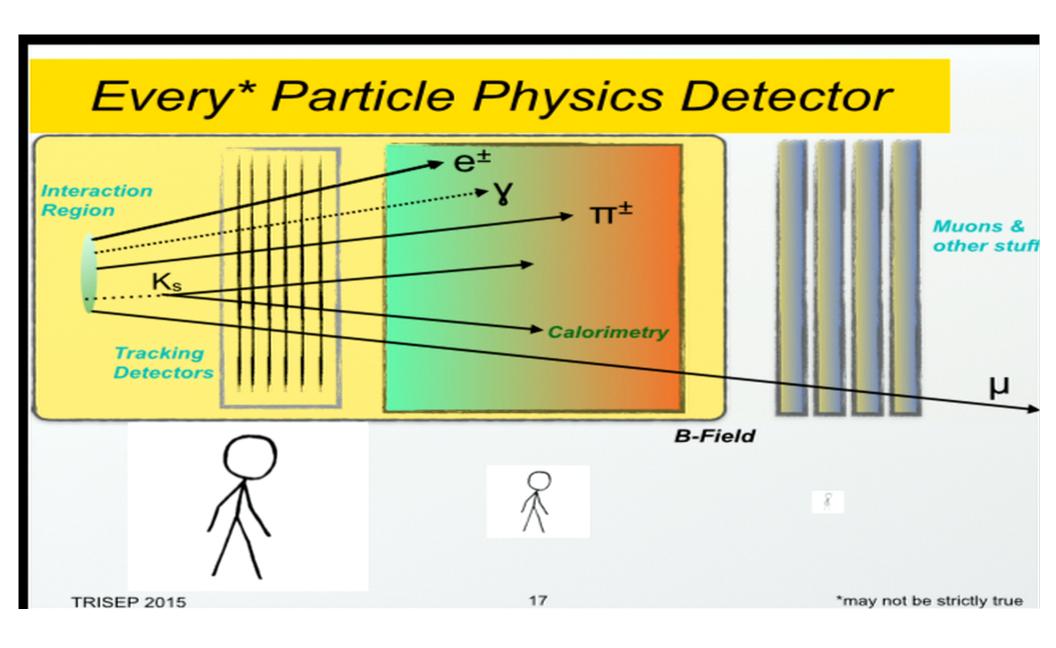
15

# **BUT FIRST!**

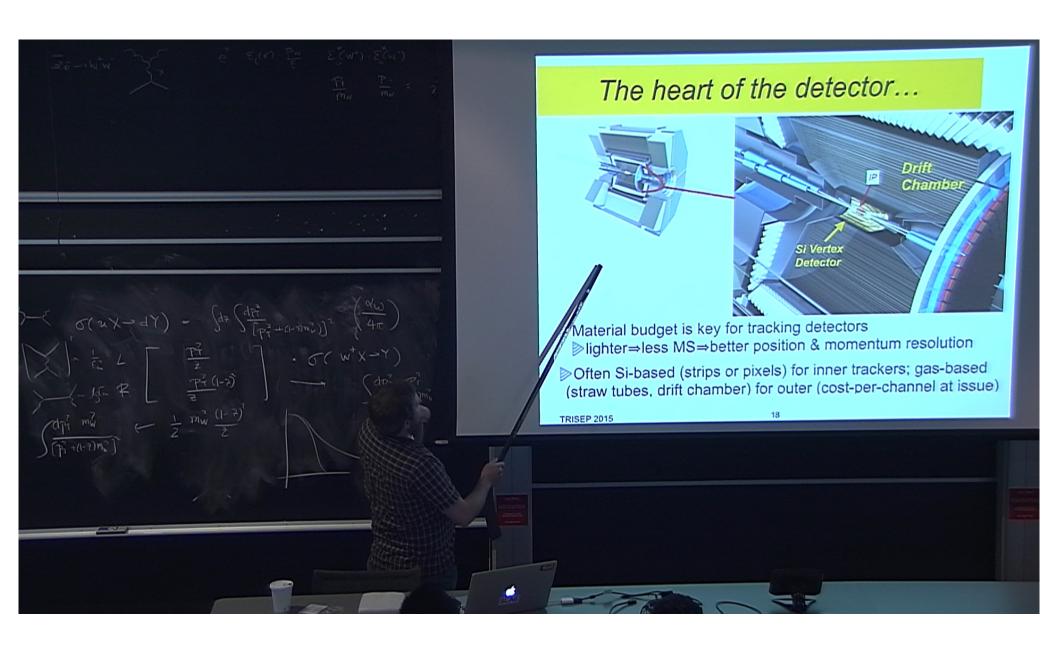
...some words about detectors...



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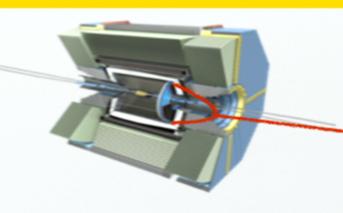


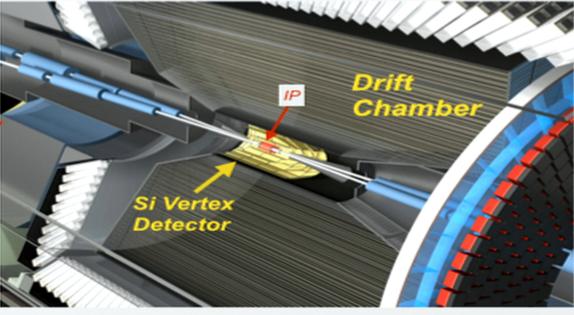
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#### The heart of the detector...



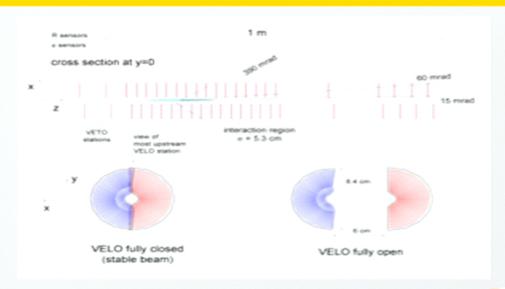


- Material budget is key for tracking detectors
  ▶lighter⇒less MS⇒better position & momentum resolution
- Often Si-based (strips or pixels) for inner trackers; gas-based (straw tubes, drift chamber) for outer (cost-per-channel at issue)

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# A very cool thing.



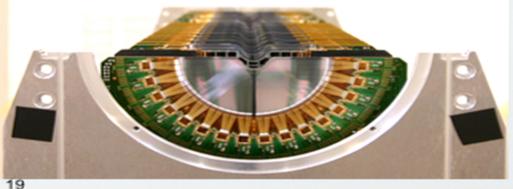
The LHCb VELO is great example of a "vertex-tracker"...

- encircles the LHC beam, aligned perpendicular to the beam direction
- designed to detect the B-meson decays slightly displaced from primary pp interaction

...retracts during beam tune and closes up for data taking

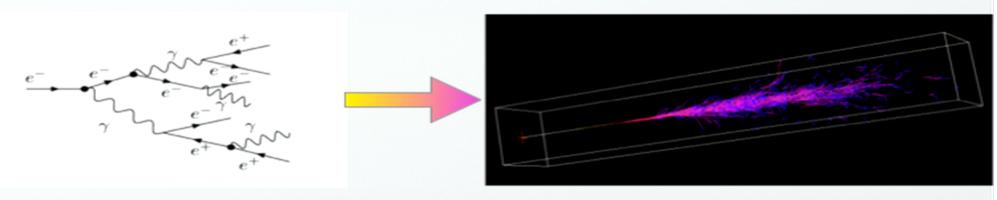
consists of pairs of radial-axial half-circle Si-strip layers

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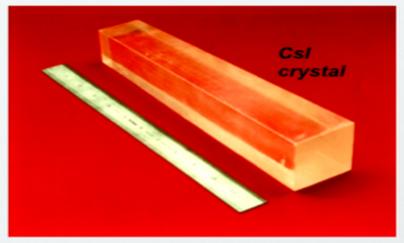


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# (EM) Calorimetry



- ...opposite of trackers, put lots of material in the way...want to *cause* showers
  - high density, high Z materials
- Pb glass, PbW, Csl, LAr
- Measure the energy either from Cerenkov radiation/scintillation light (via PMTs or APDs) or from ionization charge (in LAr detectors);



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# Two ECal Examples

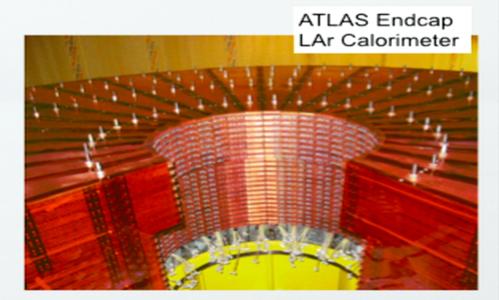
BaBar Central Calorimeter



...also key for particle ID (photons! but also electrons/muons/pions).

For hadron calorimetry, missing energy is a huge deal; need XCal systems to be ~hermetic.

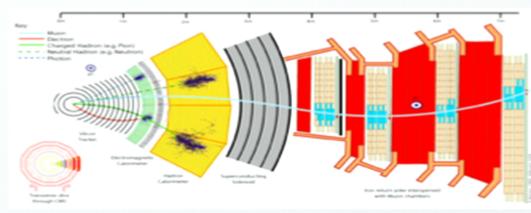
Calorimeters tend to be *fast*, both the physics and the readout...trigger systems rely heavily on EM calorimetry



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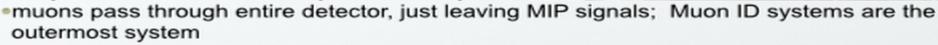
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#### Muon ID



"Muons are probably the most important particle in particle physics" — Matt Graham





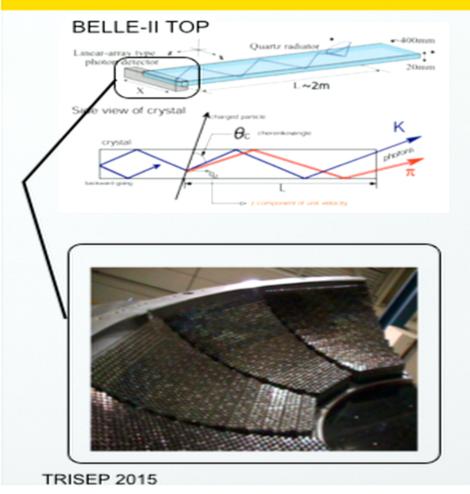
•need to be cheap because they cover large area; sandwich detectors between iron is very common…if a particle gets all the way though it's a muon

scintillating paddles; Resistive Plate Chambers; streamer tubes...etc

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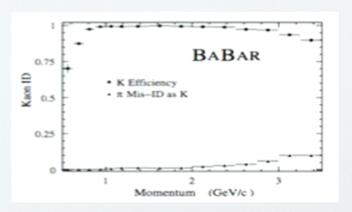


# Specialists...PID



Particularly at low-ish energies (GeV) it's important to decipher π/K(/p/d etc...);

for π/K separation at GeV, Cherenkov-based detectors are very powerful



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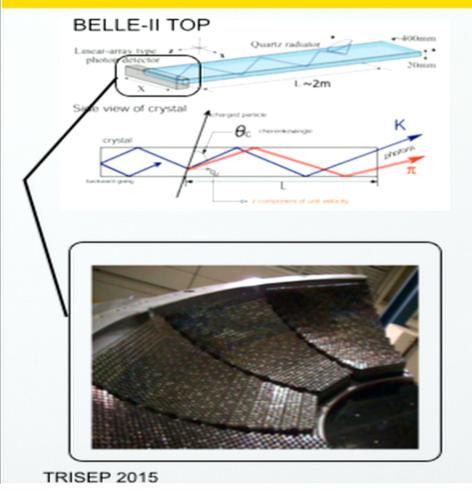
Pirsa: 15070023

# Searching Beyond the Standard Model: Precision Measurements & Rare Decays

Mathew Graham, SLAC TRISEP Summer School Perimeter Institute July 6-17, 2015

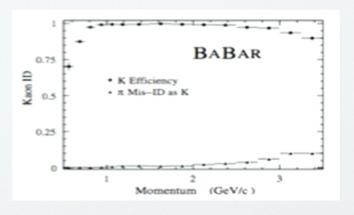
Pirsa: 15070023 Page 28/43

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23

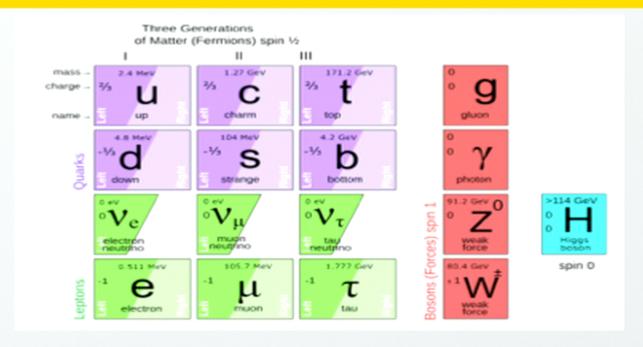
Pirsa: 15070023 Page 29/43

# Searching Beyond the Standard Model: Precision Measurements & Rare Decays

Mathew Graham, SLAC Flavo(u)r TRISEP Summer School Perimeter Institute July 6-17, 2015

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#### This talk...intro: Flavor

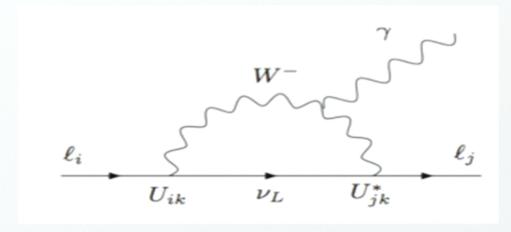


flavor, i.e. "u-ness", "μ-ness" etc, is not conserved in the SM there is no mechanism for flavor-changing neutral currents at tree-level in the SM

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# (Charged) Lepton-Flavor Violation

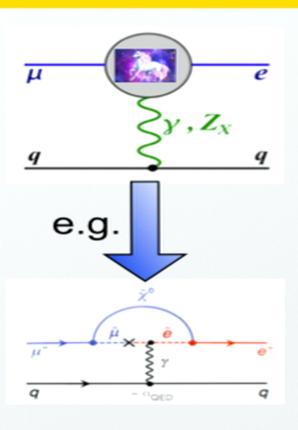


$$Br(\mu \to e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^{\star} U_{ei} \frac{\Delta m_{i1}^2}{M_W^2} \right|^2 < 10^{-54} \,,$$

this is very, very small

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Three ways to look for  $\mu \rightarrow e$ :

These are *complimentary* probes sensitive to different *classes* of models

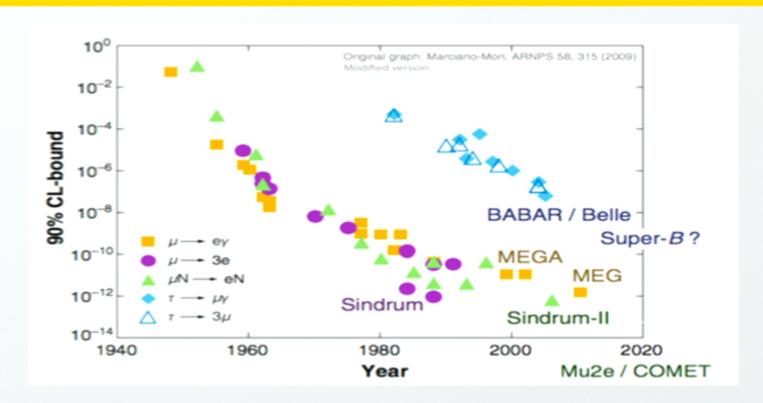
	$\mu \to 3e$	$\mu \to e \gamma$	$\mu \to e$ conversion
$O_{S,V}^{4\ell}$	✓	_	_
$O_D$	✓	✓	✓
$O_V^q$	_	_	✓
$O_S^q$	_	_	✓

Cirigliano@Beauty2014

TRISEP 2015

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# So far, so nothing.

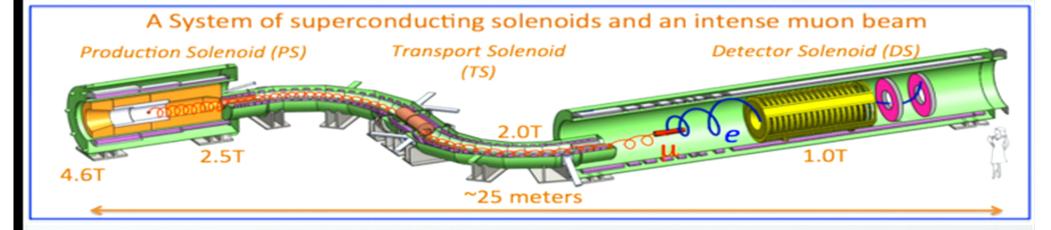


...current limits on  $\mu \rightarrow e$  probe up to **10<sup>4</sup> TeV** 

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# The µ2e Experiment @ FNAL



- crazy, curvy beam line to get rid of any line-of-sight background
- muons steered into Al "stopping" target, where the magic happens

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# Stopping muons



stopped muon replaces an electron in the Al atom...

...most of the time, the muon either is "captured" (releasing  $v_{\mu}$ ) or decays in-orbit

primary background are "DIO" decays...but energy spectrum is different from signal



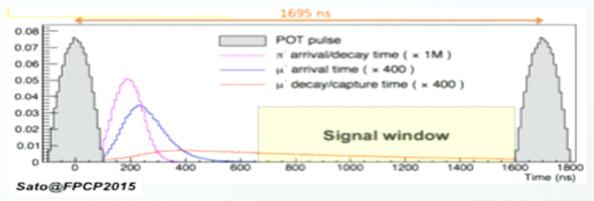
symmetry magazine

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# μ2e Signatures



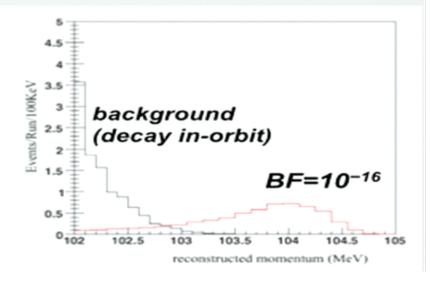
 pulsed beam→wait until all is quite

 measure energy of the electron; µ→e conversions peak at ~m(µ) while decays cut off before that



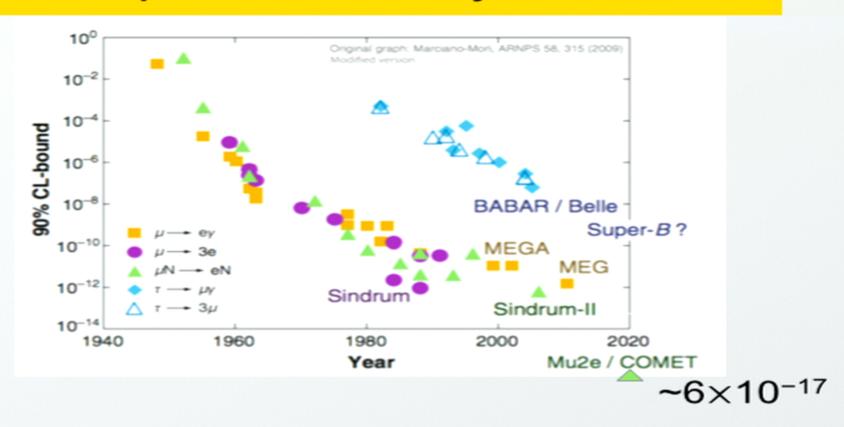
Gives ~ background free signal region

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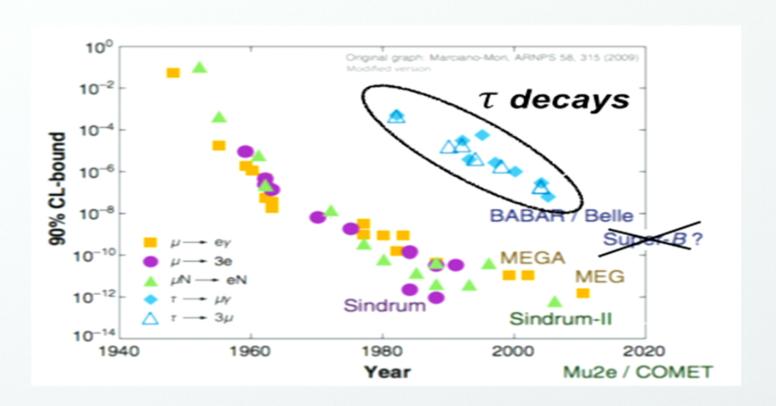
# μ2e Sensitivity



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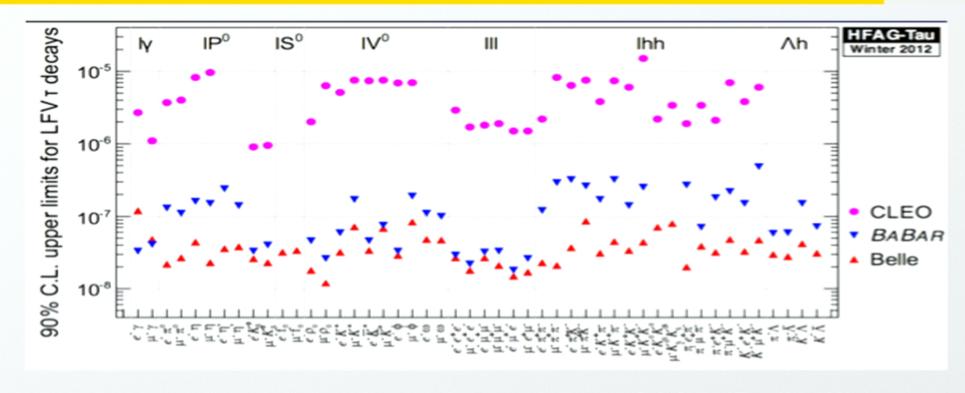
# Tau decays too...



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# ...many, many tau decays



...these should go down by x10 by ~2025 (Belle-II)

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# (C)LFV Recap

- Charged lepton flavor violating reactions are a great place to look for physics BSM…the first one we see will be a huge discovery
  - very clean theoretically AND experimentally...the limiting factor is how many muons or taus we can produce
  - can probe very high energy scales
  - good complementarity with the energy frontier
- Unfortunately, there's no indication that it's right around the corner
  - that's not a reason to not look...just keep digging!
- Upcoming experiments take a nice step forward:
  - mu2e (FNAL)
  - COMET (J-PARC)
  - mu3e (PSI)



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#### CKM Matrix & CP-Violation in the quark sector

$$\approx \begin{pmatrix} 1 - \frac{\lambda^2}{2} & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \frac{\lambda^2}{2} & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix}$$
$$\lambda = \sin \theta_C = 0.22$$

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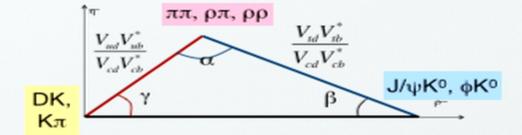
# THE Unitarity Triangle

There are 6 triangle relations but most have very un-equal sides== large/small angles == small CPV effects

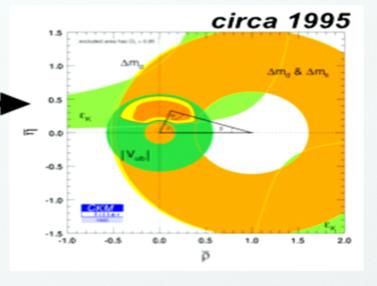
...EXCEPT!!!

THE UT = Column 1 ( $\mathbf{d}$ ) \* Column 3( $\mathbf{b}$ )

Unitarity condition:  $V_{ud}V_{ub}^* + V_{cd}V_{cb}^* + V_{sd}V_{sb}^* = 0$ 



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One of the main goals of B-Factories is to see if this triangle closes...if it doesn't, it would be sign of physics beyond the Standard Model!

CKMfitter Group (J. Charles et al.), Eur. Phys. J. C41, 1-131 (2005) [hep-ph/0406184], updated results and plots available at: <a href="http://ckmfitter.in/2p3.fr">http://ckmfitter.in/2p3.fr</a>

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